

Recap

m_D from $\bar{L} \tilde{\Phi} N_R$ (Dirac mass m_D)

(-) large suppression for all generations

Solution: $\frac{1}{2} M_R \overline{N_R^c} N_R$ with new energy scale M_R

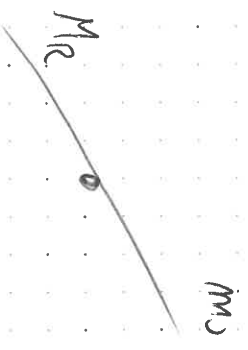
after EWSB: $m_D = \frac{m_D^2}{M_R}$ in $\mathcal{L} = \frac{1}{2} m_D \bar{\nu}_L \nu_L^c$

mixing of light "active" states with heavy "steriles"

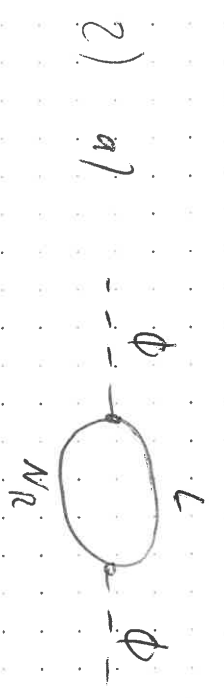
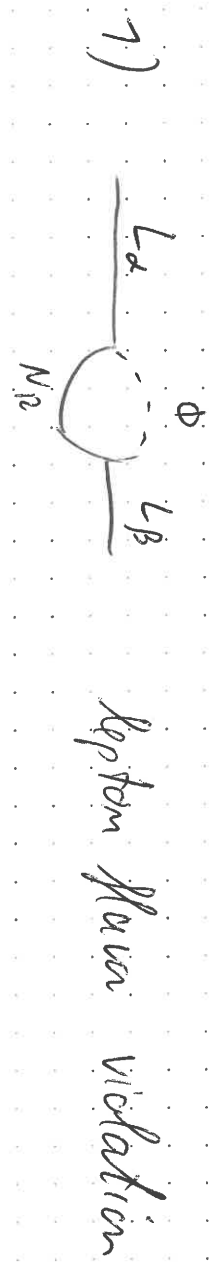
$$\mathcal{D}_{phys} = \cos \theta \nu_L + \sin \theta N_R^c \quad ; \quad \theta \approx \frac{m_D}{M_R} = \sqrt{\frac{m_D^2}{M_R^2}} = \sqrt{\frac{m_D}{M_R}}$$

for $m_D \sim m_{top} : M_R \sim 10^{15} \text{ GeV} \quad \theta \approx 10^{-12}$

$m_D \sim m_e : M_R \sim \text{TeV} \quad \theta \approx 10^{-6}$



How to test seesaw? vertex $\bar{\nu} m_D \tilde{\Phi} N_R$



$$\delta m_\beta^2 = \frac{m_D^2}{16\pi^2} M_R^2 = \frac{m_D M_R^3}{16\pi^2 v^2}$$

$\Rightarrow M_R \lesssim 70^+ \text{ GeV}$ from $\delta m_\beta^2 \leq \text{observed Higgs mass}$

"naturalness"

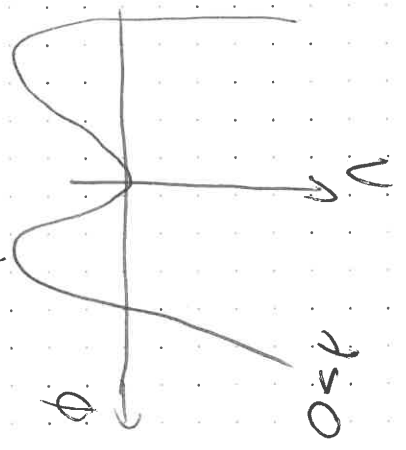
minimizing of λ

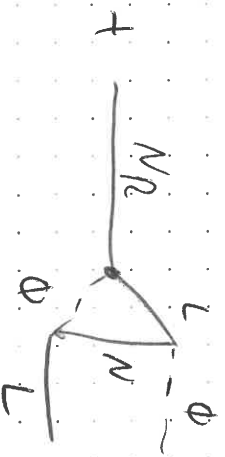
$$\lambda = -c \frac{m_D^4}{v^2}$$

λ becomes smaller with

increasing energy

"vacuum stability"





" leptogenesis "

mass) create mass
matter from
antimatter
if m_D is complex

Model-independent Effective Theories

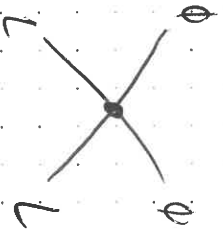
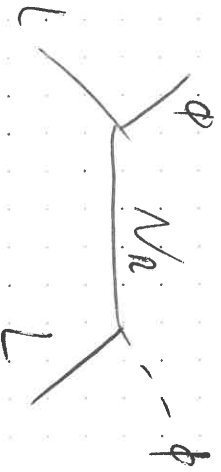
write terms constructed out of SM fields with dimension ≥ 4

$$\left(\rightarrow \text{Fermi } GF \bar{l}_1 \gamma^\mu l_2 \bar{l}_3 \gamma^\mu l_4 \quad h_i \bar{l}_i \gamma^\mu l_i \quad \text{Row } E \quad \times \right)$$

there is only one term of dim 5

$$\mathcal{L} = \frac{1}{\Lambda} \bar{l}_2 \gamma^\mu l_1 \tilde{\Phi} \tilde{\Phi}^T l_3^c \xrightarrow{\text{EWSB}} \frac{1}{2} \gamma^\mu \frac{v^2}{\Lambda} \bar{l}_2 l_3^c$$

Weinberg operator



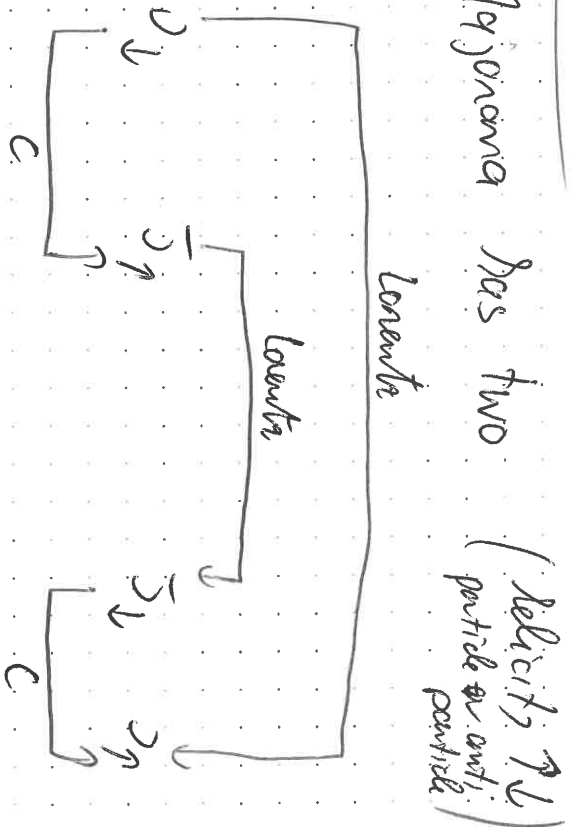
$$M_R = \Lambda$$

How to distinguish Dirac from Majorana?

Dirac has 4 degrees of freedom; Majorana has two (Relicity ↑↓
particle or anti-particle)



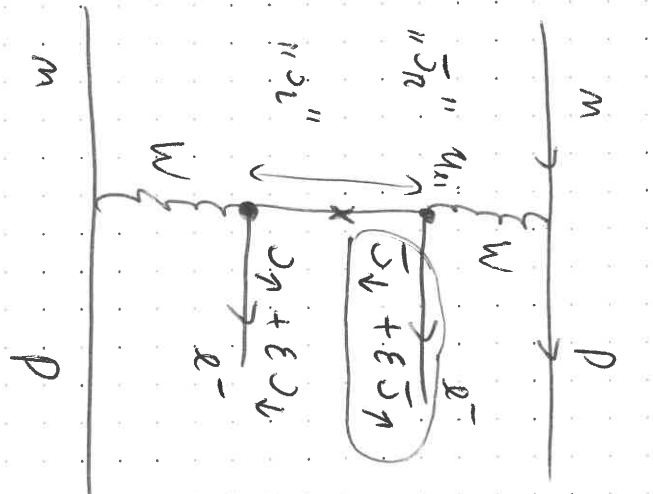
$$D_M = (D \downarrow, U \uparrow)$$



$$D_D = (D \uparrow, U \downarrow, \bar{D} \uparrow, \bar{U} \downarrow)$$

In SM, observable difference between D and M always suppressed:
 ↳ are highly relativistic and Relicity ≠ Chirality

$$E.g. \quad M \downarrow = M_L^{(m=0)} + \frac{m}{2E} M_R^{(m=0)} \quad \Rightarrow \quad P_R M \downarrow \approx m/E$$



"neutrinoless double beta decay" $(0\nu\beta\beta)$

$$(A, Z) \rightarrow (A, Z+2) + 2e^-$$

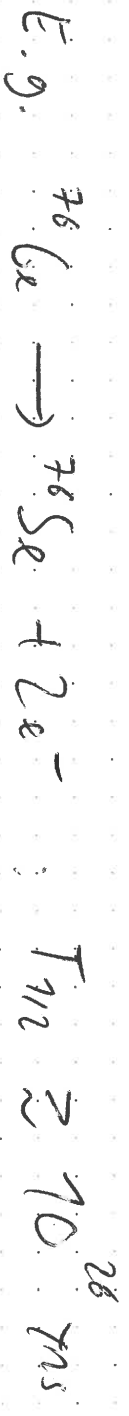
$$\epsilon \sim \frac{m^2}{E^2}$$

works if $\nu = \bar{\nu}$; absorbs ν from upper vertex in lower vertex

rate suppressed by $\left(\frac{m}{E}\right)^2$

(example sum: $\nu_e = \nu_n + \bar{\nu}_p$: antineutrino rate $\propto E^2$)

Avogadro beats $\left(\frac{m}{E}\right)^2$ suppression!



$$\Gamma(0\nu\beta\beta) \propto \frac{1}{E^2} \sum_i |U_{ei}|^2 m_i^2$$

with $E^2 \sim q^2 = \left(\frac{1}{\text{fm}}\right)^2 \sim (100 \text{ MeV})^2$

with $\left| \sum_i U_{ei}^2 m_i^2 \right| = m_{\text{eff}}^2$ effective mass

to be precise: $\Gamma(0, \beta\beta) = G(Q, Z) | \mathcal{N}(A, Z) |^2 m_{ee}^2$

Phase space factor

nuclear matrix element

particle physics

(factor 2-3 uncertainty)

Technical details: Majorana nature leads to new propagators!

usual propagator: $\langle 0 | T \{ \psi(x) \bar{\psi}(y) \} | 0 \rangle = S(x-y)$

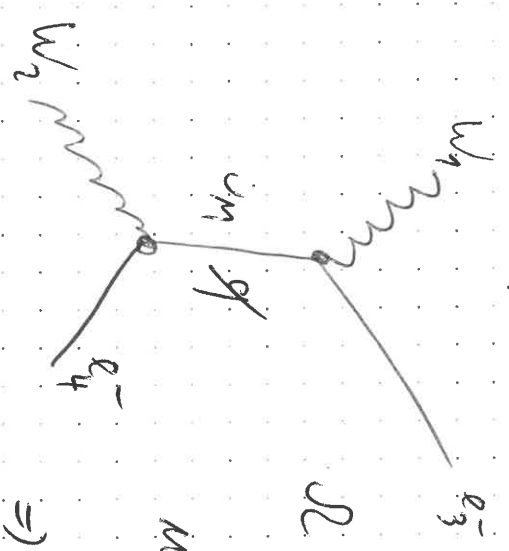
$$= -i \int \frac{d^4 p}{(2\pi)^4} \frac{\not{p} + m}{p^2 - m^2 - i\epsilon} e^{-ip(x-y)}$$

$$\langle 0 | T \{ \psi(x) \psi(y) \} | 0 \rangle = -S(x-y) C$$

$$| \langle 0 | T \{ \bar{\psi}(x) \bar{\psi}(y) \} | 0 \rangle = C^{-1} S(x-y) |$$

(-) production of two particles or antiparticles

on mass dependence comes from $\mathcal{L} = \frac{-g}{2\sqrt{3}} \bar{e} \gamma^{\mu} (1-\gamma_5) \nu_{\mu} W_{\mu}$



$$\mathcal{L} \propto \bar{e}_3 \gamma_{\mu} (1-\gamma_5) \nu_{\mu} \delta_5 (1-\gamma_5) \nu_{\mu}$$

use identity $\bar{e} \gamma_{\mu} (1-\gamma_5) \nu = -\nu^c \gamma_{\mu} (1+\gamma_5) e^c$

$$\Rightarrow \mathcal{L} \propto \bar{e}_3 \gamma_{\mu} (1-\gamma_5) \nu_{\mu} \delta_5 (1+\gamma_5) e_4^c$$

propagator

$$e_4^c \propto \gamma_{\mu} (1-\gamma_5) (q+m) \delta_5 (1+\gamma_5)$$

$$= \gamma_{\mu} (1-\gamma_5) m \delta_5$$

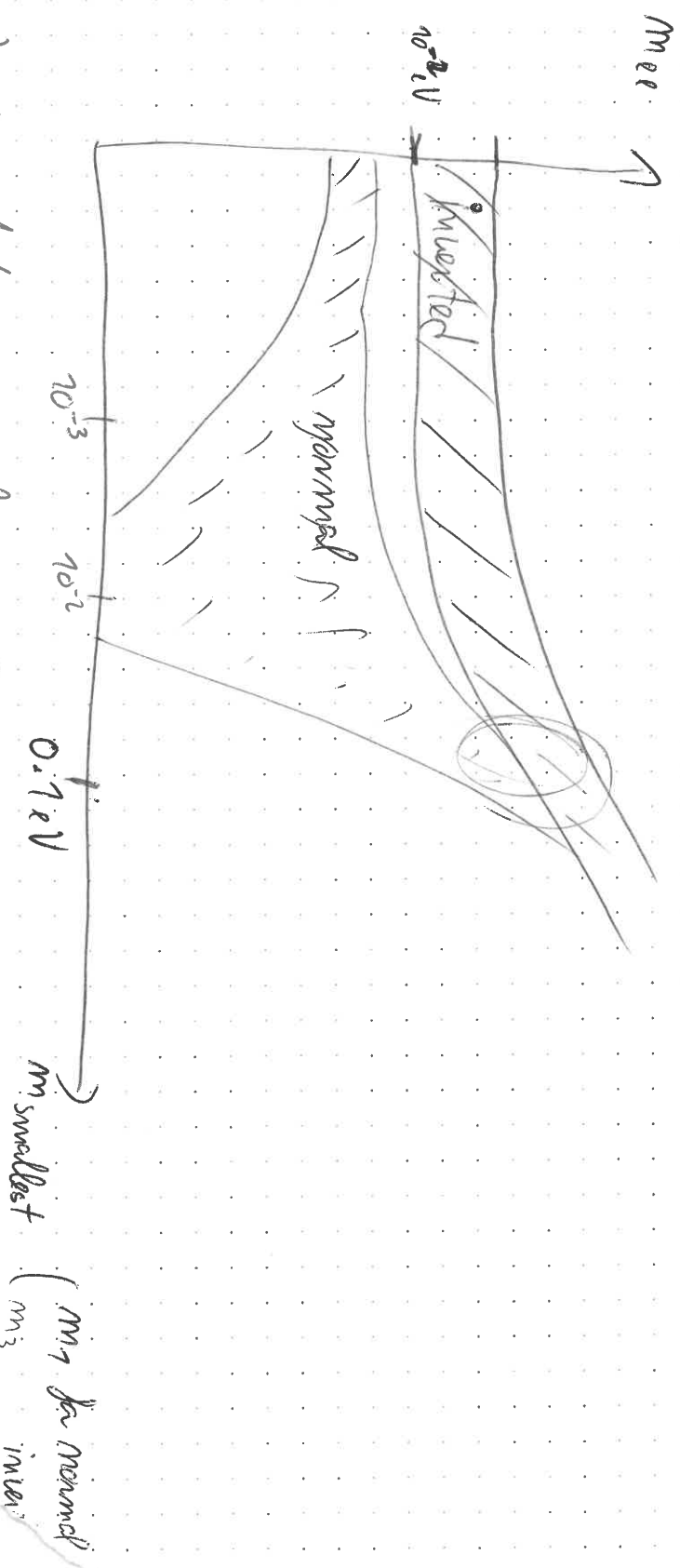
effective mass:

$$m_{ee} = \left| \sum \chi_{ei}^2 m_i^2 \right|$$

coherent sum

$$= \left| \sum |U_{ei}|^2 e^{i\phi_i} m_i \right|$$

(two new phases for Majorana neutrinos... mass term $\nu\nu$...)



1) inverted: m_{ee} always $\neq 0$

2) normal: m_{ee} can be zero

(m_1 for normal, m_3 inverted)

$T_{112} \approx 10^{26}$ yrs $\Rightarrow m_{ee} \leq 0.2$ eV $\Rightarrow m_{smallest} \leq 0.6$ eV
 vs. cosmology, $\sum m_i \leq 0.5$ eV $\Rightarrow m_{smallest} \leq 0.75$ eV
 $m_p \leq 0.8$ eV
 $m_{smallest} \leq 0.8$ eV

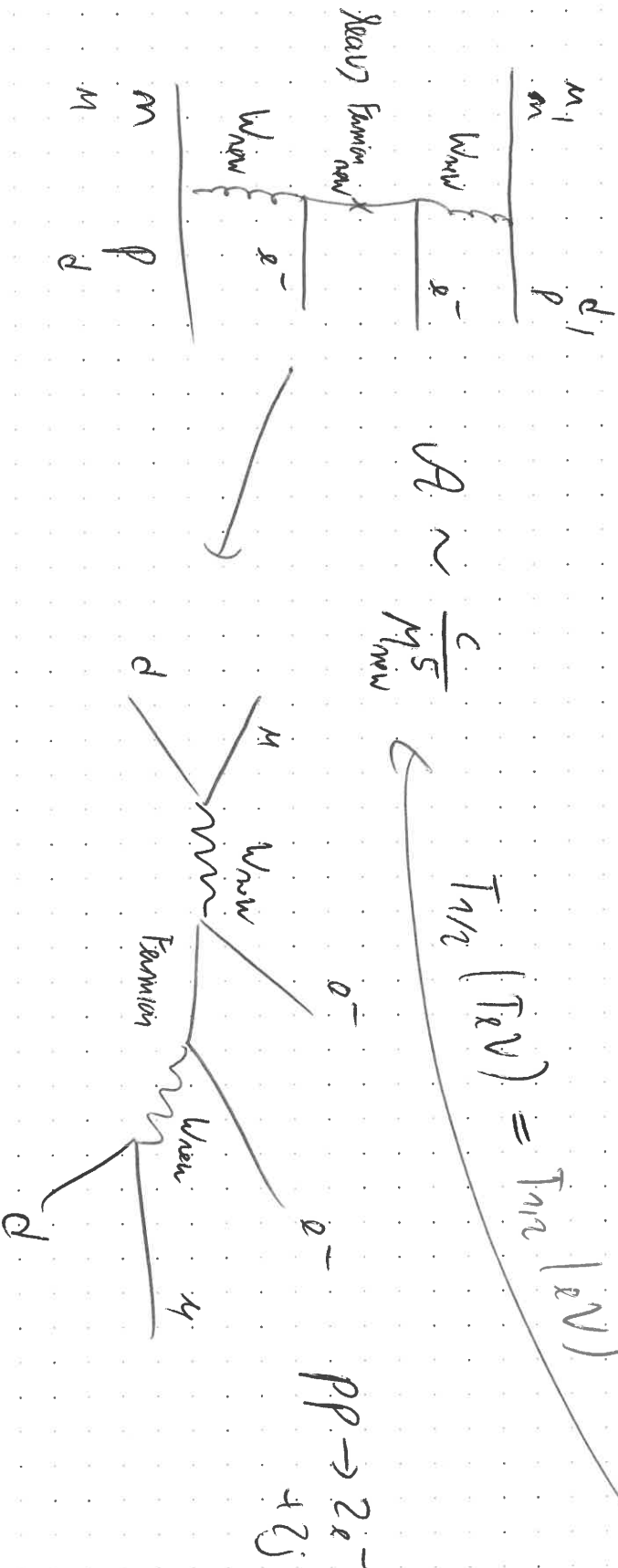
New Physics in D, B, B_s

Standard $A \approx G_F^2 \frac{m_{cc}^2}{q^2} \sim 7 \cdot 10^{-18} \left(\frac{m_{cc}}{0.5 \text{ GeV}} \right) \text{ GeV}^{-5} \sim (\text{TeV})^{-5}$

$A \sim \frac{c}{M^2}$

$T_{1/2}(\text{TeV}) = T_{1/2}(\text{eV})$

$pp \rightarrow 2e^- + \gamma$



Black-Box-Theory

