

Simplified Models for freeze-in at stronger coupling

Giorgio Arcadi

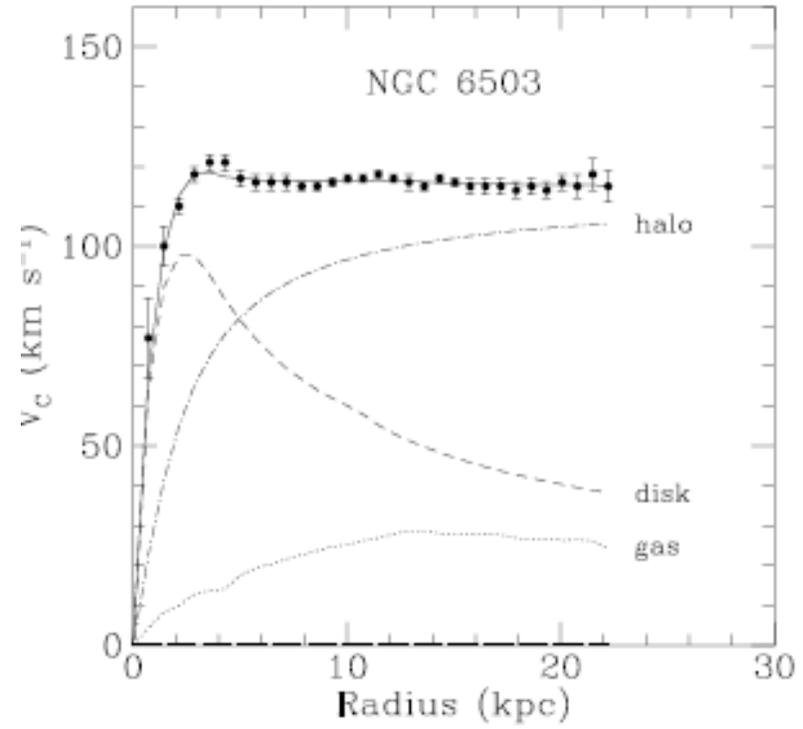
University of Messina

Based on

G.A., F. Costa, A. Goudelis, O. Lebedev JHEP 07 (2024) 044

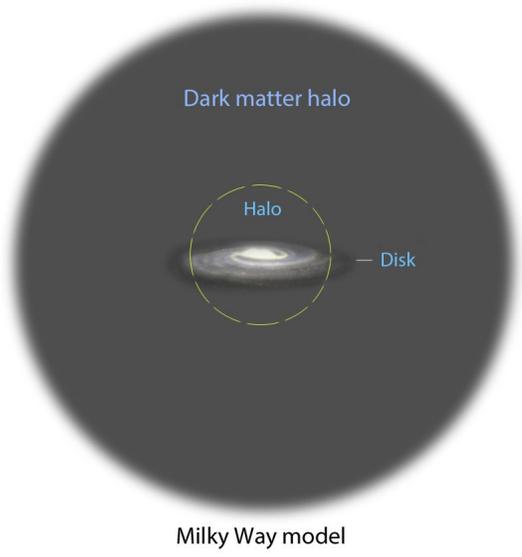
G.A. , D. Cabo-Almeida, O. Lebedev, arXiv:2409.02191

Compelling astrophysical evidence of a Dark Matter Component of the Universe



Explanation:

The galaxy is immersed into a dark matter halo with suitable mass density.

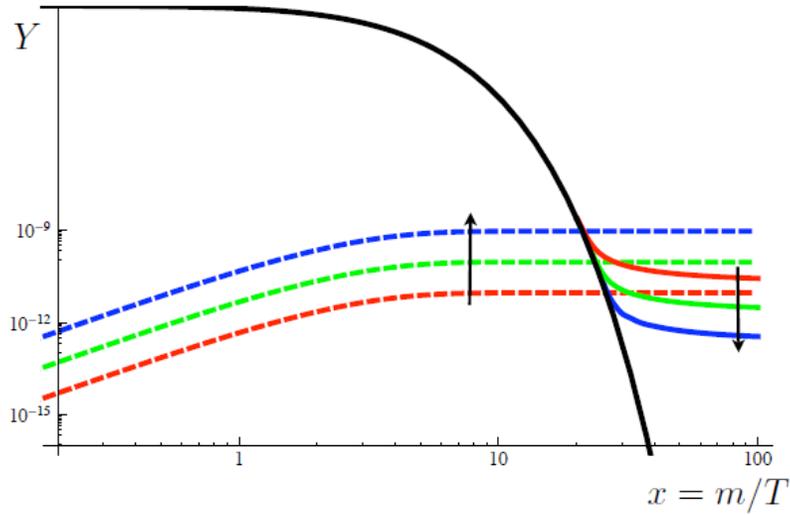


Galactic rotation curves: Data show a flat profile for the velocity well outside visible size of the galaxy

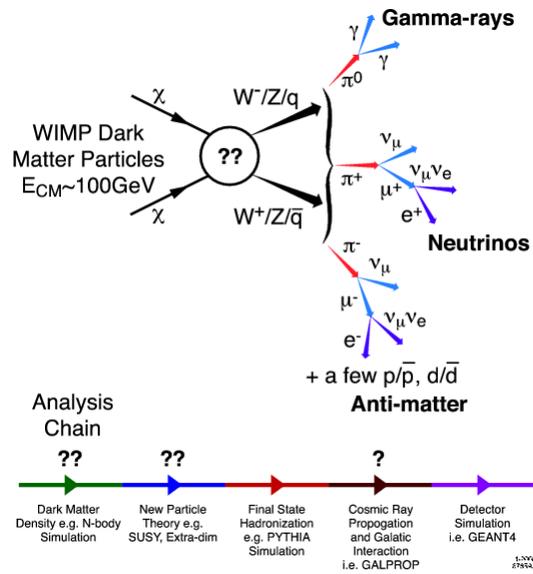


Gravitational lensing (Bullet cluster)

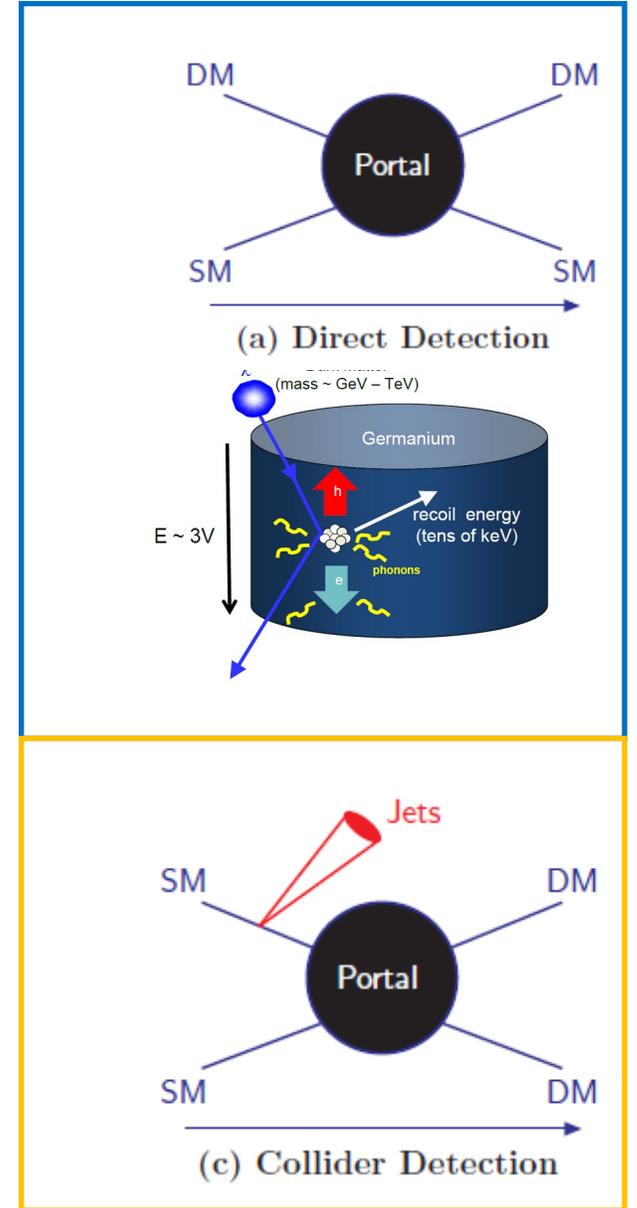
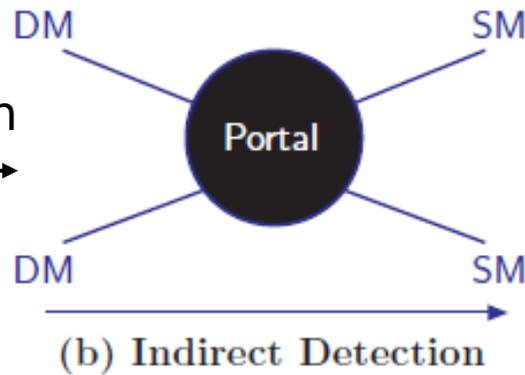
Interest in particle candidates with production mechanism connected to experimental observables



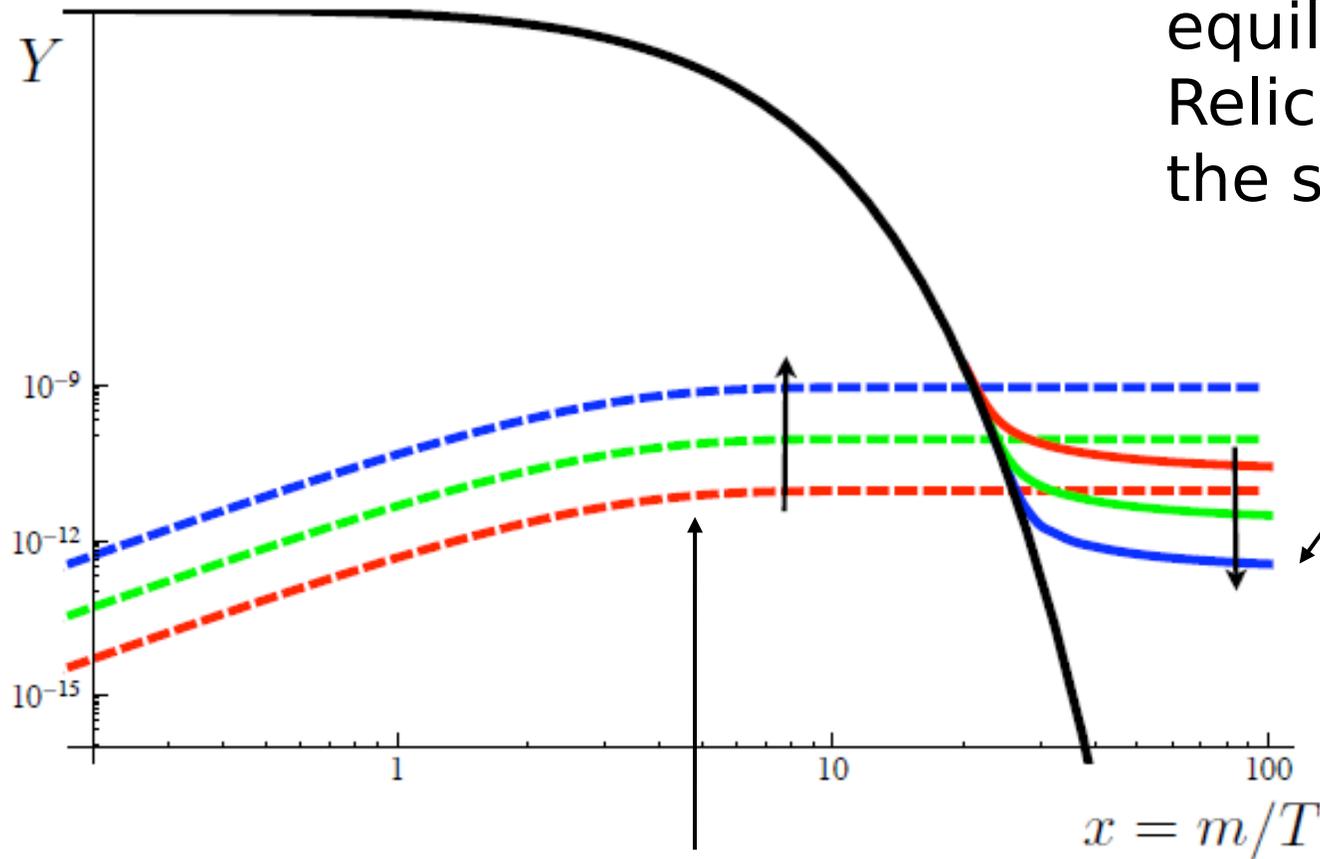
Relic Density



Indirect Detection



Freeze-out: DM initially in thermal equilibrium.
Relic density inversely proportional to the size of DM interactions.

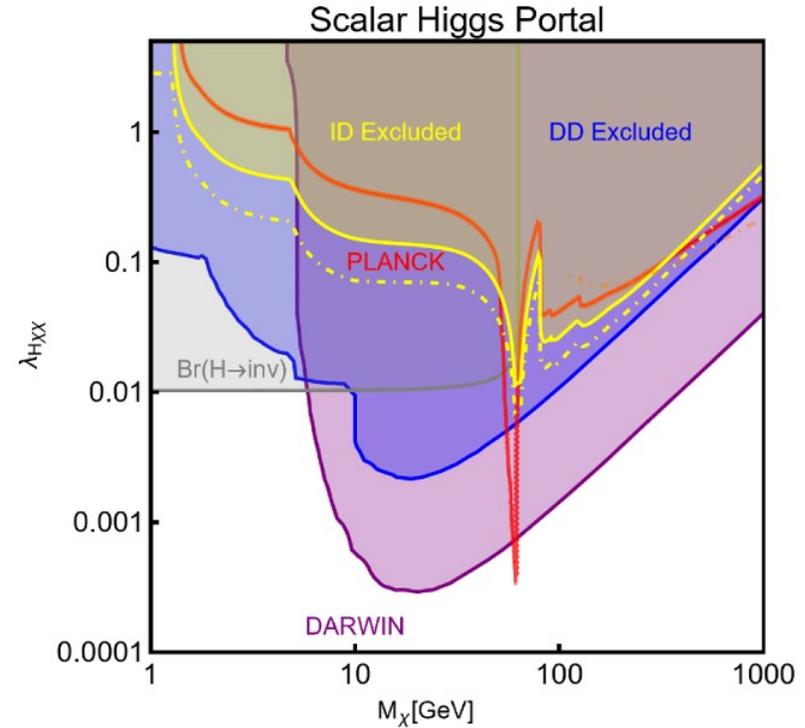


Freeze-in: DM has initially negligible abundance. DM never in thermal equilibrium and continuously produced out from thermal bath particles. Relic density proportional to size of DM interactions.

Consider the simplest Higgs portal $\Delta L_{Scalar, DM} = \frac{1}{2} \lambda_{hs} H^\dagger H s^2$

Freeze out: Correct relic density matched for

 Very constrained scenario



G.A. et al, 2403.15860

Freeze-in: Correct relic density matched for

$$\sigma_{DM,p}^{SI} \approx \frac{\mu_{DM,p}^2}{\Pi} |f_N|^2 \frac{\lambda_{hs,Fi}^2}{M_h^4} \approx 1.5 \times 10^{-59} \text{ cm}^2$$

These scenarios implicitly assume .

Freeze-in at stronger coupling

Consider . DM produced by interactions with primordial plasma (a-la-freeze-in). All the interactions are Boltzmann suppressed because of the low temperature.

$$\lambda_{hS,FI} \rightarrow \bar{\lambda}_{hS,FI} e^{-2m_{DM}/T_R}$$

Correct relic density potentially matched for



low current limits but potential benchmark for next generation experiments



Higgs Portal

$$\Delta L_{Scalar,DM} = \frac{1}{2} \lambda_{hs} H^\dagger H s^2$$

$$\Delta L_{Fermion,DM} = \frac{1}{\Lambda} H^\dagger H \bar{\chi} \chi + \frac{1}{\Lambda_5} i H^\dagger H \bar{\chi} \gamma_5 \chi$$

$$\Delta L_{Vector,DM} = \frac{1}{2} \lambda_{h\nu} H^\dagger H V^\mu V_\mu$$

$$\dot{n}_h + 3H n_h = \Gamma_{\bar{f}f \rightarrow h} - \Gamma_{h \rightarrow \bar{f}f}$$

Approximate rule of thumb

$$3H n_h \lesssim \Gamma_{\bar{f}f \rightarrow h}, \Gamma_{h \rightarrow \bar{f}f}$$

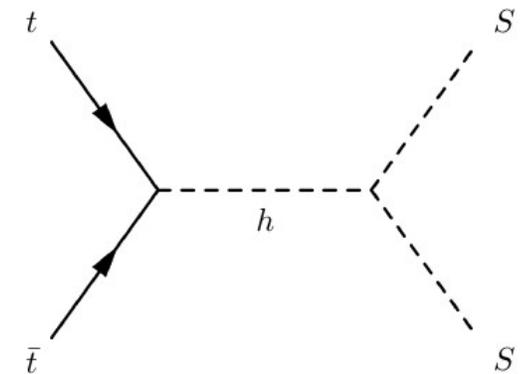
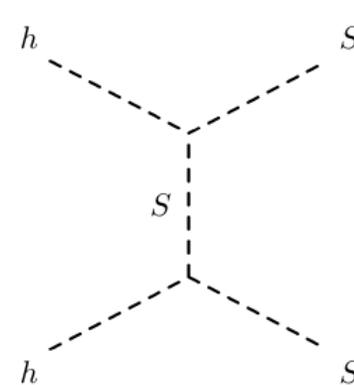
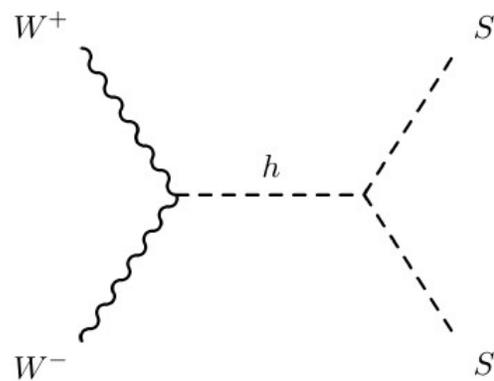
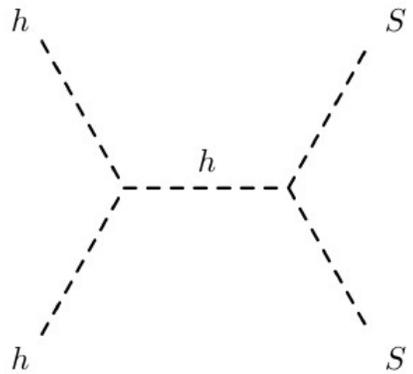
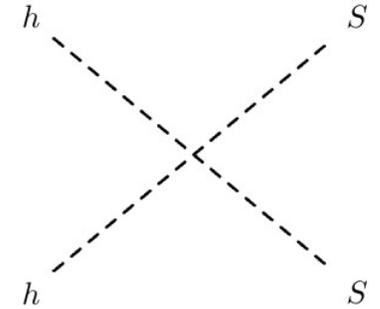
$$n_h = \left(\frac{m_h T}{2\pi} \right)^{3/2} e^{-m_h/T}$$

$$\Gamma_{\bar{f}f \rightarrow h} = \int \left(\prod_i \frac{d^3 p_i}{(2\pi)^3 2E_f} f(p_i) \right) \frac{d^3 p_f}{(2\pi)^3 2E_f} |M_{2 \rightarrow 1}|^2 (2\pi)^4 \delta^{(4)}(\sum p_i - p_f) \approx 10^{-2} \times y_f^2 m_h^{\frac{5}{2}} T^{\frac{3}{2}} e^{-m_h/T}$$

$y_f \sqrt{m_h M_{Pl}} \gtrsim T$ Always satisfied above the QCD phase transition

Detailed case of study: Scalar DM

$$\dot{n}_s + 3H n_s = 2\Gamma(SM \rightarrow ss) - 2\Gamma(ss \rightarrow SM)$$



Heavy DM regime: . DM production dominated by (inverse) annihilations into Higgs bosons.

$$\Gamma(hh \rightarrow ss) = \frac{1}{(2\pi)^6} \int \tilde{\sigma} v_r e^{-\frac{(E_1 + E_2)/T}{d^3 p_1 d^3 p_2}} = \frac{2\pi^2 T}{(2\pi)^6} \int_{4m_s^2}^{+\infty} ds \tilde{\sigma}(s) (s - 4m_h^2) \sqrt{s} K_1(\sqrt{s}/T)$$

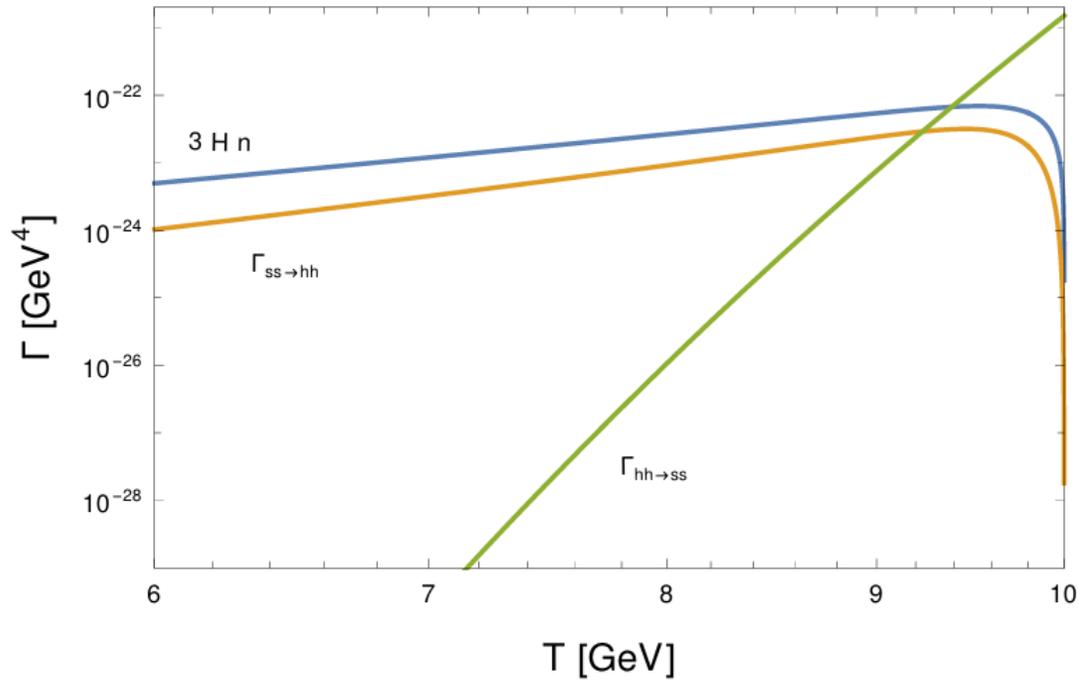
$$\tilde{\sigma} = \frac{\lambda_{hs}^2}{16\pi s} \sqrt{\frac{s - 4m_s^2}{s - 4m_h^2}}$$

$m_s \gg T$ \rightarrow $\Gamma(hh \rightarrow ss) \simeq \frac{\lambda_{hs}^2 T^3 m_s}{256 \pi^4} e^{-2m_s/T}$

annihilation rate is computed in the hypothesis of kinetic equilibrium

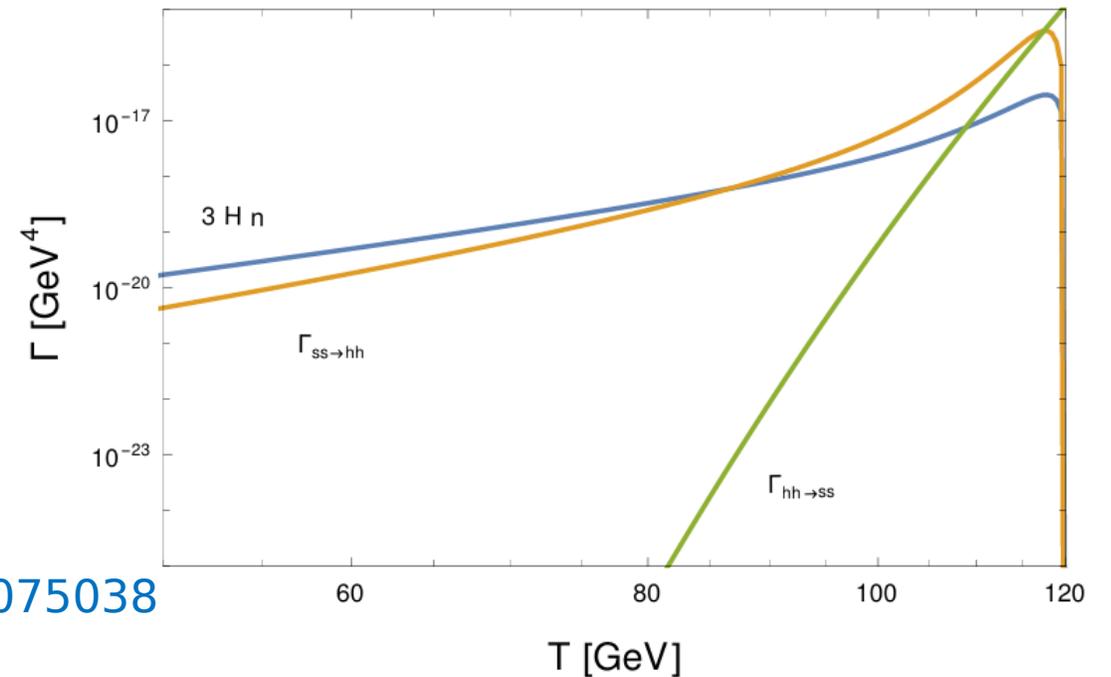
$$\Gamma(ss \rightarrow hh) = \sigma(ss \rightarrow hh) v_r n_{DM}^2$$


$$\sigma(ss \rightarrow hh) v_r = \frac{\lambda_{hs}^2}{16 \pi m_s^2}$$



$$T_R = 10 \text{ GeV}, m_s = 223.5 \text{ GeV}, \lambda_{hs} = 0.05$$

$$T_R = 120 \text{ GeV}, m_s = 2.81 \text{ TeV}, \lambda_{hs} = 0.82$$



For small enough couplings one can neglect (pure freeze-in regime)

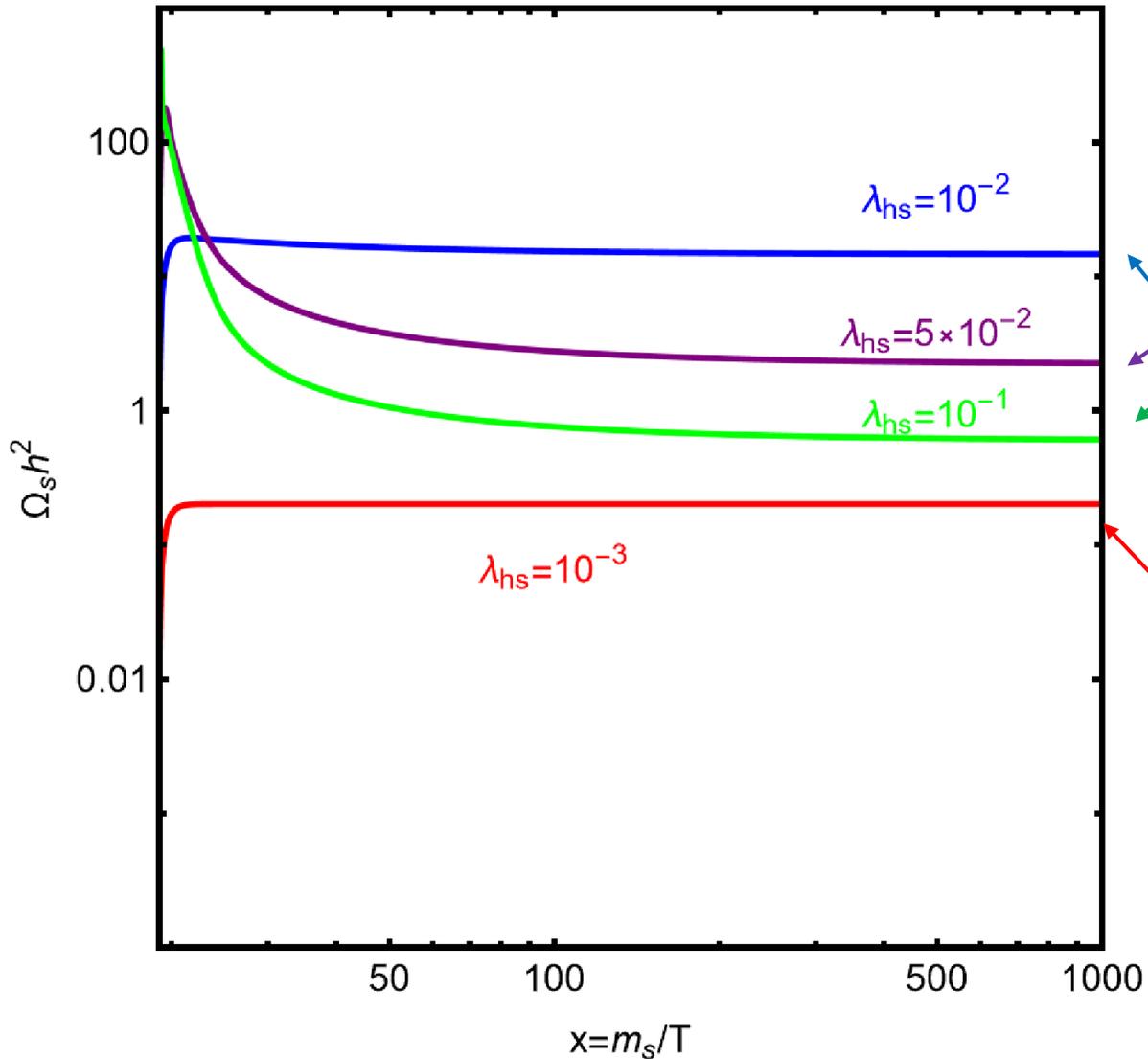
$$\frac{dY_{DM}}{dT} = -2 \frac{\Gamma(hh \rightarrow ss)}{HT s_{SM}} \quad s_{SM} = \frac{2\pi^2}{45} g_* T^3 \quad H = 1.66 \sqrt{g_*} \frac{T^2}{M_{Pl}}$$

→ $Y_{DM} \approx 2.5 \times 10^{-7} \frac{\lambda_{hs}^2 M_{Pl}}{T_R} e^{-2m_s/T_R}$

Correct relic density achieved for:

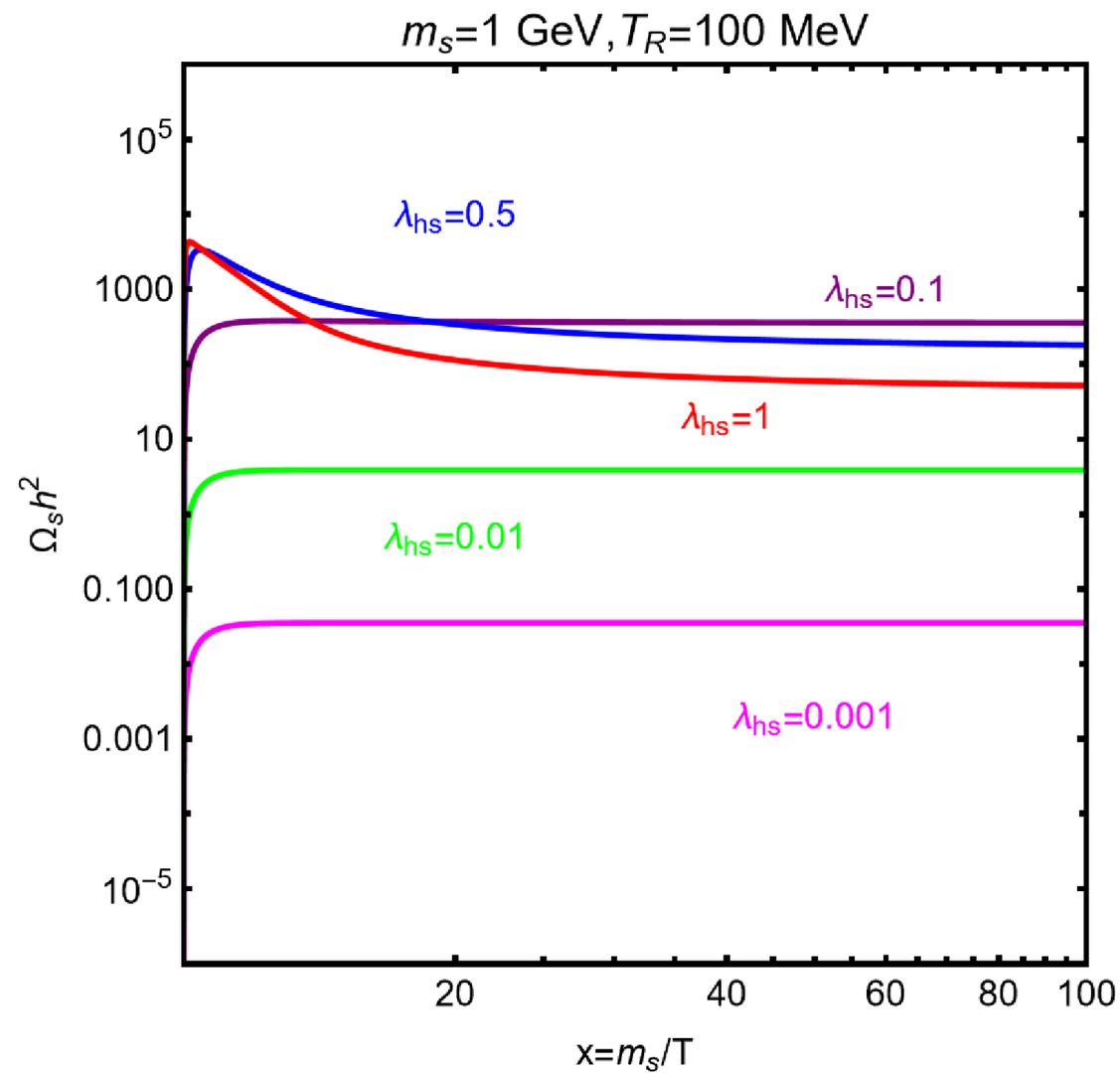
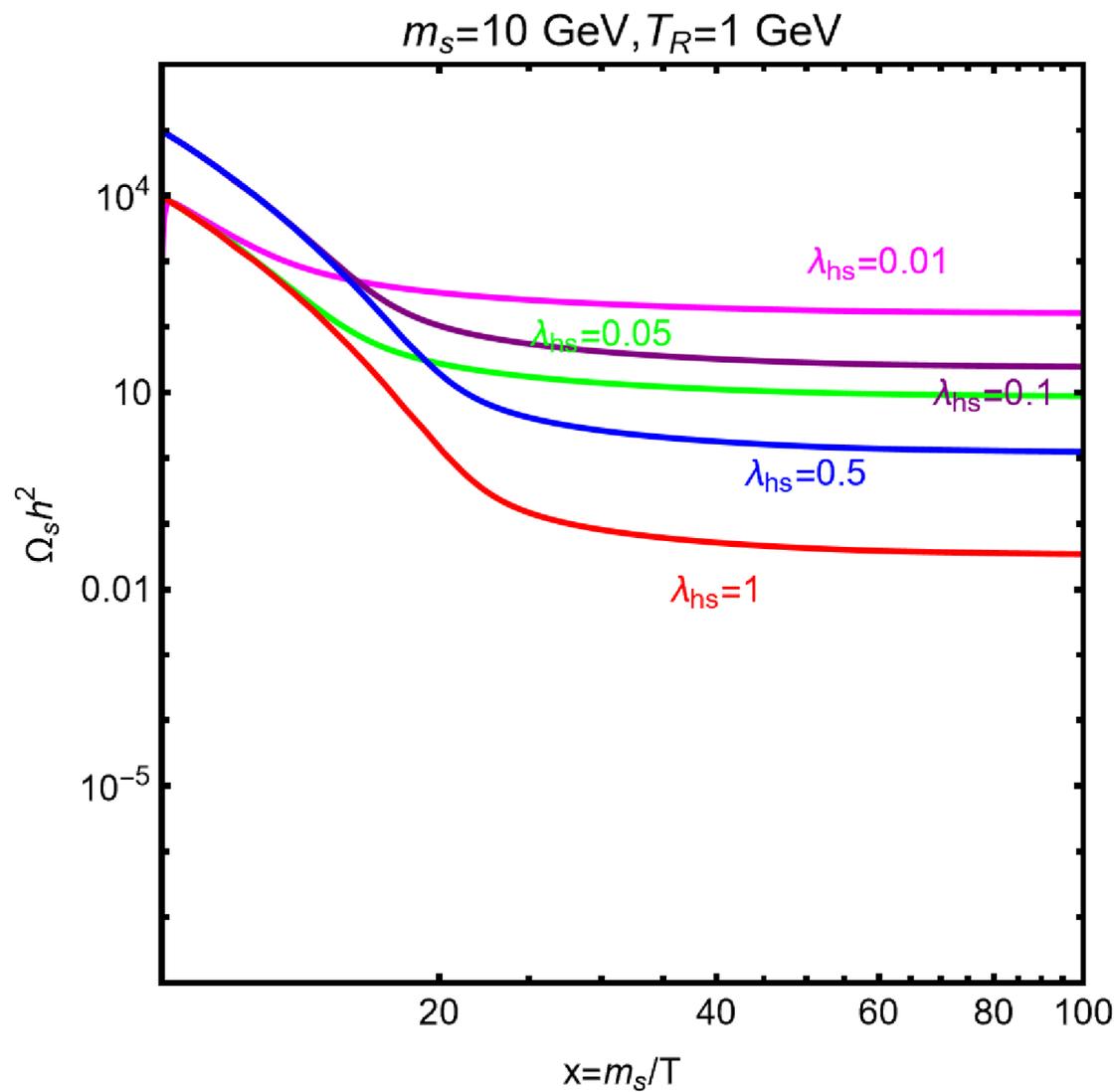
$$\lambda_{hs} \approx 3 \times 10^{-11} e^{m_s/T_R} \sqrt{\frac{T_R}{m_s}}$$

$m_s=570 \text{ GeV}, T_R=30 \text{ GeV}$



Higher coupling: DM production is efficient enough to activate annihilation processes. The relic density decreases with the coupling.

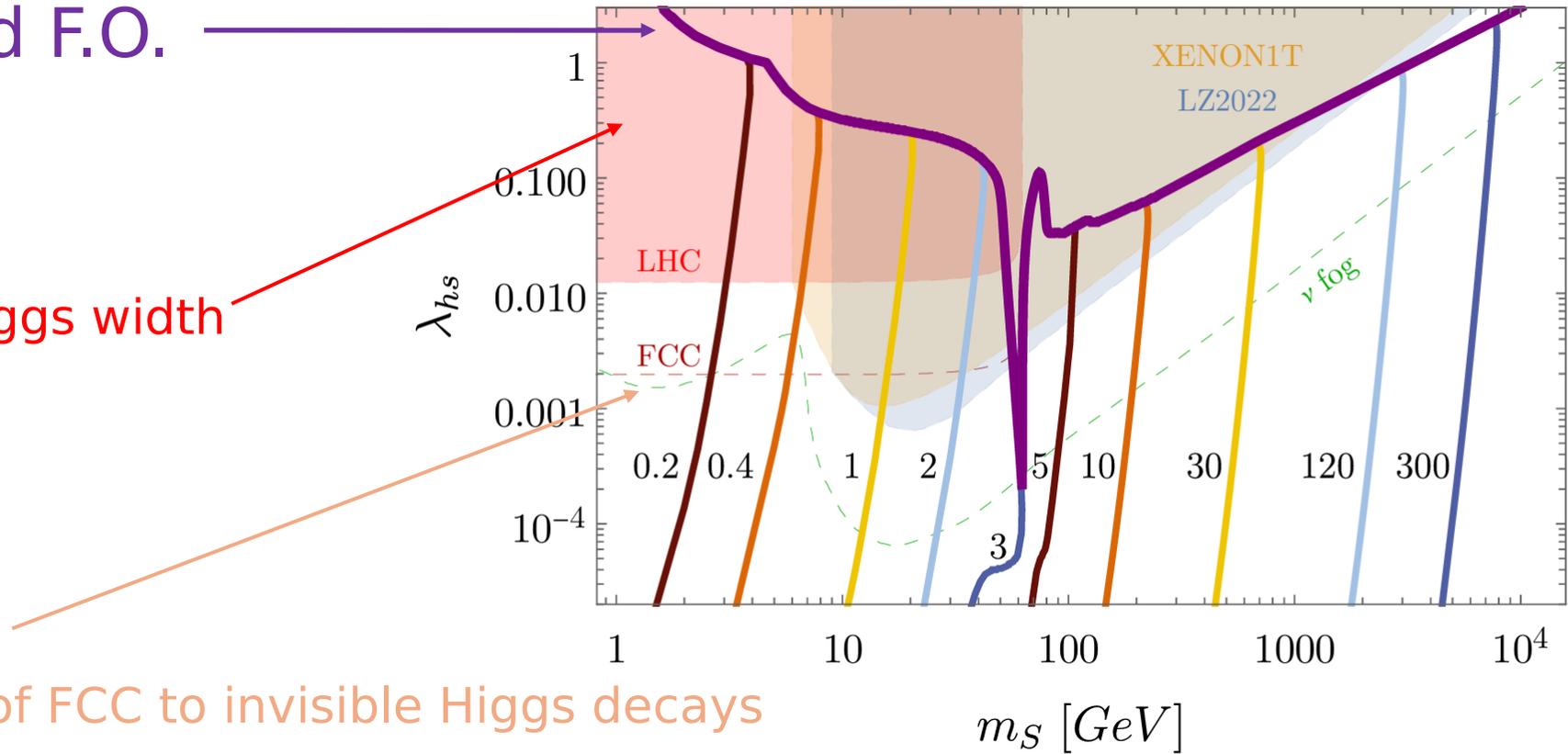
Low couplings, pure freeze-in regime. The relic density increases with the coupling.



Scalar Higgs Portal

Standard F.O. →

Limit on Invisible Higgs width →



Projected sensitivity of FCC to invisible Higgs decays →

Fermionic Higgs Portal

$$\frac{1}{\Lambda} i \bar{\chi} \chi H^\dagger H$$

$$\Gamma(hh \rightarrow \chi\chi) \simeq \frac{3}{16\pi^4} \frac{m_\chi^2 T^4}{\Lambda^2} e^{-\frac{2m_\chi}{T}}$$

$$Y \simeq 1.2 \times 10^{-5} \frac{m_\chi M_{Pl}}{\Lambda^2} e^{-\frac{2m_\chi}{T_R}}$$

Correct Relic Density for:

$$\frac{m_\chi}{\Lambda} e^{-m_\chi/T_R} \simeq 4 \times 10^{-12}$$

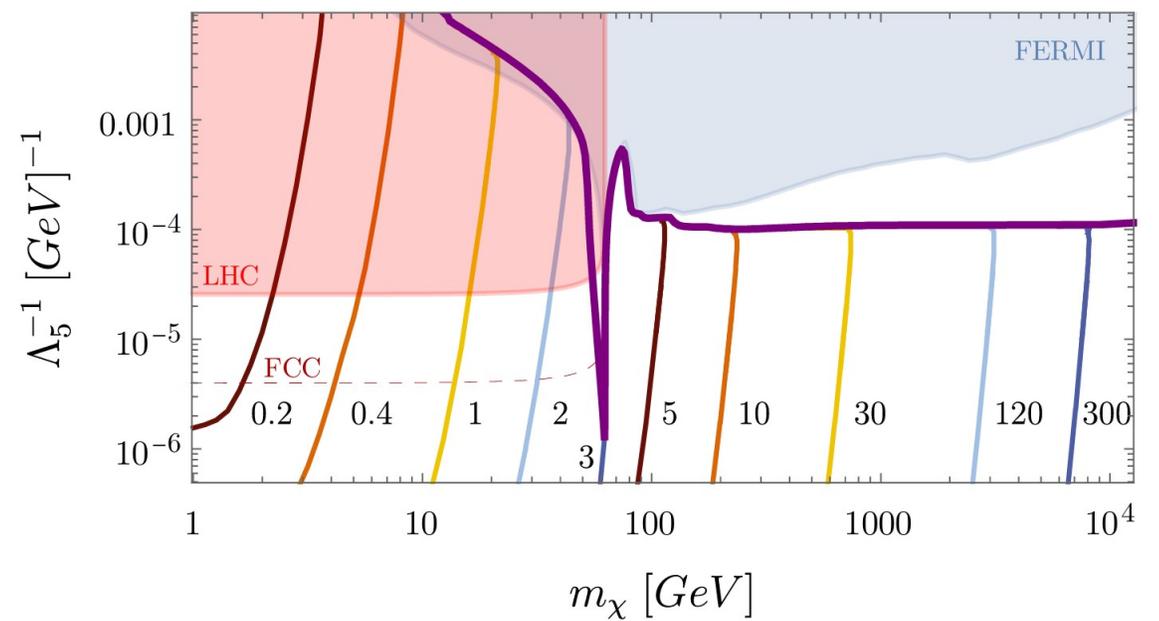
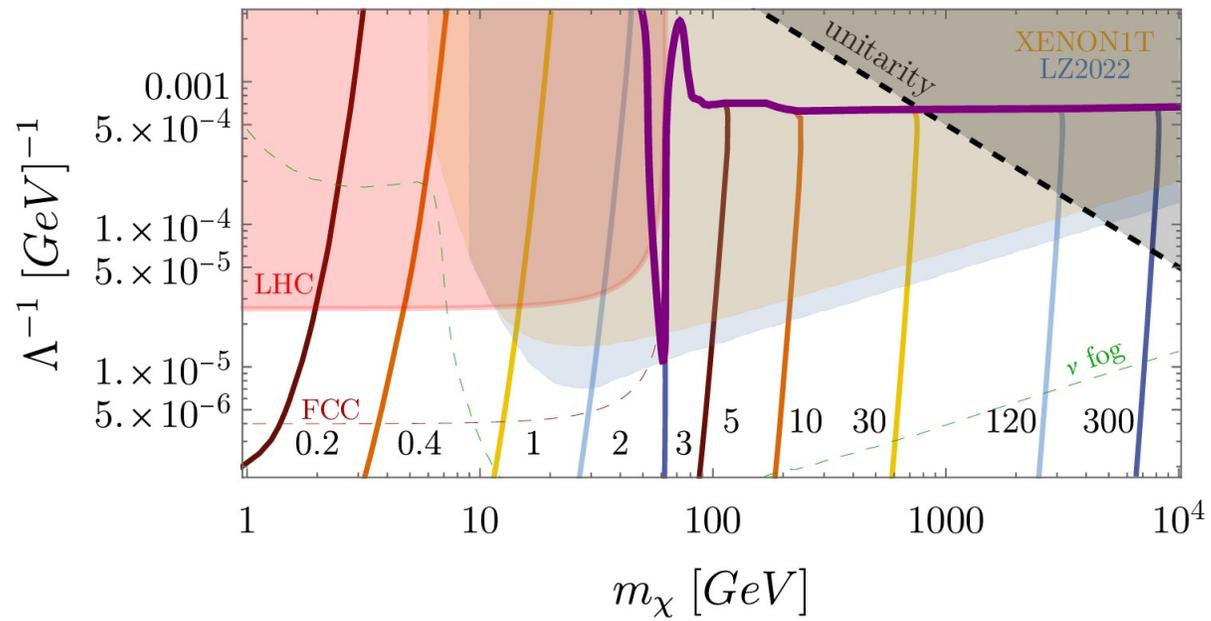
$$\frac{1}{\Lambda_5} i \bar{\chi} \gamma_5 \chi H^\dagger H$$

$$\Gamma(hh \rightarrow \chi\chi) \simeq \frac{1}{8\pi} \frac{m_\chi^3 T^3}{\Lambda_5^2} e^{-\frac{2m_\chi}{T}}$$

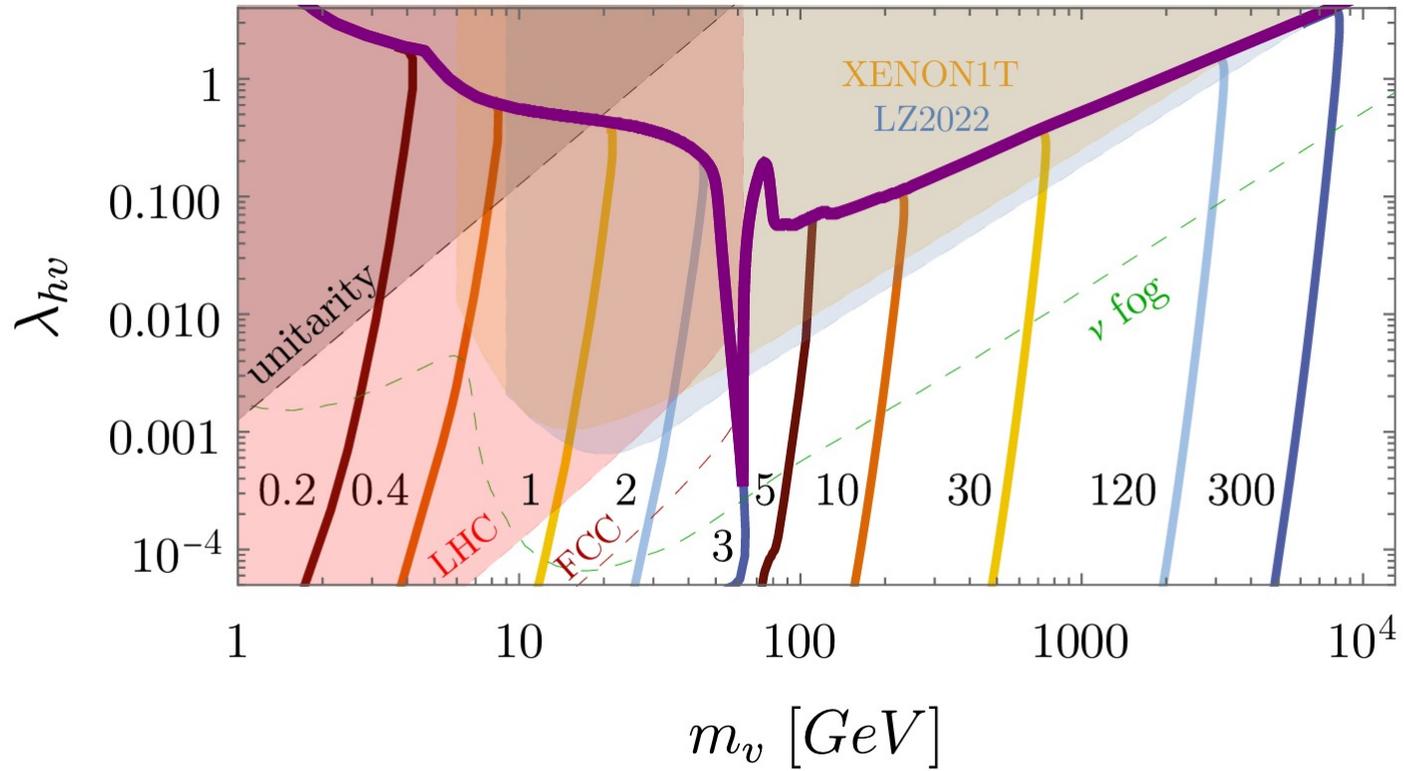
$$Y \simeq 8 \times 10^{-6} \frac{m_\chi^2 M_{Pl}}{\Lambda_5^2 T_R} e^{-\frac{2m_\chi}{T_R}}$$

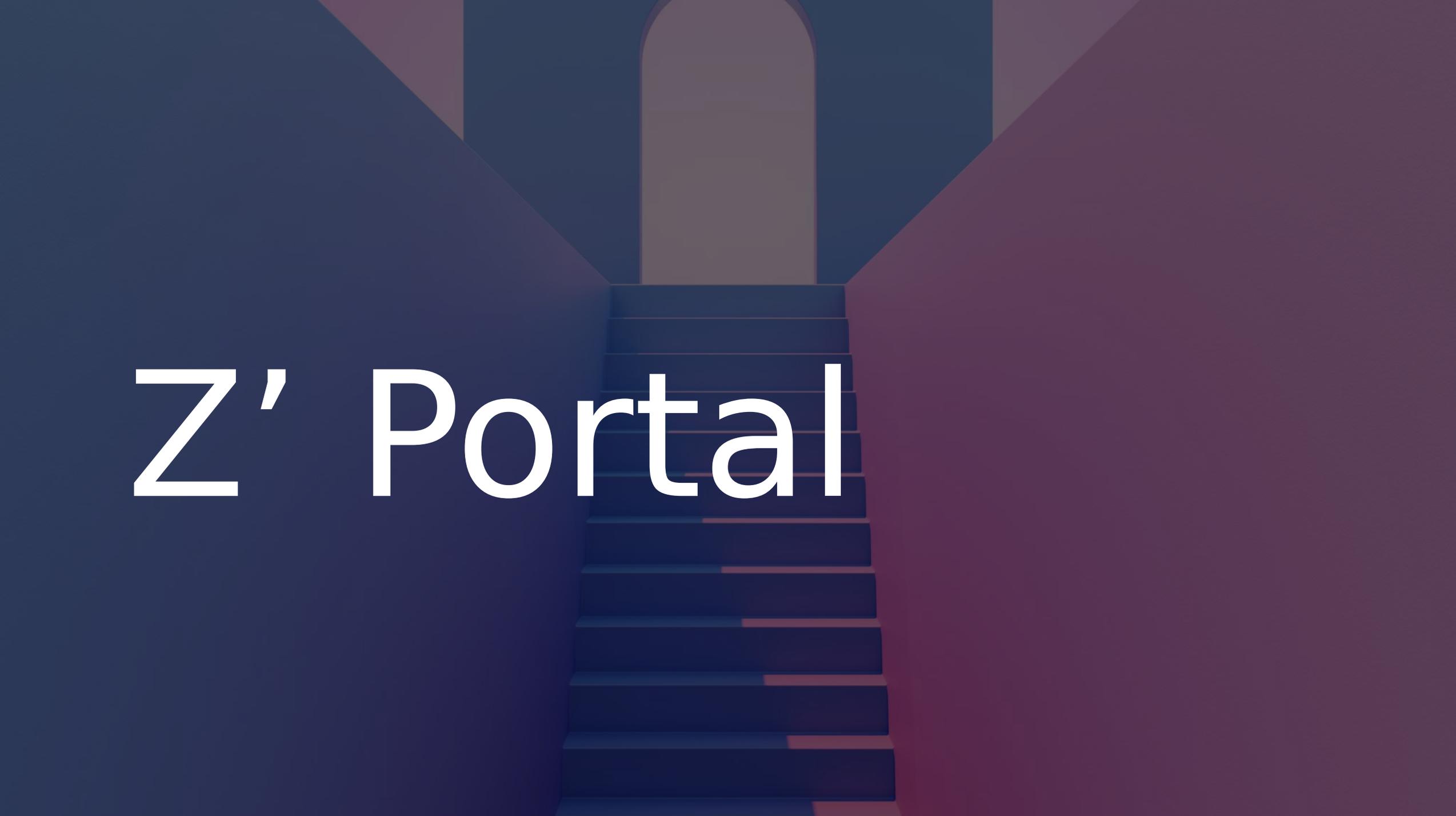
Correct Relic Density for:

$$\frac{m_\chi^{3/2}}{\Lambda_5 T_R^{1/2}} e^{-m_\chi/T_R} \simeq 5 \times 10^{-12}$$



Vector Higgs Portal





Z' Portal

$$\dot{n}_\chi + 3H n_\chi = 2\Gamma(\bar{f} f \rightarrow \bar{\chi} \chi) - 2\Gamma(\bar{\chi} \chi \rightarrow \bar{f} f)$$

only vectorial couplings pure freeze-in regime

$$\Gamma(\bar{f} f \rightarrow \bar{\chi} \chi) = \frac{V_f^2 V_\chi^2 m_\chi^5 T^3}{2 \pi^4 M_{Z'}^4} e^{-2m_\chi/T_R}$$

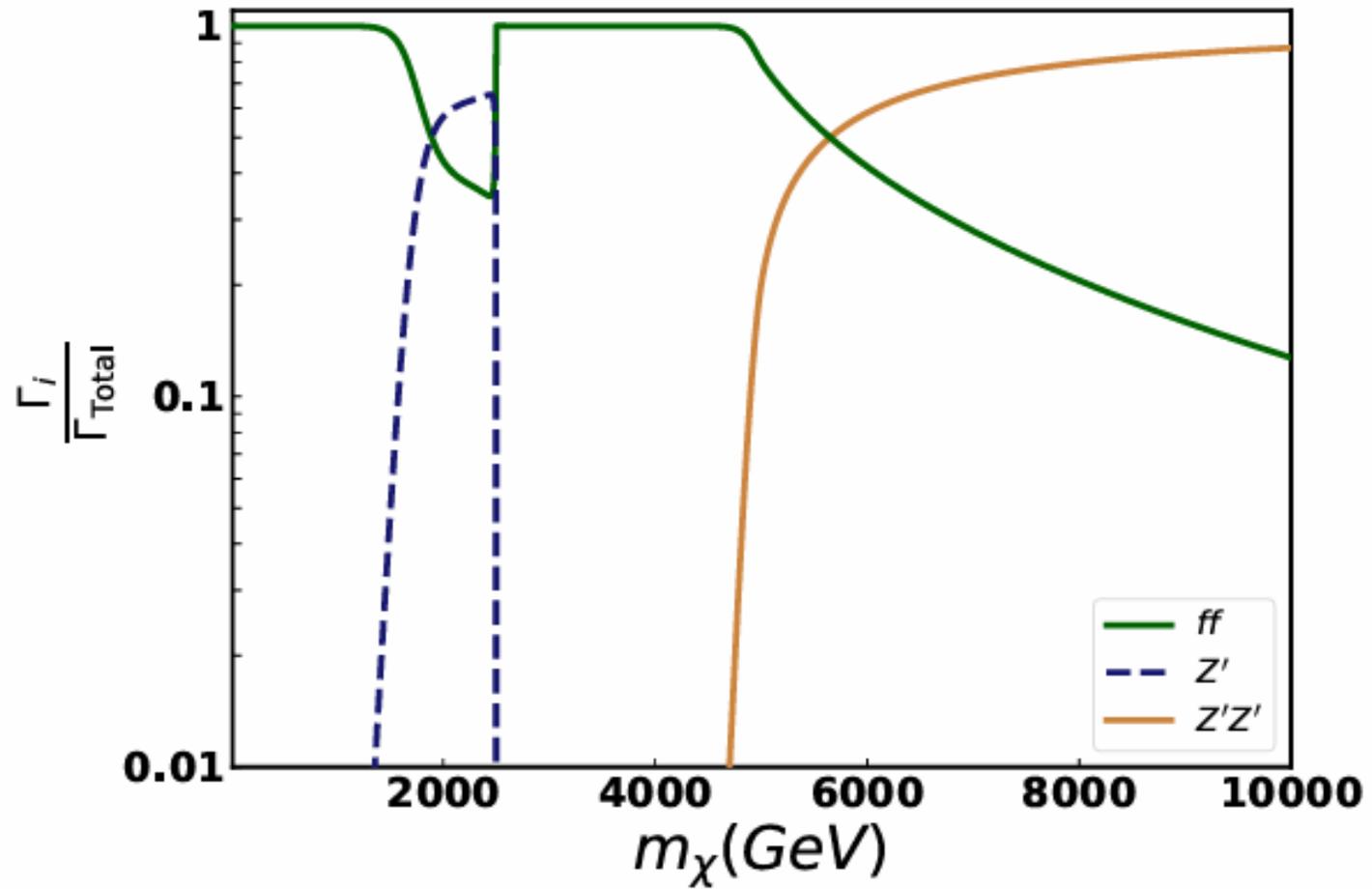
$$\longrightarrow Y_\chi \approx 3.3 \times 10^{-5} \sum_f \frac{V_f^2 V_\chi^2 m_\chi^4}{M_{Z'}^4} \frac{M_{Pl}}{T_R} e^{-2m_\chi/T_R}$$

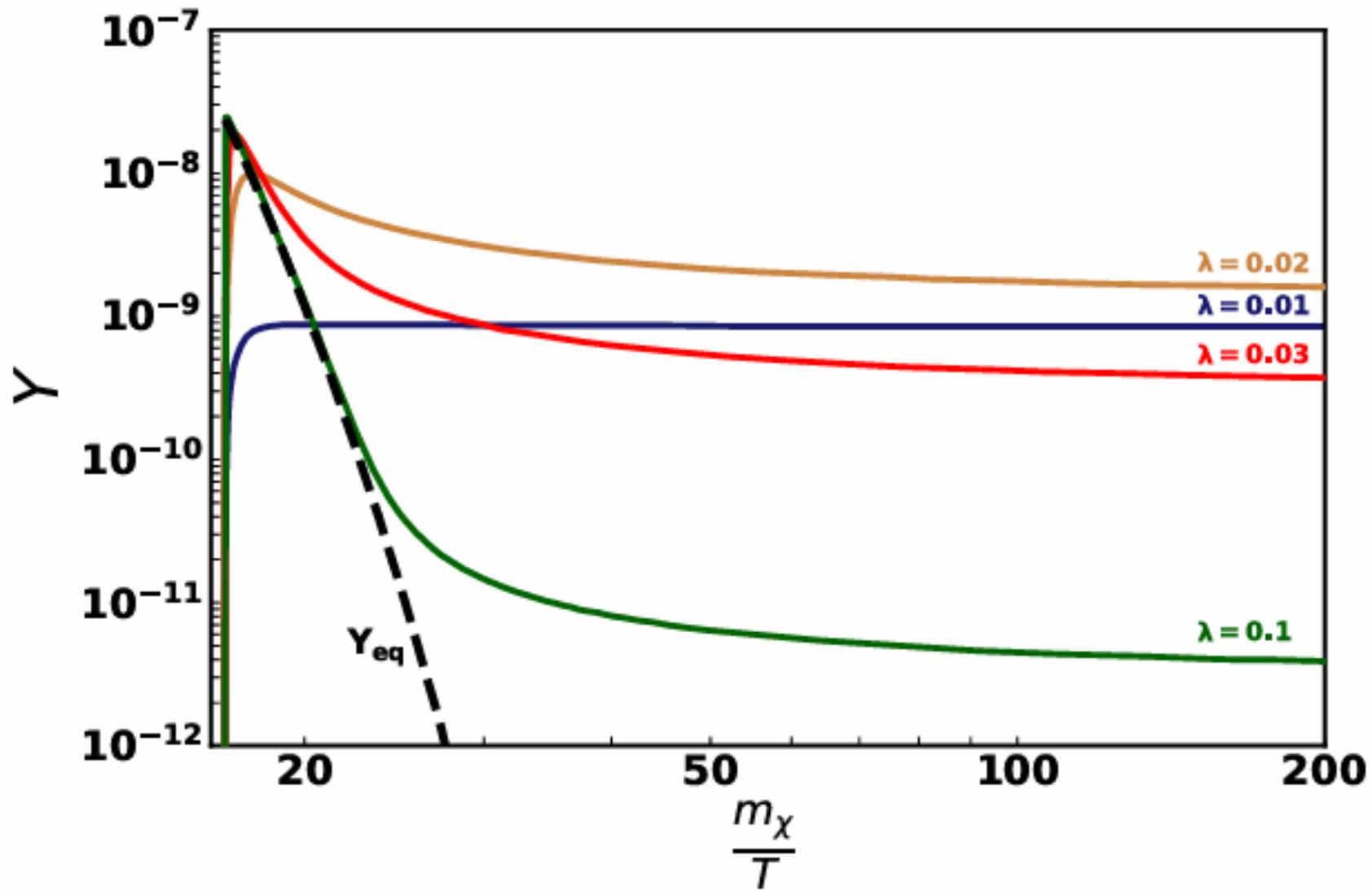
Assuming $V_\chi = V_f = \lambda$

$$Y_{\chi,obs} = 4.4 \times 10^{-10} \left(\frac{GeV}{m_\chi} \right) \longrightarrow \lambda \approx 10^{-6} \frac{M_{Z'} T_R^{\frac{1}{4}}}{m_\chi^{\frac{5}{4}}} e^{m_\chi/(2T_R)}$$

Analogous result for purely axial couplings.

$$M_{Z'} = 5 \text{ TeV}$$

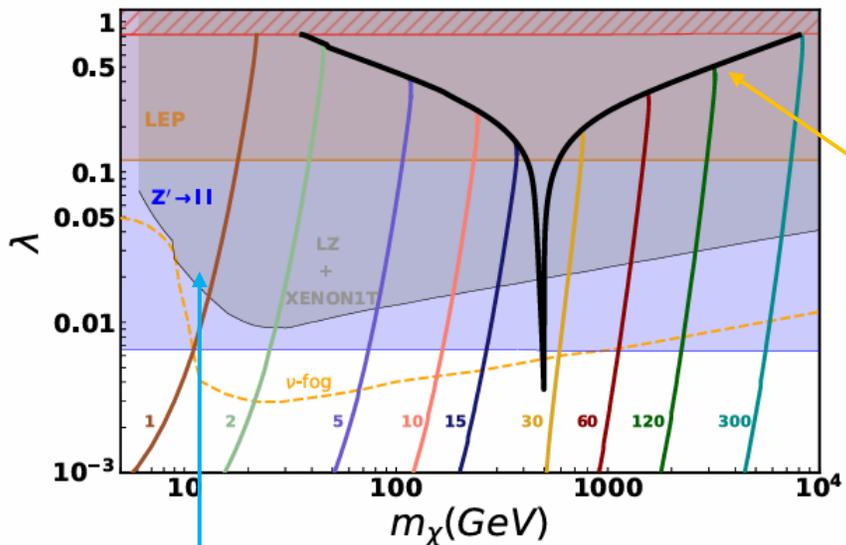




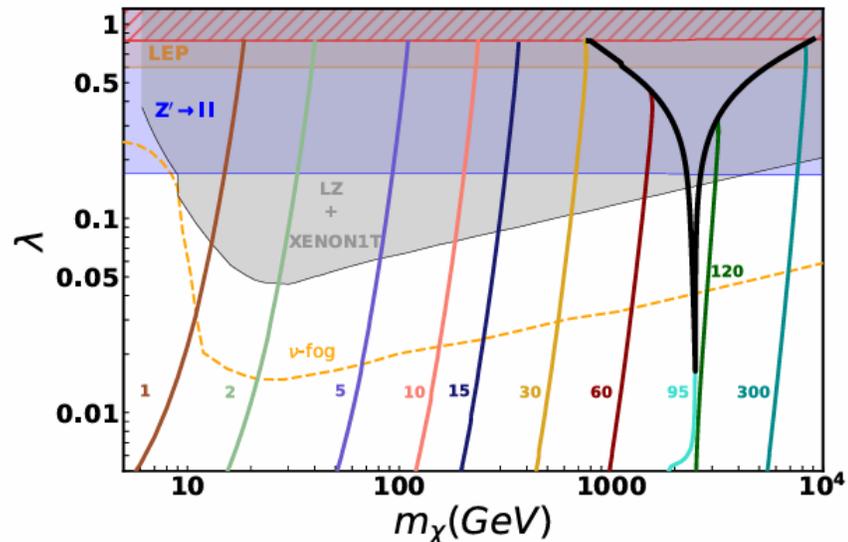
$M_{Z'}=1\text{TeV}$

Vectorial Couplings

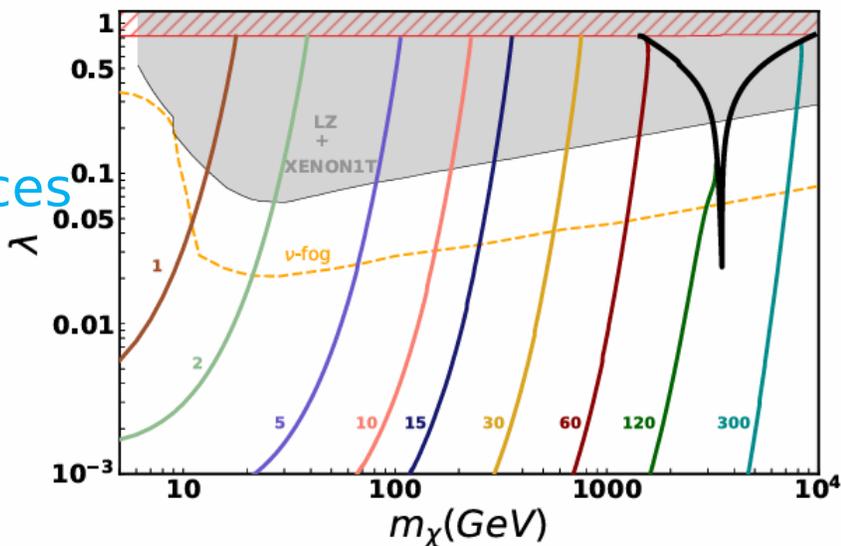
$M_{Z'}=5\text{TeV}$



$$\frac{\lambda}{M_{Z'}} < \frac{0.12}{\text{TeV}} \quad (\text{LEP})$$



$M_{Z'}=7\text{TeV}$

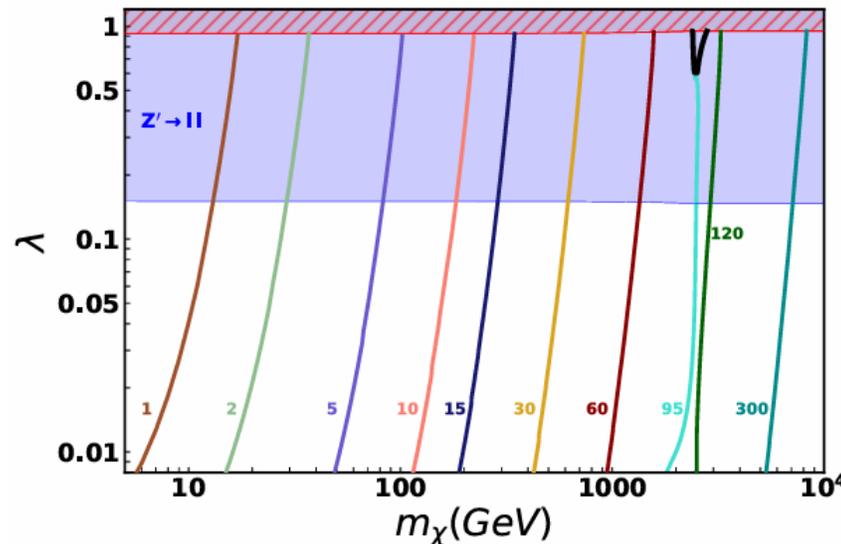
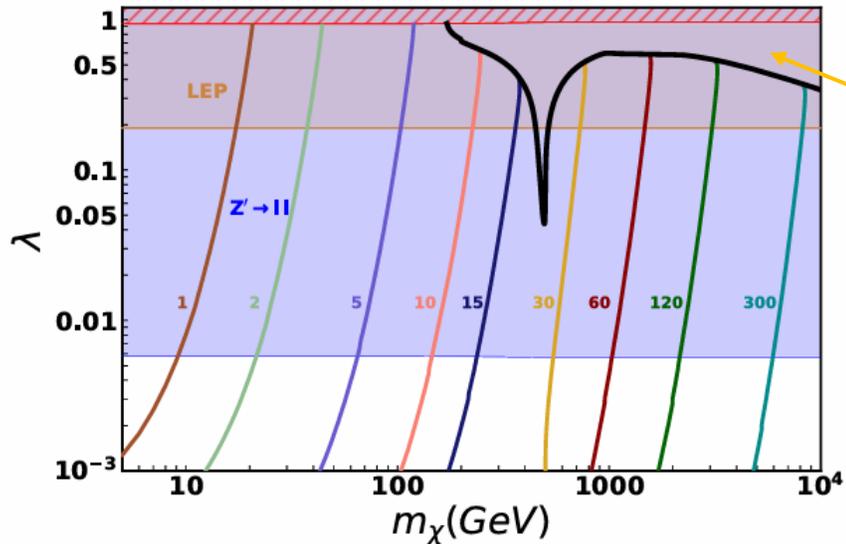


LHC searches of dilepton resonances

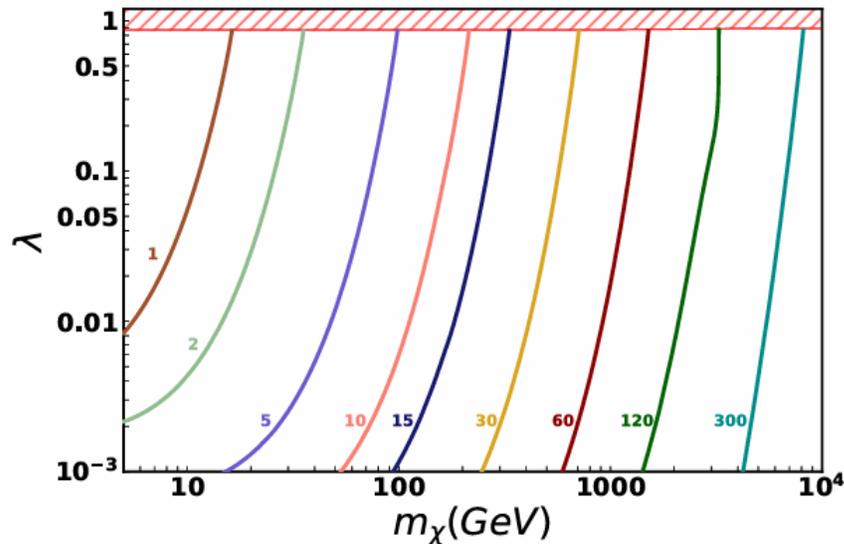
$M_{Z'}=1\text{TeV}$

Axial Couplings

$M_{Z'}=5\text{TeV}$



$M_{Z'}=7\text{TeV}$



Conclusions

We have studied the scenario of «freeze-in at stronger couplings» in the case of Higgs and Z' portals.

Freeze-in at stronger coupling allows to a viable relic density solutions testable at future detectors.

The parameter space is enlarged, with respect to the conventional freeze-out, allowing for light DM masses. In the case of Higgs portals we have benchmarks for future updates in the sensitivity to invisible decays of the Higgs boson.