

Top mass from asymptotic safety

Astrid Eichhorn
University of Heidelberg

recent work with:

Aaron Held ([1705.02342](#), [1707.01107](#))

& Jan Pawłowski ([1604.02041](#); Phys.Rev. D94 (2016) no.10, 104027)

Nic Christiansen ([1702.07724](#); Phys.Lett. B770 (2017) 154-160)

Stefan Lippoldt ([1611.05878](#); Phys.Lett. B767 (2017) 142-146)

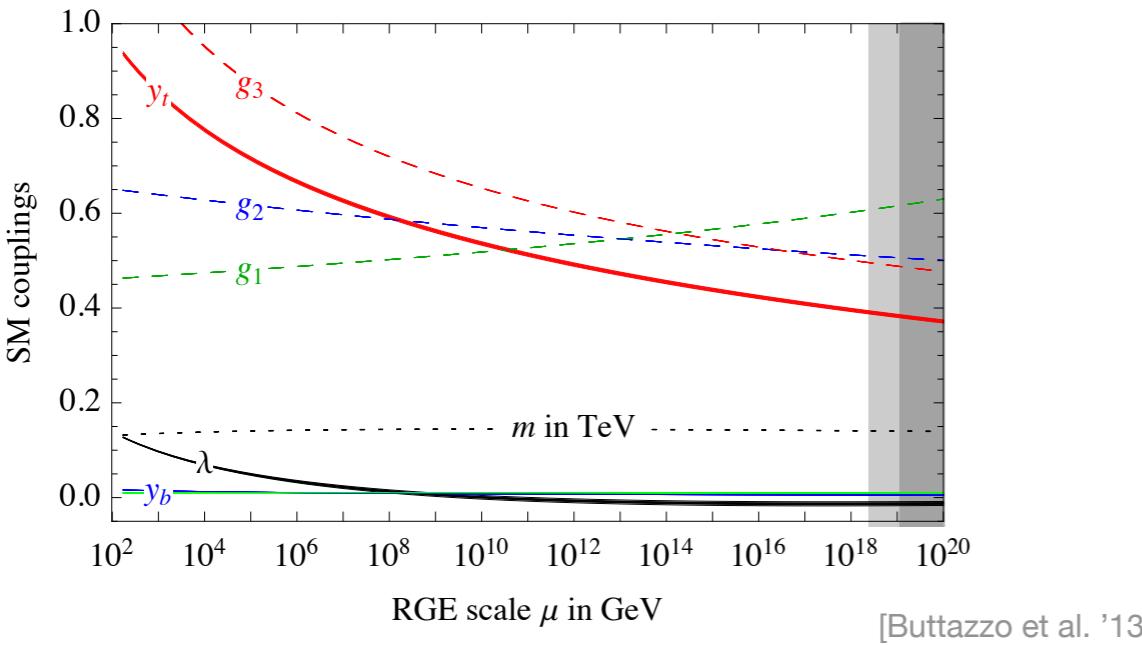


July 24, 2017

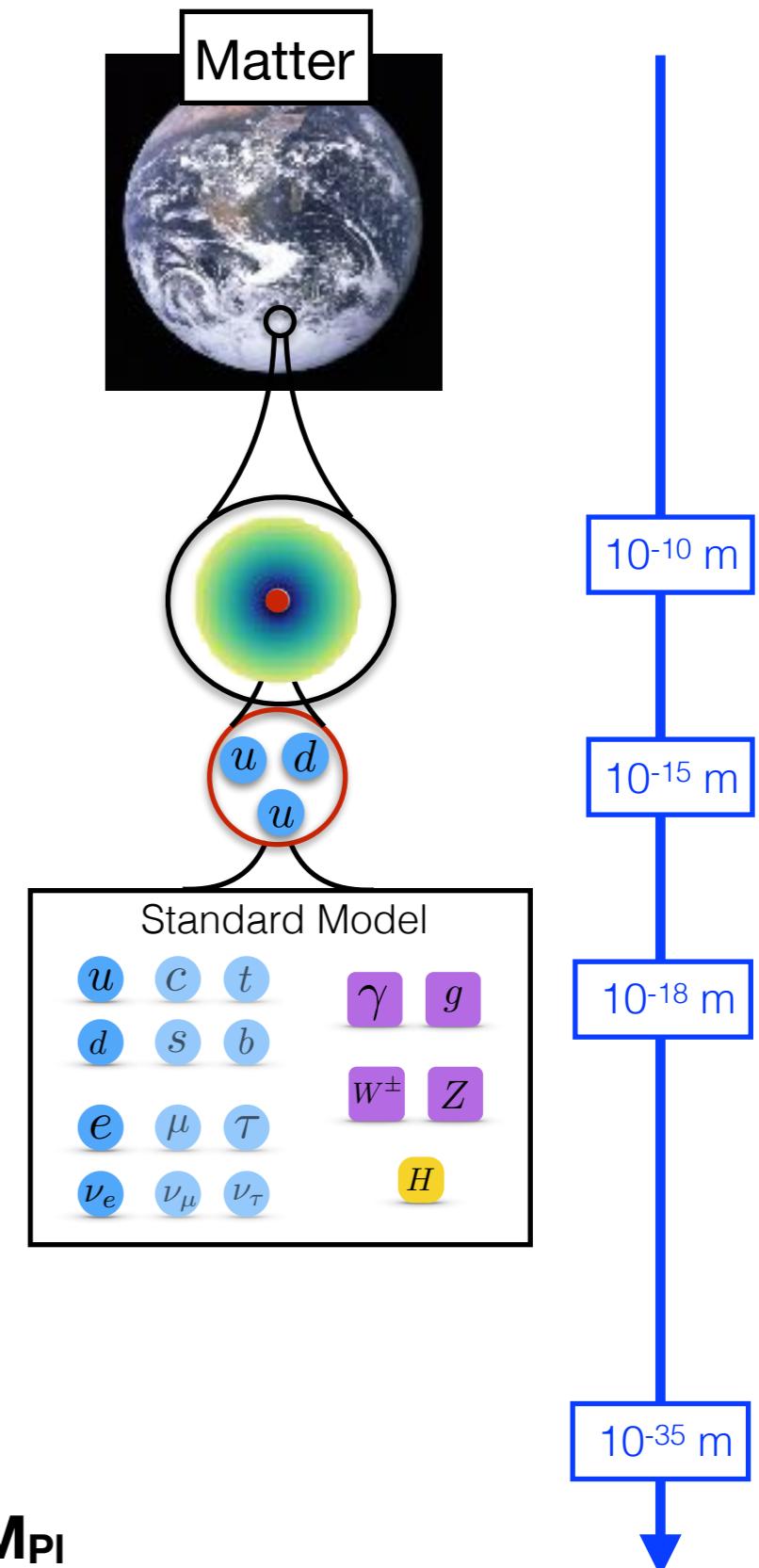
Particle and Astroparticle Theory Seminar
Max-Planck Institut für Kernphysik , Heidelberg

What are the fundamental building blocks of our universe?

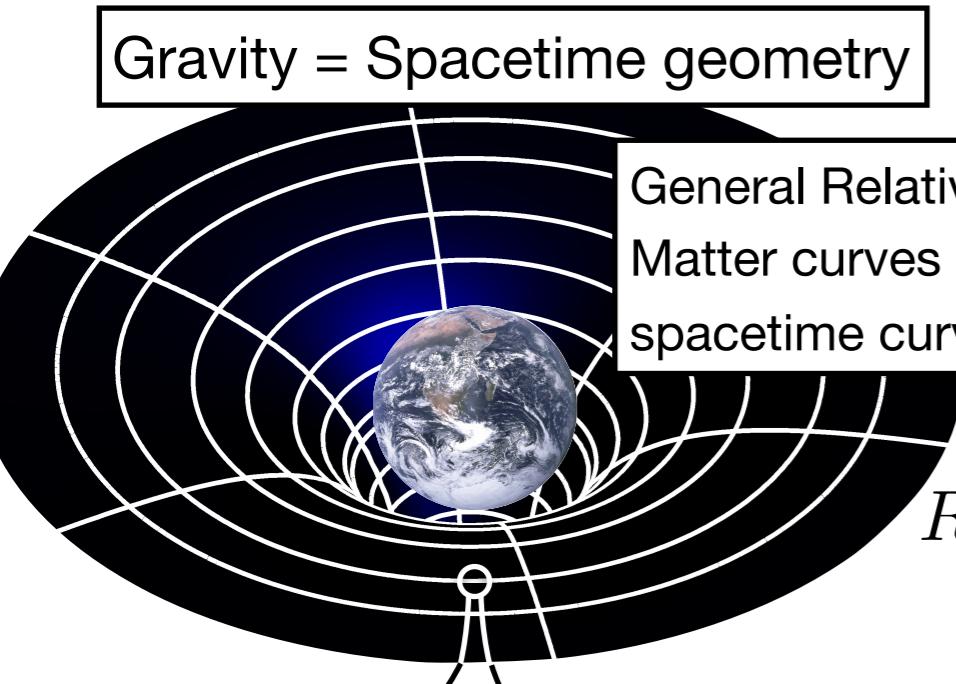
Status of the Standard Model of particle physics



- can extend up to M_{Pl} (@ $M_h \sim 125$ GeV)
- breakdown in transplanckian regime
(Landau pole/ triviality problem in
Abelian hypercharge & Higgs-Yukawa)
- fails to include quantum gravity
- 19 free parameters
 - highly successful effective field theory
 - new physics (quantum gravity?) required beyond M_{Pl}



What are the fundamental building blocks of our universe?

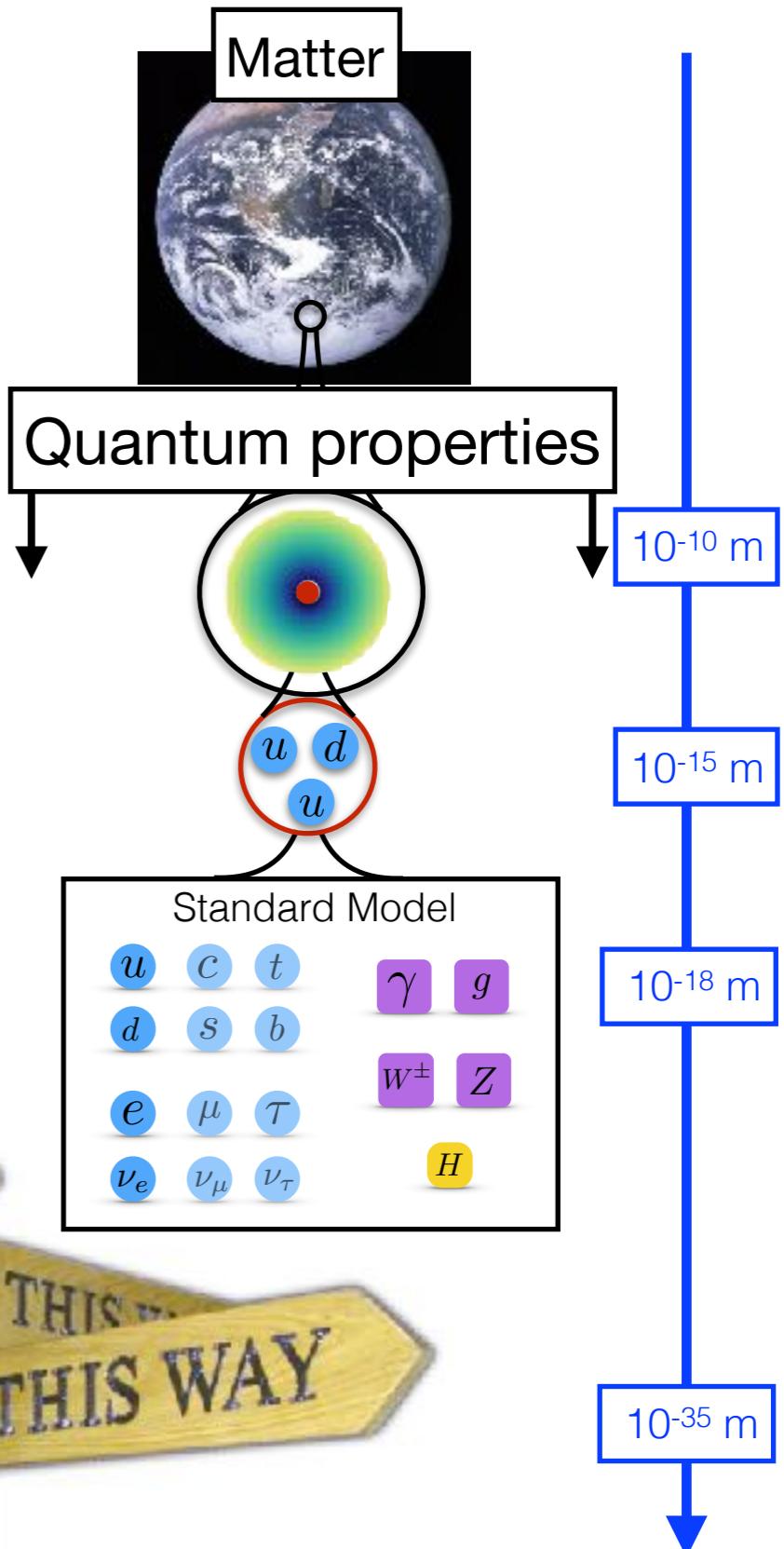
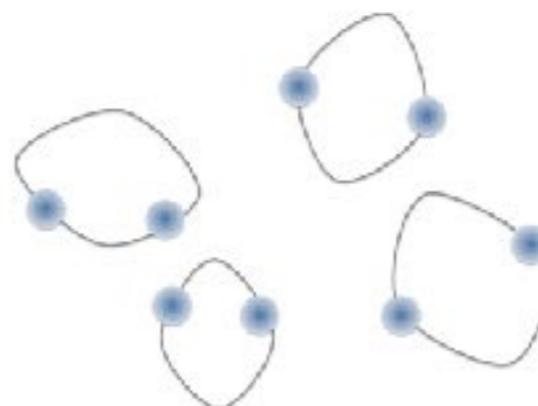


General Relativity:
Matter curves spacetime &
spacetime curvature tells matter how to move

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi G_N T_{\mu\nu}$$

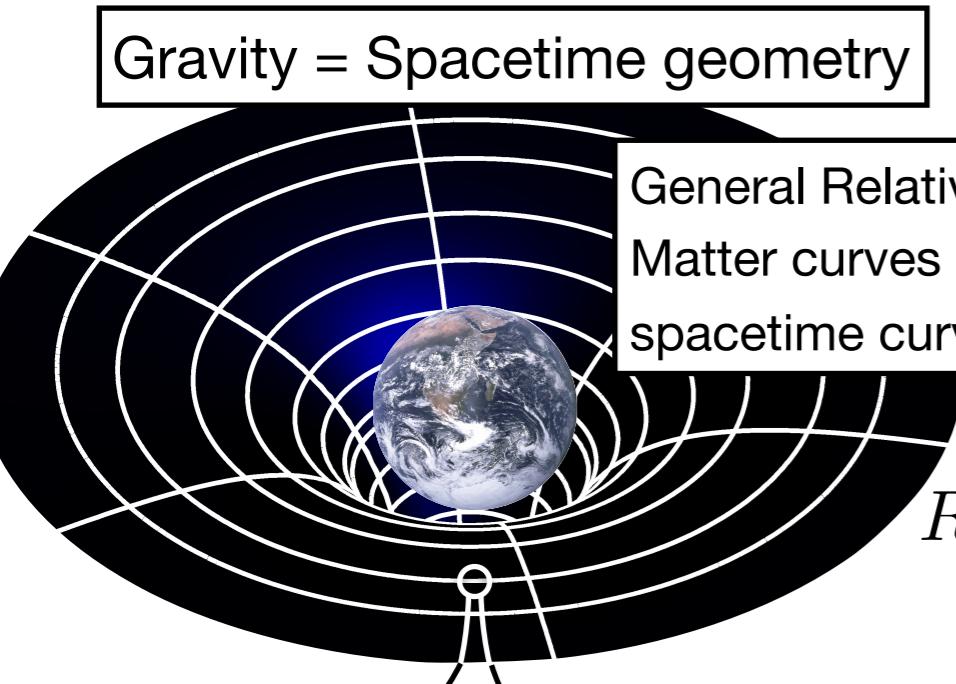
Quantum properties

- string theory
- supergravity
- asymptotically safe gravity
- loop quantum gravity
- causal sets
- non-commutative spacetime



Which way to quantum gravity?

What are the fundamental building blocks of our universe?

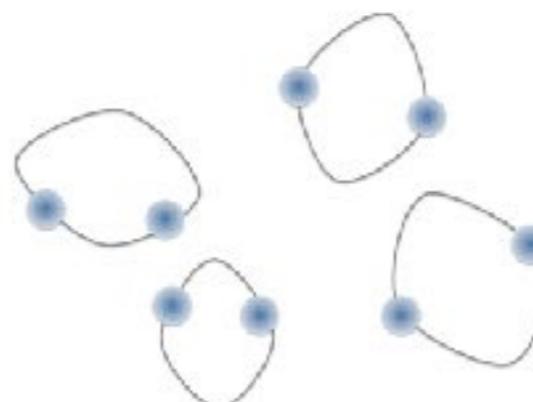


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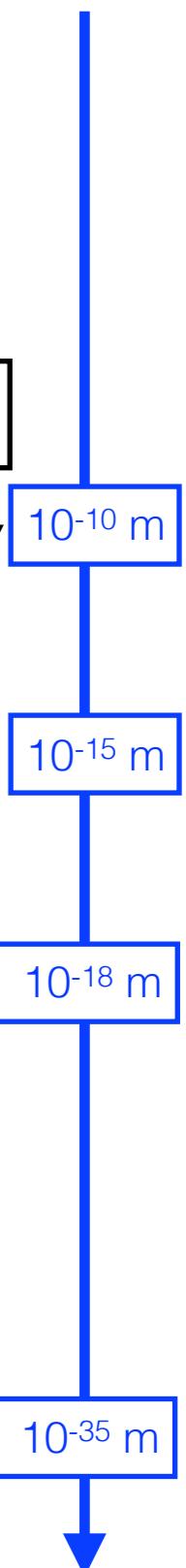
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Quantum properties

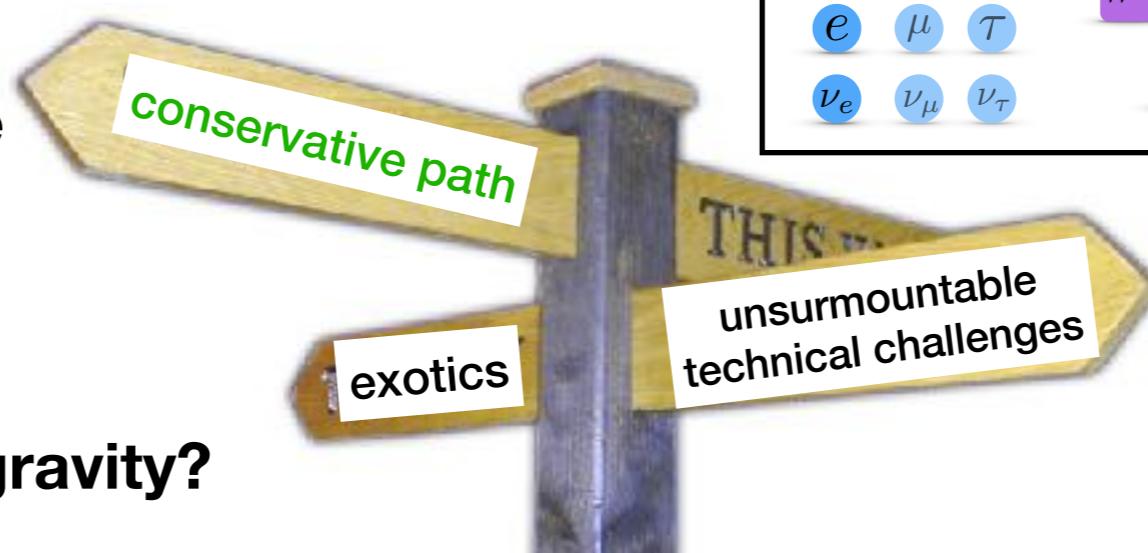
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Quantum properties

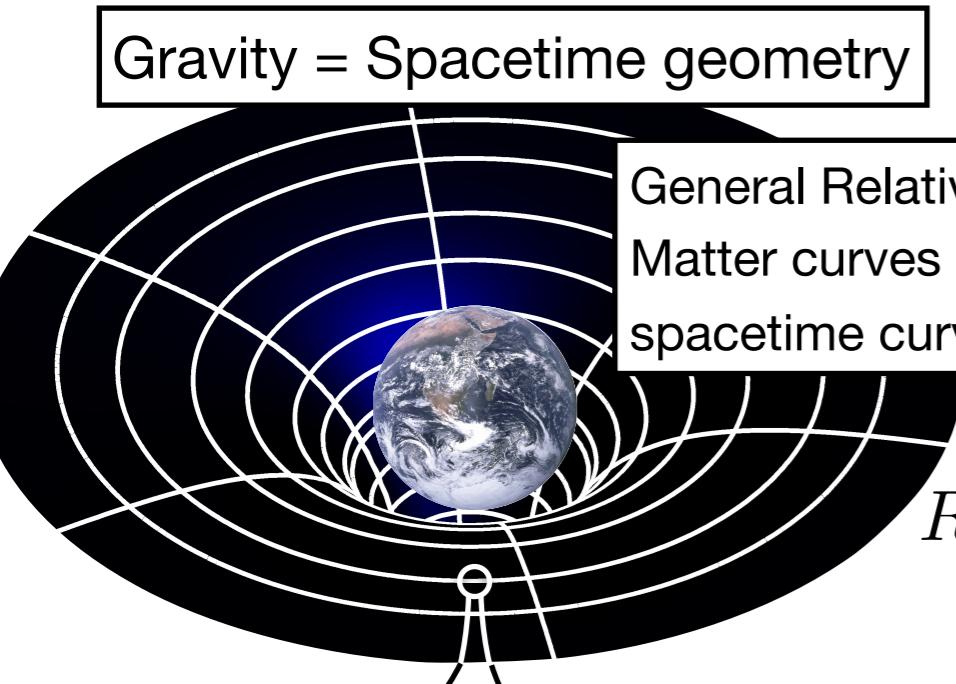


Standard Model					
u	c	t	γ	g	
d	s	b	W^\pm	Z	
e	μ	τ			H
ν_e	ν_μ	ν_τ			



Which way to quantum gravity?

What are the fundamental building blocks of our universe?

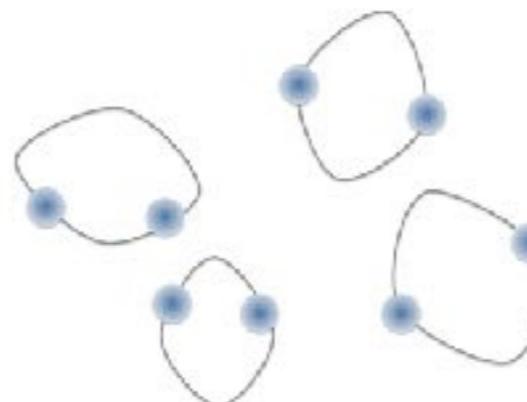


Gravity = Spacetime geometry

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Matter curves spacetime &
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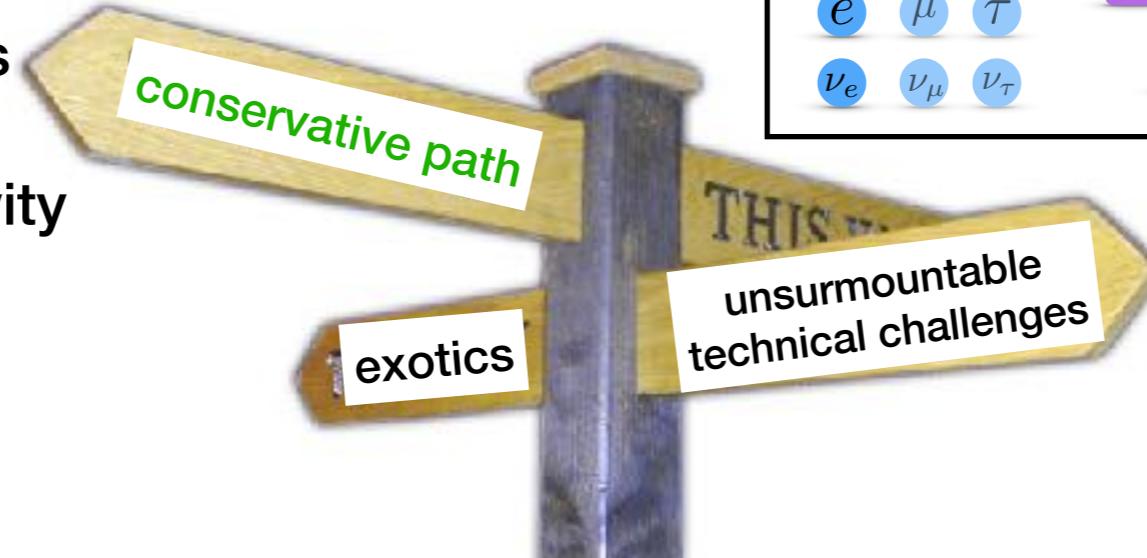
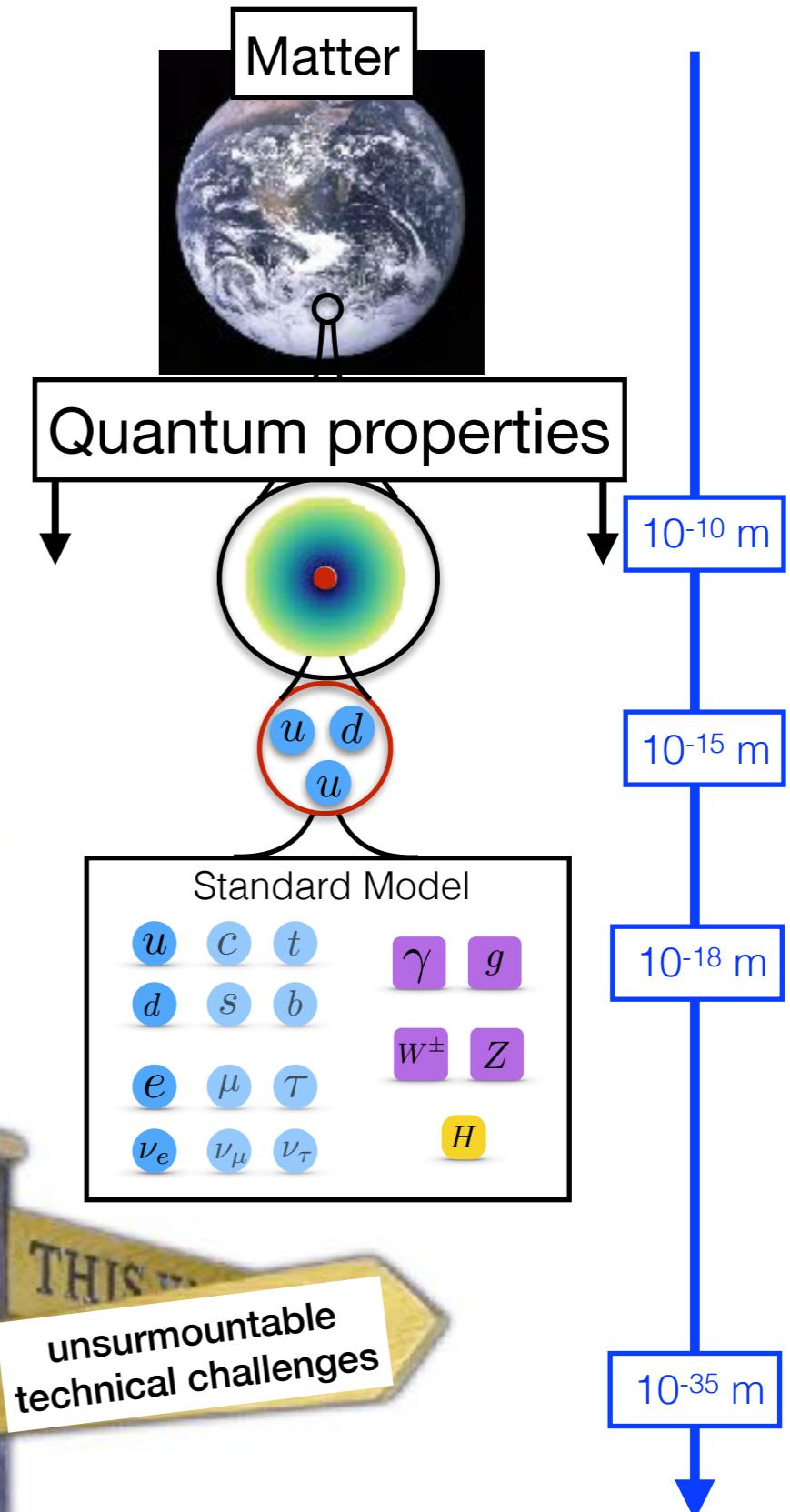
$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi G_N T_{\mu\nu}$$

Quantum properties



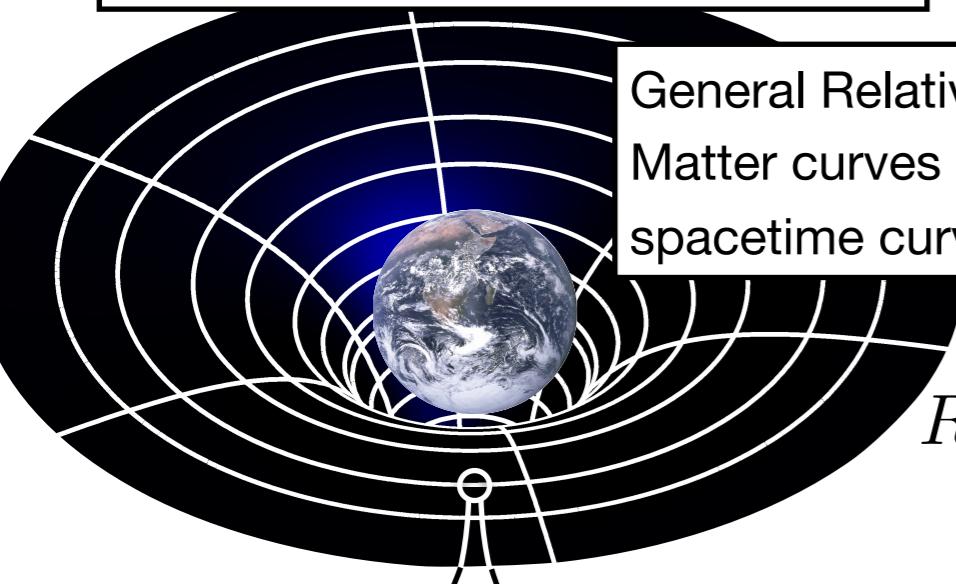
→ **asymptotically safe gravity**

- local QFT framework
 - works for all other forces
 - degrees of freedom: metric
 - good description of gravity at low energies
- could induce predictive UV completion of SM



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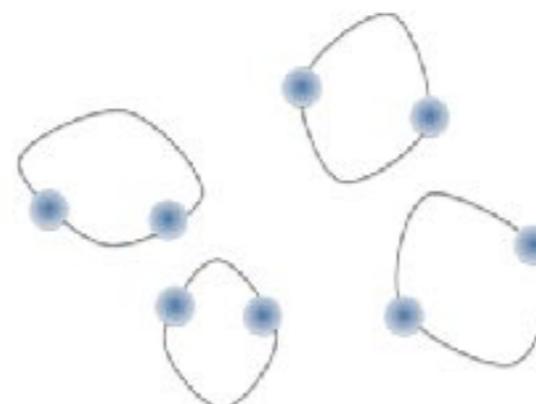
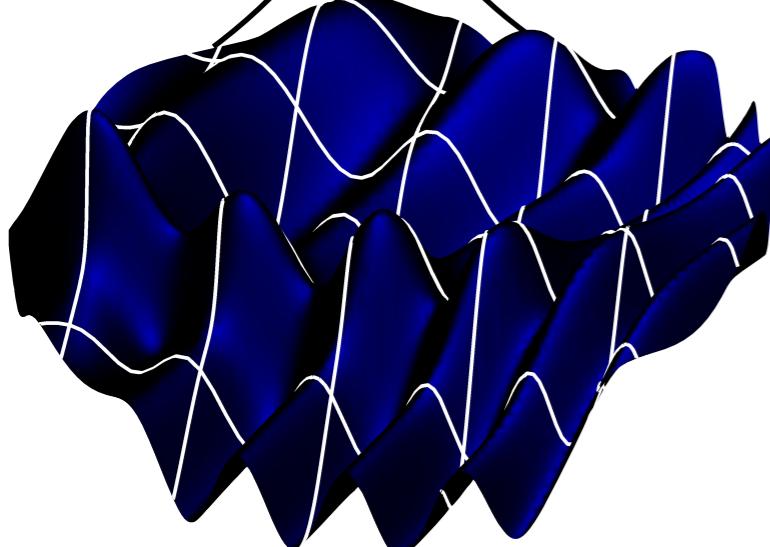
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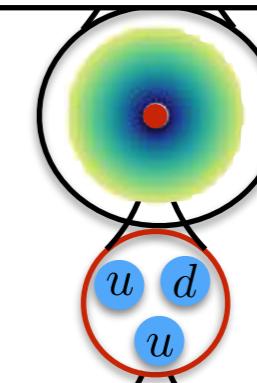


$$Z = \int \mathcal{D}g_{\mu\nu} e^{iS[g_{\mu\nu}]}$$

Matter

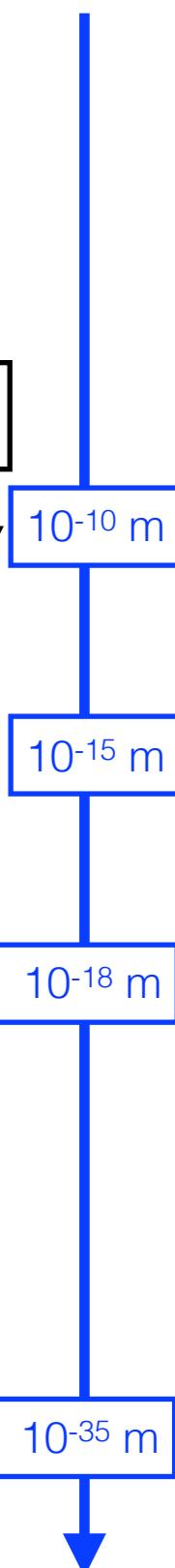


Quantum properties



Standard Model

u	c	t	γ	g
d	s	b	W^\pm	Z
e	μ	τ		H
ν_e	ν_μ	ν_τ		



Quantum field theory for gravity

$$Z = \int \mathcal{D}g_{\mu\nu} e^{-S[g_{\mu\nu}]}$$

Which microscopic action?

Einstein-Hilbert action & perturbative quantisation: $S = -\frac{1}{16\pi G_N} \int d^4x \sqrt{g} R$

$$g_{\mu\nu} = \eta_{\mu\nu} + \sqrt{16\pi G_N} h_{\mu\nu}$$

spin-2-field on flat background

counterterms:

1-loop: $R^2, R_{\mu\nu}R^{\mu\nu}$ 't Hooft, Veltman '74;
Deser, Nieuwenhuizen '74

2-loop: $C_{\mu\nu\kappa\lambda}C^{\kappa\lambda\rho\sigma}C_{\rho\sigma}^{\mu\nu}$ Goroff, Sagnotti '86;
Van de Ven '92

...

breakdown of predictivity

consistent choice of S with finite number of free parameters?

Quantum field theory for gravity

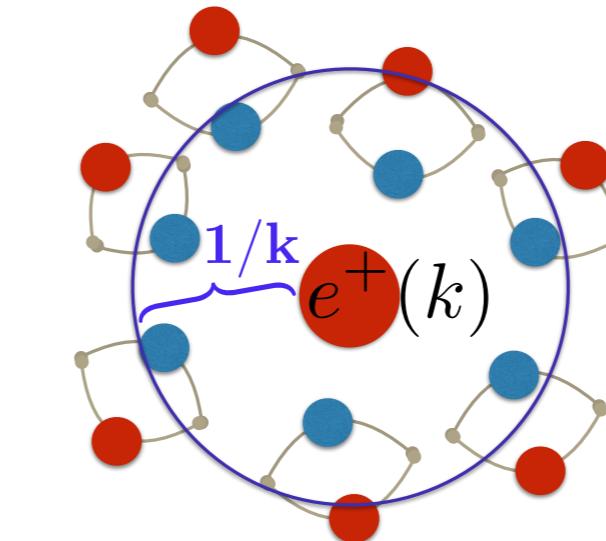
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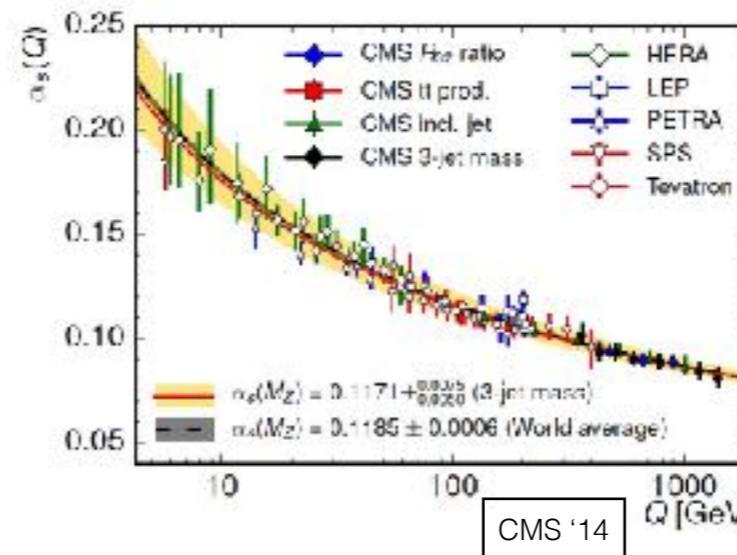
Quantum fluctuations generate running (scale-dependent) couplings

Asymptotic freedom in non-Abelian gauge theories

[Gross, Wilczek '73; Politzer '73]



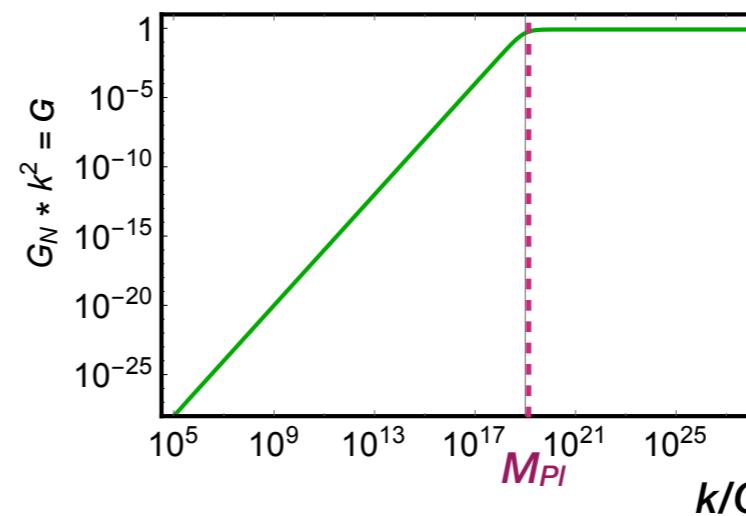
microscopic regime
in fundamental theory
(viable w/o "new physics"):
scale- invariance



$$\beta_{\alpha_s} = -\frac{7}{2\pi} \alpha_s^2 + \dots$$

quark+ gluon fluctuations
(universal @ 1loop)

$$\alpha_s^* = 0$$



$$\beta_G = 2G - \frac{23G^2}{3\pi} + \dots$$

metric fluctuations

$$G^* = \frac{6\pi}{23}$$

asymptotically safe beyond M_{Pl}

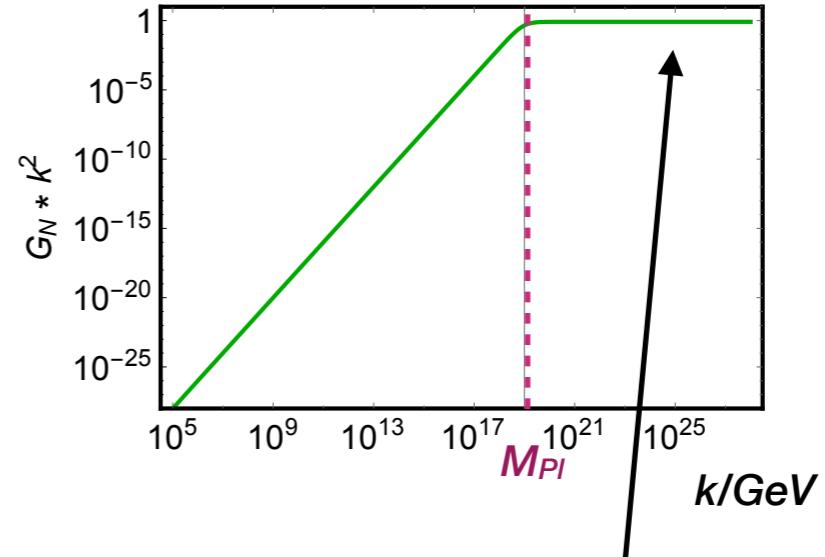
[Weinberg '76, '79; Reuter '96]

[Codello, Percacci, Rahmede '08]

Asymptotic safety

Quantum gravity:
 Newton coupling grows towards UV
 asymptotically safe beyond M_{Pl}

[Weinberg '76, '79; Reuter '96]



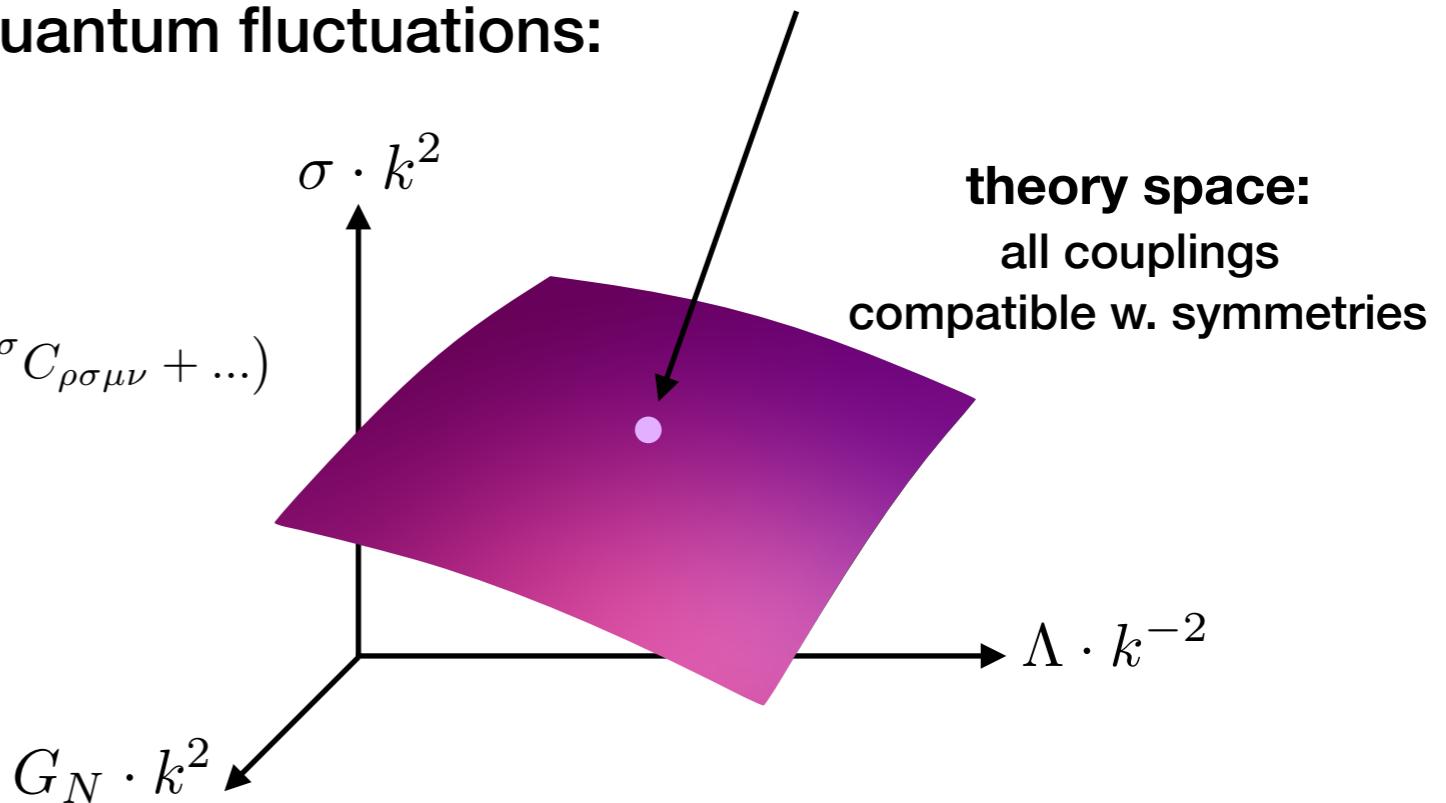
**scale invariant
 fixed-point regime**
 (microscopic dynamics @ $k \rightarrow \infty$)

higher-order interactions generated by quantum fluctuations:

$$\Gamma_k = -\frac{1}{16\pi G_N} \int d^4x \sqrt{g} (R - 2\Lambda)$$

$$+ \int d^4x \sqrt{g} (a_1 R^2 + a_2 R^{\mu\nu} R_{\mu\nu} + \sigma C^{\mu\nu\kappa\lambda} C_{\kappa\lambda}^{\rho\sigma} C_{\rho\sigma\mu\nu} + \dots)$$

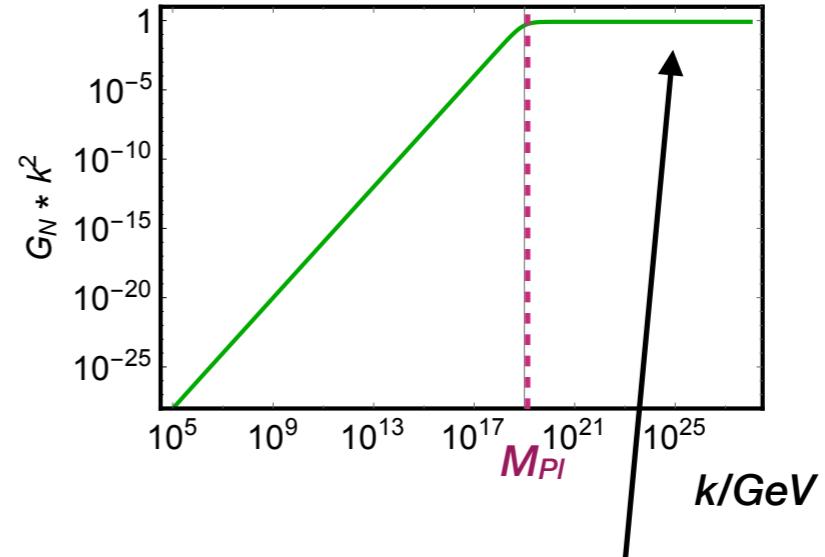
But at large scales our world
 is clearly not scale-invariant?!



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[Weinberg '76, '79; Reuter '96]



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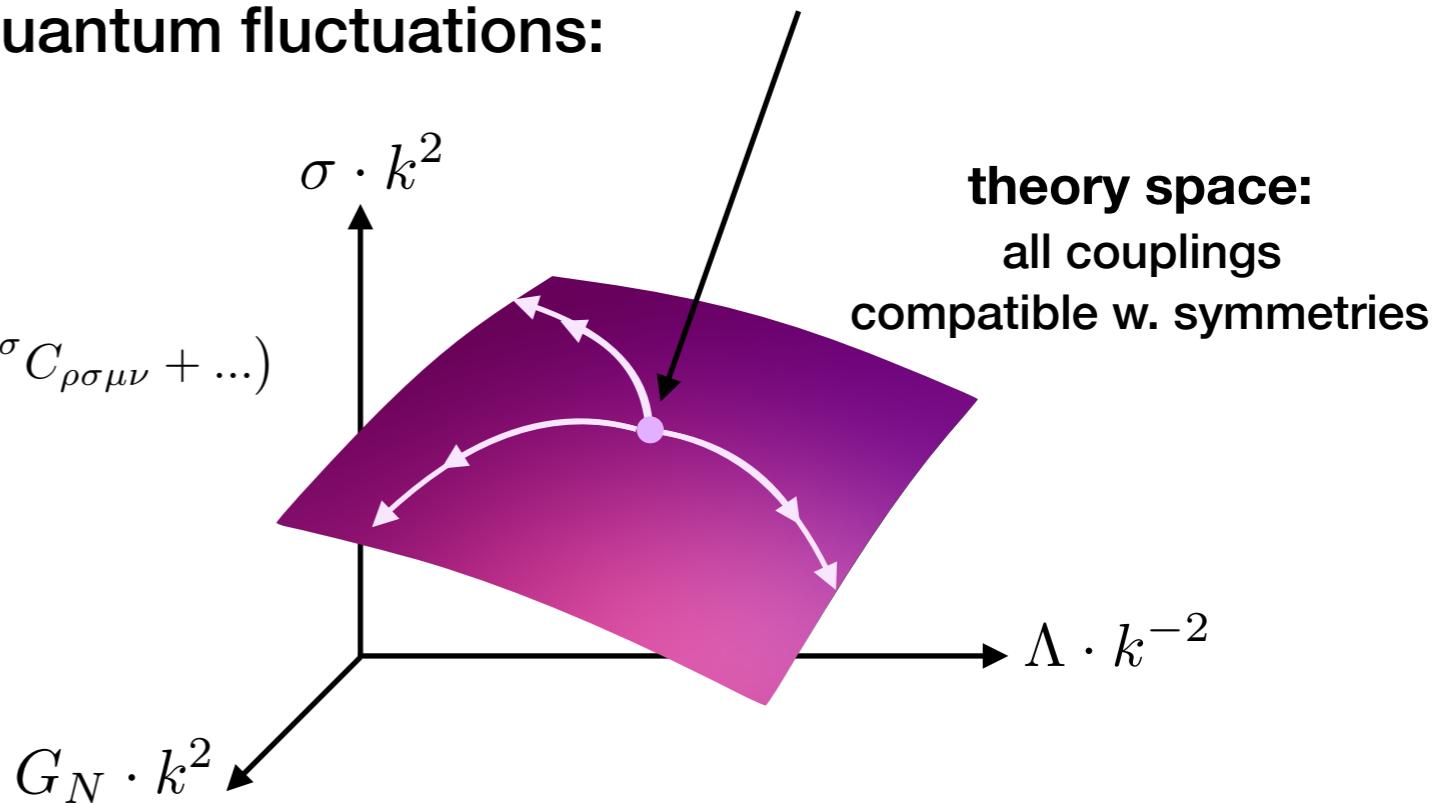
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But at large scales our world
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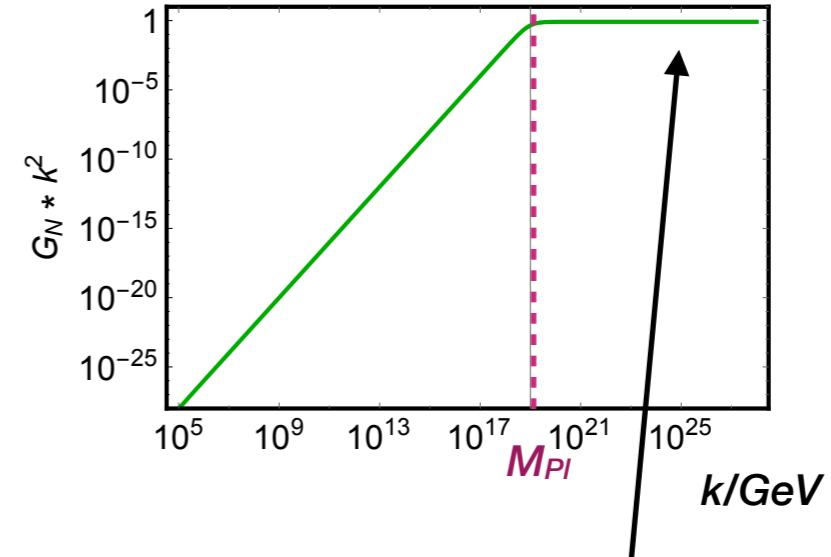
→ Flow away from fixed point



Asymptotic safety

Quantum gravity:
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[Weinberg '76, '79; Reuter '96]



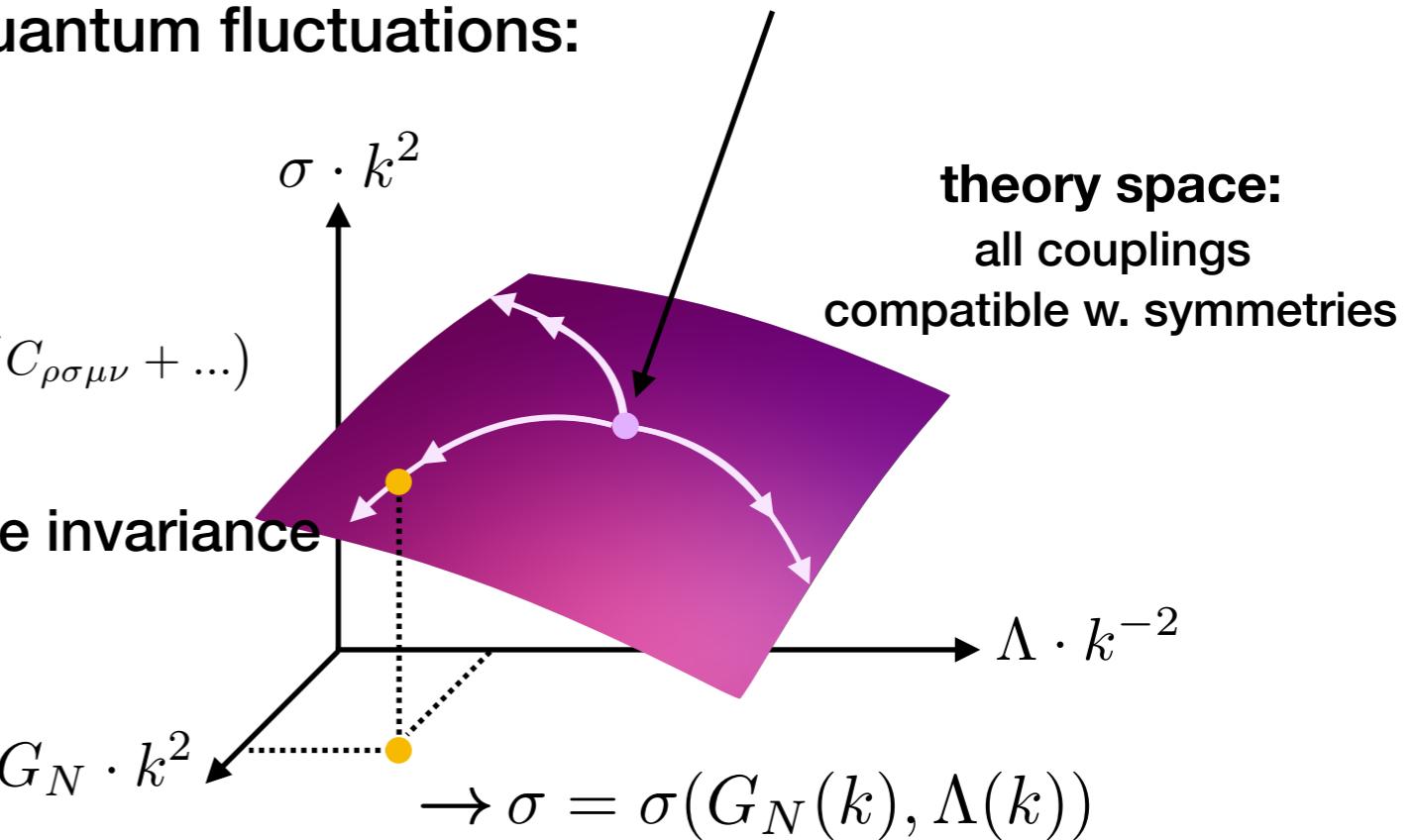
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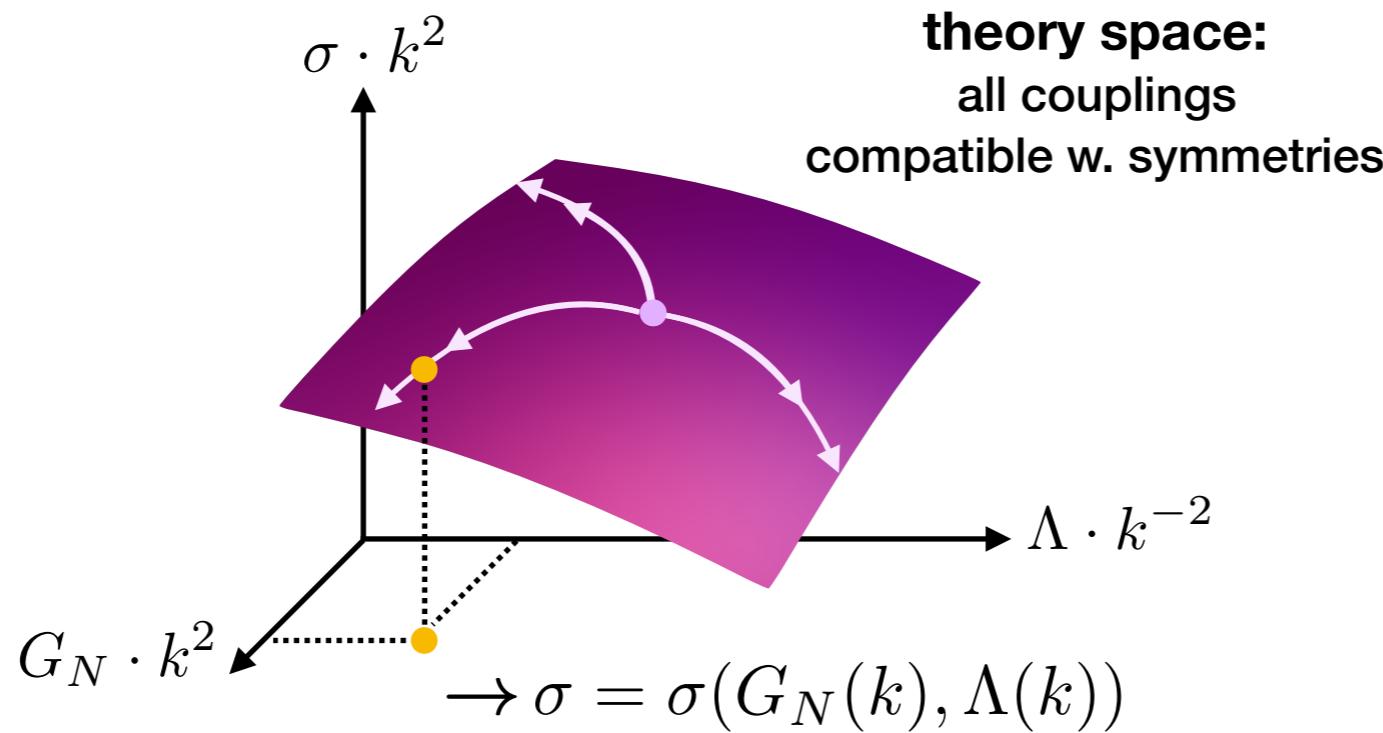
$$\Gamma_k = -\frac{1}{16\pi G_N} \int d^4x \sqrt{g} (R - 2\Lambda) + \int d^4x \sqrt{g} (a_1 R^2 + a_2 R^{\mu\nu} R_{\mu\nu} + \sigma C^{\mu\nu\kappa\lambda} C_{\kappa\lambda}^{\rho\sigma} C_{\rho\sigma\mu\nu} + \dots)$$

low-energy dynamics:

- free parameters encode deviation from scale invariance (relevant couplings)
- irrelevant couplings ***predicted*** (qm fluc's force flow to stick to **critical hypersurface**)



Asymptotic safety in a nutshell



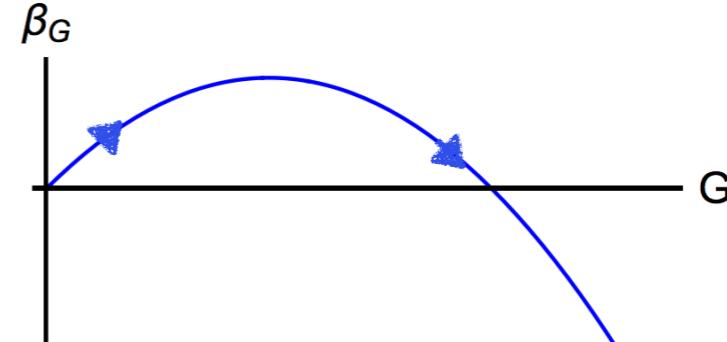
Theory space features an interacting fixed point with a finite number of relevant directions.
(At least) one trajectory emanating from the fixed point reaches a phenomenologically viable IR regime.

- UV complete
- predictive
(finite # free parameters)
- predictions for irrelevant couplings match observations

Hints for asymptotic safety of gravity?

- **ϵ expansion in $d=2+\epsilon$** $\beta_G = \epsilon G - \frac{38}{3}G^2$

Weinberg '76; Christensen, Duff '78
Gastmans, Kallosh, Truffin '78



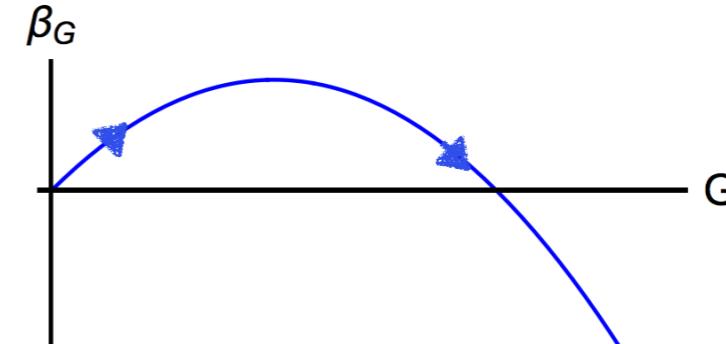
- **lattice simulations (Euclidean/Causal Dynamical Triangulations)**
continuum limit not conclusively established

Ambjorn, Jurkiewicz, Loll '01, '04...
Coumbe, Laiho '11, '17

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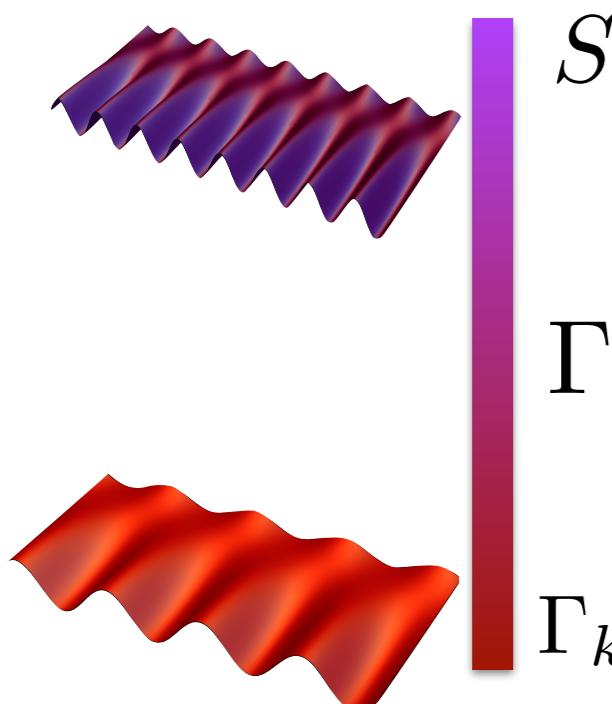
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- **Functional Renormalization Group**

probe scale dependence of QFT

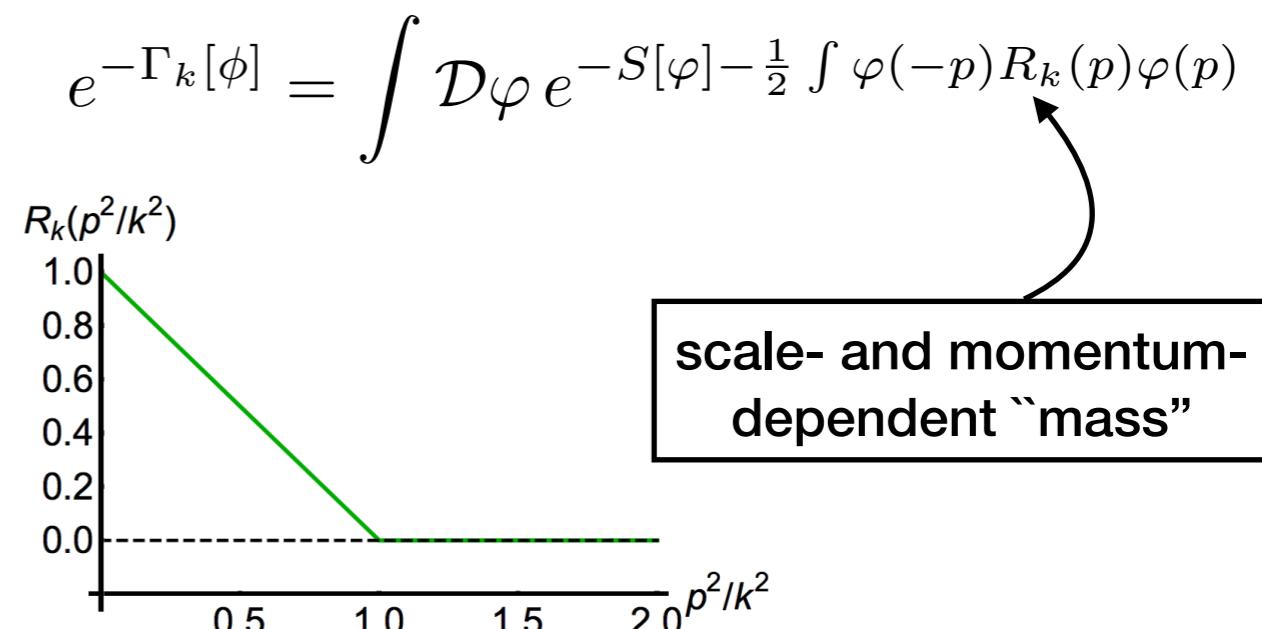
Wetterich '93, Reuter '96



Γ_k contains effect of quantum fluctuations above k

$$\Gamma_k = \sum_i g_i(k) \int d^d x \mathcal{O}^i \rightarrow k \partial_k \Gamma_k = \sum_i \beta_{g_i} \int d^d x \mathcal{O}^i$$

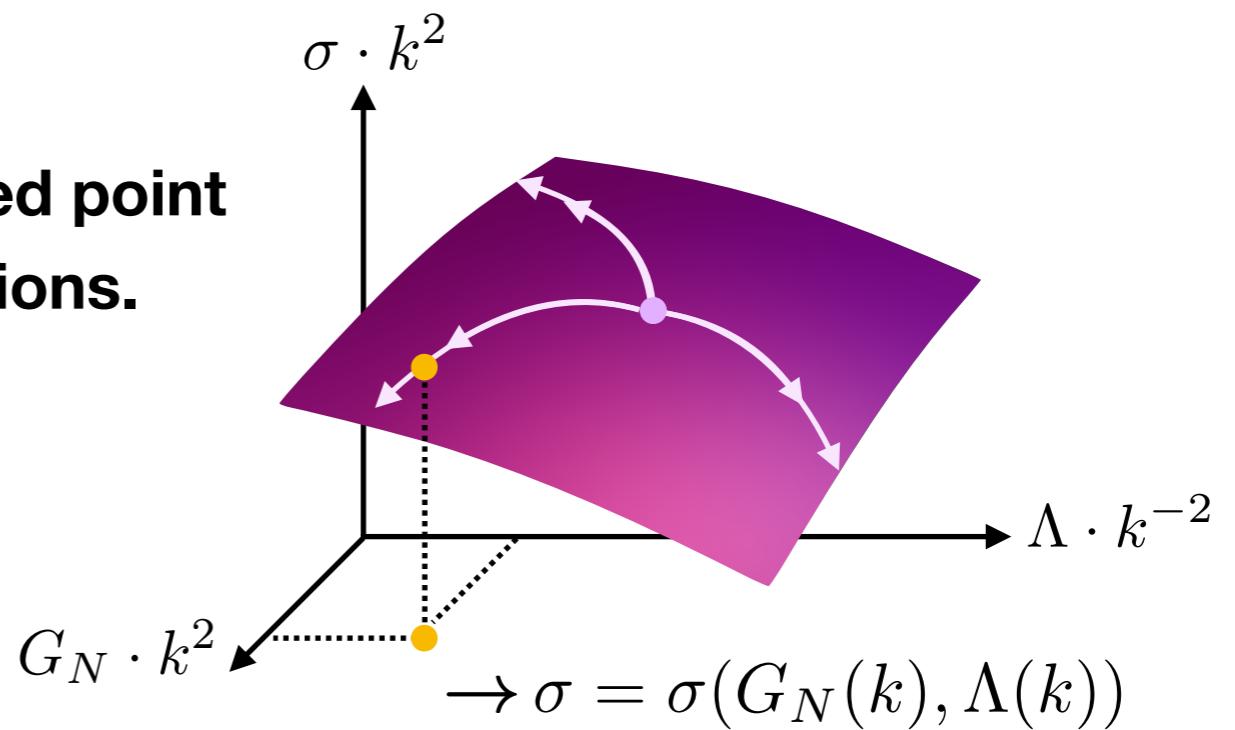
Wetterich equation: $\partial_k \Gamma_k = \frac{1}{2} S \text{Tr} \left(\Gamma_k^{(2)} + R_k \right)^{-1} \partial_k R_k =$



Hints for asymptotic safety of gravity

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fixed
point



operators

$$\sqrt{g}$$

[Reuter '96, Lauscher, Reuter '01;
Reuter, Saueressig '02;

Becker, Reuter '14;

Christiansen, Knorr, Meibohm, Pawłowski, Reichert '15]

$$\sqrt{g}R$$

$$G$$

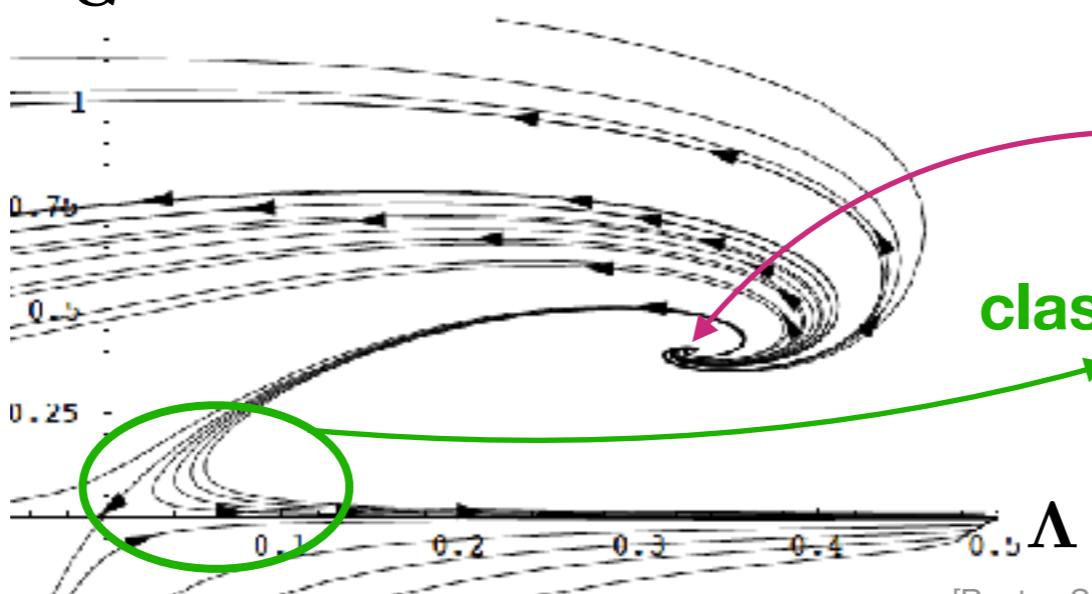
relevant irrelevant

X

X

tool: functional RG equation
in truncations of theory space
(well-tested for interacting fixed points)

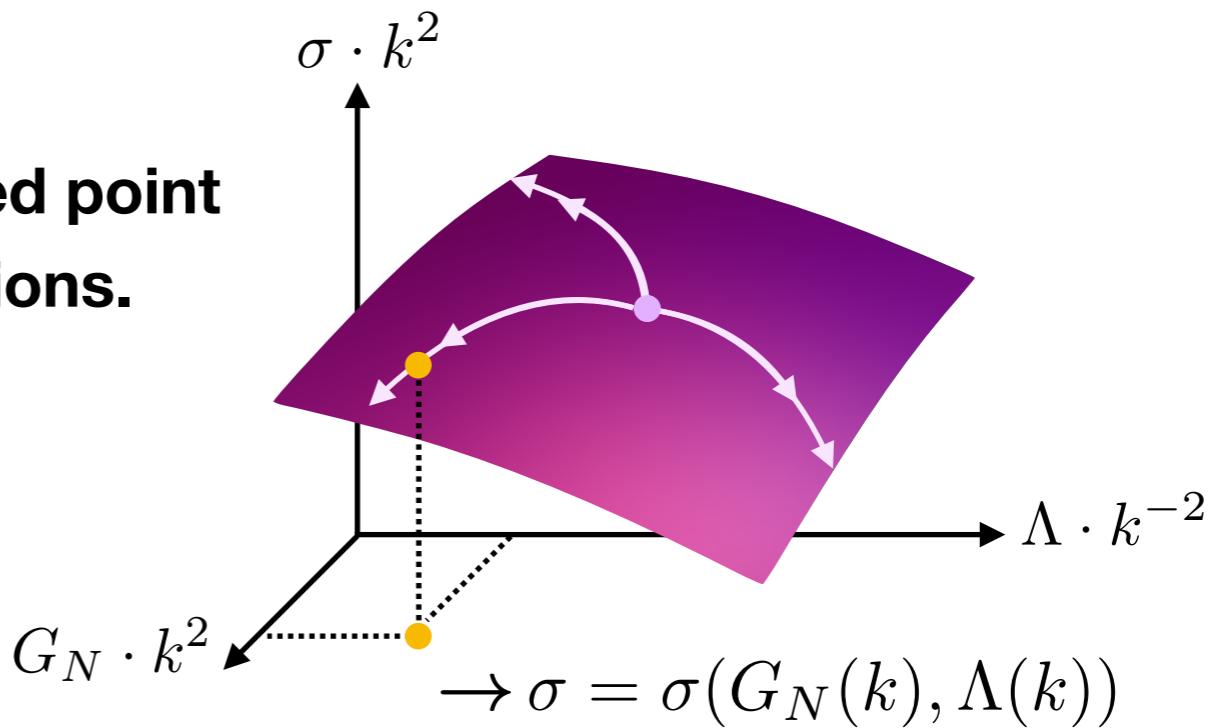
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✓ $\sqrt{g}R$

✓ $\sqrt{g}R^2, \sqrt{g}R^{\mu\nu} R_{\mu\nu}$

[Benedetti, Machado, Saueressig '09;
Denz, Pawłowski, Reichert '17]

✓ $\sqrt{g}R^3$

✓ $\sqrt{g}R^{34}$

[Codello, Percacci, Rahmede '07, '08
Machado, Saueressig '07;
Eichhorn '15]

✓ $\sqrt{g}C^{\mu\nu\kappa\lambda} C_{\kappa\lambda}^{\rho\sigma} C_{\rho\sigma\mu\nu}$

[Falls, Litim, Nikolopoulos, Rahmede '13 '14]
[Gies, Knorr, Lippoldt, Saueressig '16]

relevant

irrelevant

✗

✗

✗

✗

✗

✗

✗

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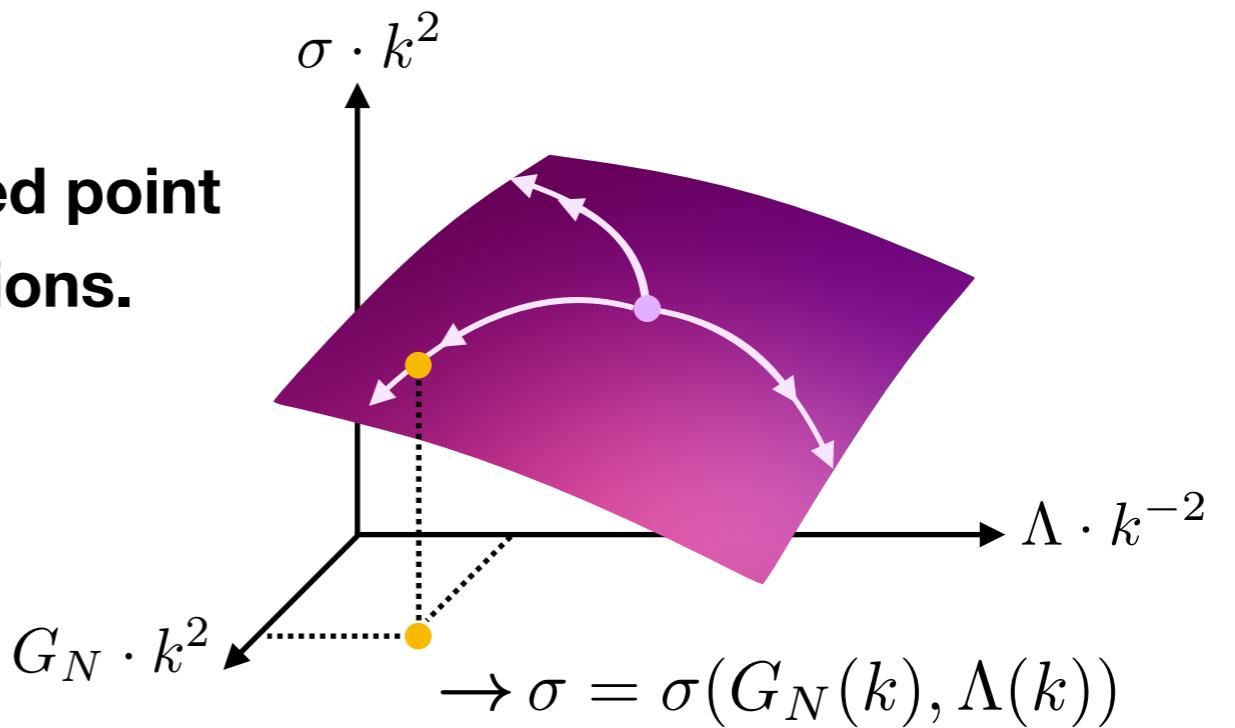
ordering principle:
canonical mass dimension
of couplings

→ control over approximations
with finitely many couplings

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✗

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✗

[Wetterich '93, Reuter '96]

✗

status: compelling hints
for asymptotic safety
in pure gravity

✗

.

.

✗

→ what about matter?

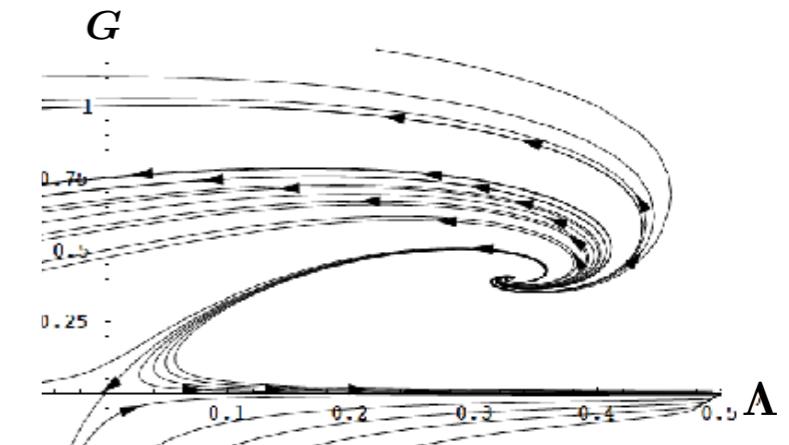
✗

Asymptotic safety for gravity & matter

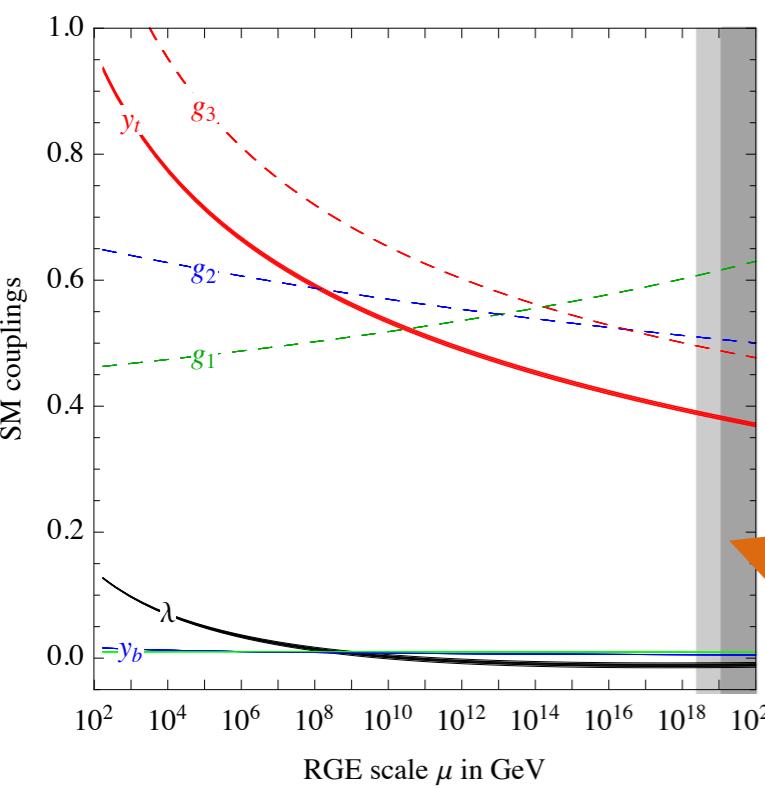
quantum gravity dynamics

matter dynamics

Can quantum fluctuations
of matter destroy a
consistent quantum gravity model?



Can quantum fluctuations
of gravity generate
a **viable UV completion for matter?**



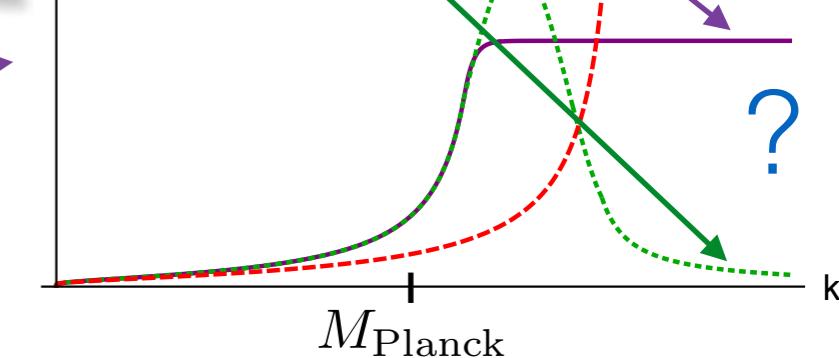
match onto SM
at Planck scale

QG generated
fixed point

triviality in Higgs-Yukawa & U(1)

QG-induced
asymptotic safety

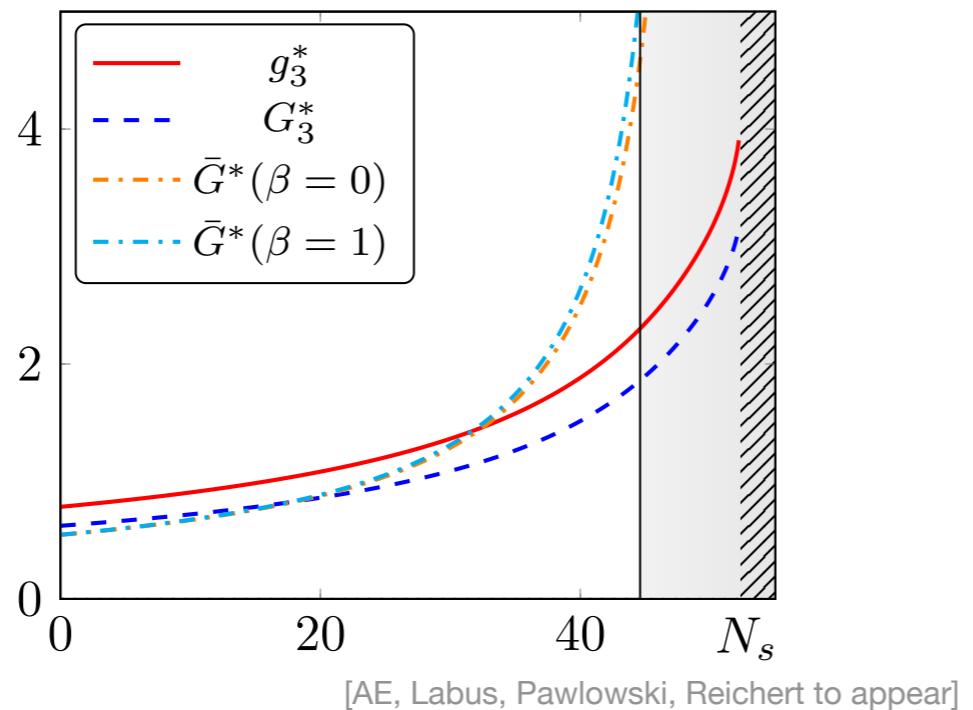
QG-induced
asymptotic freedom



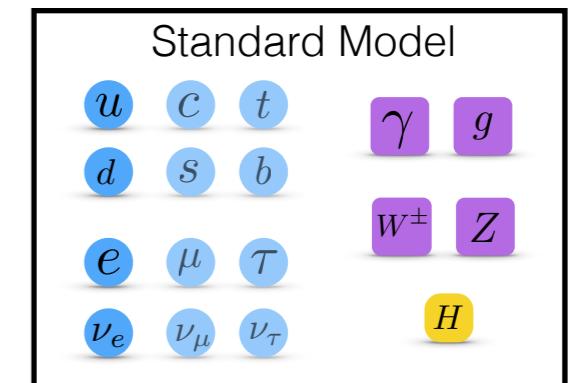
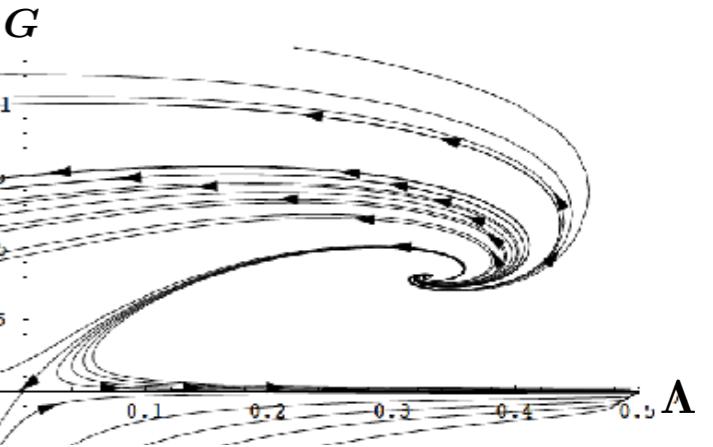
Asymptotic safety for gravity & matter

Can quantum fluctuations
of matter destroy a
consistent quantum gravity model?

matter matters:



**minimally
coupled
SM fields**



**matter content of SM (& small extensions)
admits a gravity fixed point in Einstein-Hilbert truncation**

[Dona, AE, Percacci '13, '14]

[Meibohm, Pawłowski, Reichert '15
Dona, AE, Labus, Percacci '15]

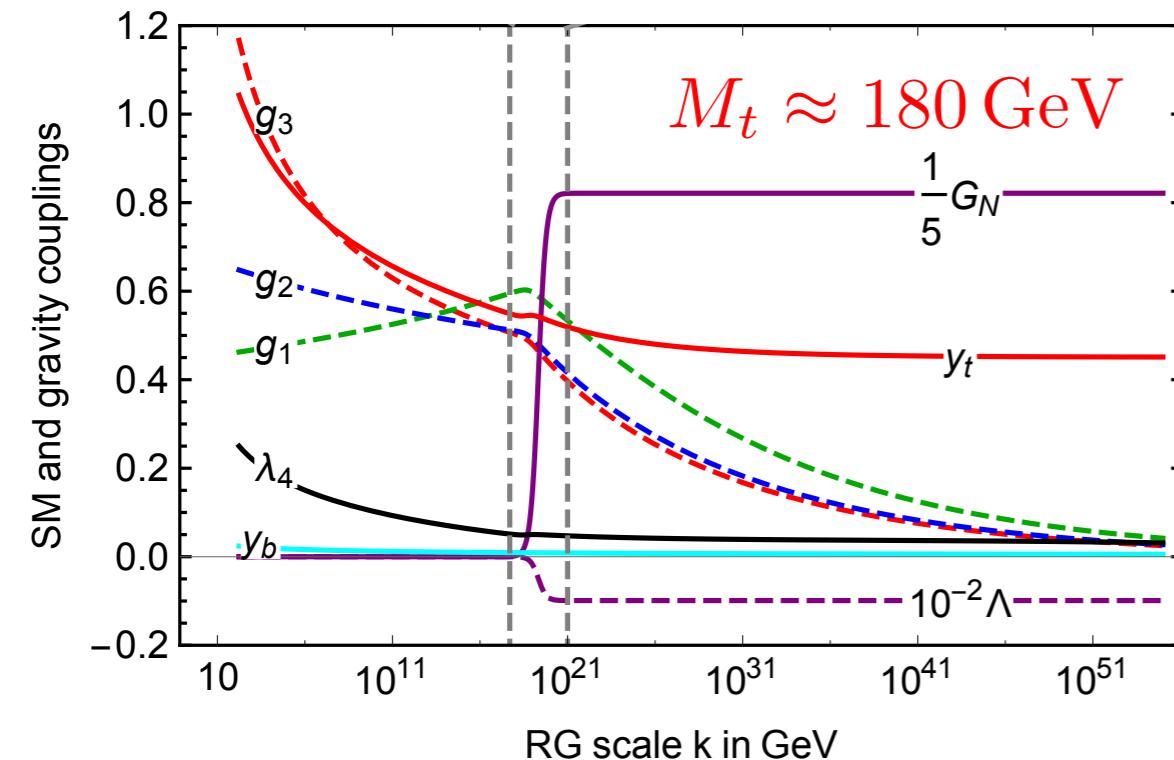
Quantum-gravity induced UV completion for the SM



results within simple truncations

→ **convergence of results in extended truncations:
stay tuned...**

Quantum-gravity induced UV completion for the SM



within simple truncations:

- **asymptotic freedom in all gauge couplings (incl. Abelian hypercharge)**

[Daum, Harst, Reuter '10; Folkerts, Litim, Pawłowski '11; Harst, Reuter '11, Christiansen, AE '17]

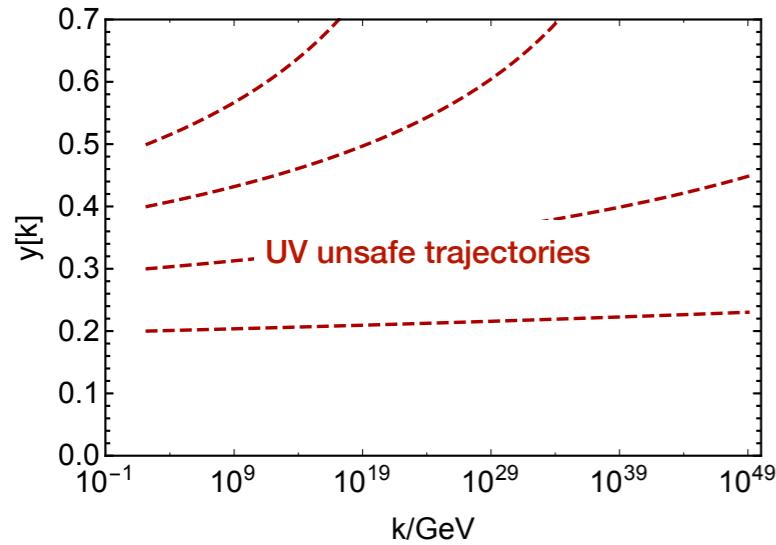
- **asymptotic safety in top Yukawa coupling with $M_t \gg M_b$ fixed uniquely**

[AE, Held, Pawłowski '16; AE, Held 05/17, 07/17]

Top mass from asymptotic safety

two regimes in grav. coupling space: **asymptotic safety in top Yukawa coupling**

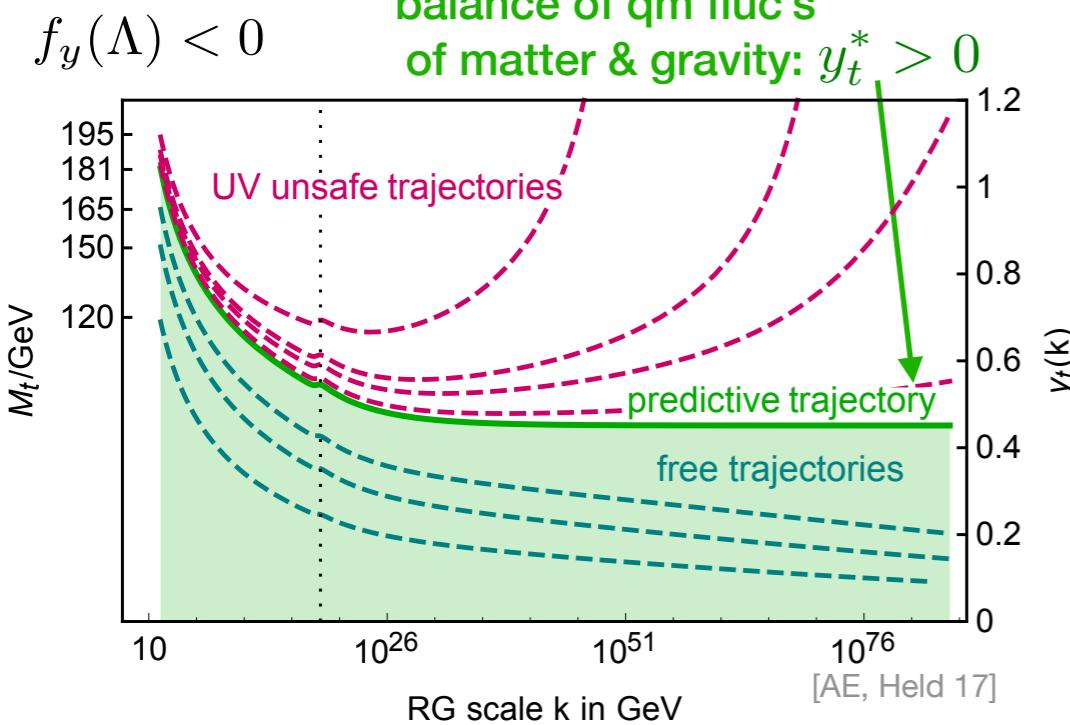
$$f_y(\Lambda) > 0$$



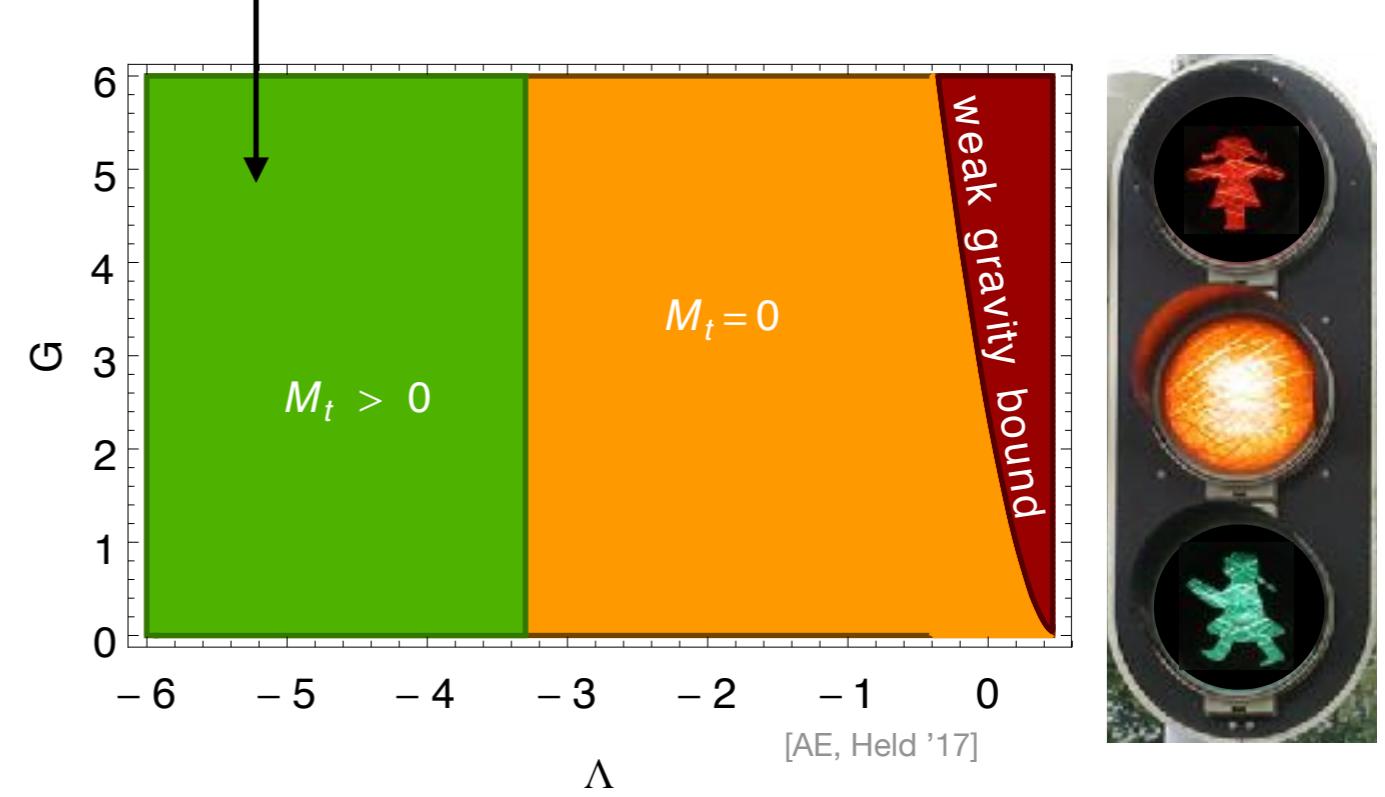
y_t : free parameter in SM $M_t \approx 172.4 \text{ GeV}$ [CMS '16]

$$\beta_{y_t} = \frac{1}{32\pi^2} (9y_t^3 + G y_t f_y(\Lambda))$$

[AE, Held, Pawłowski '16; AE, Held 17]



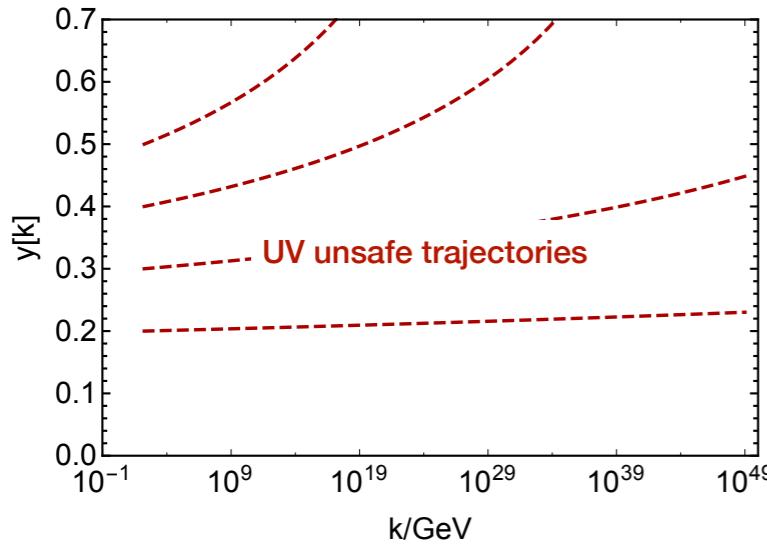
$M_t(G, \Lambda)$ uniquely fixed



Top mass from asymptotic safety

two regimes in grav. coupling space:

$$f_y(\Lambda) > 0$$



asymptotic safety in top Yukawa coupling

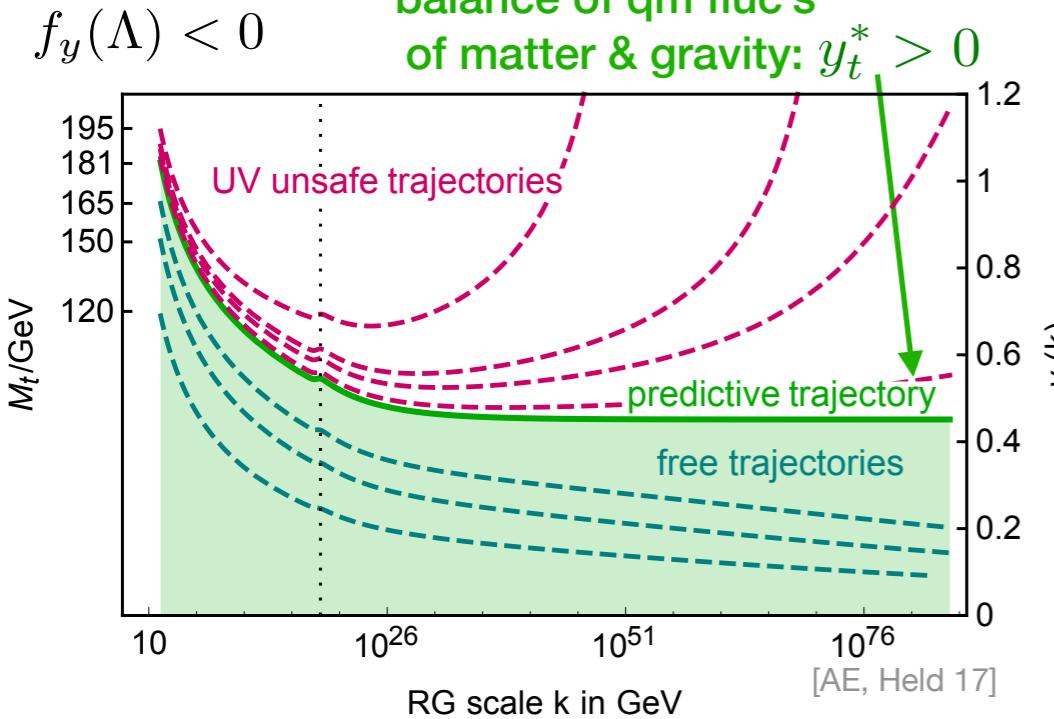
$$y_t: \text{free parameter in SM} \quad M_t \approx 172.4 \text{ GeV} \quad [\text{CMS '16}]$$

→ Why is $M_t > M_b$?

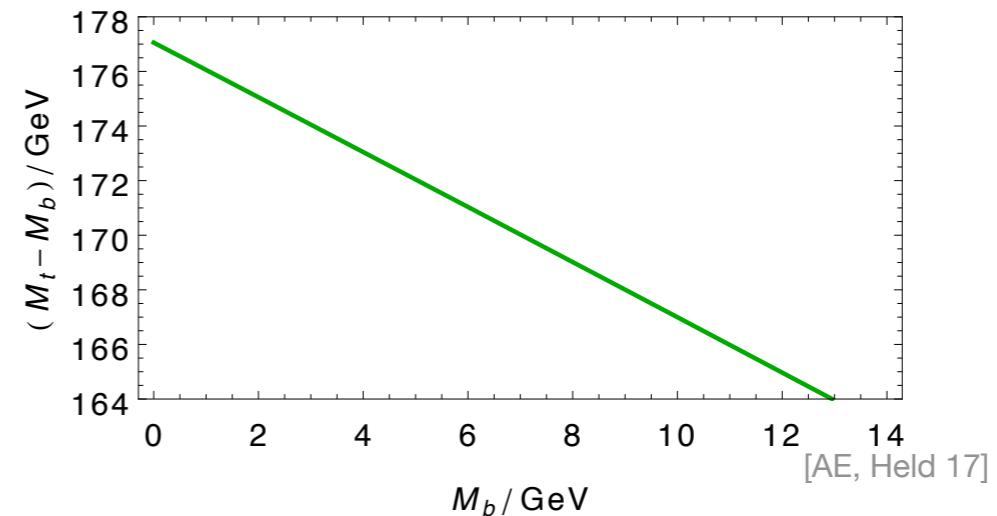
$$\beta_{y_t} = \frac{1}{32\pi^2} (9y_t^3 + 3y_b^2 y_t) + G y_t f_y(\Lambda) \quad [\text{AE, Held 17}]$$

$$\beta_{y_b} = \frac{1}{32\pi^2} (9y_b^3 + 3y_t^2 y_b) + G y_b f_y(\Lambda)$$

$y_t^* > 0, y_b^* = 0$
irrelevant: in IR: $y_t(G, \Lambda, y_b)$ relevant



M_t fixed in terms of M_b : mass difference from asymptotic safety!



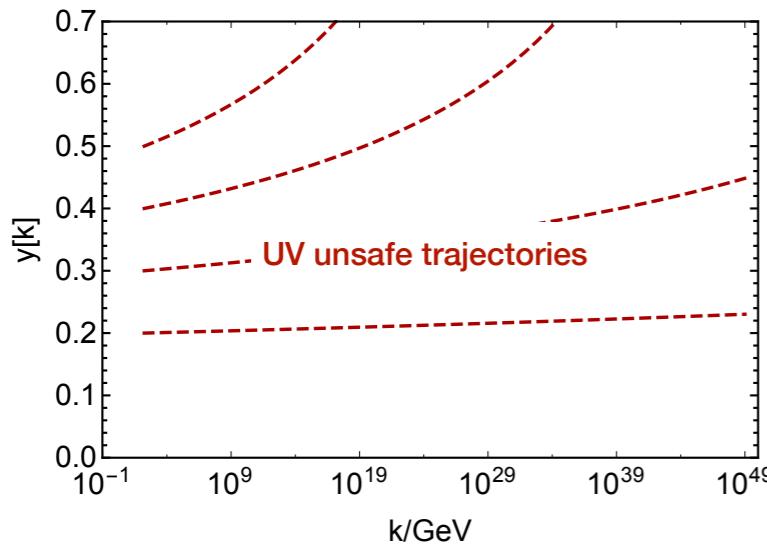
$M_t, M_b < M_{Pl}$: chiral symmetry unbroken by QG

[AE, Gies 11; Meibohm, Pawłowski '15; AE, Lippoldt '16]

Top mass from asymptotic safety

two regimes in grav. coupling space:

$$f_y(\Lambda) > 0$$



asymptotic safety in top Yukawa coupling

$$y_t: \text{free parameter in SM} \quad M_t \approx 172.4 \text{ GeV} \quad [\text{CMS '16}]$$

→ Why is $M_t > M_b$?

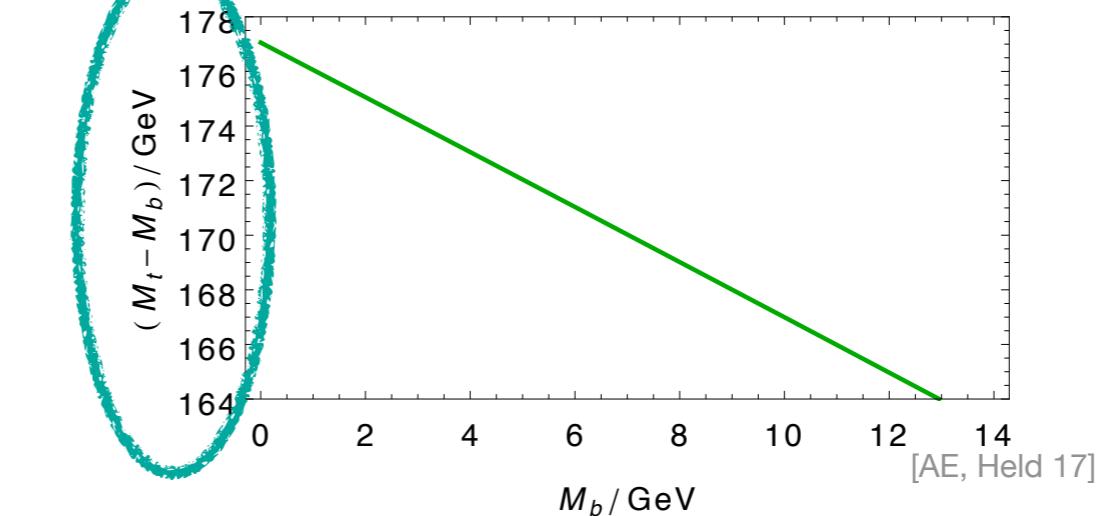
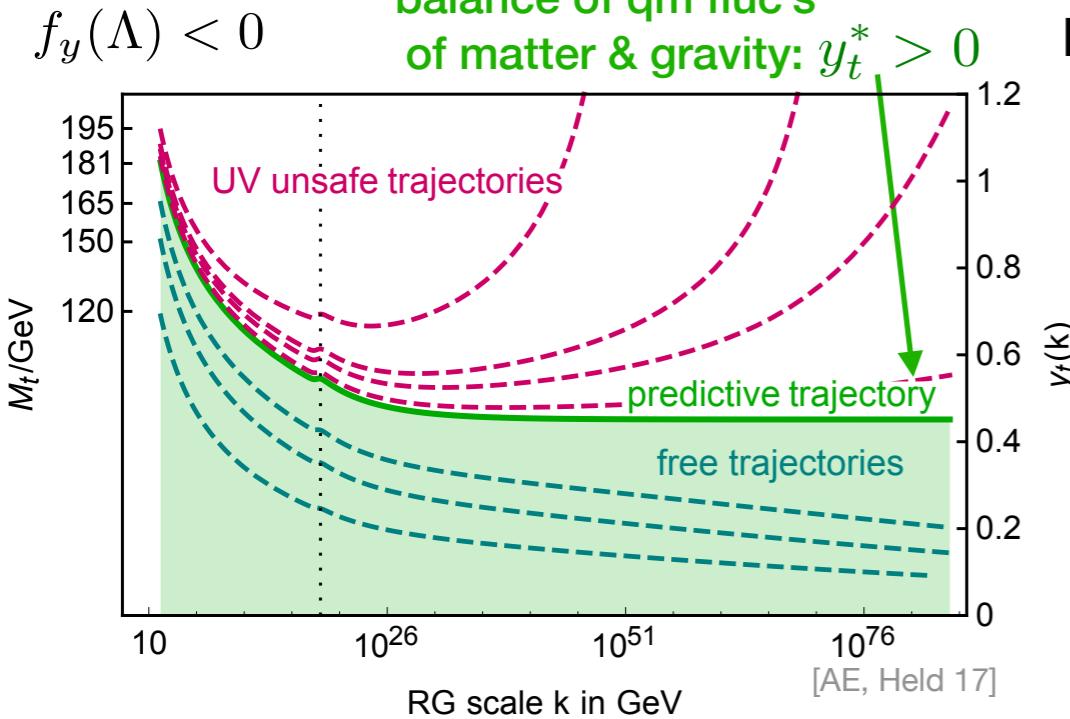
$$\beta_{y_t} = \frac{1}{32\pi^2} (9y_t^3 + \text{SM}) + G y_t f_y(\Lambda) \quad [\text{AE, Held 17}]$$

contains running gauge couplings

$$\beta_{y_b} = \frac{1}{32\pi^2} (9y_b^3 + \text{SM}) + G y_b f_y(\Lambda)$$

$y_t^* > 0, y_b^* = 0$
irrelevant: in IR: $y_t(G, \Lambda, y_b)$ relevant

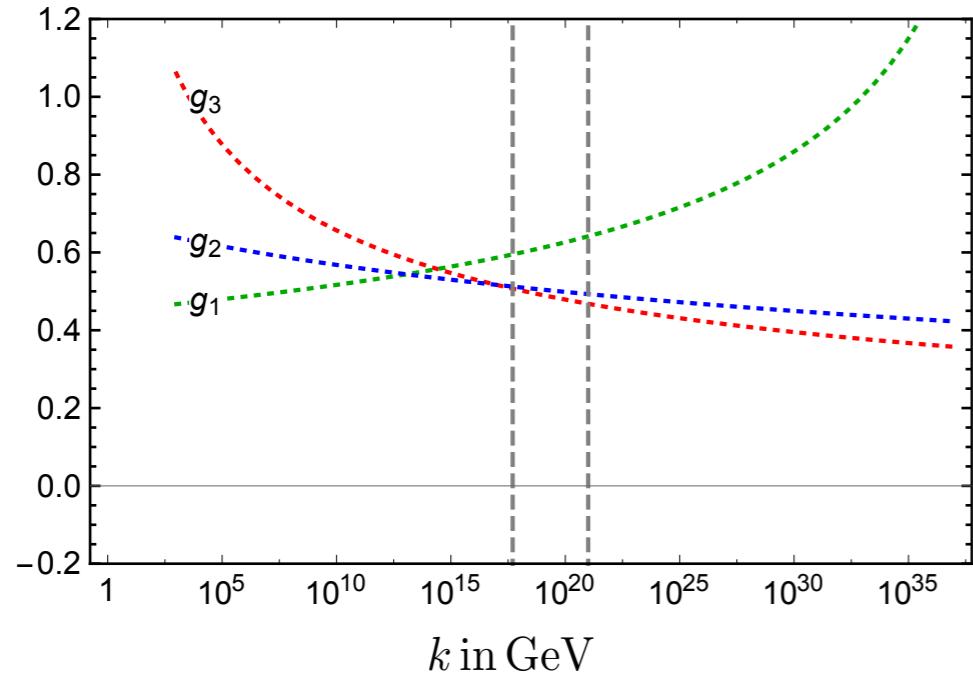
M_t fixed in terms of M_b : mass difference from asymptotic safety!



$M_t, M_b < M_{Pl}$: chiral symmetry unbroken by ASQG

[AE, Gies 11; Meibohm, Pawłowski '15; AE, Lippoldt '16]

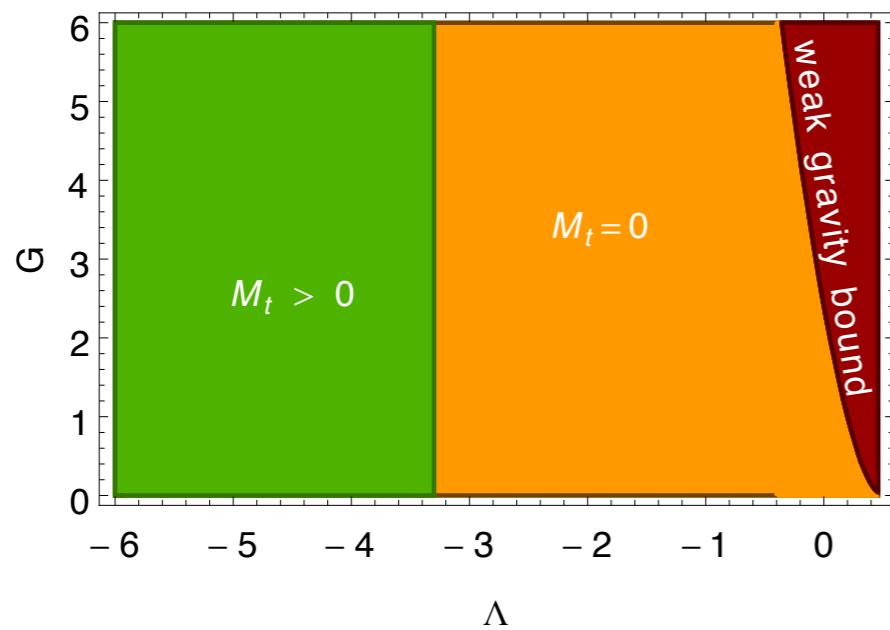
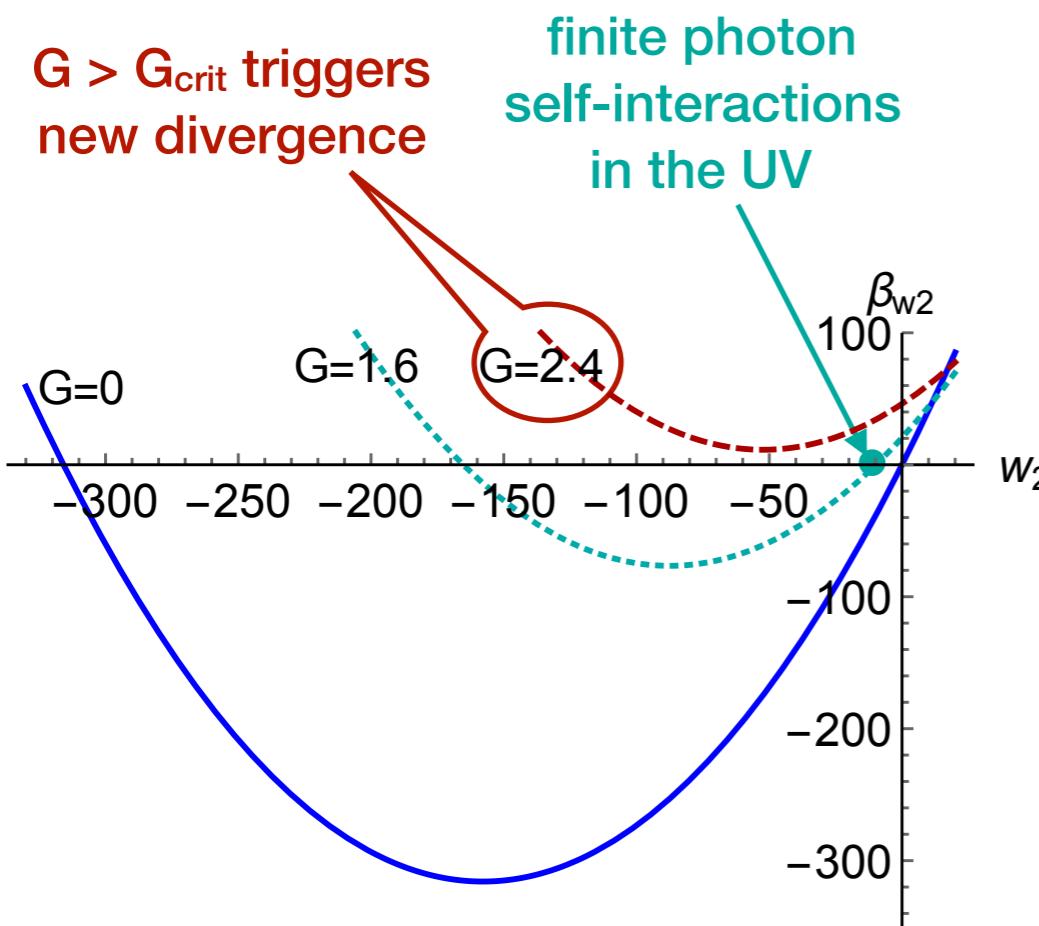
An asymptotically safe solution to the U(1) triviality problem?



triviality problem in Abelian hypercharge

[Gell-Mann, Low '54; Gockeler et al. 98; Gies, Jaeckel '04]

An asymptotically safe solution to the U(1) triviality problem?



within simple truncations ASQG induces:

- **asymptotic freedom in all gauge couplings (incl. Abelian hypercharge)**

truncation: Einstein- Hilbert + F^2

[Daum, Harst, Reuter '10; Folkerts, Litim, Pawłowski '11; Harst, Reuter '11]

$$\beta_{g_1} = \frac{g_1^3}{16\pi^2} \frac{41}{10} - G g_1 f(\Lambda)$$

$+ w_2 F^4$ QG-induced photon interactions

[Christiansen, AE '17]

weak-gravity bound:

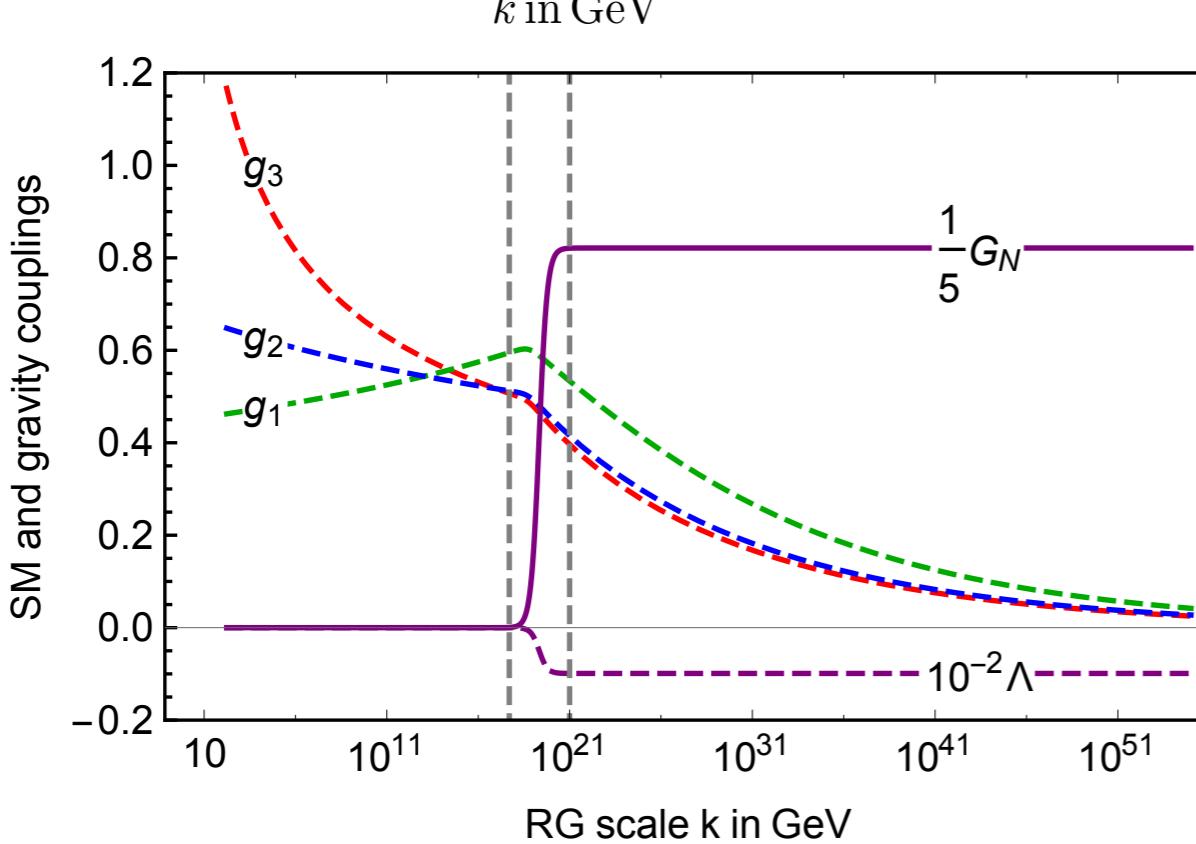
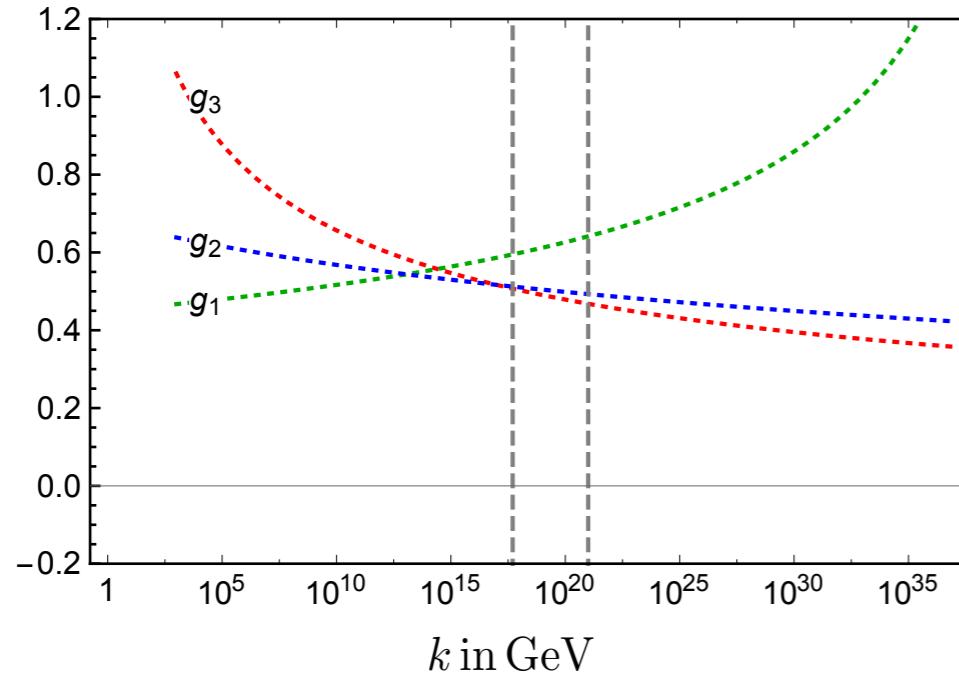
[AE, Held, Pawłowski '16;
Christiansen, AE '17]

“Strong” quantum gravity triggers new divergences in matter interactions

$G < G_{\text{crit}}$ in truncations

see [AE, Held 17]

An asymptotically safe solution to the U(1) triviality problem?



- within simple truncations ASQG induces:**
- **asymptotic freedom in all gauge couplings (incl. Abelian hypercharge)**
- **power-law running towards free fixed point**

truncation: Einstein- Hilbert + F^2

[Daum, Harst, Reuter '10; Folkerts, Litim, Pawłowski '11; Harst, Reuter '11]

$$\beta_{g_1} = \frac{g_1^3}{16\pi^2} \frac{41}{10} - G g_1 f(\Lambda)$$

+ $w_2 F^4$ **QG-induced photon interactions**

[Christiansen, AE '17]

weak-gravity bound:

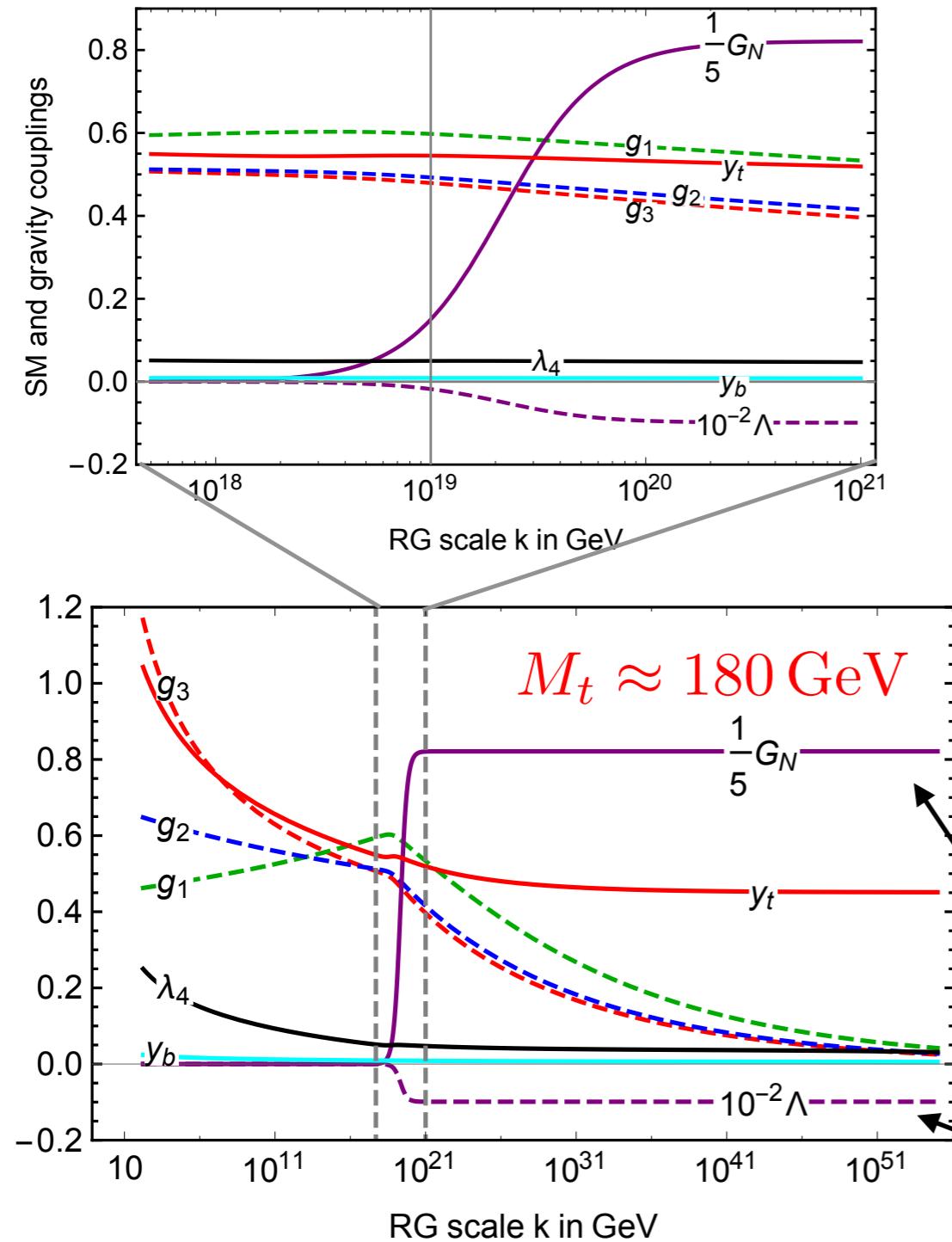
[AE, Held, Pawłowski '16;
Christiansen, AE '17]

“Strong” quantum gravity triggers
new divergences in matter interactions

$G < G_{\text{crit}}$ in truncations

see [AE, Held 17]

Quantum-gravity induced UV completion for the SM



mechanism for Higgs mass:

[Shaposhnikov, Wetterich '09]

β_G, β_Λ including SM matter fields *

[Dona, AE, Percacci '13]

* convergence in fixed-point values: to be tested

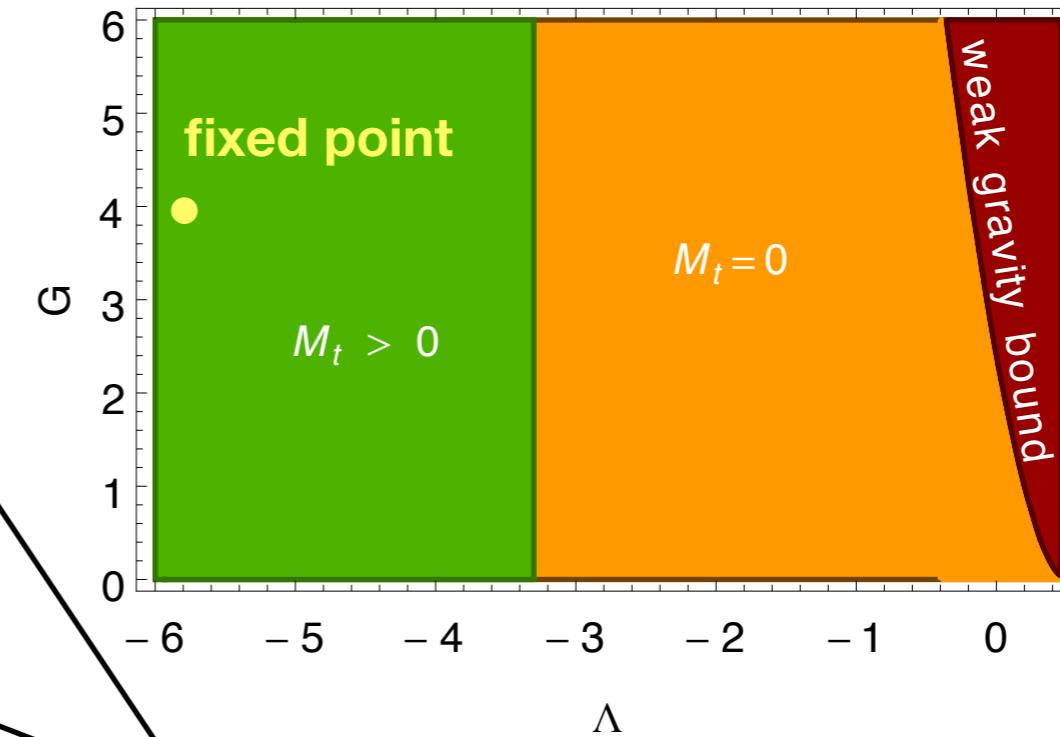
within simple truncations:

- asymptotic freedom in all gauge couplings (incl. Abelian hypercharge)

[Daum, Harst, Reuter '10; Folkerts, Litim, Pawłowski '11; Harst, Reuter '11, Christiansen, AE '17]

- asymptotic safety in top Yukawa coupling in part of grav. coupling space

[AE, Held, Pawłowski '16; AE, Held 05/17, 07/17]

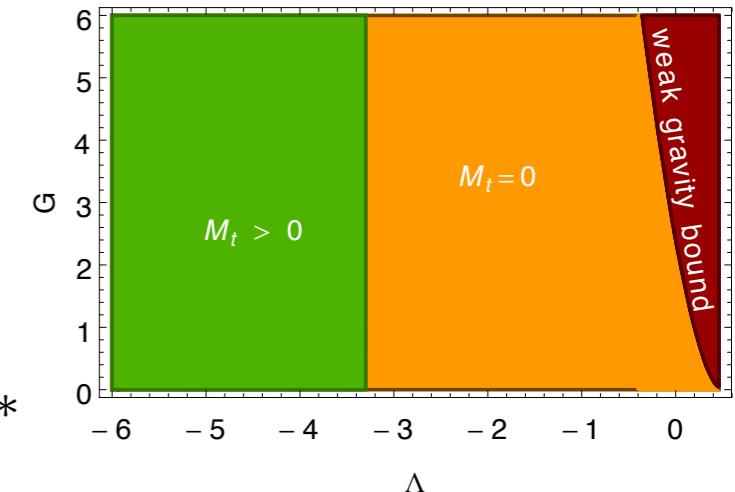


Learning about the dark sector from asymptotic safety

dark sector might only couple gravitationally

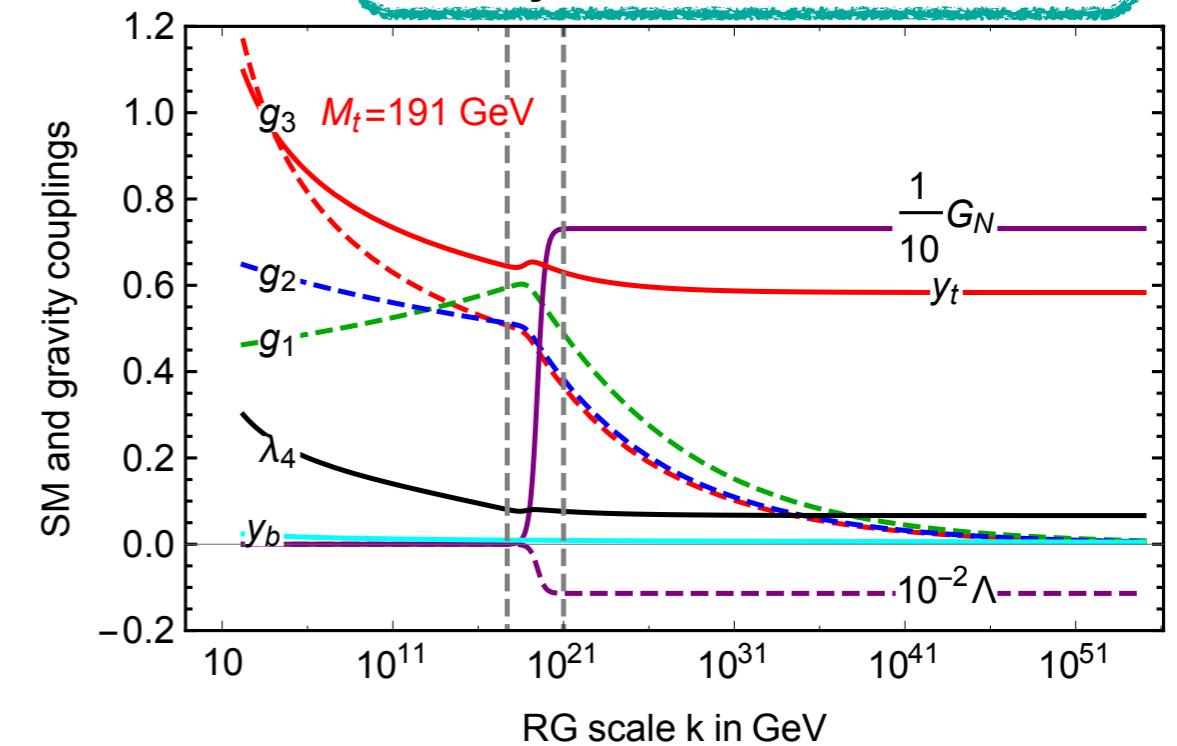
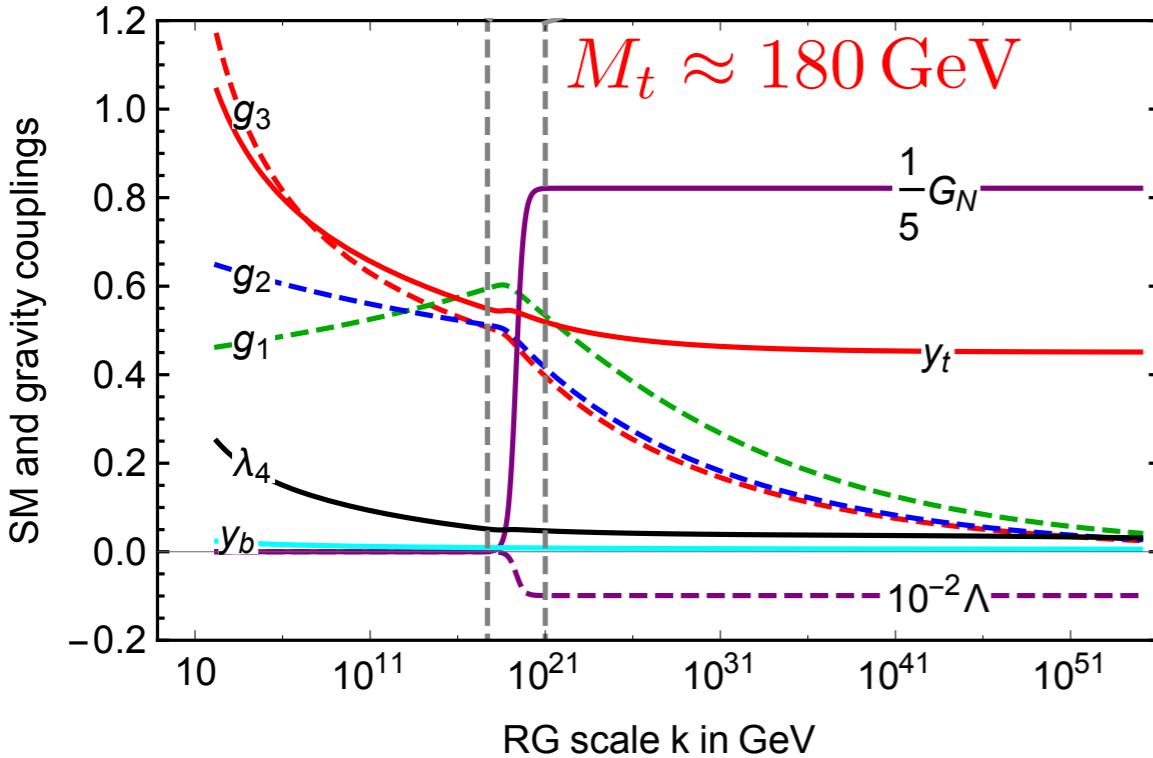
→ **direct detection very challenging**

asymptotic safety: ALL degrees of freedom affect G^* , Λ^*



→ **top-mass value depends on dark sector!** only gravitationally coupled

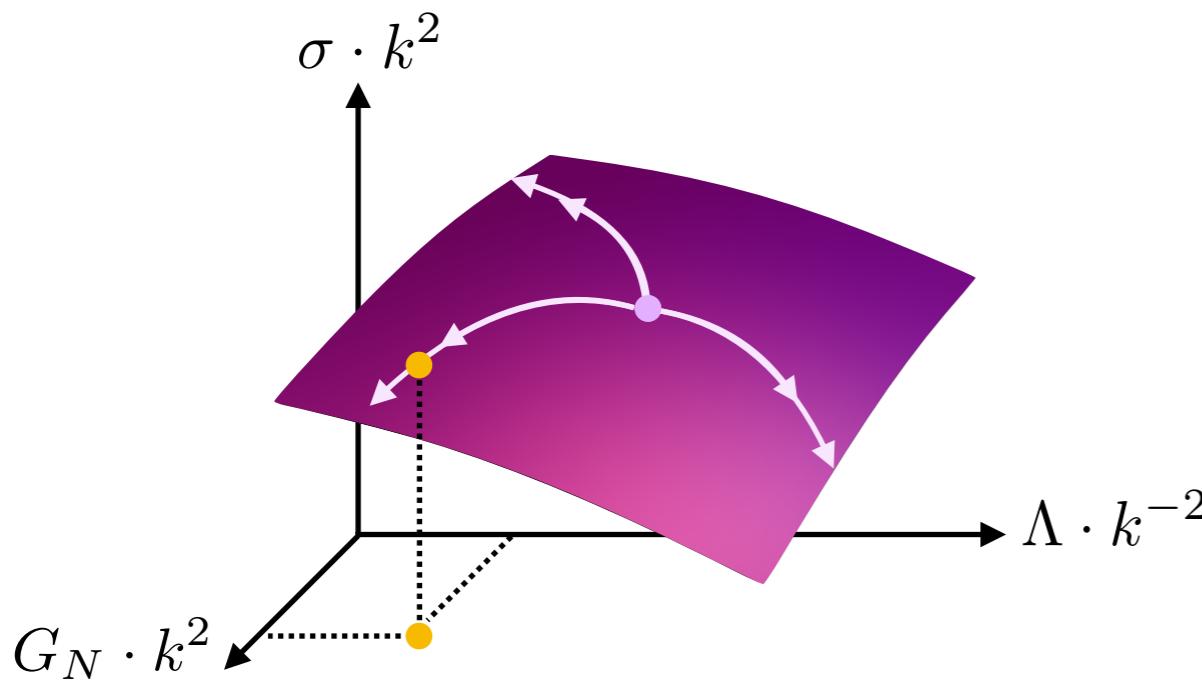
toy examples: convergence in fixed-point results: to be tested! **SM + 3 Weyl fermions + 1 scalar**



outlook: constrain dark sector by matching top-mass value from AS to measured value

Conclusions

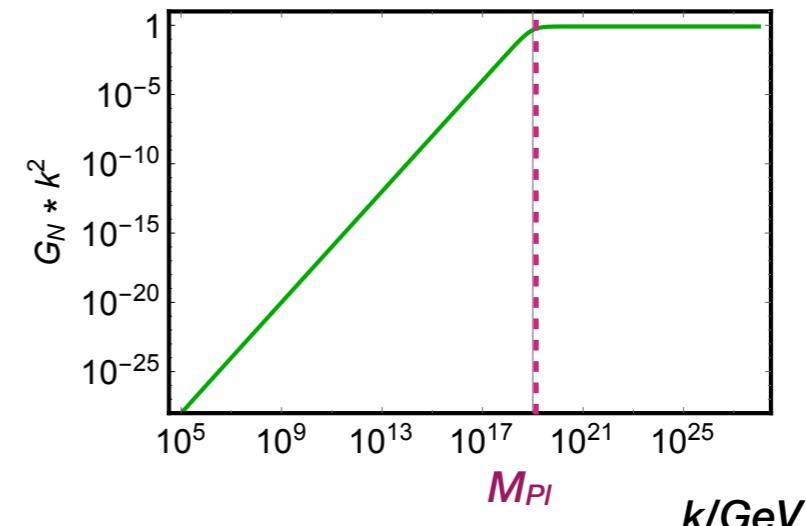
Asymptotic safety:
Quantum field theory for gravity & matter
on all scales



potential consequences:
UV completion for Standard Model
with fewer free parameters:
top-mass value explained,
mass-difference to bottom generated

outlook:

- quantitative convergence
- what about the other parameters of the SM?
- global stability of Higgs potential & link to Higgs inflation



microscopically:
quantum scale- invariance

macroscopically:
predictions for irrelevant couplings

