

Internal calibration of the Stereo Detector



MAX-PLANCK-INSTITUT
FÜR KERNPHYSIK
HEIDELBERG

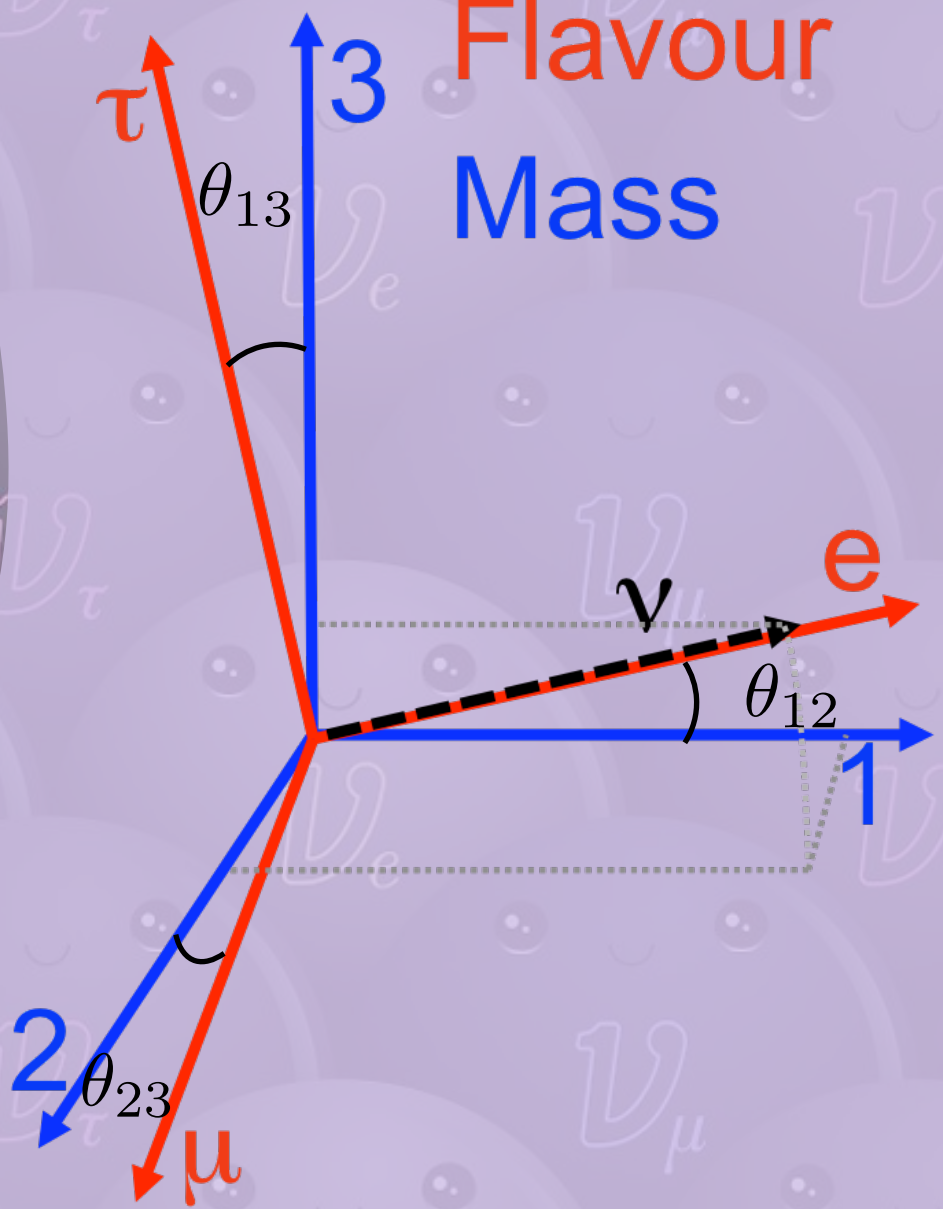
IMPRS Seminar 03.VII.2017

CHRISTIAN ROCA
MAX-PLANCK-INSTITUT FÜR KERNPHYSIK

Neutrino Oscillations 101



Flavour Mass



* angles not at scale

Flavour eigenstates do not correspond to **mass** eigenstates

$$\begin{pmatrix} \nu_e & \nu_\mu & \nu_\tau \end{pmatrix} \neq \begin{pmatrix} \nu_1 & \nu_2 & \nu_3 \end{pmatrix}$$

Defined flavour
(production / detection)

Defined kinematics
(propagation)

Transformation of basis through unitary matrix U

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \cdot \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Matrix elements in U contain:

Oscillation angles

$$\begin{matrix} \theta_{23} & \theta_{12} \\ & \theta_{13} \end{matrix}$$

Are oscillation amplitudes

Squared mass splittings

$$\begin{matrix} m_{23}^2 & m_{12}^2 \\ & m_{13}^2 \end{matrix}$$

Are oscillation frequencies

Reactor plants are powerful sources of $\bar{\nu}_e$ and are commonly used to study neutrino oscillations

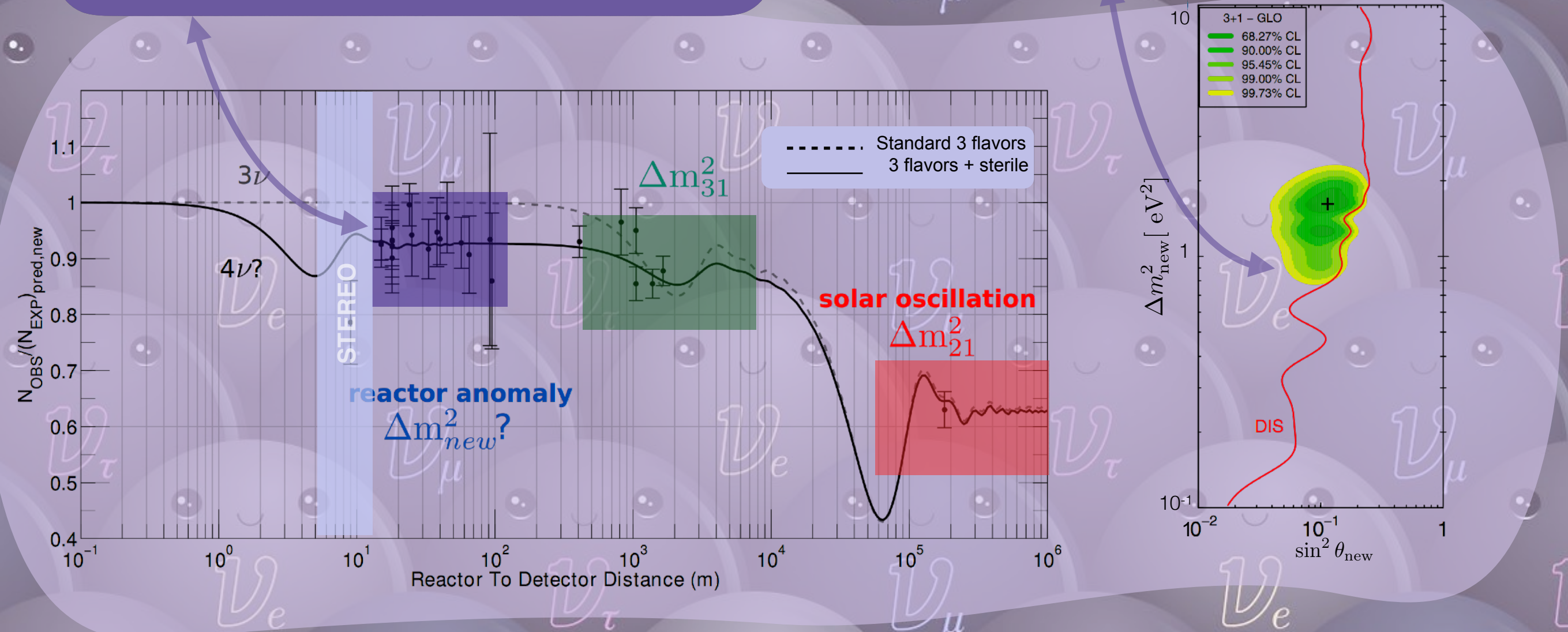
- Reactor Antineutrino Anomaly -

- Reevaluation of reactor $\bar{\nu}_e$ spectra +6% [1]
- Survival $P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = 0.940 \pm 0.024$ is $\sim 3\sigma$ from 1
- Why the deficit? New oscillation into a sterile state?

- Best Fit Parameters -

$$\sin^2(2\theta_{\text{new}}) = 0.11$$

$$|\Delta m_{\text{new}}^2| = 1.6 \text{ eV}^2$$



Stereo Detector

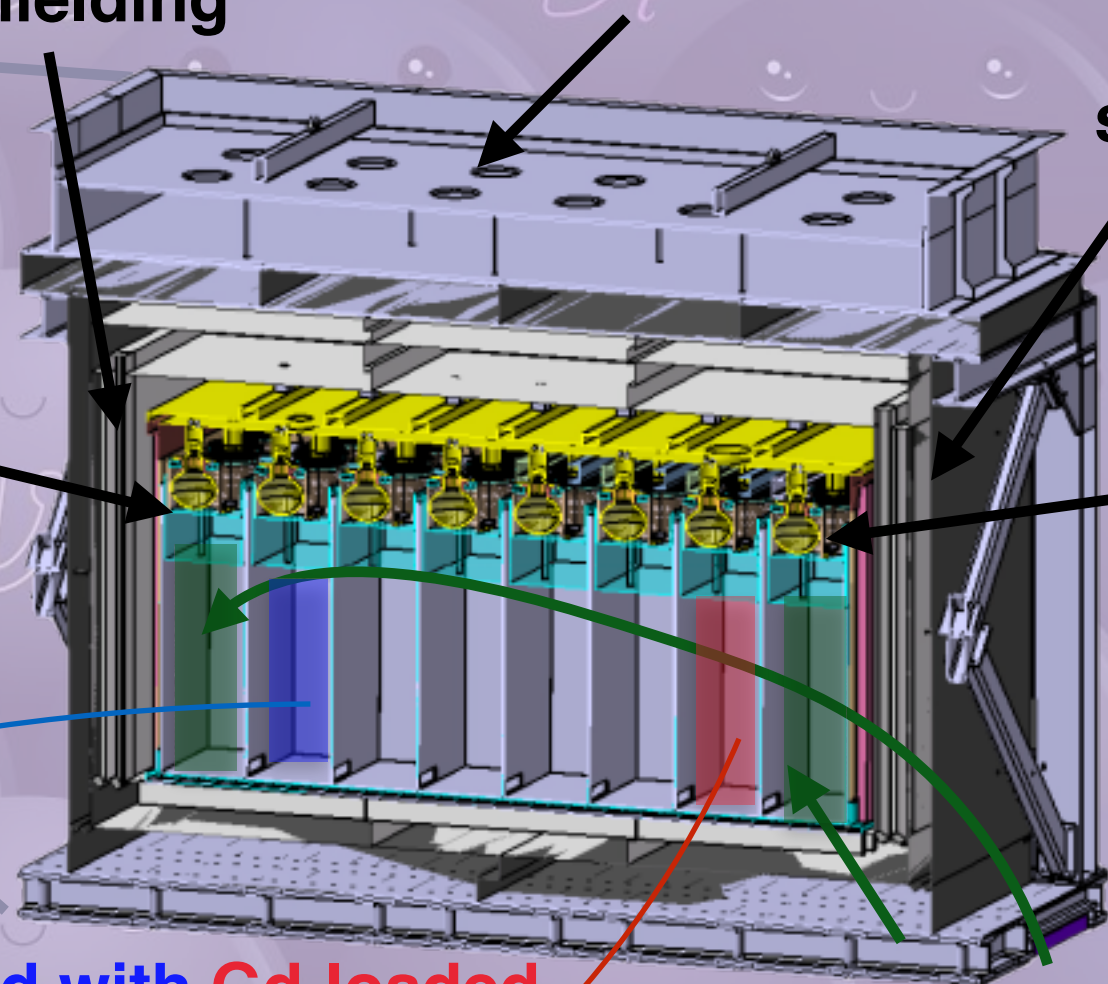
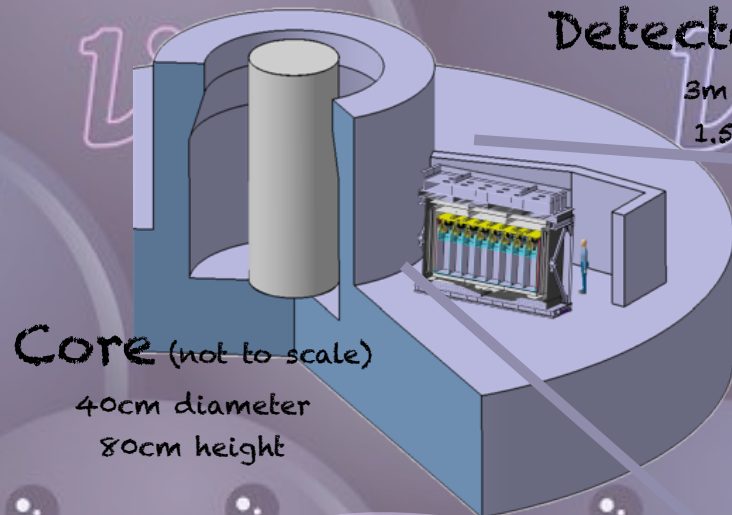
ILL nuclear reactor

Detector (not to scale)
3m diameter
1.5m height

borated polyethylene
shielding

water Cherenkov
muon veto

lead
shielding

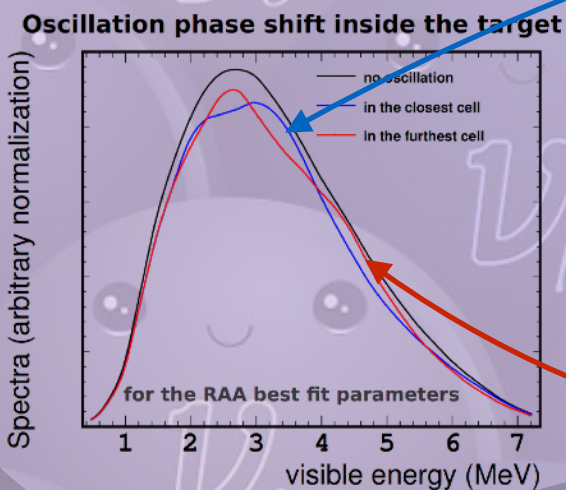
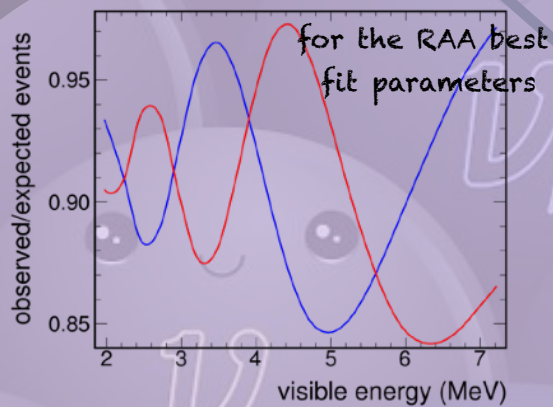


acrylic
buffer

48
PMTs

6 cells filled with Gd-loaded
liquid scintillator (2m³)

outer crown with unloaded
liquid scintillator



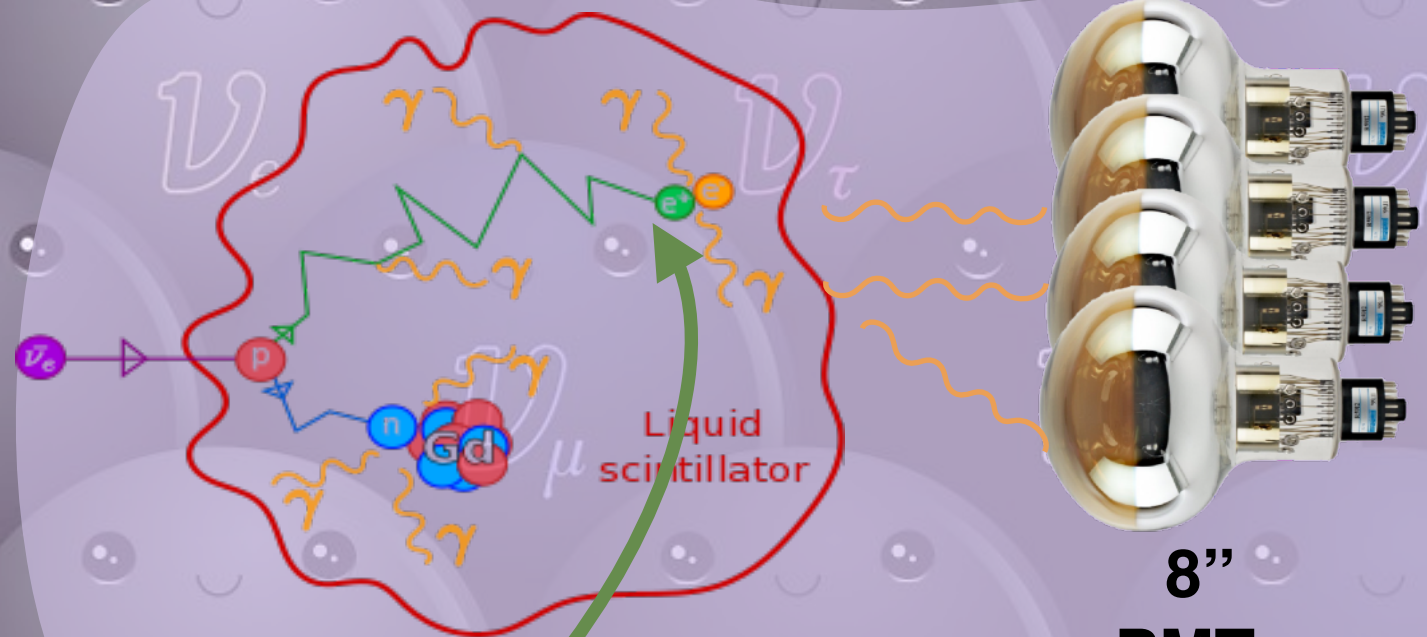
Detection process - Inverse Beta Decay (IBD)

Neutrinos interact via IBD:



Two signals in coincidence:

- Prompt signal: energy reconstruction
- Delayed signal: coincidence with prompt guarantees IBD



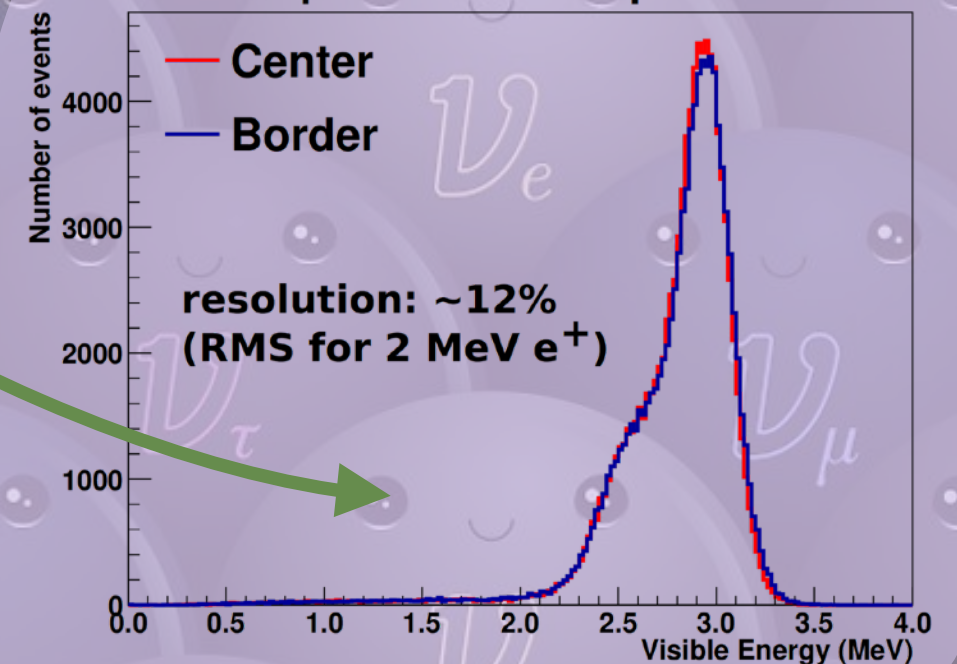
- Prompt Signal -

- e^+ ionisation + annihilation $E_{e^+} = [1, 8]$ MeV
- Oscillation depends on $E_\nu = E_{e^+} + 0.8$ MeV

- Delayed Signal -

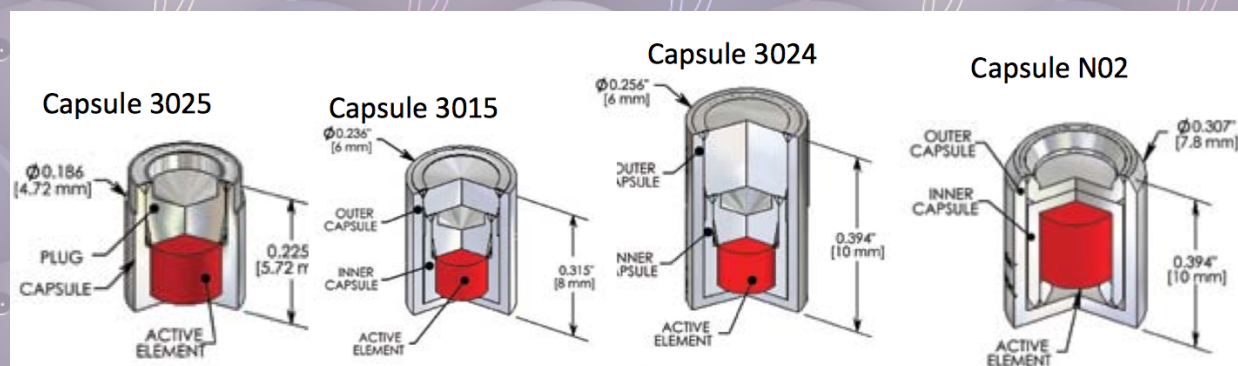
- n-capture on Gd releases γ_s : $\sum E_\gamma = 8$ MeV
- time separation from prompt $\Delta t \sim 15\mu s$

Response to 2 MeV positrons



Calibration in Stereo

Different gamma emitters are used in Stereo as calibration sources



- Three different calibration systems -

- Internal tubes inside target cells.
- External mechanical system.
- Rail underneath the detector.

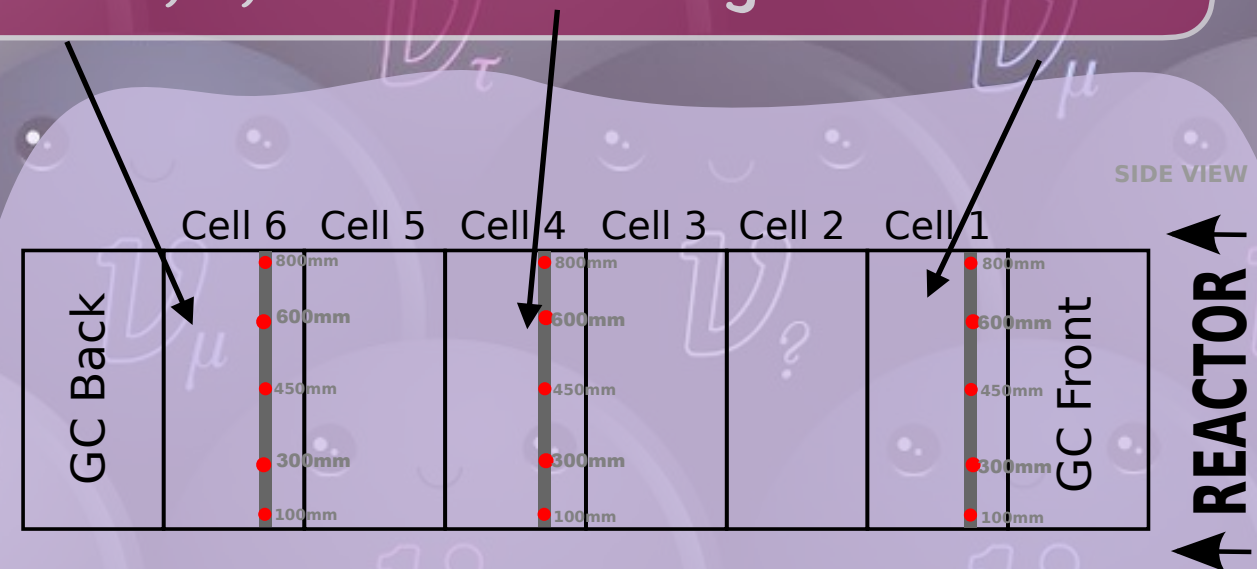
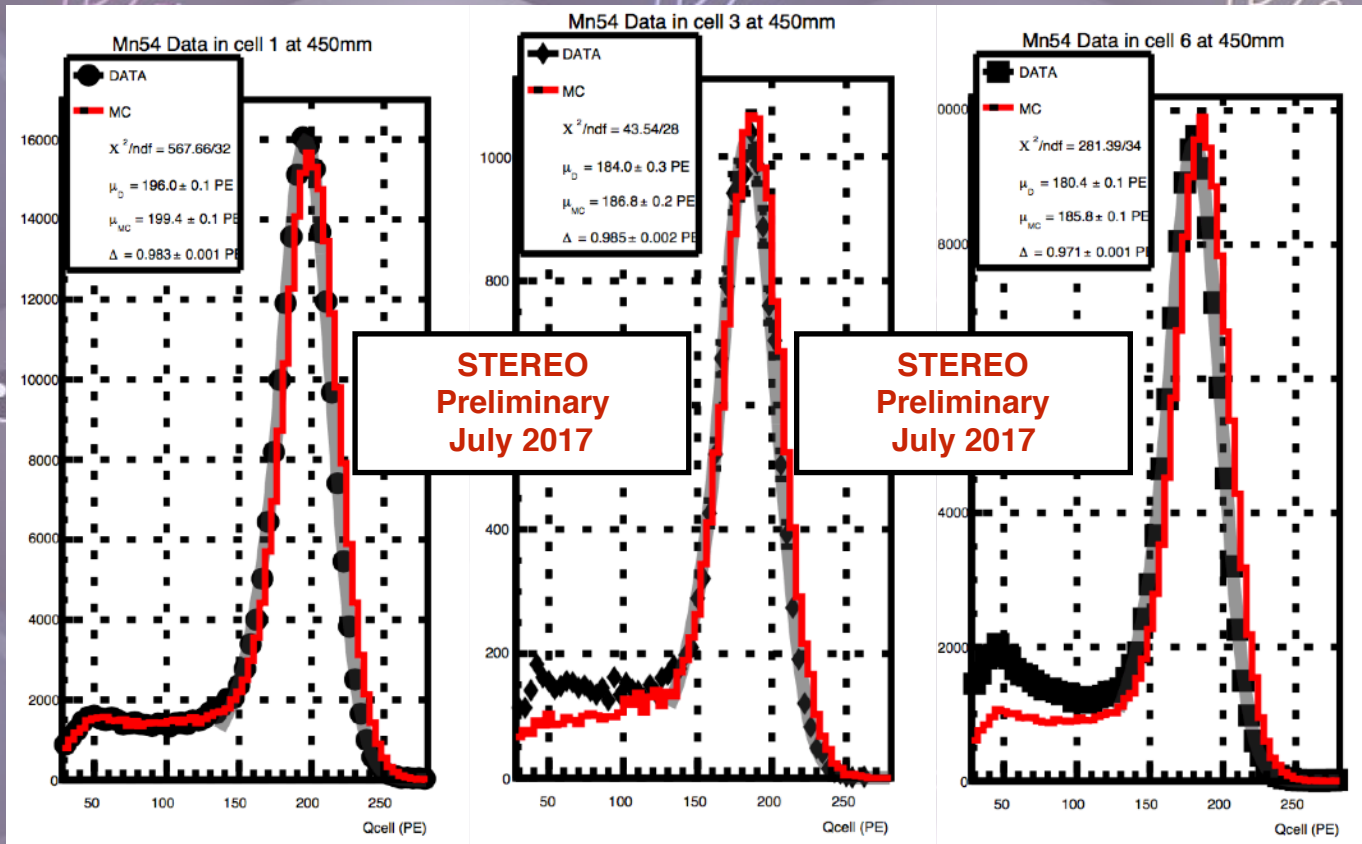
Today's topic

• Broad range of energies allow a good coverage of energy range.

- Low energy sources to account for quenching.
- High activity neutron emitter to calibrate for coincidence signals.

Isotope	Activity	γ - energy (keV)
Ge-68	90 kBq	2x511
Cs-137	37 kBq	660
Mn-54	90 kBq	830
Co-60	50 kBq	1170 & 1330
Zn-65	3.3 kBq	1120
Sb-124	2.4 kBq	600 & 1690
Na-24	5.9 kBq	1370 & 2750
AmBe	250 MBq (n ~kHz)	4400 + n
Cf-252	50 kBq (n ~kHz)	fission γ + n

3 tubes are placed in 3 different cells (1,4,6) inside the Target



For every tube, 5 different heights used as positions for the calibration sources: study z-axis

Response

- Attenuation length of the liquid scintillator above 6 meters - small z dependence.
- Target cells with very similar response, except for cell 4 which has a ~50% lower light collection efficiency due to leak of the buffer aquarium affecting the optical coupling of the PMTs.
- Intermediate cells (2,3,5) can be calibrated as well.



Energy reconstruction

To translate from Q collected at the PTMs to actual E, STEREO needs a dedicated energy reconstruction

The charge collected in a cell i is proportional to the light produced in that cell and the light leaks from neighbouring cells:

$$Q_i = \sum_j E_j^{dep} \times C_j \times L_{ji} = \sum_j E_j^{dep} M_{ji}$$

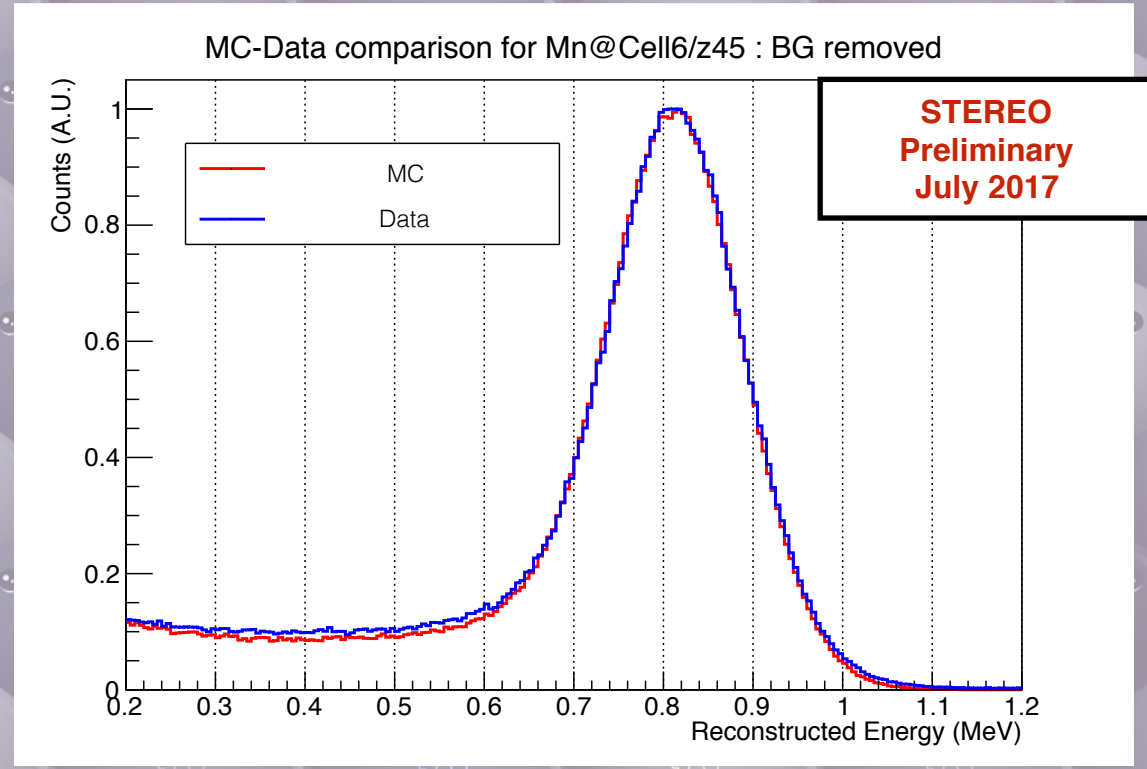
- C_j are the calibration coefficients for Cell j and L_{ji} are the Light Leaks from Cell j to Cell i.
- Inverting M_{ji} and solving the equation gives the final E_{dep} as a sum of contributions of all cells.

- Advantages of Energy over Charge -

- Stabilise detector response
- Direct comparison Data-MC
- Allows topological cuts

How do we calculate L_{ji} ?

How do we calculate C_j ?



1. Use Light Leaks to create an Energy Leaks Cut (ELC) for Data and MC charge distributions. The main idea is to filter as many Full Energy Deposition events as possible.

2. Apply ELC and fit Data and MC cell's deposited charge spectra. Get deposited charge for Full Energy Deposition events

$$\bar{Q}_D, \bar{Q}_{MC}$$



4. If MC and Data match, calculating calibration coefficients is defined as

$$CC = \bar{Q}_{MC} / E_{T,dep}$$

3. From MC, obtain True Edep ($E_{T,dep}$) corresponding to the deposited charge

\bar{Q}_{MC} calculated in 2.

1. Use Light Leaks to create an Energy Leaks Cut (ELC) for Data and MC charge distributions. The main idea is to filter as many Full Energy Deposition events as possible.

2. Apply ELC and fit Data and MC cell's deposited charge spectra. Get deposited charge for Full Energy Deposition events

4. If MC and Data match, calculating calibration coefficients is defined as

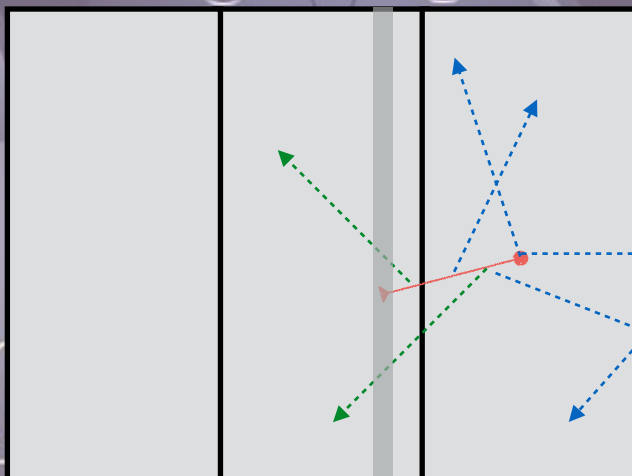
$$CC = \bar{Q}_{MC} / E_{T,dep}$$

3. μ From MC, obtain True E_{dep} ($E_{T,dep}$) corresponding to the deposited charge \bar{Q}_{MC} calculated in 2.

Charge collection in a cell can be affected by two different processes:

1. Gammas produced in cell A can escape to cell B and deposit energy there - **Energy Leaks (EL)**
2. Scintillation light from produced in cell A can be collected in cell B - **Light Leaks (LL)**

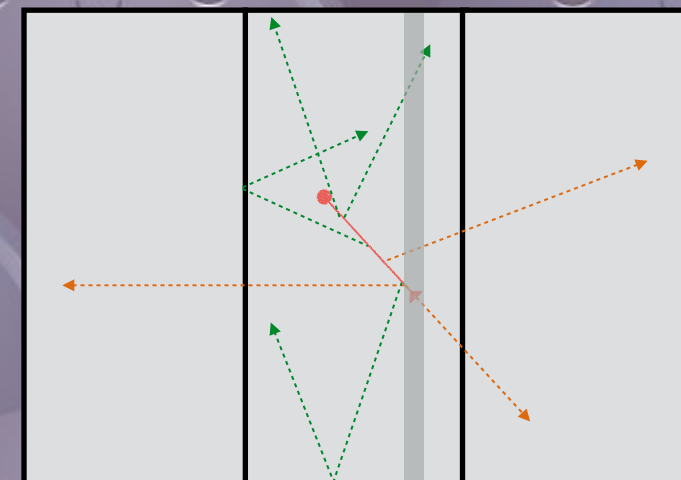
Escaping gamma



No Full Energy Deposition!

Mn54 Gamma
Leaked light
Collected light
Leaked Energy

Full Energy Deposition (FED)



1. When Energy Leaks happen there is no Full Energy Deposition (FED) in the cell

2. When only light leaks happen there is FED in the cell

A cut on Energy leaks helps rejection of events with no FED

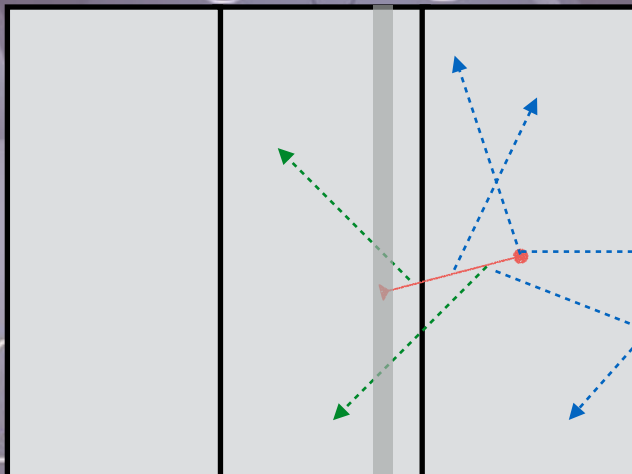
ENERGY LEAK CUT

Energy Leak Cut

Charge collection in a cell can be affected by two different processes:

1. Gammas produced in cell A can escape to cell B and deposit energy there - **Energy Leaks (EL)**
2. Scintillation light from produced in cell A can be collected in cell B - **Light Leaks (LL)**

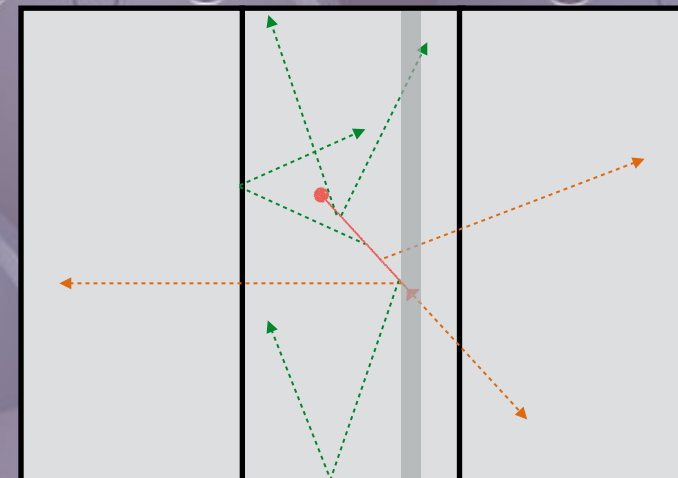
Escaping gamma



No Full Energy Deposition!

Mn54 Gamma
 Leaked light
 Collected light
 Leaked Energy

Full Energy Deposition (FED)



Pros of using a cut

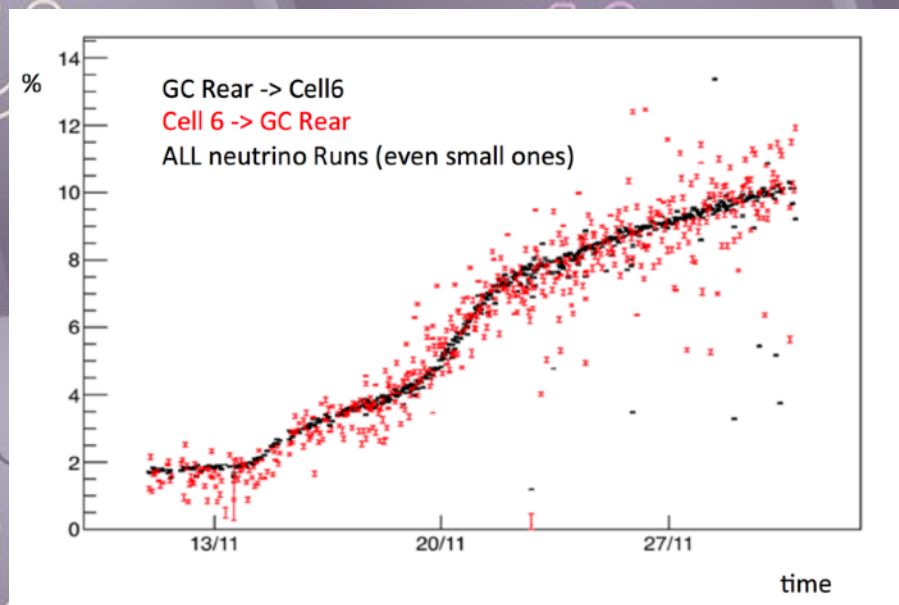
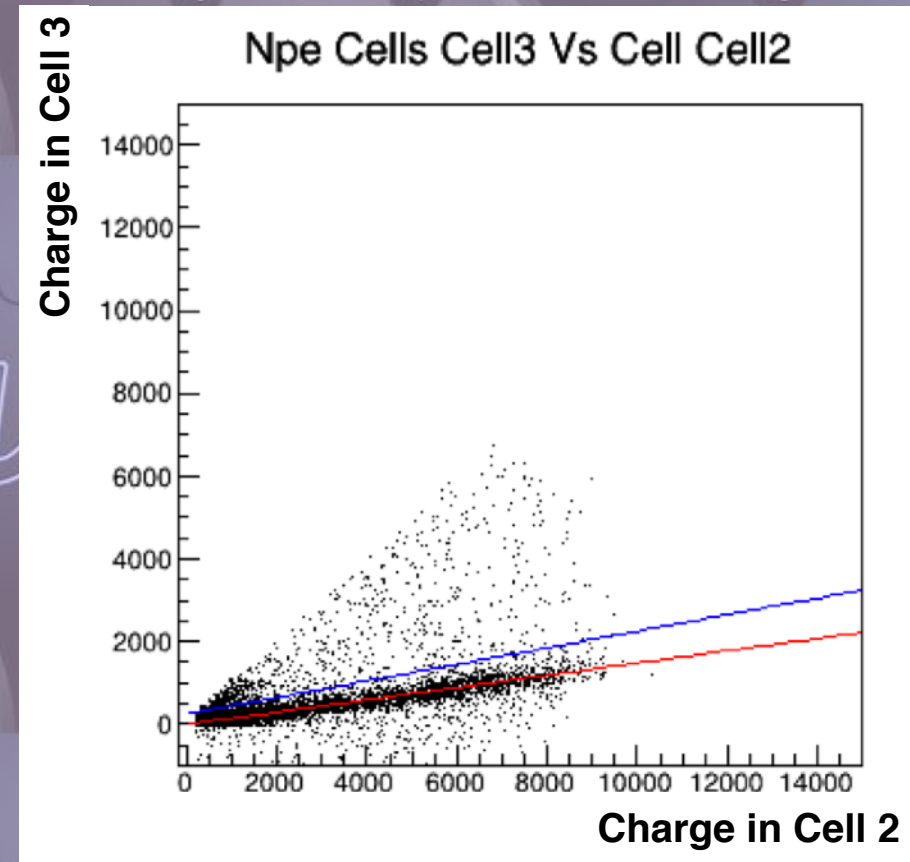
- Cells with different LY do not contaminate each other.
- Direct comparison with FED events MC.
- Allows intermediate cells to be calibrated.

Cons of using a cut

- Since LLs and ELs are technically indistinguishable:
- Cutting too strongly will cut FED events that leak light away. Bias on the final spectrum.
 - How large are Light Leaks in Stereo?

How do we obtain a value for the Light Leaks in Stereo?

- Cosmic rays crossing one cell leak light to neighbours proportionally.
- Fiteable correlation between light produced in FED cell and neighbour.
- Events below cutoff (blue line) correspond to Light Leaks.



1. Non-calibration runs of the detector are partially used to monitor cosmic rays. Light Leaks information over time can be used to estimate time evolution.

1. Use Light Leaks to create an Energy Leaks Cut (ELC) for Data and MC charge distributions. The main idea is to filter as many Full Energy Deposition events as possible.

2. Apply ELC and fit Data and MC cell's deposited charge spectra. Get deposited charge for Full Energy Deposition events

$$\bar{Q}_D, \bar{Q}_{MC}$$

4. If MC and Data match, calculating calibration coefficients is defined as

$$CC = \bar{Q}_{MC} / E_{T,dep}$$

3. From MC, obtain True E_{dep} ($E_{T,dep}$) corresponding to the deposited charge

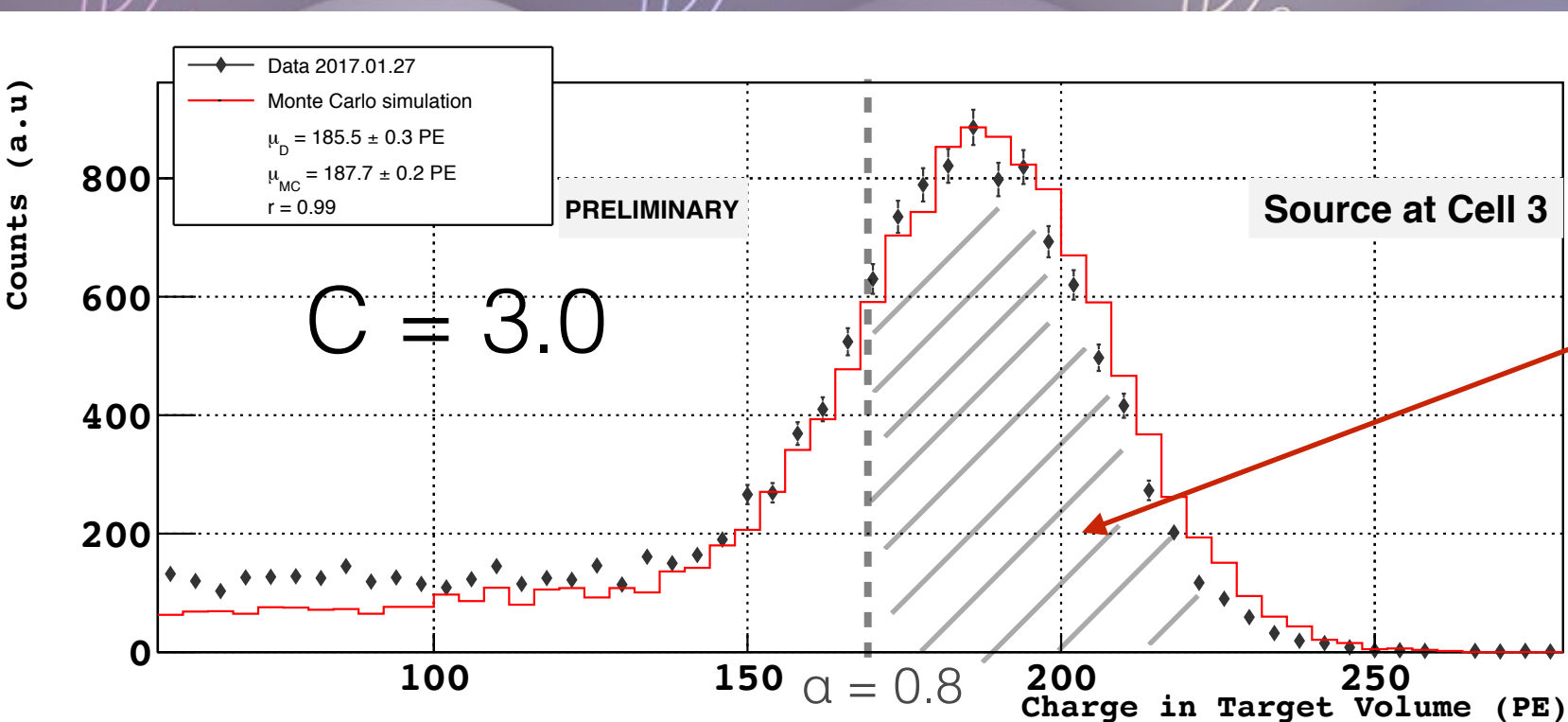
\bar{Q}_{MC} calculated in 2.

1. Choose as Energy Leak Cut to reject events:

$$Q_i > \frac{Q_{det}}{(1 + \sum_{i \neq j} LL_{ij} + C\sigma_{ij})}$$

- Where i is the cell being calibrated.
- C is a variable parameter to scan.
- LL_{ij} are the correspondent LL from cell i to cell j

- The ELC depends majorly on the amount of LLs that Stereo has at the moment of calibration.
- Use normal statistical deviation $\sigma_{ij} = \sqrt{LL_{ij}}$
- As C increases, the cut loosens up allowing more events. Tuning C is crucial.



2. Calculate $\bar{Q}_D(C, \alpha)$ and $\bar{Q}_{MC}(C, \alpha)$ as average of charge within the range $\{\alpha \cdot \mu_D, \infty\}$, with $\alpha = [0.1, 1.0]$ to be scanned.

General concepts to find CC

1. Use Light Leaks to create an Energy Leaks Cut (ELC) for Data and MC charge distributions. The main idea is to filter as many Full Energy Deposition events as possible.

2. Apply ELC and fit Data and MC cell's deposited charge spectra. Get deposited charge for Full Energy Deposition events

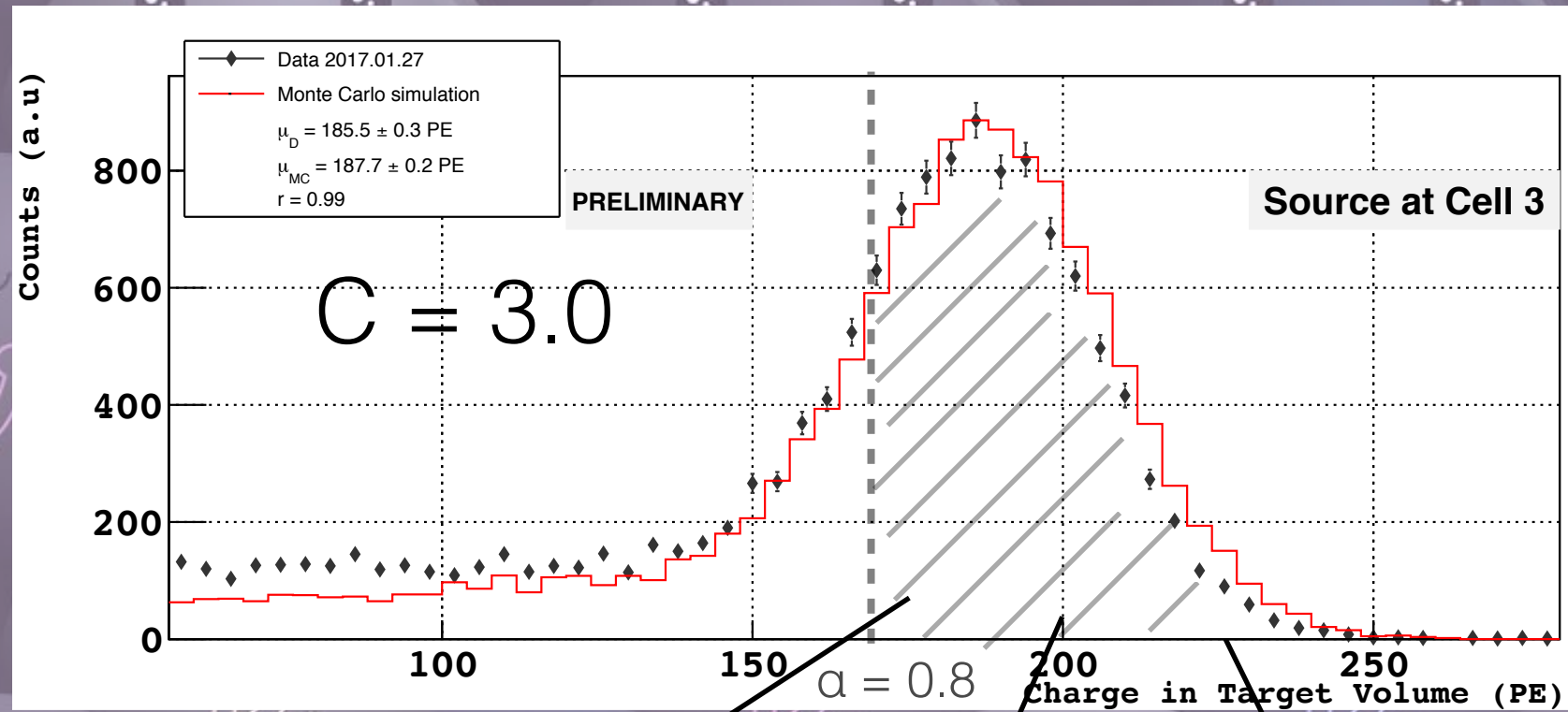
4. If MC and Data match, calculating calibration coefficients is defined as

$$CC = \bar{Q}_{MC} / E_{T,dep}$$

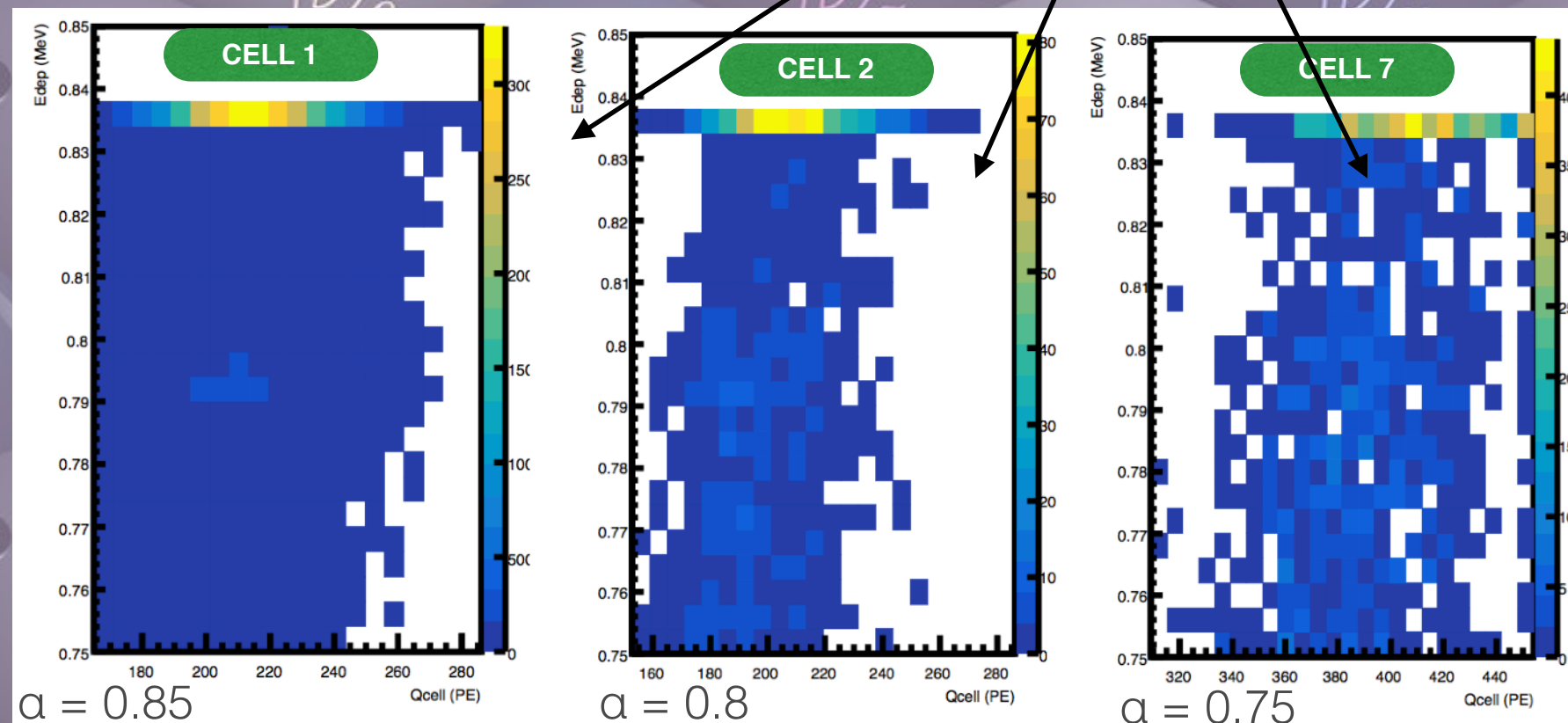
3. From MC, obtain True Edep ($E_{T,dep}$) corresponding to the deposited charge \bar{Q}_{MC} calculated in 2.

Obtaining True Deposited Energy

3. From MC Charge distribution, select the corresponding True Deposited energy distribution.



3. Obtain $E_{T,dep}$ as the average of MC True Energy Deposition.



General concepts to find CC

1. Use Light Leaks to create an Energy Leaks Cut (ELC) for Data and MC charge distributions. The main idea is to filter as many Full Energy Deposition events as possible.

 $\nu_e \rightarrow$

2. Apply ELC and fit Data and MC cell's deposited charge spectra. Get deposited charge for Full Energy Deposition events

 \bar{Q}_D, \bar{Q}_{MC}

4. If MC and Data match, calculating calibration coefficients is defined as

$$CC = \bar{Q}_D(C, \alpha) / E_{T,dep}$$

3. μ From MC, obtain True E_{dep} ($E_{T,dep}$) corresponding to the deposited charge \bar{Q}_{MC} calculated in 2.

$$CC(C, \alpha) = \bar{Q}_D(C, \alpha) / E_{T,dep}$$

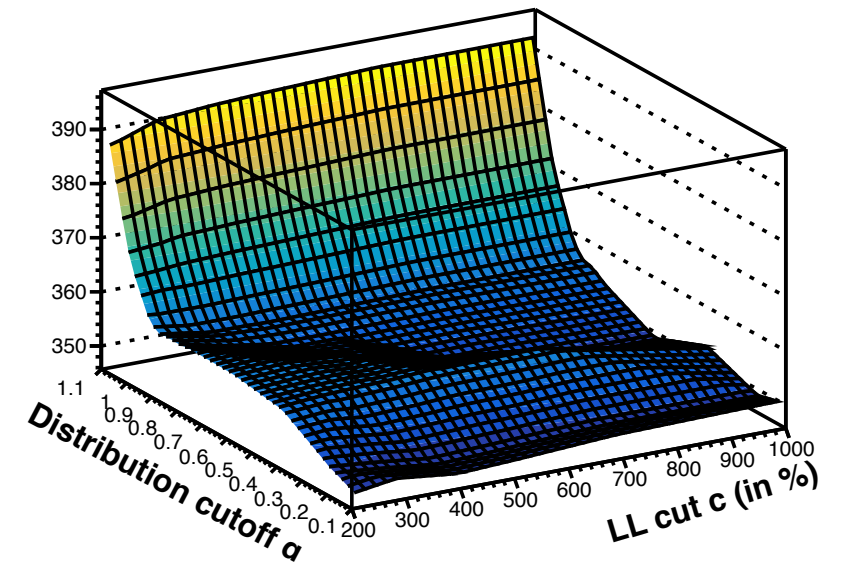
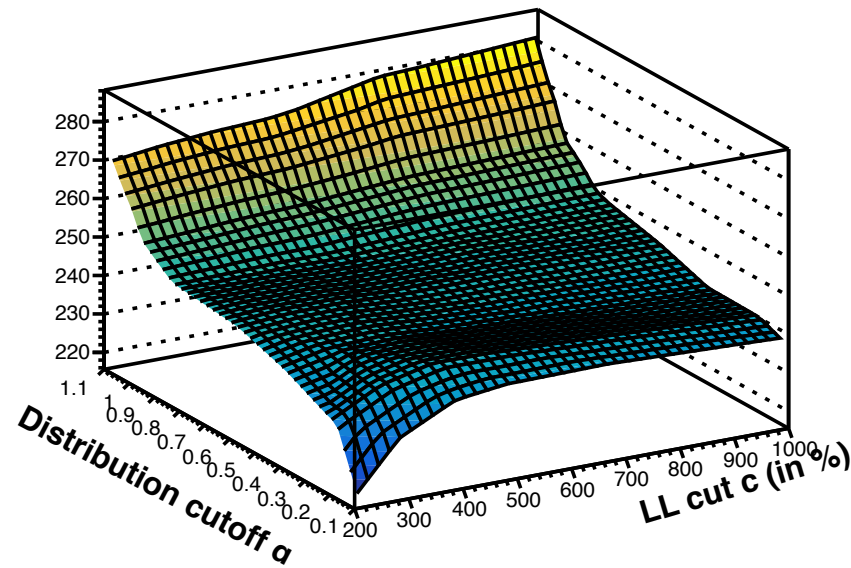
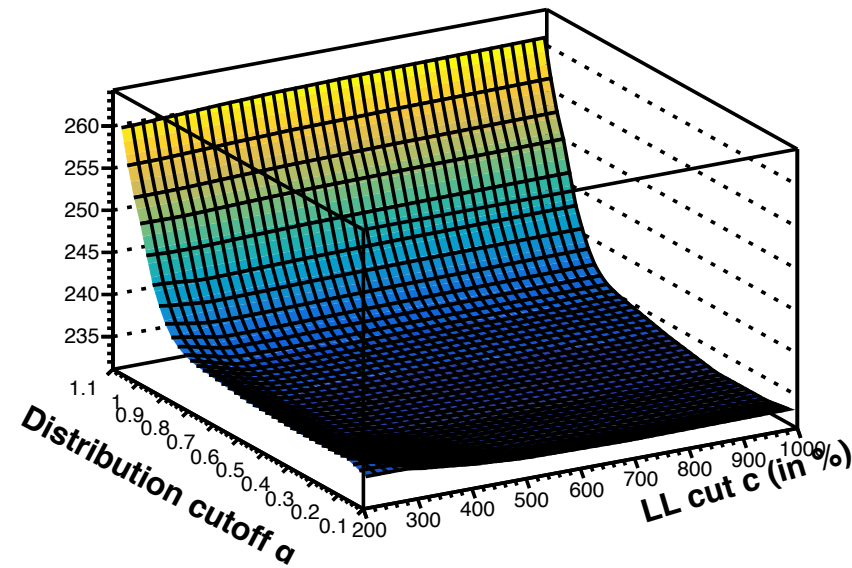
4. To obtain CC find the most stable combination of $\{C, \alpha\}$, follow some heuristic rules:

- (I) Too strict ELC ($C < 2.0$) biases $\bar{Q}_D(C, \alpha)$ to high energies - cutting LLs asymmetrically.
- (II) Too loose ELC ($C > 5.0$) accepts too much Energy Leaks. Specially bad for intermediate cells where source is not deployed.
- (III) Short cutoff on charge ($\alpha > 0.9$) will cut FED peak and generates unstable CCs
- (IV) Long cutoff on charge ($\alpha < 0.6$) includes too much non removed Energy leaks that is not present in true energy distribution

Cell 1 - 22.II.2017

Cell 2 - 22.II.2017

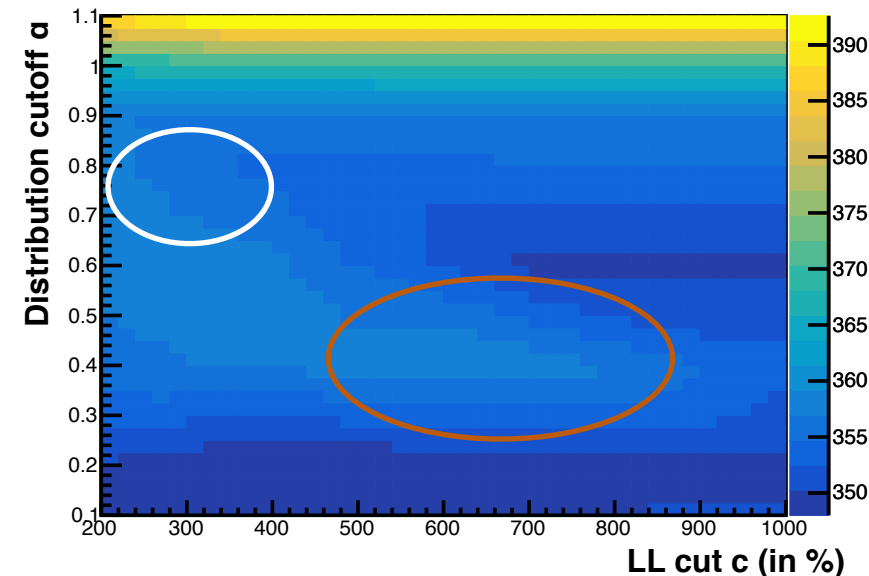
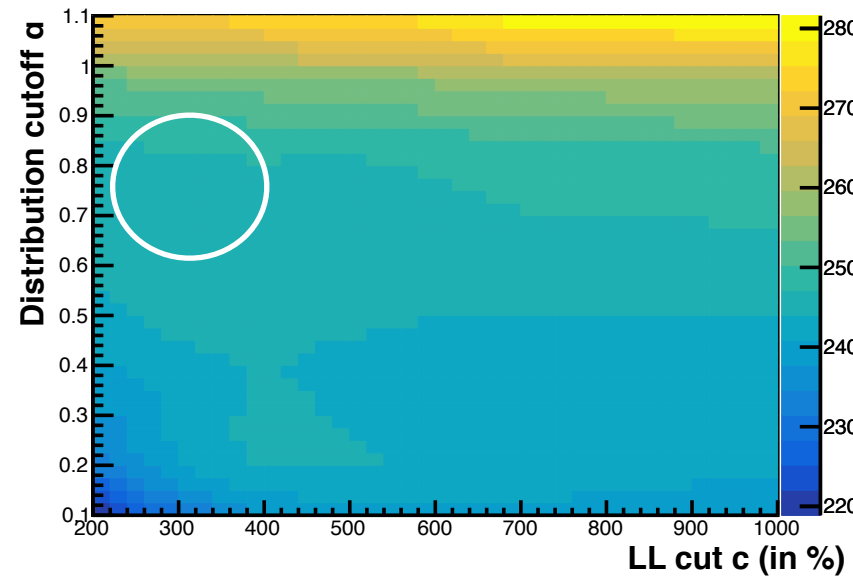
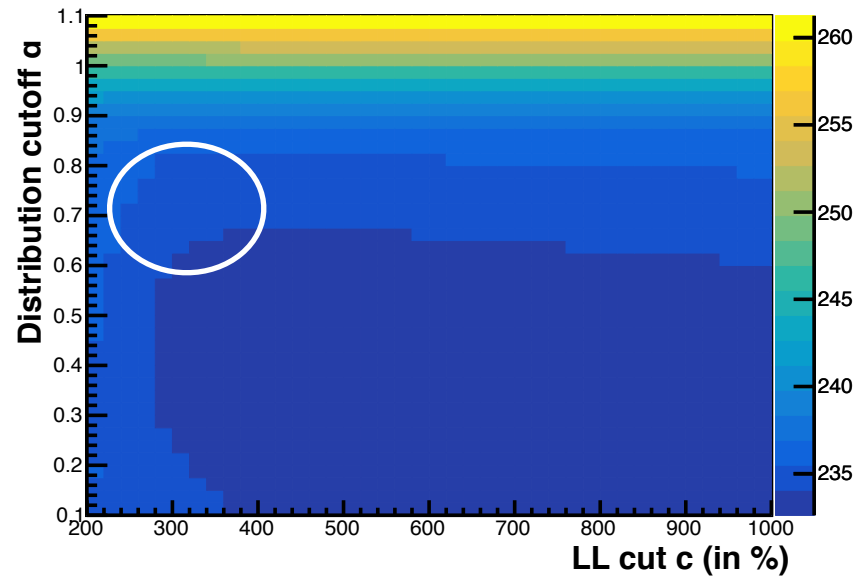
Cell 7 - 22.II.2017



Cell 1

Cell 2

Cell 7



- Look for the regions where CCs are most stable.
- Large a gives unstable CC for the different LL cuts c .
- Large LL cut c gives unstable CC for different cutoffs a .

Relatively stable areas in parameter space

Contribution from lower scintillation cells

Conclusions

Stereo aims to observe a sterile neutrino oscillation signal. For this purpose, a dedicated Energy Reconstruction needs to be performed.

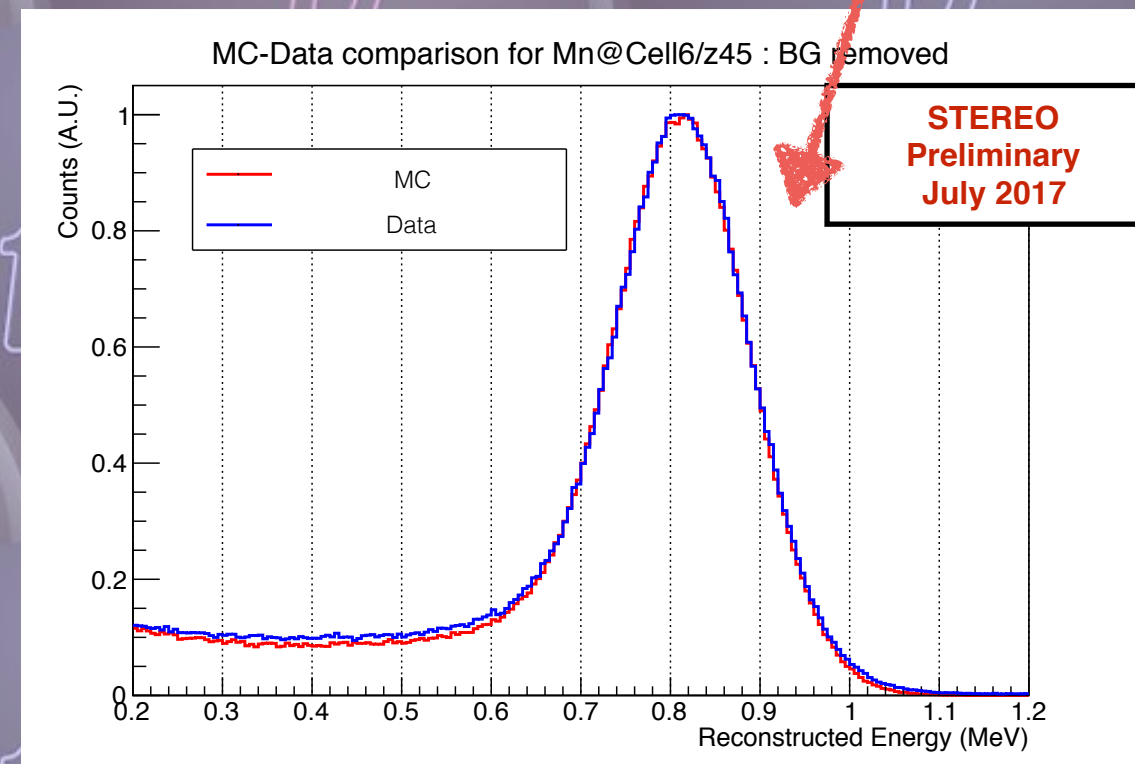
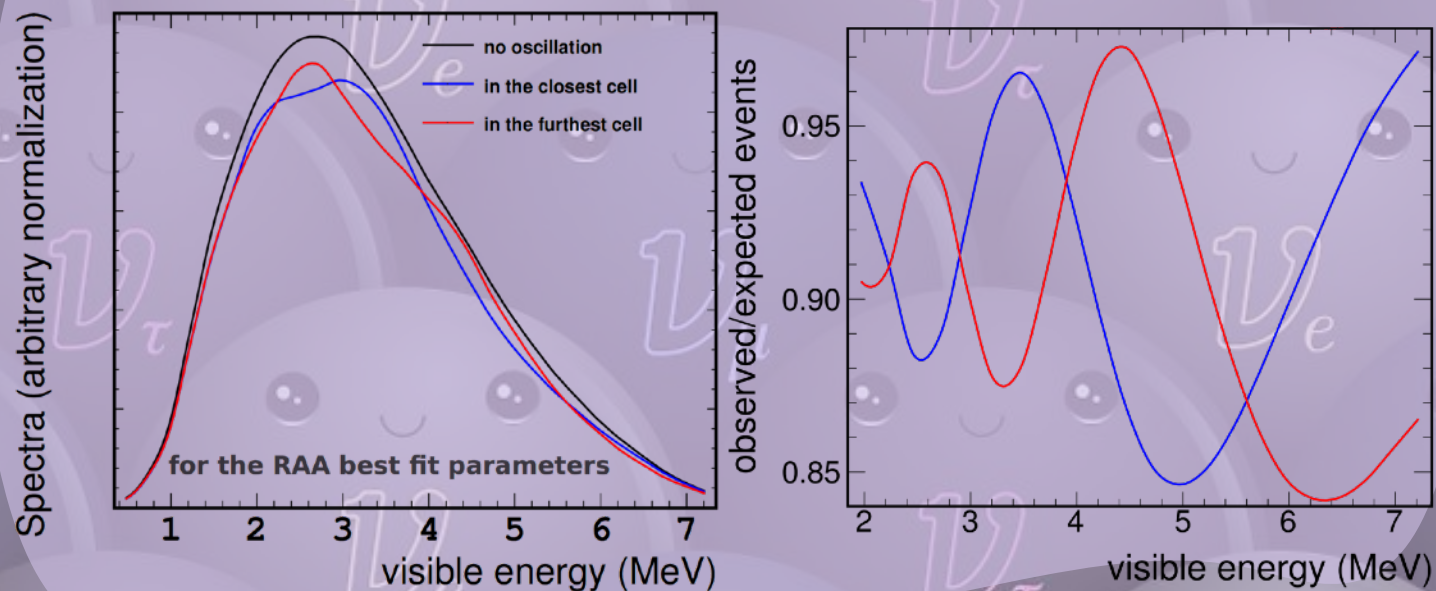
$$Q_i = \sum_j E_j^{dep} \times CC_j \times LL_{ji}$$

Heuristic arguments applied to find stable area in CC parameter space. Relative stable calibration coefficients found.

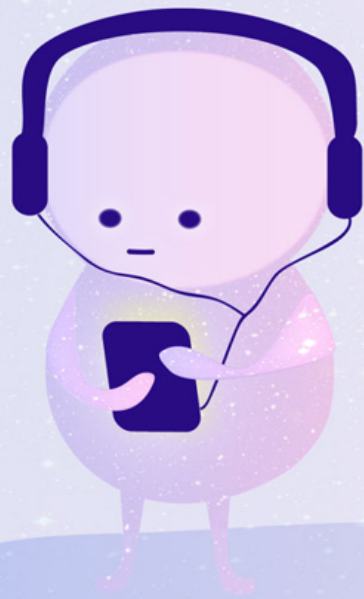
Calculated CCs give accurate Energy Reconstruction and low discrepancy Data / MC.

Good Energy Reconstruction is crucial

Oscillation phase shift inside the target



STERILE NEUTRINOS



featuring **STEREO** experiment

Thanks for your attention - now BBQ is a little
bit closer

