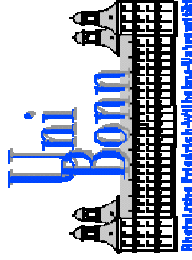


Direct neutrino mass experiments: Present and Future

XXth International Conference on Neutrino Physics and Astrophysics, Munich, May 25–30, 2002



Christian Weinheimer

Institut für Strahlen- und Kernphysik, Rheinische Friedrich-Wilhelms-Universität, 53115 Bonn

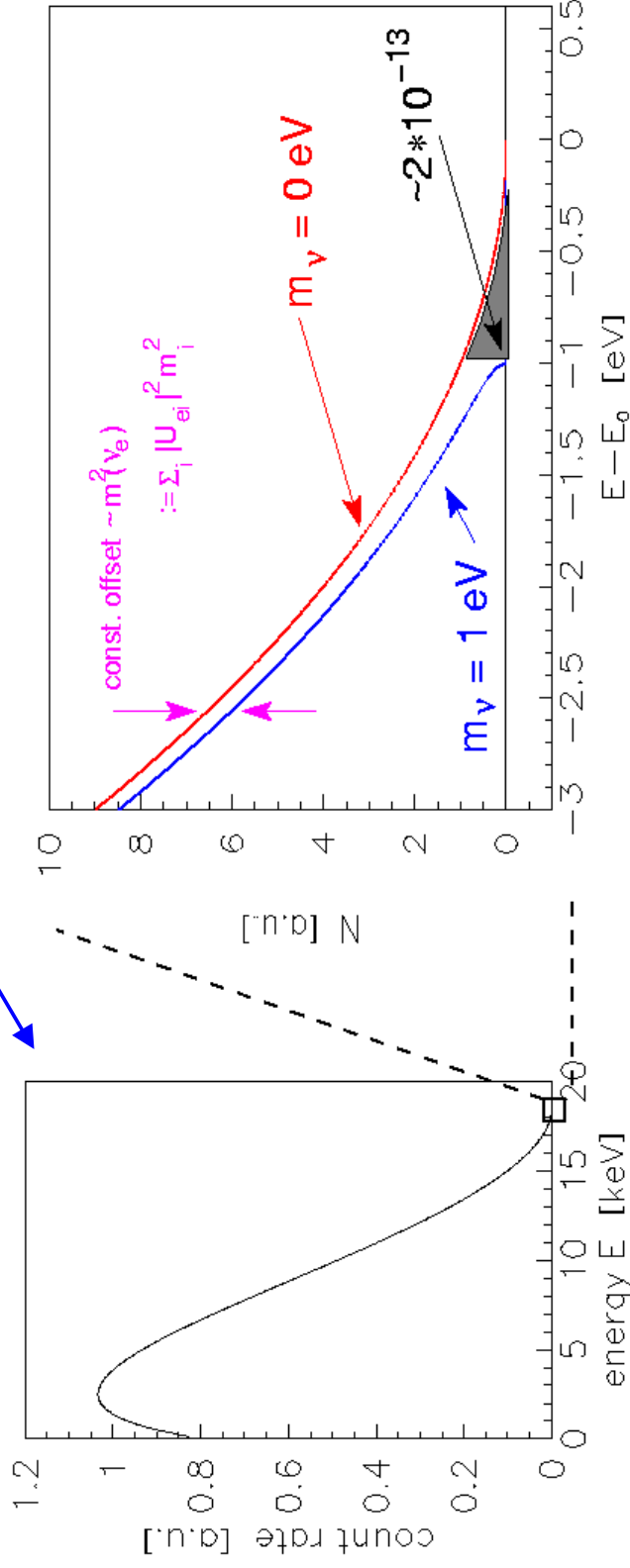
Email: weinheimer@iskp.uni-bonn.de

- ⇒ Tritium β decay experiments at Mainz and Troitsk
- ⇒ How to proceed?
future projects
- ⇒ **KATRIN**
a large tritium β decay experiment with sub-eV sensitivity
- ⇒ **Conclusion**

Direct measurement of $m(\nu_e)$

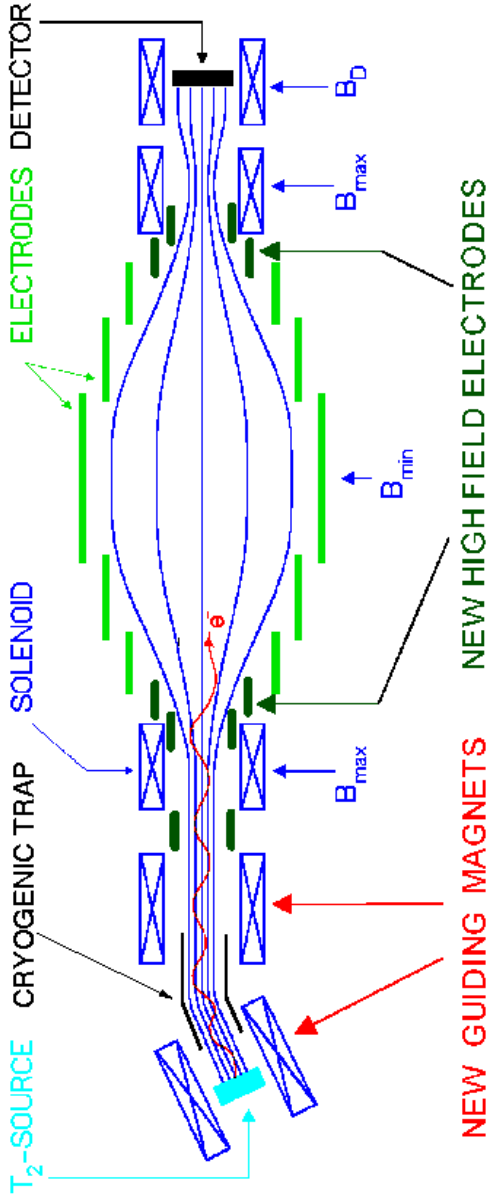


super-allowed
 $E_0 = 18.6 \text{ keV}$
 $t_{1/2} = 12.3 \text{ a}$



Need very high energy resolution & very high signal rate & very low background

Mainz Neutrino Mass Experiment since 1997

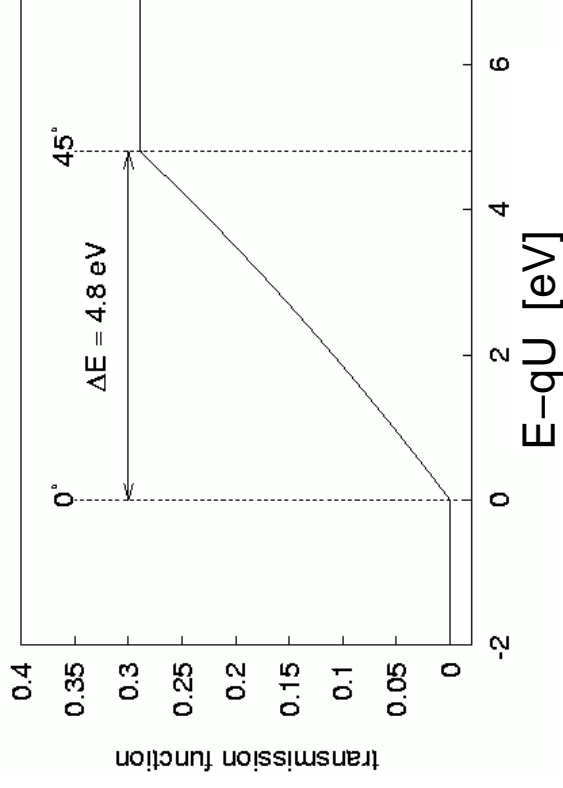


Magnetic Adiabatic Collimation + Electrostatic Filter (MAC-E-Filter)

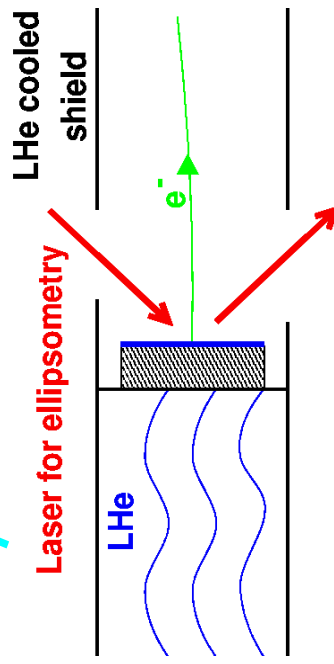
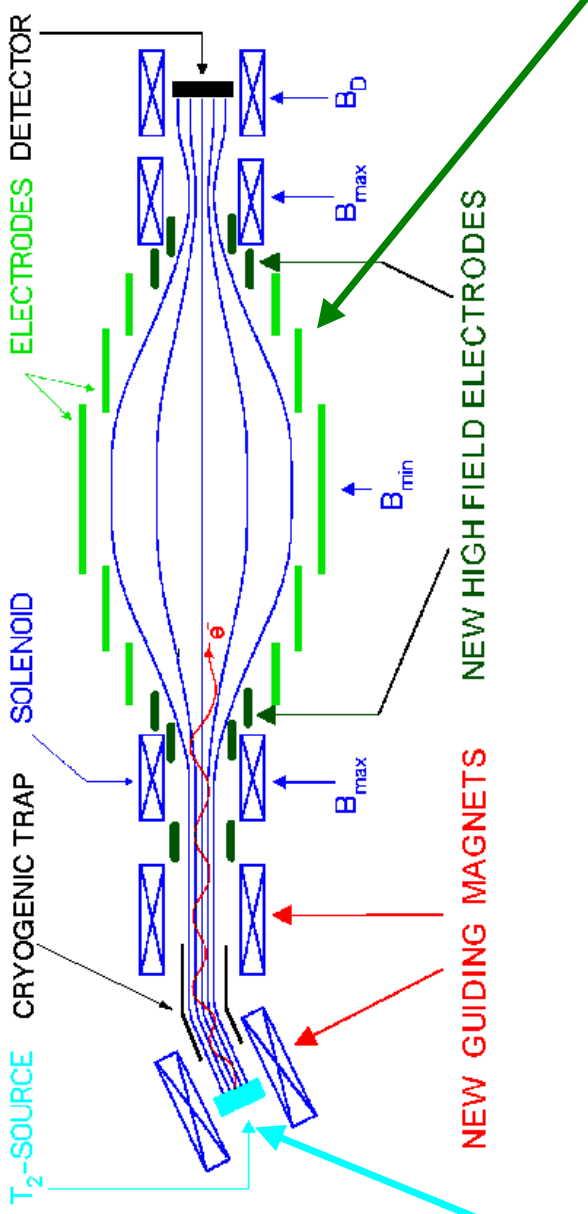
⇒ sharp integrating transmission function

without tails:

$$\Delta E = E \cdot B_{\min} / B_{\max} = E \cdot A_{s,\text{eff}} / A_{\text{analysis}} \approx 4.8 \text{ eV}$$



Mainz Neutrino Mass Experiment since 1997



- T₂ Film at 1.86 K
- quench-condensed on graphite (HOPG)
- 45 nm thick (≈130ML), area 2cm²
- Thickness determination by ellipsometry

Mainz

ν group

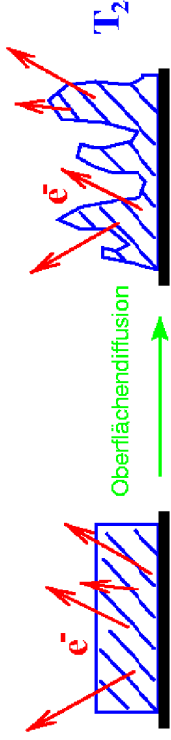
2001:

- J. Bonn
- B. Bornschein*
- L. Bornschein
- B. Flatt
- Ch. Kraus
- B. Müller
- E.W. Otten
- J.P.Schall
- Th. Thümmler**
- Ch. Weinheimer**

* → FZ Karlsruhe
 ** → Univ. Bonn

Investigation and improvement of systematics

- Roughening transition of T_2 film



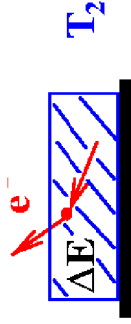
Determination of dynamics: $\Delta E = (45 \pm 6) k_B K$

\Rightarrow no roughening transition below 2 K

L. Fleischmann et al., J. Low Temp Phys. **119** (2000) 615, (with P. Leiderer

L. Fleischmann et al., Eur. Phys. J. **B16** (2000) 521 Konstanztz)

- Inelastic scattering



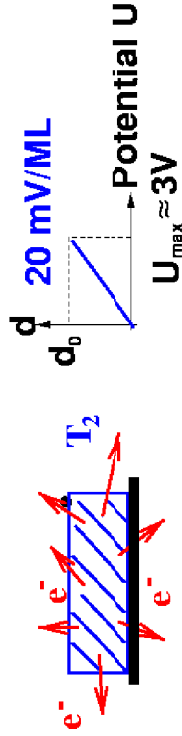
Determination of cross section:

$$\sigma_{tot} = (2.98 \pm 0.16) \cdot 10^{-18} \text{ cm}^2$$

Determination of energy loss function:

V.N. Aseev et al., Eur. Phys. J. **D10** (2000) 39

- Self charging of T_2 film



Determination of critical field:

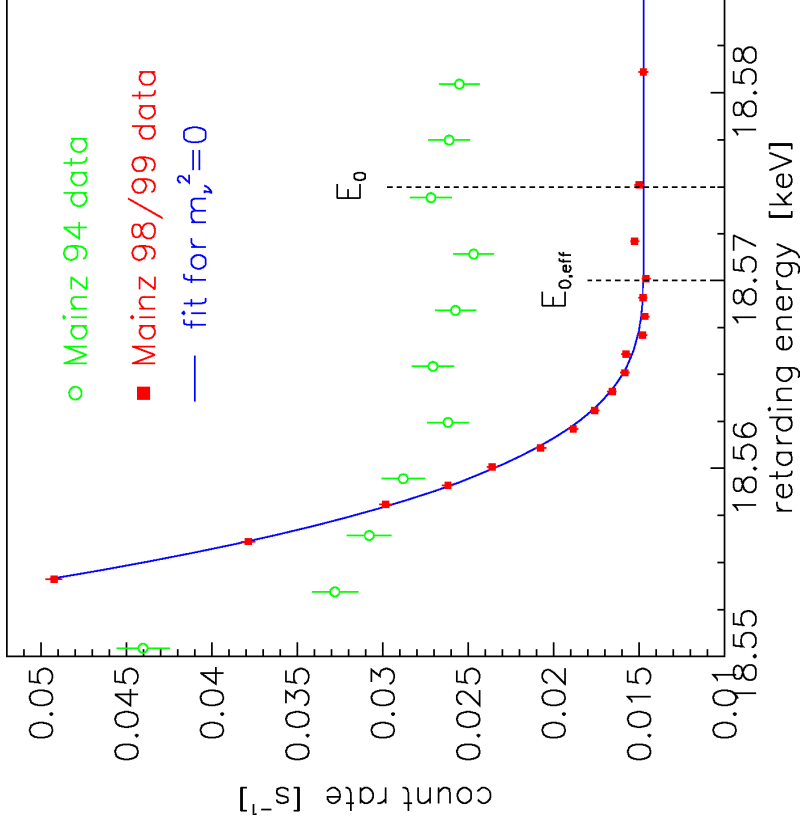
$$E_c = (63 \pm 4) \text{ MV/m} = 20 \text{ mV/monolayer}$$

\Rightarrow slight broadening of energy resolution

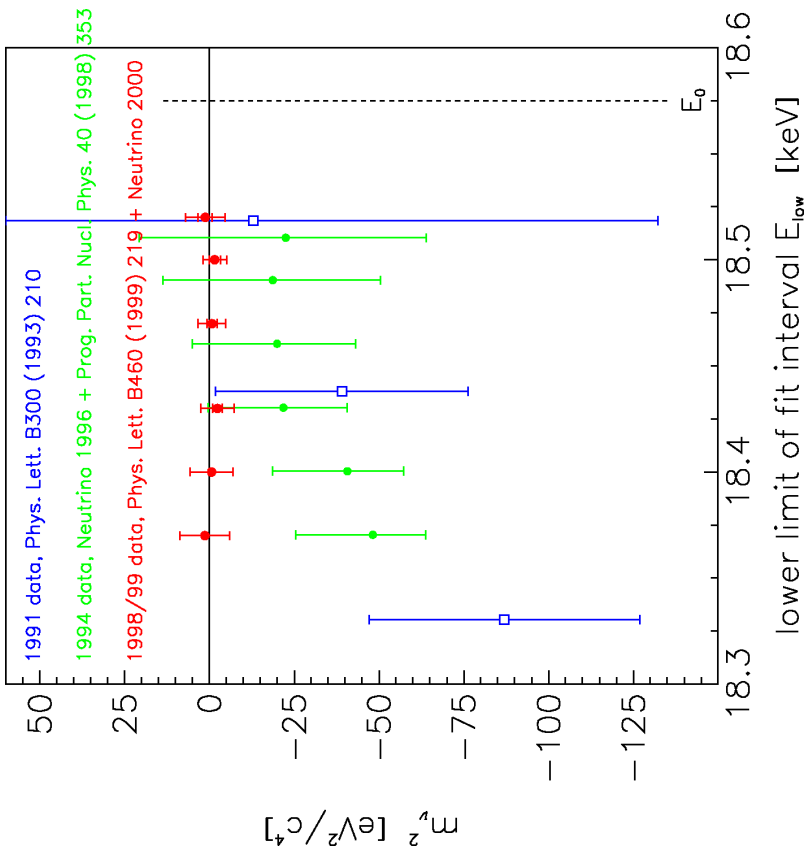
H. Barth et al., Prog. Part. Nucl. Phys. **40** (1998) 353,

B. Bornschein, PhD thesis, publication in preparation

Mainz data of 1998, 1999



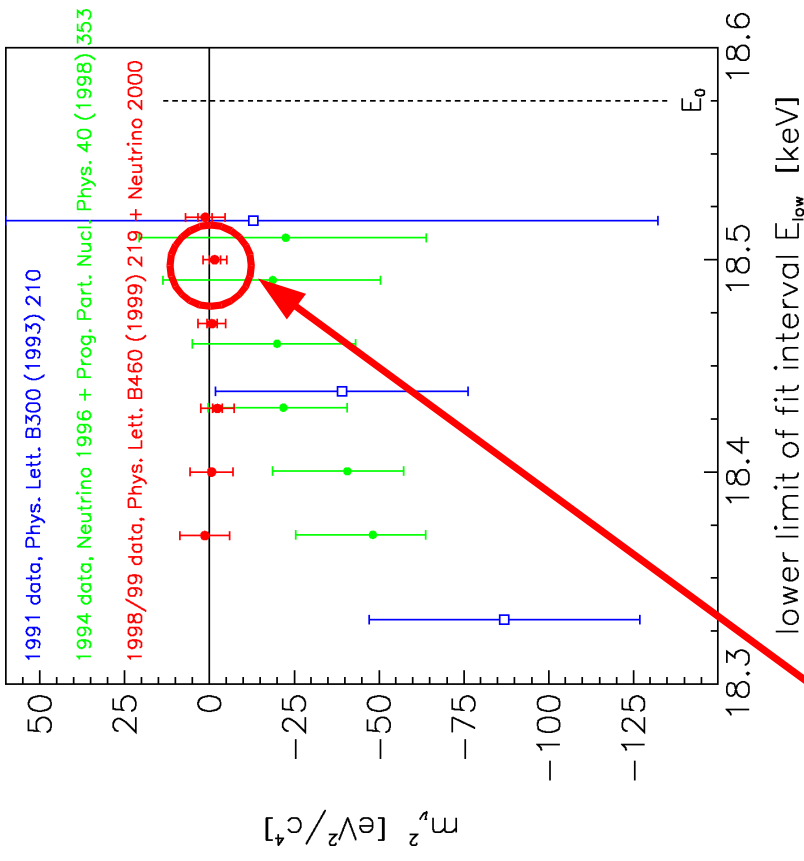
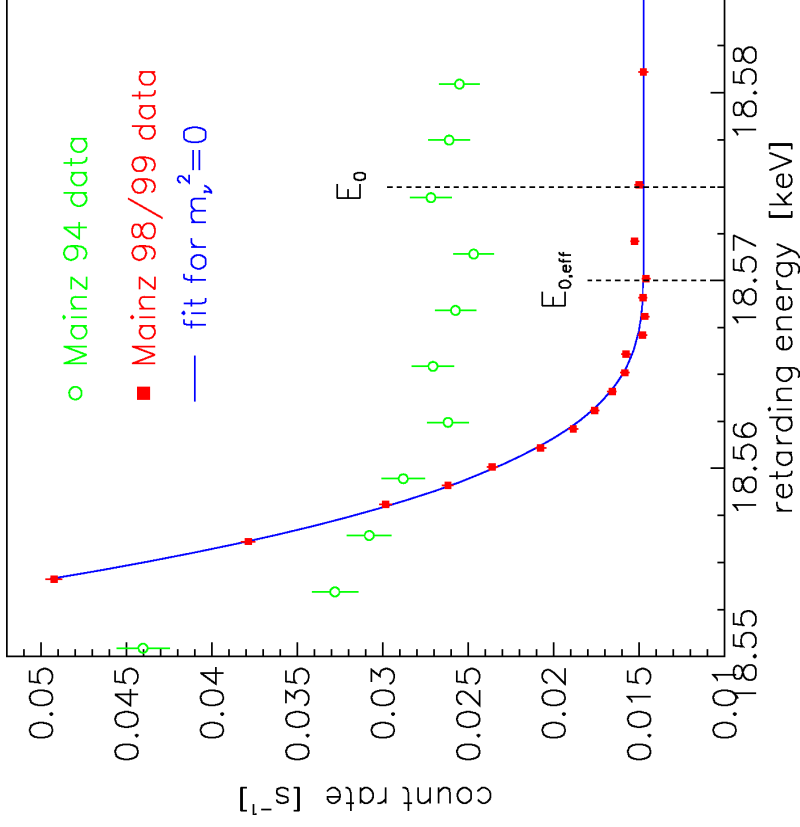
← lower limit of fit“
 ← Fit range →



1998+1999: roughening transition
 avoided by $T < 2$ K

1998+1999: Signal/background 10 x higher

Mainz data of 1998, 1999



$$m^2(\nu) = -1.6 \pm 2.5 \pm 2.1 \text{ eV}^2 \quad (\chi^2/\text{d.o.f.} = 125/121)$$

$$\Rightarrow m(\nu) < 2.2 \text{ eV} \quad (95\% \text{ C.L.})$$

(J. Bonn et al., Nucl. Phys. B (Proc. Suppl.) 91 (2001) 273)

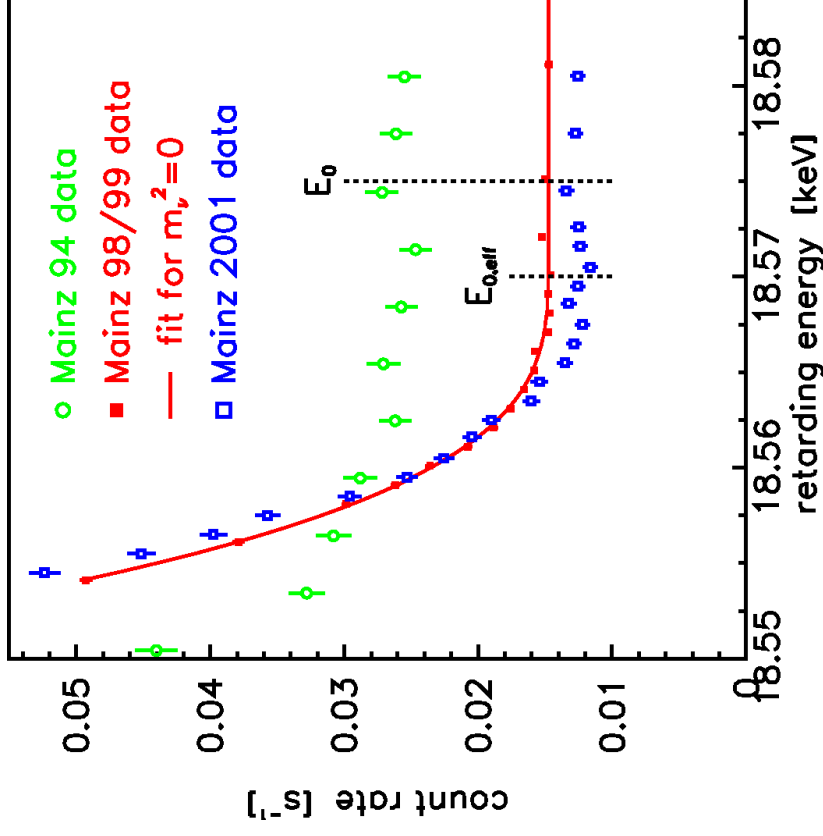
Mainz measurements in 2001

(In 2000 two measurements (Q9,Q10) with background problems due to trapped particles: hysteresis effect in dependence of scanning direction)

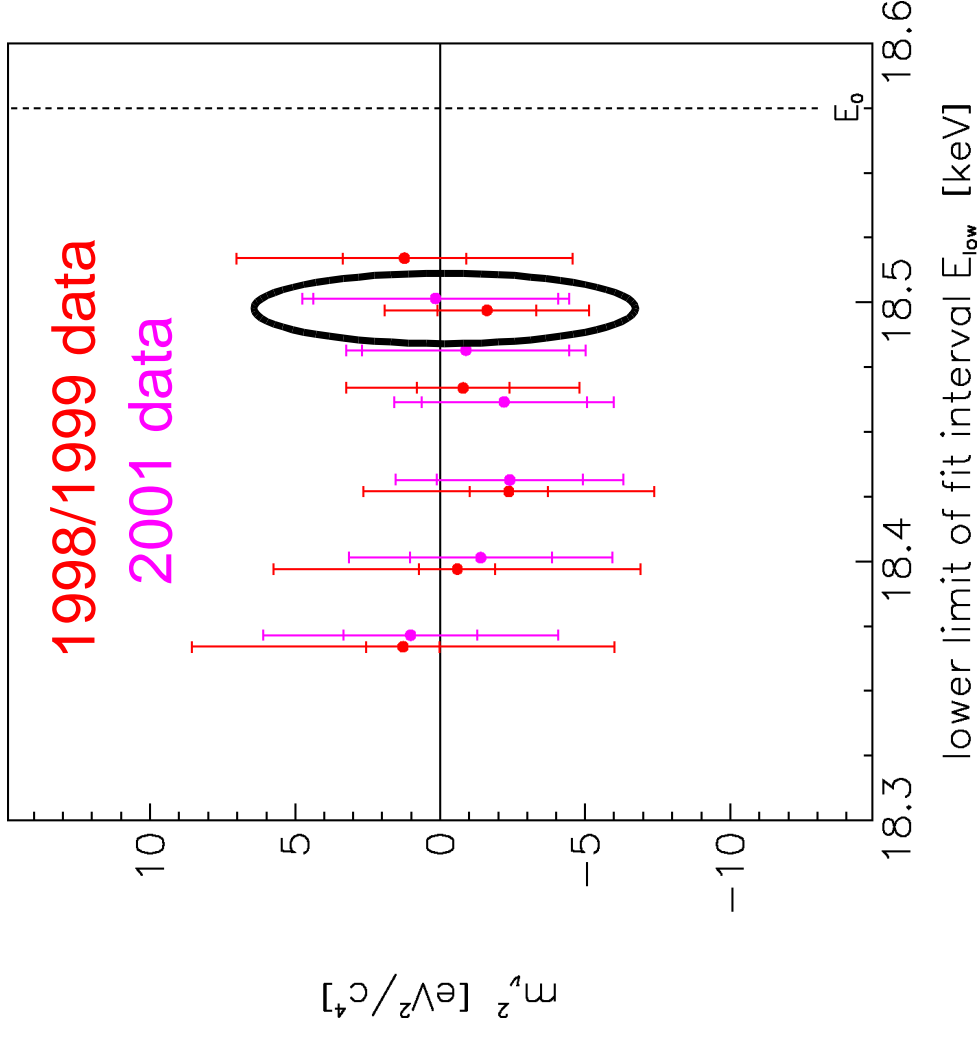
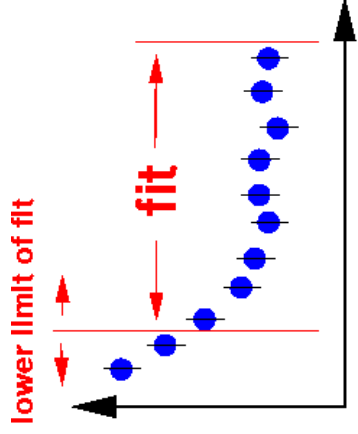
Before Q11,Q12: optimal preparation

- Vacuum: baken out, getter activated:
tritium source
spectrometer
detector
- T₂ source, HV: new:
source substrate
tritium ampula
HV divider oil
- Conditioning at high pressure (10⁻⁷mbar)

⇒ **Low and stable background,
no Troitsk-like anomalies**



Results of 1998/1999, 2001 data



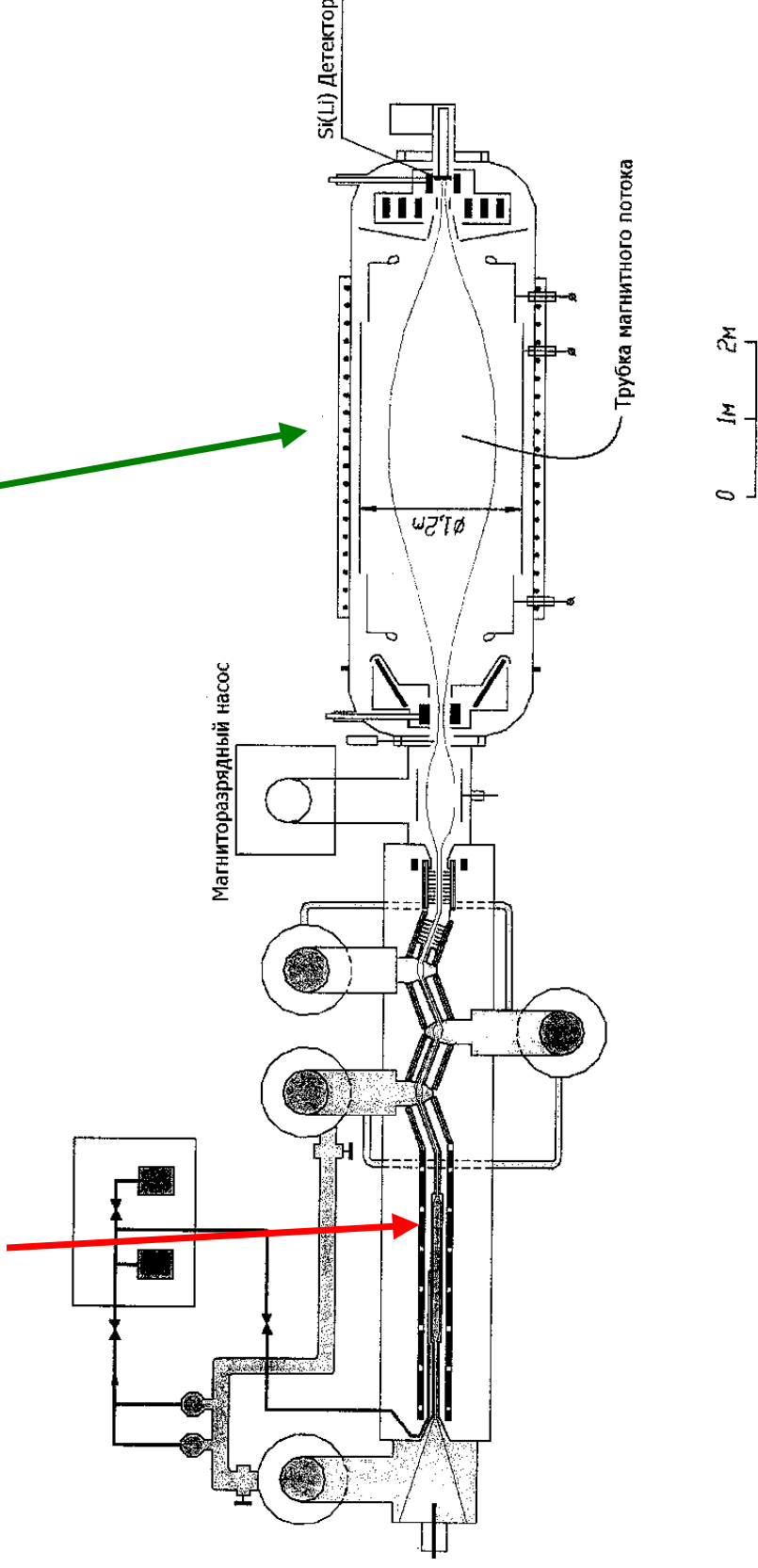
(see poster E2)

- 1998/1999:** $m^2(\nu) = -1.6 \pm 2.5 \pm 2.1 \text{ eV}^2 \Rightarrow m(\nu) < 2.2 \text{ eV (95\% C.L.)}$
- 2001:** $m^2(\nu) = +0.1 \pm 4.2 \pm 2.0 \text{ eV}^2$
- 1998/1999/2001:** $m^2(\nu) = -1.2 \pm 2.2 \pm 2.1 \text{ eV}^2 \Rightarrow m(\nu) < 2.2 \text{ eV (95\% C.L.)}$
- \Rightarrow **Mainz sensitivity limit reached, final analysis of all Mainz data soon**

The Troitsk Neutrino Mass Experiment

Gaseous T_2 source

MAC-E-Filter



column density: 10^{17} cm^{-2}

energy resolution: $\Delta E = 3.5eV$

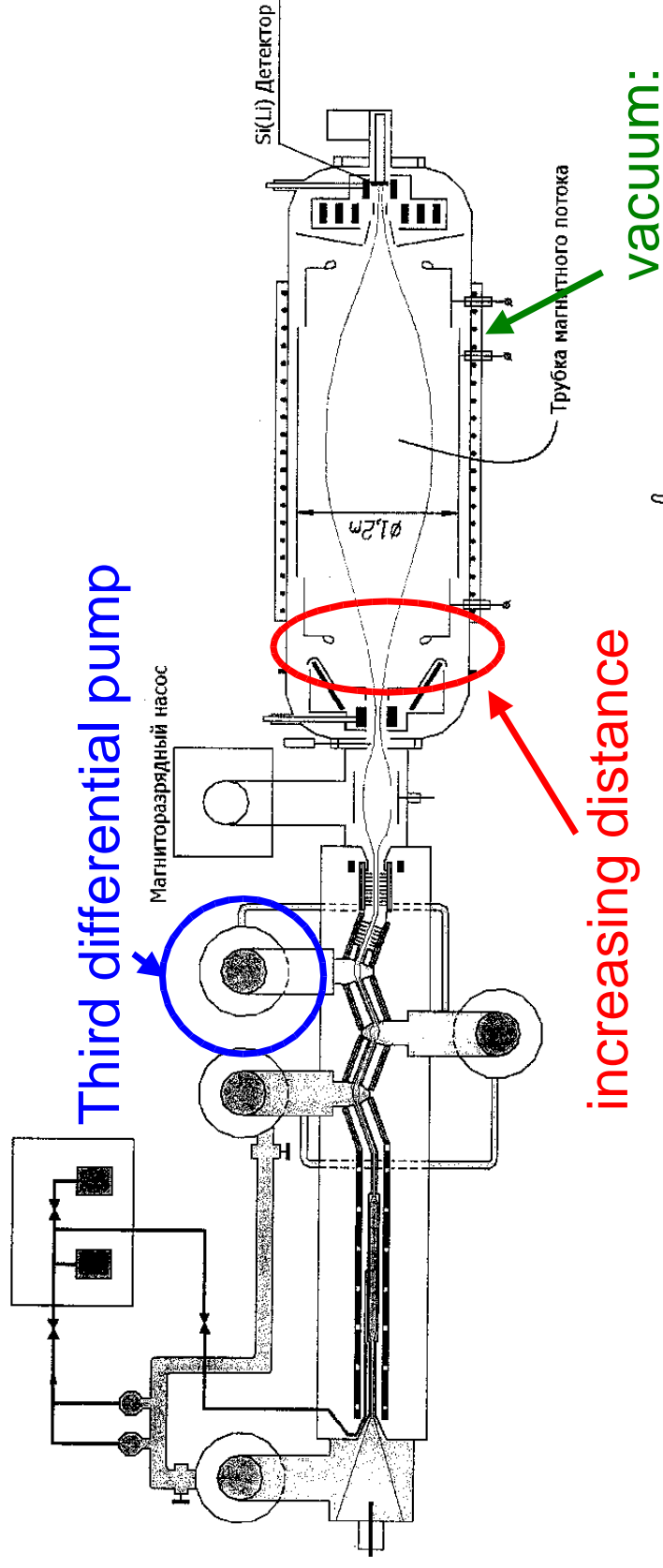
luminosity: $L = 0.6cm^2$

3 electrode system in 1.5m

$$(L = \Delta\Omega/2\pi * A_{\text{source}})$$

diameter UHV vessel ($p < 10^{-9}$ mbar)

Troitsk improvements in 2001

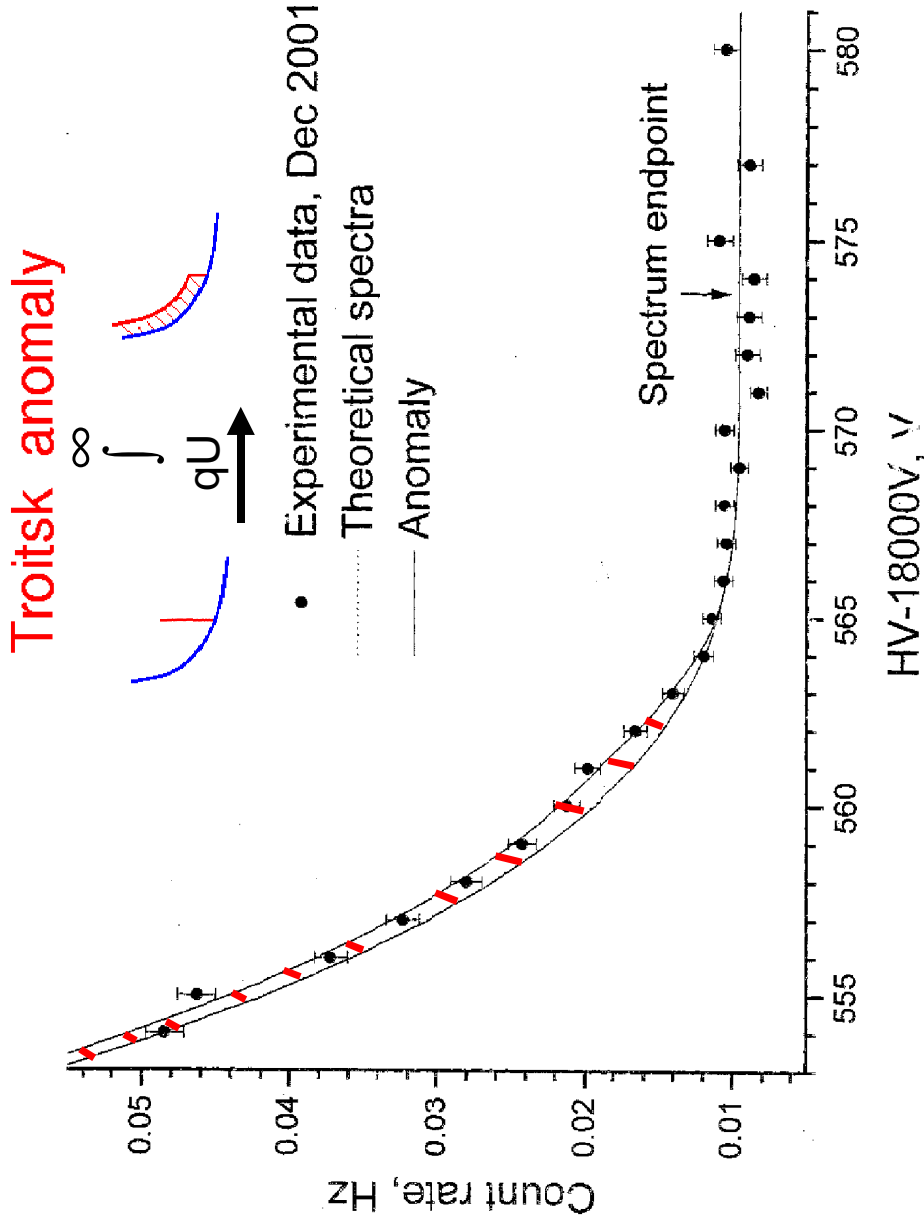


lowering temperature
of spectrometer:

30°C → 13°C

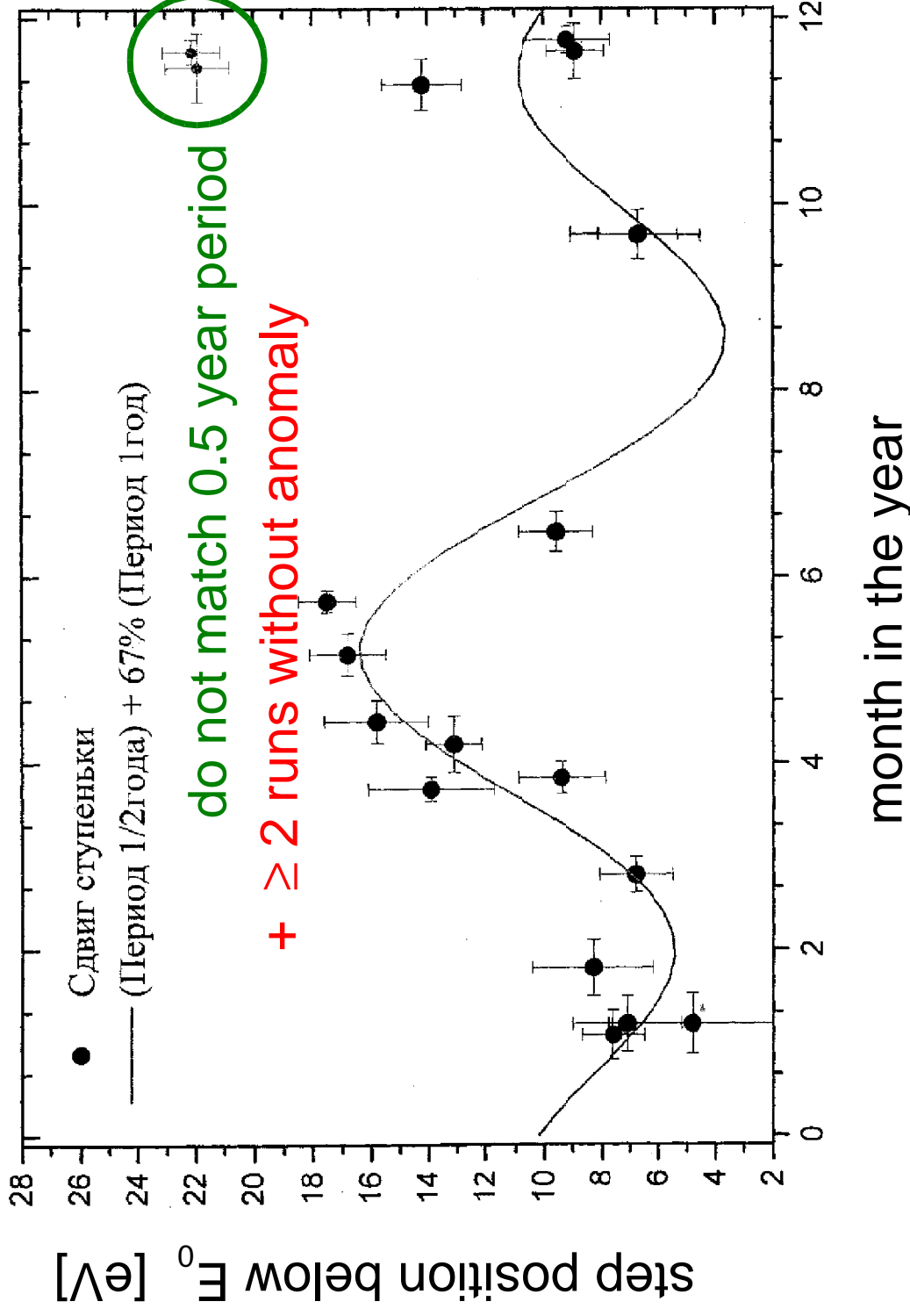
⇒ 3 times lower background
less dependent on tritium source activity

Troitsk anomaly



Troitsk anomaly: monoenergetic line in β spectrum of 10^{-10} amplitude origin ?

Periodicity of Troitsk anomaly



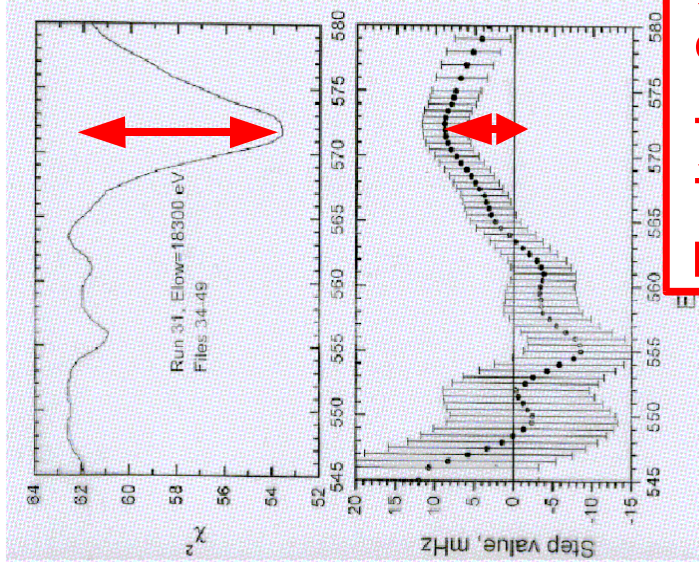
Describing anomaly phenomenologically by additional line, different run-by-run
 Troitsk 1994–1999,2001 data: $m^2(\nu) = -2.3 \pm 2.5 \pm 2.0 \text{ eV}^2 \Rightarrow m(\nu) < 2.2 \text{ eV}$ (95% C.L.)

Coincident measurements Troitsk and Mainz

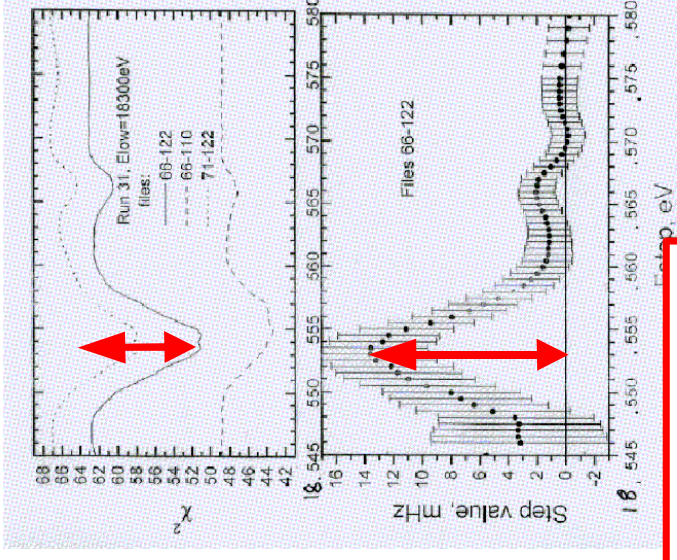
06. – 13. Dec. 2000

22.–28. Dec. 2000

Troitsk

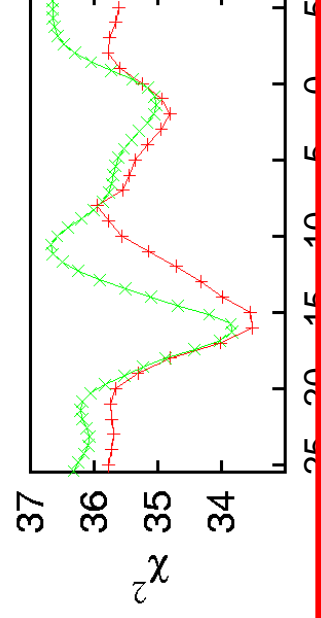
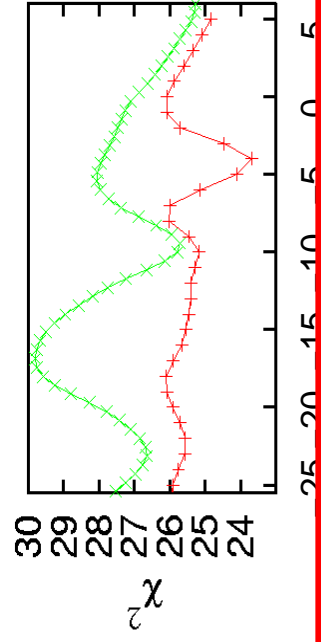


V.M. Lobashev
NANP2001



Troitsk: 2 times sizeable anomaly

Mainz



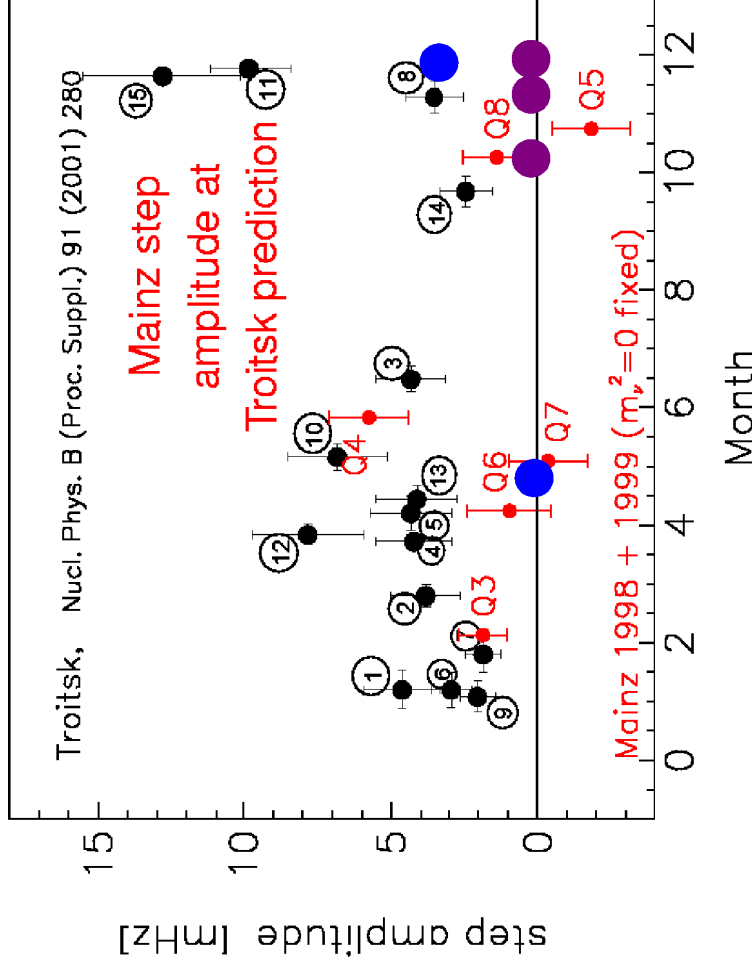
$m_\nu^2 = 0$ fixed

m_ν^2 free

Mainz: no significant change of $\chi^2 \Rightarrow$ no indication of an anomaly

Status of Troitsk anomaly

Amplitude of anomaly: Troitsk, Mainz



Troitsk 2001:

- No anomaly in May 2001
- Only small anomaly in Dec. 2001

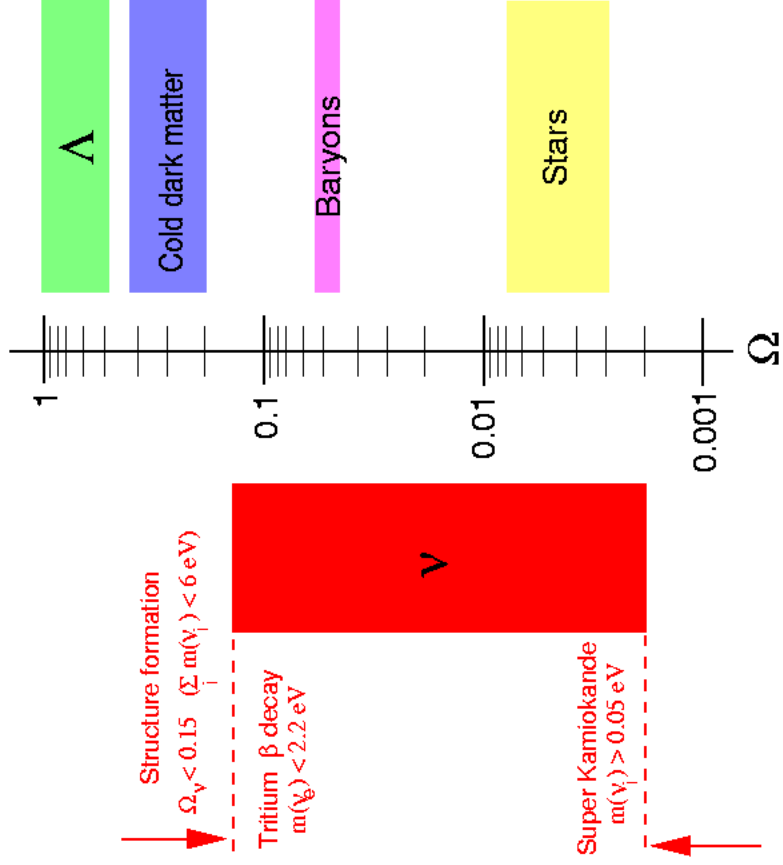
Mainz:

- Clear contradiction to 0.5 y period
- Similiar effect observed only once (Q4 1998)
- Does not show up in in newest Mainz data: of 2000 (Q9,Q10, partially in parallel with Troitsk) and of 2001 (Q11,Q12)

⇒ Troitsk anomaly is very likely experimental artefact, which can be avoided (Mainz)

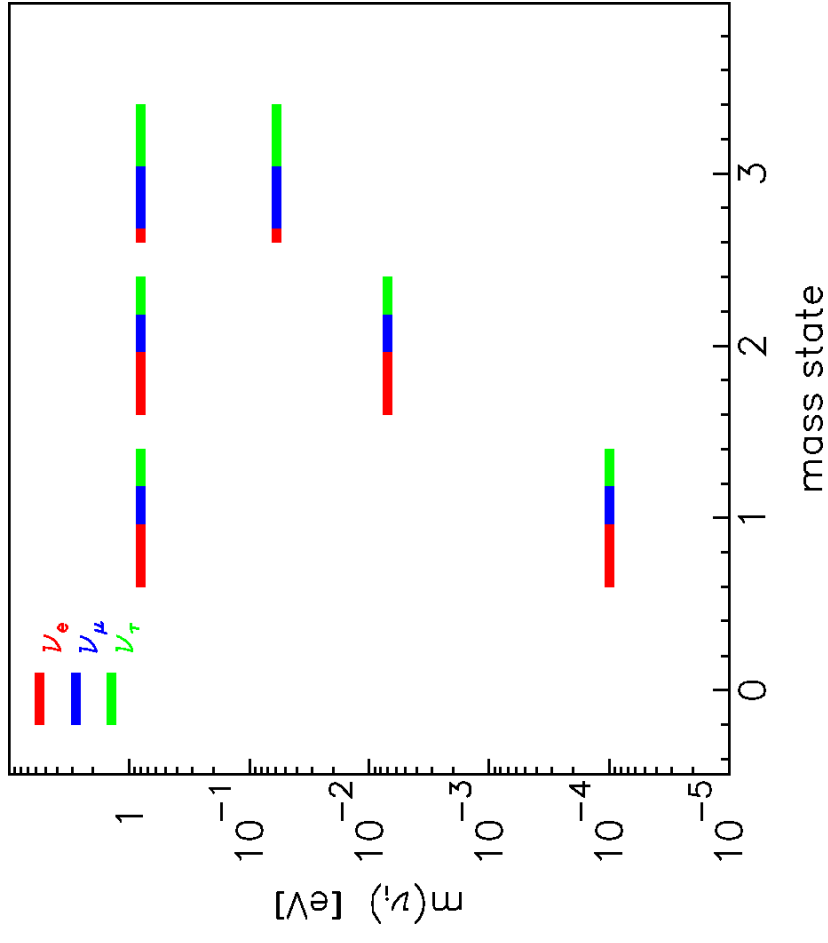
Motivation for a sub-eV direct neutrino mass search

Cosmological relevant neutrino mass



Neutrino mass scale

quasi-degenerate vs hierarchical



What can we learn from β decay w.r.t. $0\nu\beta\beta$?

Direct neutrino mass determination

If neutrino masses are not resolved \Rightarrow average neutrino mass, e.g.:

$$m^2(\nu_e) = \sum |U_{ei}|^2 m^2(\nu_i) \quad (\text{incoherent sum, real average, since } 0 \leq |U_{ei}|^2 \leq 1)$$

Sensitivity of β decay: check degenerate neutrino mass scenarios

$0\nu\beta\beta$ (only possible for Majorana neutrinos)

$$m_{ee}(\nu) = \left| \sum |U_{ei}|^2 e^{i\alpha(i)} m(\nu_i) \right| \quad (\text{coherent sum})$$

\Rightarrow partial cancelation possible (not fully since SNO says: no max solar mixing)

Future experiments very sensitive $m_{ee}(\nu) < 0.1$ eV

$m(\nu_e)$ vs $m_{ee}(\nu)$: complementary information, differences due to:

- Dirac neutrino
- CP-phases
- Other processes (right-handed currents, Susy-particles, ...)
- Problems with nuclear matrix elements

Cryogenic Bolometer Rhenium Experiments

Multi-purpose, scalable new detector technology

Basic idea: β emitting crystal = cryodetector

\Rightarrow single final state: excitation by excited electronic states and inelastic scattering is collected

free choice of β emitter: ^{187}Re : $E_0 = 2.5\text{keV}$ ($t_{1/2} = 5 \cdot 10^{10}\text{y}$)

Current experiments:

MANU2 (F. Gatti et al., Genova)

– Re metallic crystal (1.5 mg)

– BEFS measured (F.Gatti et al., Nature 397 (1999) 137)

– Sensitivity:

current: $m(\nu) < 26\text{ eV}$

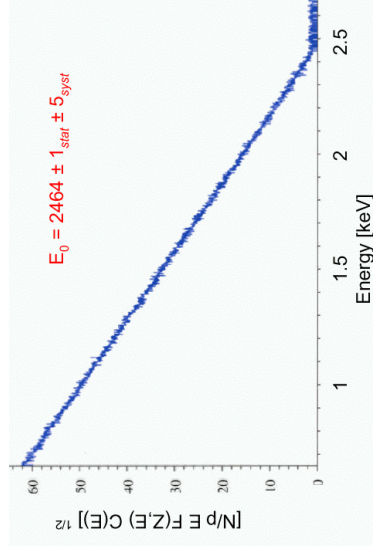
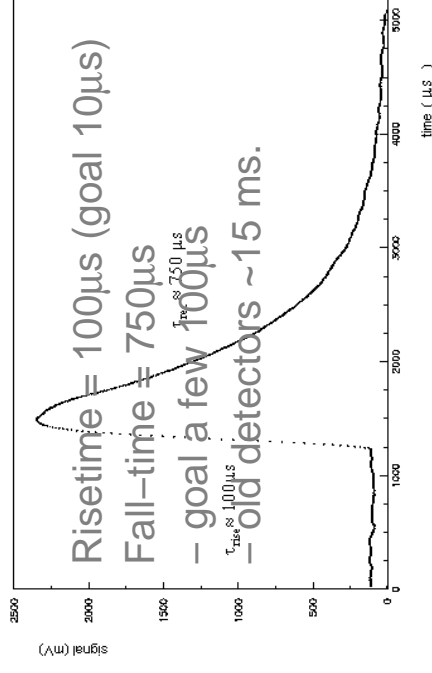
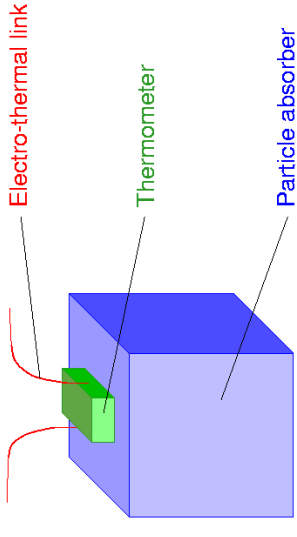
near future: sensitivity of 10 eV expected

future: eV resolution by s.c. sensors

MiBeta (E. Fiorini et al., Milano, Como)

– AgReO_4 (250 – 350 μg)

– Sensitivity: similar to MANU2
(see poster E4)



UTA neutrino rest mass experiment

M. Fink et al., University of Texas, Austin (see poster E5)

Double deflecting electrostatic

deflector aiming for $\Delta E \geq 1\text{eV}$ (FWHM)

Windowless gaseous tritium source

$\rho d \approx 10^{17}/\text{cm}^2$ (as Mainz/Troitsk)

Luminosity: at least one order of magnitude less w.r.t. Mainz/Troitsk

Very low background (cnts/day) ?

Aim: $\Delta m(\nu) = 1.4\text{ eV}$ (3σ) ?

Response function:
no tails beyond Gaussian?

Background from tritium?

Competition for Mainz and Troitsk

2

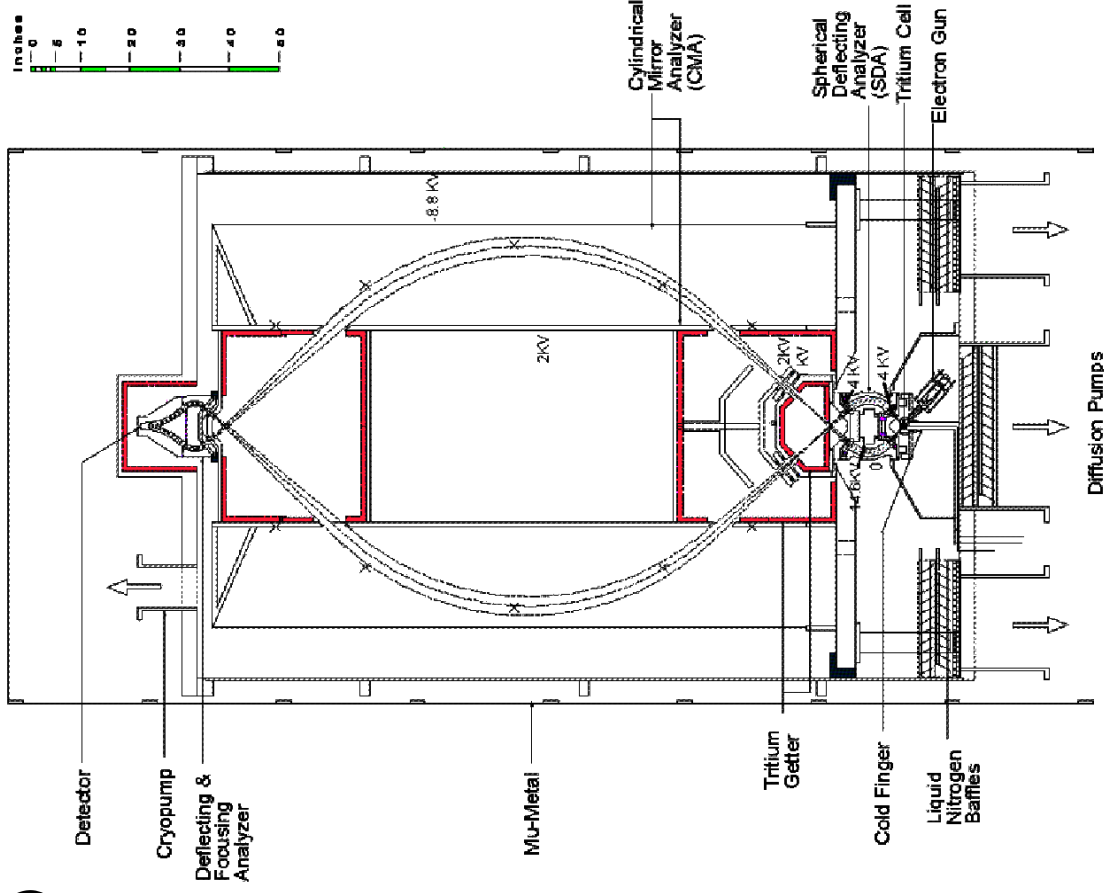


Fig. 1a. The whole electron spectrometers with T₂ cell and detector

Which way to sub-eV neutrino masses?

Neutrinos of galactic supernova (s. talk by J. Beacom)

- galactic SN only every 40 years
- not sensitive below 1 eV (uncertainty in time spectrum of neutrino emission)

Large scale structure (s. talk by S. Hannestad)

- model dependent
- neutrino mass from lab can serve as input for astrophysics

Search for $0\nu\beta\beta$ (s. talk by O. Cremonesi)

Direct neutrino mass determination

- ^{187}Re β decay with cryogenic bolometers
- Tritium β decay with electrostatic deflector (UTA-exp.)
- Tritium β decay with MAC-E-Filter
 - first MAC-E-Filters (Mainz/Troitsk) are very successful
 - no material in beam line, no tails of resolution function
 - quasi „single final state“ experiment
 - also: not-integrating MAC-E-TOF mode

complementary



The Karlsruhe Tritium Neutrino experiment

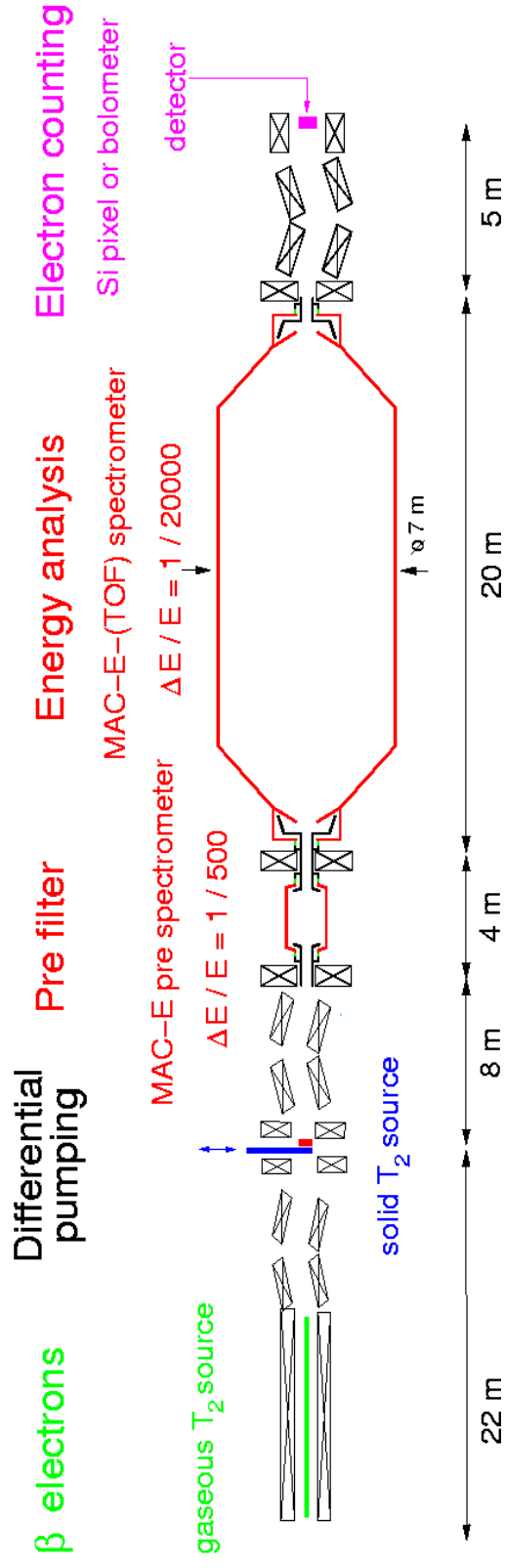
KATRIN

Physics aim:

Sensitivity on neutrino mass scale: $m(\nu) \ll 1\text{eV}$

- **Higher energy resolution: $\Delta E \approx 1\text{eV}$**
since $E/\Delta E \sim A_{\text{spectrometer}}$ \Rightarrow larger spectrometer
- **Relevant region below endpoint is smaller**
even less count rate $dN/dt \sim A_{\text{spectrometer}}$ \Rightarrow larger spectrometer

$\} \text{ } \varnothing 7\text{m}$

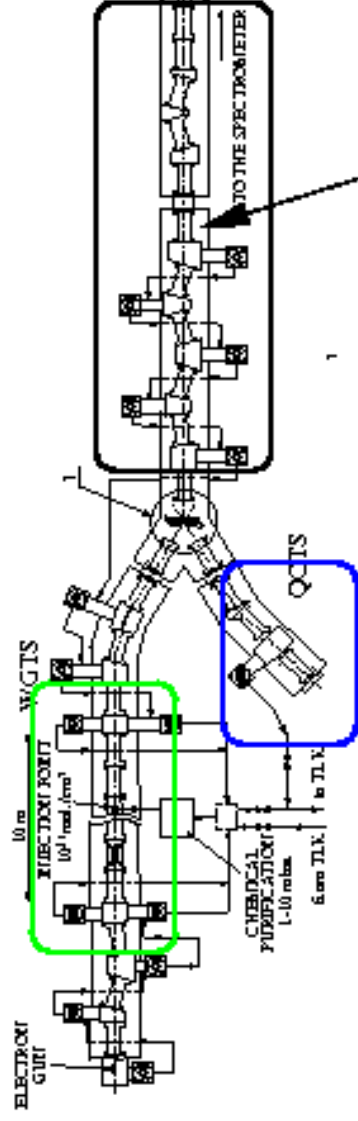


Molecular tritium sources

Windowless gaseous molecular tritium source

(pioneered by LANL (Seattle), adapted to MAC-E by Troitsk)

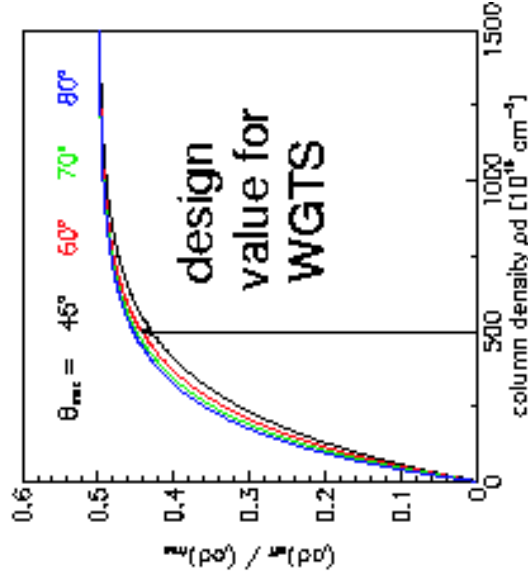
10 m length, 70 mm diameter, $5 \cdot 10^{17} \text{ T}_2 / \text{cm}^2$



Quench condensed tritium source (QCTS) (pioneered by Mainz)

larger but complementary sys. uncert.

WGTS allows almost maximum count rate without large sys. uncertainties:

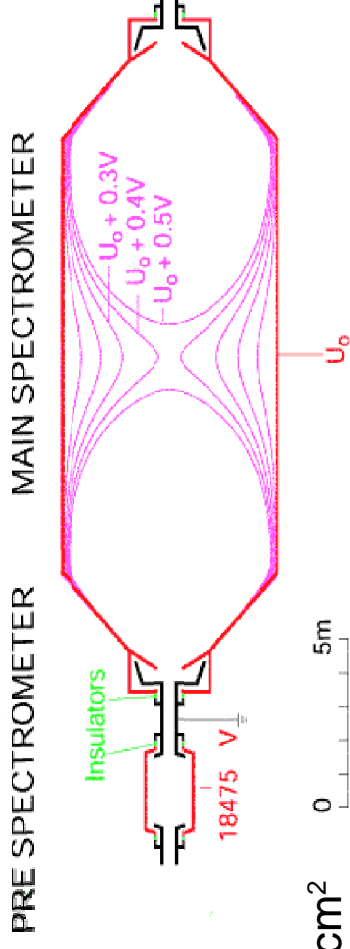


Differential pumping and cryo stations

Pre and main spectrometer

Main spectrometer

- Energy resolution:
 $\Delta E = 1\text{eV}$
- high luminosity:
 $L = A_{\text{Seff}} \Delta\Omega/4\pi = A_{\text{analysier}} \Delta E/(2E) = 10 \text{ cm}^2$
- Ultrahigh vacuum requirements (Background) $p < 10^{-11}$ mbar
- „simple“ construction: vacuum vessel at HV = electrode



Pre spectrometer:

- Transmission of electron with highest energy only
(10^{-7} part in last 100 eV)
 \Rightarrow Reduction of scattering probability in main spectrometer
 \Rightarrow Reduction of background
 - only moderate energy resolution required:
 $\Delta E = 50 \text{ eV}$
 - Test of new ideas (shape of electrodes, removal of trapped particles, ...)
- (see poster E3)

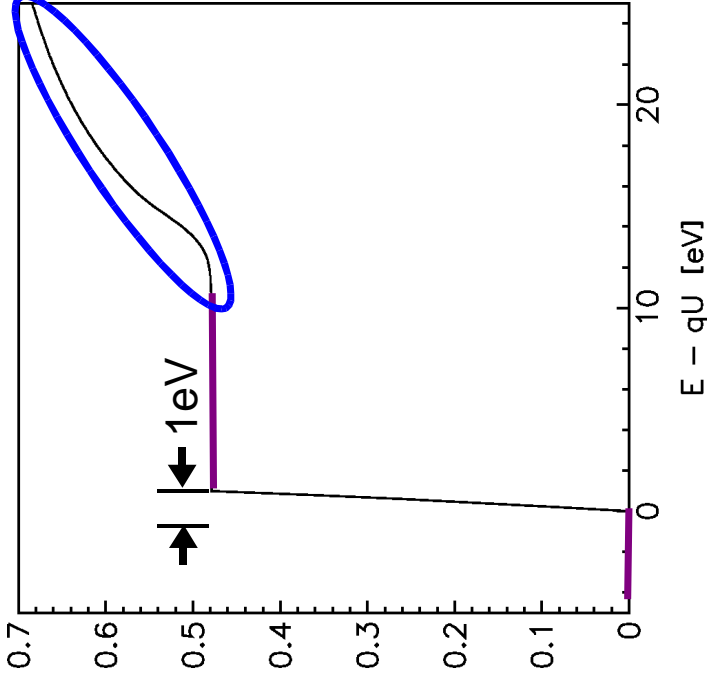
Systematic uncertainties

As smaller $m(\nu)$,
as smaller the region of interest below endpoint

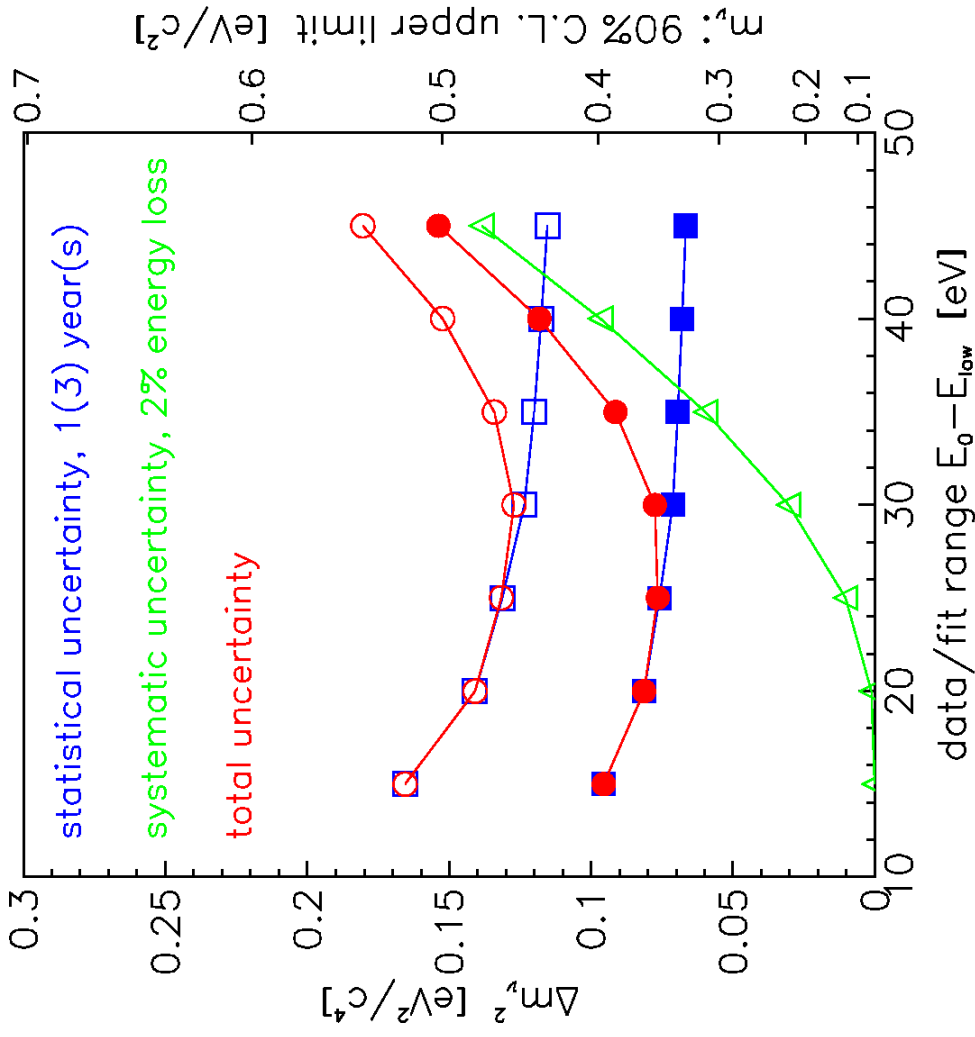
- ⇒ Excited electronic final states does not play a role ($\Delta E_{\text{exc}} > 27 \text{ eV}$)
 - ⇒ Inelastic scattering in T_2 is small ($\Delta E_{\text{inel.}} > 12 \text{ eV} \Rightarrow$ largest interval 25eV: 2%)
 - ⇒ One well-defined final state (similar to cryo detectors)
- Is only true, since MAC-E-Filter response function has no tails

Systematic uncertainties

- Rotation-vibration excitation of final state
- Inelastic scattering (systematic uncertainty: Troitsk 2%, Mainz 6%)
- electrical potential distribution over source
- Solid state effects (for QCTS only)
- Stability of parameters (HV, T_2 partial pressure, T_2 purity, ...)



Estimation of sensitivity



energy resolution: 1eV
 source area: 29 cm²
 gaseous source column density: 5 10¹⁷/cm²
 max accepted starting angle: 51°
 background rate: 11 mHz

First simulations with
 conservative assumptions

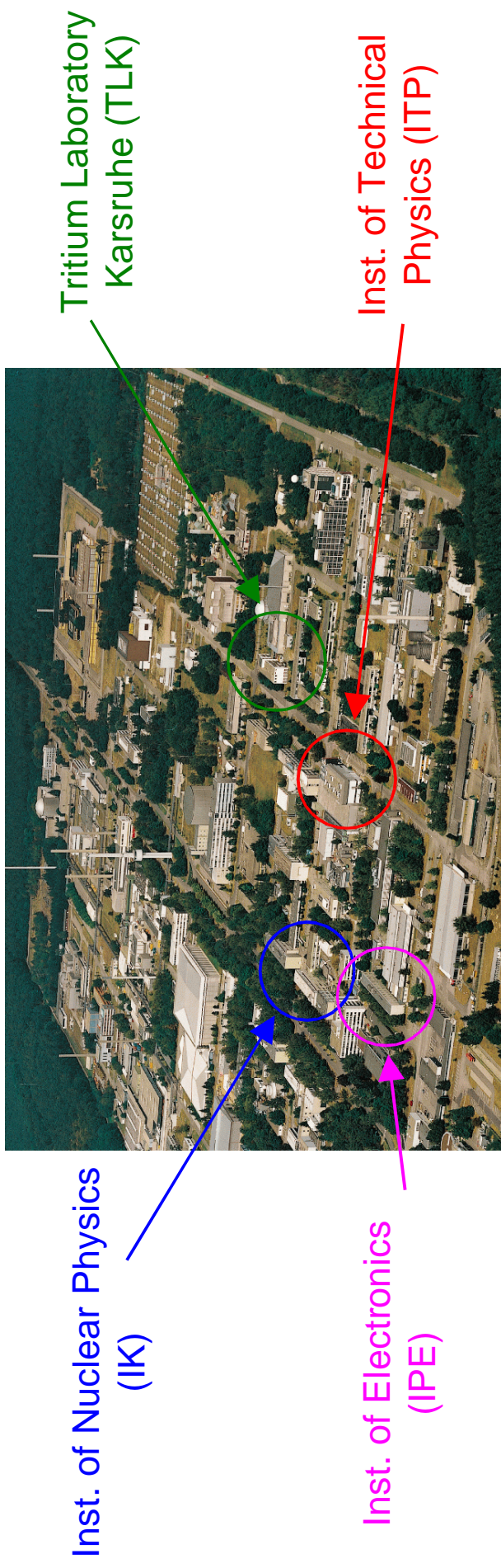
⇒ Sensitivity on $m(\nu_e)$
 ≈ 0.35eV/c²

m_ν : 90% C.L. upper limit [eV/c²]


Technical challenges

- Recirculation and purification of tritium to a large extent (kCi)
- ≈ 30 superconducting solenoids
- UHV ($< 10^{-11}$ mbar) in huge volume (10000m²)
- HV calibration and stability on ppm level
- High resolution detectors
-

⇒ ideal place: Forschungszentrum Karlsruhe/Germany

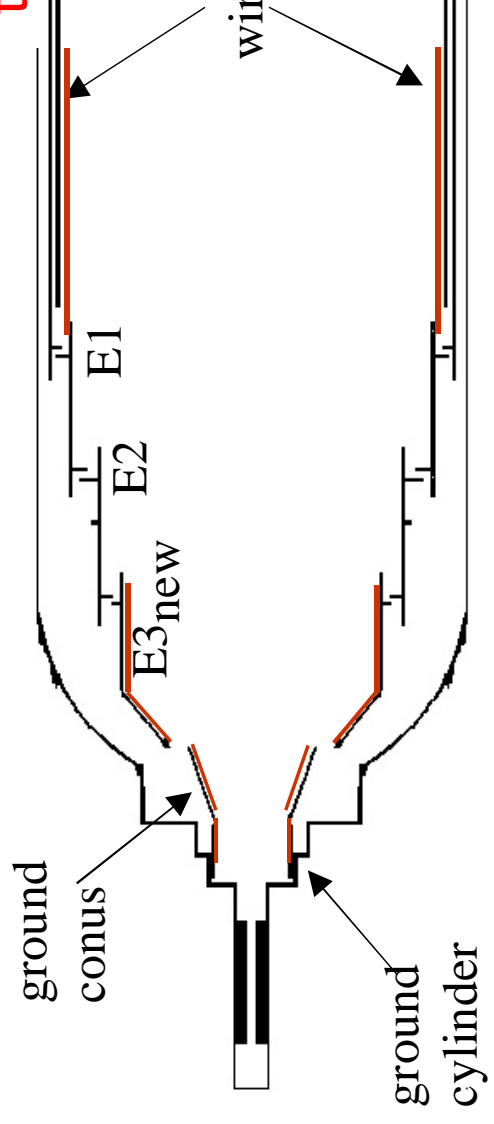
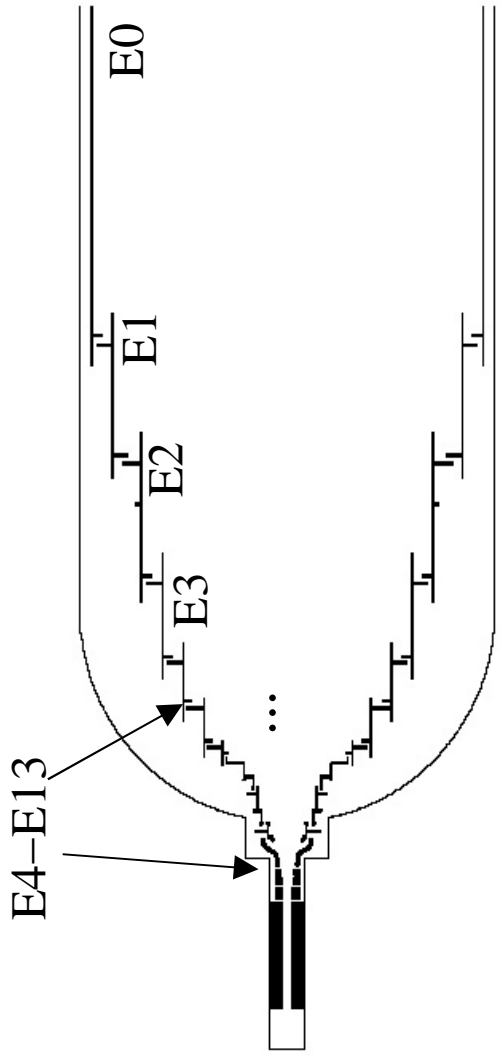


Status and schedule of Katrin

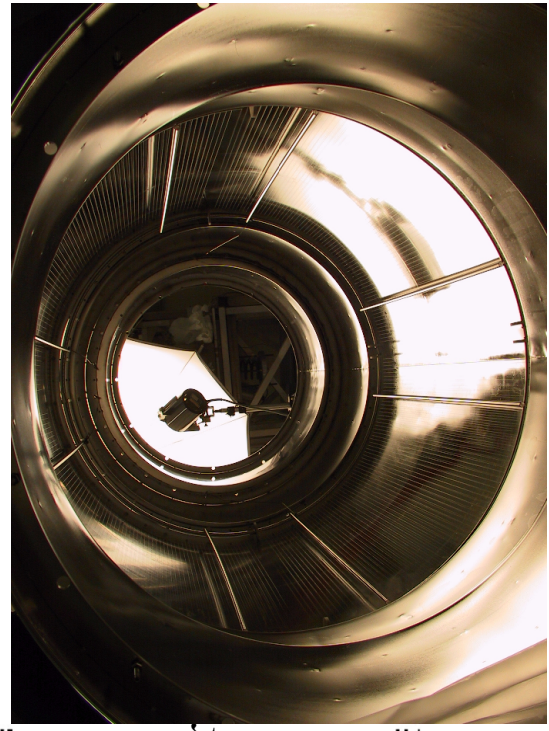
- 1999/2000 First discussions
- 1/2001 **Neutrino Masses in the sub-eV Range**
International workshop on future direct measurements of the electron neutrino mass and their implications
Bad Liebenzell / Schwarzwald , Germany, January 18 – 21, 2001
- 
- transperencies:
<http://www-ik1.fzk.de/tritium/liebenzell>
- 6/2001 KATRIN collaboration founded:
Fulda, Karlsruhe, Mainz, Prag, Seattle, Troitsk + Bonn (12/01)
- 9/2001 Letter of Intent (hep-ex/0109033)
First founding by BMBF/Germany
- 5/2002 International review panel
- 2002 Discussion with other groups to enlarge collaboration
Proposal and application for major funding
Construction of pre spectrometer at FZK
Background investigations at Mainz
- 2007 Start of data taking

Modification of the Mainz Spectrometer 01-05/02

(see poster E1)



electrical dipoles to remove trapped particles by ExB drift



Summary

Current tritium β experiments: Final sensitivity reached

- Mainz: All problems solved by intensive studies of systematics and optimal preparation of experiment
 $\Rightarrow m(\nu_e) < 2.2 \text{ eV}$
- Troitsk: Experimental improvements \Rightarrow lower background anomaly is becoming less stable, probably experimental artefact by subtraction of anomaly run-by-run: $m(\nu_e) < 2.2 \text{ eV}$

How to proceed?

Need future direct sub-eV neutrino mass experiment

- UTA experiment: electrostatic deflector
- Cryogenic detectors with Rhenium: challenging new approach
how far can they go?
- KATRIN: A large tritium β decay neutrino mass experiment at FZ Karlsruhe
with sub-eV sensitivity ($< 0.35 \text{ eV}$)
probes degenerate neutrino masses and cosmologically relevant mass
 \Rightarrow key experiment w.r.t. neutrino mass