

Gauge Theories for Baryon and Lepton Numbers

IMPRS-PTFS Seminar, 7 November 2013

Based on arXiv:1304.0576 [hep-ph],
arXiv:1306.0568 [hep-ph],
arXiv:1309.3970 [hep-ph].

With P. Fileviez Pérez (MPIK), M. Lindner (MPIK),
M. B. Wise (Caltech).

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FOR PRECISION TESTS
OF FUNDAMENTAL
SYMMETRIES



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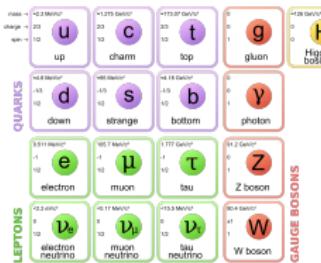
Outline

- ▶ Introduction
- ▶ Gauging Baryon and Lepton Numbers
- ▶ Fermionic Leptoquarks
- ▶ Summary

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The Standard Model of Particle Physics



Wikipedia

► Standard Model gauge group:

$$G_{\text{SM}} = SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$$

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$	$U(1)_L$
Q_L	3	2	1/6	1/3	0
u_R	3	1	2/3	1/3	0
d_R	3	1	-1/3	1/3	0
ℓ_L	1	2	-1/2	0	1
e_R	1	1	-1	0	1
H	1	2	1/2	0	0

Standard Model Features

- ▶ Renormalizable SM couplings conserve B and L , e.g.,

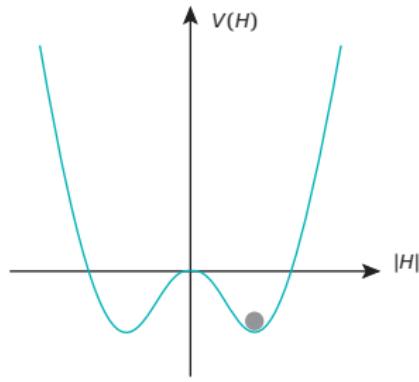
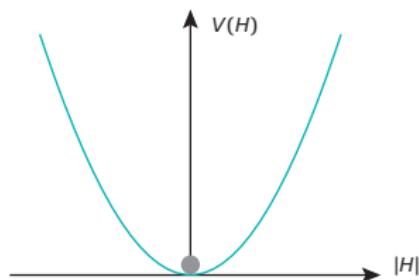
$$\mathcal{L}_{\text{SM}} \supset \overline{\ell_L} \not{D} \ell_L + Y_Q \overline{Q_L} H d_R$$

- ▶ Yukawa couplings of massless charged leptons:

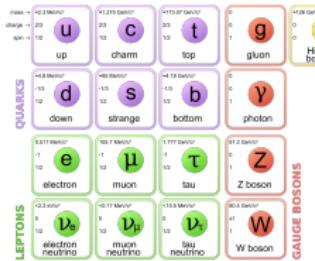
$$\mathcal{L}_Y = -Y_L \overline{\ell_L} H e_R + \text{h.c.}$$

- ▶ Spontaneous symmetry breaking
 $SU(2)_L \otimes U(1)_Y \xrightarrow{\langle H^0 \rangle = v/\sqrt{2}} U(1)_{\text{em}}$:

$$\mathcal{L}_Y \rightarrow -\frac{v}{\sqrt{2}} Y_L \overline{e_L} e_R + \text{h.c.}$$



The Standard Model of Particle Physics



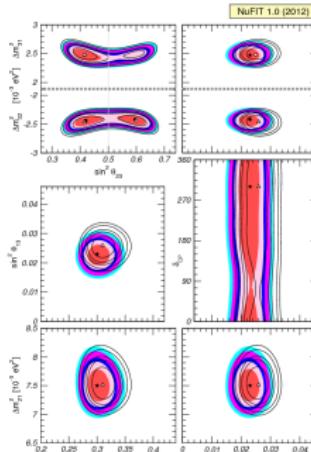
Wikipedia

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u_R	3	1	$2/3$	$1/3$	0
d_R	3	1	$-1/3$	$1/3$	0
l_L	1	2	$-1/2$	0	1
e_R			No right-handed neutrinos: neutrinos massless in the SM!		1
H					0

Neutrinos Have Mass



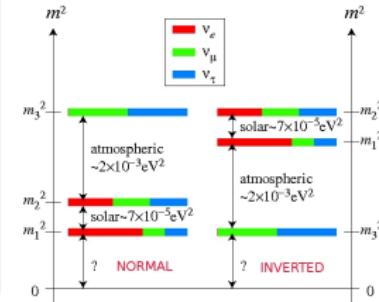
M. C. Gonzalez-Garcia et al., arXiv:1209.3023 [hep-ph]

Neutrino oscillations

$$\nu_i = \sum_{\alpha} U_{i\alpha}^* \nu_{\alpha}$$

$$\nu_i(t) = e^{-i(Et - p_i x)} \nu_i$$

$$P_{\alpha \rightarrow \beta} = \sin^2(2\theta) \sin^2\left(1.27 \frac{\Delta m^2}{\text{eV}^2} \frac{L/\text{km}}{E/\text{GeV}}\right)$$



R. N. Mohapatra et al., arXiv:hep-ph/0412099

Oscillation parameters

- Is lepton number conserved in Nature? → $0\nu\beta\beta$ experiments

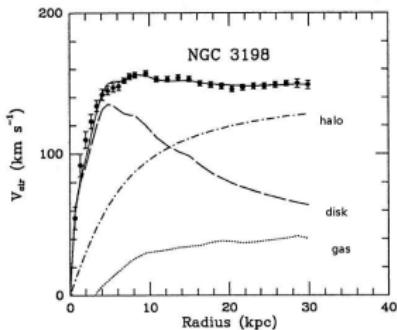


- Weinberg:

$$\mathcal{O}_5 = \frac{c_5}{\Lambda_L} LLHH$$

Hints for DM

Galaxy
rotation
curves



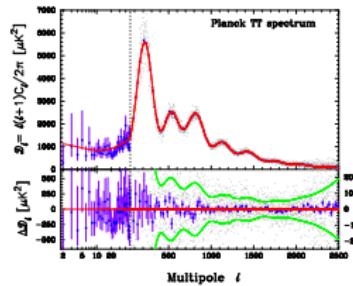
Begeman et al.,
MNRAS **249** (1991) 523

Bullet cluster



NASA

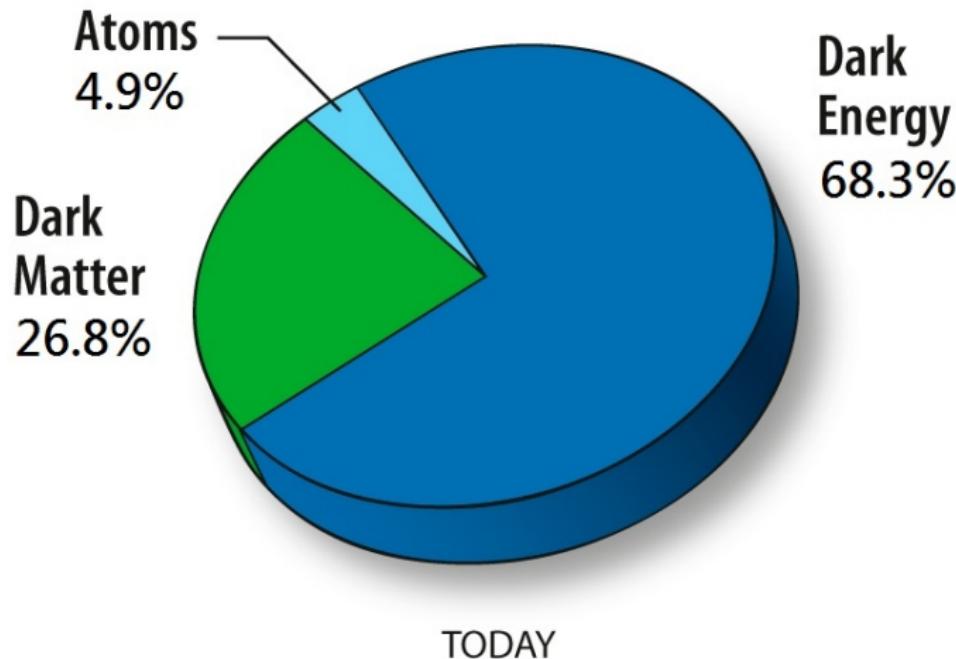
CMB



Planck Collaboration,
arXiv:1303.5076 [astro-ph.CO]

Consistent hints on all scales.

Content of the Universe



Wikipedia

What about Baryon Number?

B and L accidental global symmetries in the SM.

- ▶ Violation of B :

- ▶ Matter-antimatter asymmetry of the Universe.
- ▶ Proton decay ($\Delta B = 1$, $\Delta L = \text{odd}$):

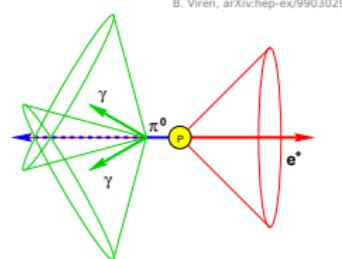
$$\tau_p \geq 10^{32-34} \text{ yrs.}$$

- ▶ Add non-renormalizable operators to the SM, e.g.,

$$\mathcal{O}_6 = \frac{c_6}{\Lambda_B^2} QQLQ$$

- ▶ The scale Λ_B must be large:

$$\Lambda_B \geq 10^{15} \text{ GeV}$$



B. Viren, arXiv:hep-ex/9903029

The Big Desert

S. Weinberg, Phys. Rev. Lett. **49** (1979) 1566

fermions	up	u	down	d	s	b	top	t	bottom	b	Higgs	H
quarks	up	c	down	s	strange	b	bottom	c	bottom	b	gluon	g
leptons	electron	e	muon	u	tau	t	neutrino	l	neutrino	l	photon	γ
gauges	U(1)	$U(1)_Y$	U(1)	$U(1)_L$	U(1)	$U(1)_B$	U(1)	$U(1)_Y$	U(1)	$U(1)_L$	Z boson	Z
											W boson	W

Wikipedia



Wikipedia

Low scale
Electroweak scale
($\Lambda_{EW} \sim 10^2$ GeV)

Michael Duerr (MPIK)

Gauge Theories for B and L

$$\frac{C_5}{\Lambda_L} LLHH$$

$$\frac{C_6}{\Lambda_B^2} QQQL$$

High scale
e.g. GUT scale
($\Lambda_{GUT} \sim 10^{15}$ GeV)

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The Big Desert

S. Weinberg, Phys. Rev. Lett. **49** (1979) 1566

$$\frac{C_5}{\Lambda_L} LLHH$$



A photograph of a desert landscape with large sand dunes under a clear blue sky. Six question marks are overlaid on the image, pointing towards different features of the dunes.

Wikipedia

QUARKS		LEPTONS		GAUGE BOSONS	
up	down	electron	electron neutrino	gluon	Higgs boson
u	d	e	ν_e	g	H
charm	strange	muon	ν_μ	b	γ
c	s	tau	ν_τ	t	Z boson
top	bottom				W boson
t	b				
gluon					
Higgs boson					

Low scale
Electroweak scale
($\Lambda_{EW} \sim 10^2$ GeV)

$$\frac{C_5}{\Lambda_L} LLHH$$

~~$$\frac{C_6}{\Lambda_B^2} QQQL$$~~

High scale
e.g. GUT scale
($\Lambda_{GUT} \sim 10^{15}$ GeV)

Aim of this Talk

Define a consistent gauge theory
for baryon and lepton numbers
that can be broken at a low scale.

- ▶ Neutrino masses
- ▶ Baryogenesis?
- ▶ Dark matter
- ▶ Signals at the LHC

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B and L in the Standard Model

- SM: B and L are **accidental symmetries**, not free of anomalies.

⇒ we need **additional fermions** for anomaly cancellation.

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$	$U(1)_L$
Q_L	3	2	$\frac{1}{6}$	$\frac{1}{3}$	0
u_R	3	1	$\frac{2}{3}$	$\frac{1}{3}$	0
d_R	3	1	$-\frac{1}{3}$	$\frac{1}{3}$	0
ℓ_L	1	2	$-\frac{1}{2}$	0	1
ν_R	1	1	0	0	1
e_R	1	1	-1	0	1
H	1	2	$\frac{1}{2}$	0	0

- B and L may be gauged to obtain the gauge group:

$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_B \otimes U(1)_L$$

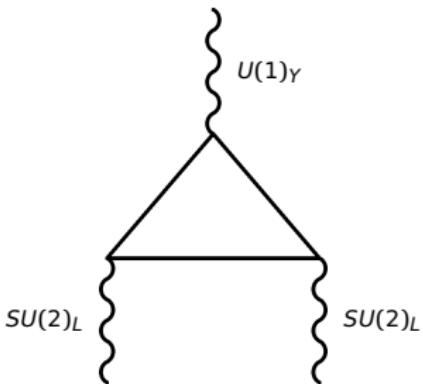
Standard Model Anomalies

Example: $SU(2)_L^2 \otimes U(1)_Y$

$$\mathcal{A} = \text{Tr}(t^a t^b Y) = \frac{1}{2} \delta^{ab} \cdot \sum_i Y_i$$

$$\sum_i Y_i = -\frac{1}{2} + 3 \cdot \frac{1}{6} = 0$$

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$	$U(1)_L$
Q_L	3	2	$\frac{1}{6}$	$\frac{1}{3}$	0
u_R	3	1	$\frac{2}{3}$	$\frac{1}{3}$	0
d_R	3	1	$-\frac{1}{3}$	$\frac{1}{3}$	0
ℓ_L	1	2	$-\frac{1}{2}$	0	1
ν_R	1	1	0	0	1
e_R	1	1	-1	0	1
H	1	2	$\frac{1}{2}$	0	0



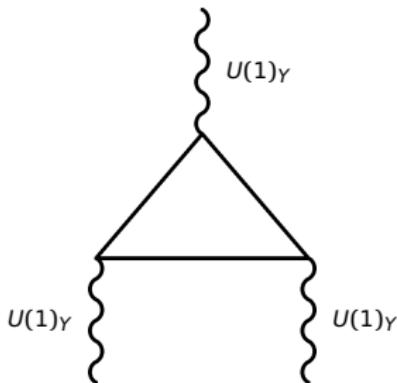
Standard Model Anomalies

Another example: $U(1)_Y^3$

$$\mathcal{A} = \text{Tr}(Y^3) = \sum_i Y_i^3$$

$$\sum_i Y_i^3 = 2\left(-\frac{1}{2}\right)^3 - (-1)^3 + 3\left[2\left(\frac{1}{6}\right)^3 - \left(\frac{2}{3}\right)^3 - \left(-\frac{1}{3}\right)^3\right] = 0$$

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$	$U(1)_L$
Q_L	3	2	$\frac{1}{6}$	$\frac{1}{3}$	0
u_R	3	1	$\frac{2}{3}$	$\frac{1}{3}$	0
d_R	3	1	$-\frac{1}{3}$	$\frac{1}{3}$	0
ℓ_L	1	2	$-\frac{1}{2}$	0	1
ν_R	1	1	0	0	1
e_R	1	1	-1	0	1
H	1	2	$\frac{1}{2}$	0	0



Baryonic and Leptonic Anomalies

► Purely baryonic anomalies:

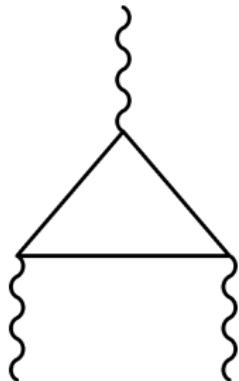
$$\begin{aligned} \mathcal{A}_1 & (SU(3)^2 \otimes U(1)_B), \quad \mathcal{A}_2 (SU(2)^2 \otimes U(1)_B), \\ \mathcal{A}_3 & (U(1)_Y^2 \otimes U(1)_B), \quad \mathcal{A}_4 (U(1)_Y \otimes U(1)_B^2), \\ \mathcal{A}_5 & (U(1)_B), \quad \mathcal{A}_6 (U(1)_B^3). \end{aligned}$$

► Purely leptonic anomalies:

$$\begin{aligned} \mathcal{A}_7 & (SU(3)^2 \otimes U(1)_L), \quad \mathcal{A}_8 (SU(2)^2 \otimes U(1)_L), \\ \mathcal{A}_9 & (U(1)_Y^2 \otimes U(1)_L), \quad \mathcal{A}_{10} (U(1)_Y \otimes U(1)_L^2), \\ \mathcal{A}_{11} & (U(1)_L), \quad \mathcal{A}_{12} (U(1)_L^3). \end{aligned}$$

► Mixed anomalies:

$$\begin{aligned} \mathcal{A}_{13} & (U(1)_B^2 \otimes U(1)_L), \quad \mathcal{A}_{14} (U(1)_L^2 \otimes U(1)_B), \\ \mathcal{A}_{15} & (U(1)_Y \otimes U(1)_L \otimes U(1)_B). \end{aligned}$$



Baryonic and Leptonic Anomalies

► Purely baryonic anomalies:

$$\begin{aligned} \mathcal{A}_1 & (SU(3)^2 \otimes U(1)_B), \quad \mathcal{A}_2 (SU(2)^2 \otimes U(1)_B), \\ \mathcal{A}_3 & (U(1)_Y^2 \otimes U(1)_B), \quad \mathcal{A}_4 (U(1)_Y \otimes U(1)_B^2), \\ \mathcal{A}_5 & (U(1)_B), \quad \mathcal{A}_6 (U(1)_B^3). \end{aligned}$$

► Purely leptonic anomalies:

$$\begin{aligned} \mathcal{A}_7 & (SU(3)^2 \otimes U(1)_L), \quad \mathcal{A}_8 (SU(2)^2 \otimes U(1)_L), \\ \mathcal{A}_9 & (U(1)_Y^2 \otimes U(1)_L), \quad \mathcal{A}_{10} (U(1)_Y \otimes U(1)_L^2), \\ \mathcal{A}_{11} & (U(1)_L), \quad \mathcal{A}_{12} (U(1)_L^3). \end{aligned}$$

► Mixed anomalies:

$$\begin{aligned} \mathcal{A}_{13} & (U(1)_B^2 \otimes U(1)_L), \quad \mathcal{A}_{14} (U(1)_L^2 \otimes U(1)_B), \\ \mathcal{A}_{15} & (U(1)_Y \otimes U(1)_L \otimes U(1)_B). \end{aligned}$$

SM +
right-handed ν

$$\mathcal{A}_2 = -\mathcal{A}_3 = \frac{3}{2},$$

$$\mathcal{A}_8 = -\mathcal{A}_9 = \frac{3}{2}$$

Possible Solutions

- ▶ Sequential/Mirror family:

P. Fileviez Pérez, M. B. Wise, [arXiv:1002.1754 \[hep-ph\]](https://arxiv.org/abs/1002.1754)

Ruled out: new quarks change gluon fusion production;
Landau poles of the new Yukawas near the weak scale.

- ▶ Vector-Like fermions:

P. Fileviez Pérez, M. B. Wise, [arXiv:1106.0343 \[hep-ph\]](https://arxiv.org/abs/1106.0343)

Ruled out: new charged leptons reduce BR of $H \rightarrow \gamma\gamma$ by a factor of 3.

- ▶ One family of leptoquarks:

P. V. Dong, H. N. Long, [arXiv:1010.3818 \[hep-ph\]](https://arxiv.org/abs/1010.3818)

$$F_L \sim (3, 2, 0, -1, -1), j_R \sim (3, 1, \frac{1}{2}, -1, -1), k_R \sim (3, 1, -\frac{1}{2}, -1, -1).$$

Ruled out: stable charged fields.

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New Solution: Vectorlike Leptoquarks

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$	$U(1)_L$
Ψ_L	1	2	$-\frac{1}{2}$	$-\frac{3}{2}$	$-\frac{3}{2}$
Ψ_R	1	2	$-\frac{1}{2}$	$+\frac{3}{2}$	$+\frac{3}{2}$
η_R	1	1	-1	$-\frac{3}{2}$	$-\frac{3}{2}$
η_L	1	1	-1	$+\frac{3}{2}$	$+\frac{3}{2}$
χ_R	1	1	0	$-\frac{3}{2}$	$-\frac{3}{2}$
χ_L	1	1	0	$+\frac{3}{2}$	$+\frac{3}{2}$

M. Duerr, P. Fileviez Pérez, M. B. Wise, [arXiv:1304.0576 \[hep-ph\]](https://arxiv.org/abs/1304.0576)

Interactions

- ▶ Responsible for the new fermion masses:

$$\begin{aligned} -\mathcal{L} \supset & h_1 \bar{\Psi}_L H \eta_R + h_2 \bar{\Psi}_L \tilde{H} \chi_R + h_3 \bar{\Psi}_R H \eta_L + h_4 \bar{\Psi}_R \tilde{H} \chi_L \\ & + \lambda_1 \bar{\Psi}_L \Psi_R S_{BL} + \lambda_2 \bar{\eta}_R \eta_L S_{BL} + \lambda_3 \bar{\chi}_R \chi_L S_{BL} \\ & + a_1 \chi_L \chi_L S_{BL} + a_2 \chi_R \chi_R S_{BL}^\dagger + \text{h.c.} \end{aligned}$$

with $S_{BL} \sim (1, 1, 0, -3, -3)$

- ▶ Neutrino masses:

$$-\mathcal{L}_\nu = Y_\nu \bar{\ell}_L \tilde{H} \nu_R + \frac{\lambda_R}{2} \nu_R \nu_R S_L + \text{h.c.}$$

with $S_L \sim (1, 1, 0, 0, -2)$

Further Aspects

- ▶ Symmetry breaking:
 - ▶ S_{BL} breaks $U(1)_B$ and $U(1)_L$ ($\Delta B = \pm 3$, $\Delta L = \pm 3$),
 - ▶ S_L contributes to breaking of $U(1)_L$ ($\Delta L = \pm 2$).
→ no proton decay
- ▶ Fermionic sector:
 - ▶ 4 neutral and 4 charged new chiral fermions after symmetry breaking.
 - ▶ No coupling to the SM fermions → no new source of flavor violation.
 - ▶ Lightest new fermion automatically stable → DM candidate.

Simple Example: Baryon Number Only

- ▶ Gauge group: $G_{\text{SM}} \otimes U(1)_B$
- ▶ Additional fields for an anomaly-free theory:

$$\begin{array}{ll} \Psi_L \sim (\mathbf{1}, \mathbf{2}, -1/2, B_1), & \Psi_R \sim (\mathbf{1}, \mathbf{2}, -1/2, B_2), \\ \eta_R \sim (\mathbf{1}, \mathbf{1}, -1, B_1), & \eta_L \sim (\mathbf{1}, \mathbf{1}, -1, B_2), \\ \chi_R \sim (\mathbf{1}, \mathbf{1}, 0, B_1), & \chi_L \sim (\mathbf{1}, \mathbf{1}, 0, B_2), \end{array}$$

- ▶ New Higgs for spontaneous breaking of baryon number:
 $S_B \sim (\mathbf{1}, \mathbf{1}, 0, -3)$
- ▶ Condition from anomaly cancellation: $B_1 - B_2 = -3$.

M. Duerr, P. Fileviez Pérez, [arXiv:1309.3970 \[hep-ph\]](https://arxiv.org/abs/1309.3970)

Spontaneous Symmetry Breaking

- ▶ Relevant interactions of the new fields (for $B_1 \neq -B_2$):

$$-\mathcal{L} \supset \lambda_1 \bar{\Psi}_L \Psi_R S_B + \lambda_2 \bar{\eta}_R \eta_L S_B + \lambda_3 \bar{\chi}_R \chi_L S_B + \text{h.c.}$$

- ▶ $\langle S_B \rangle \neq 0$:

$$-\mathcal{L} \supset M_\Psi \bar{\Psi}_L \Psi_R + M_\eta \bar{\eta}_R \eta_L + m_\chi \bar{\chi}_R \chi_L + \text{h.c.}$$

- ▶ Remnant Z_2 stabilizes the DM candidate:

$$\Psi_{L,R} \rightarrow -\Psi_{L,R}, \quad \eta_{L,R} \rightarrow -\eta_{L,R}, \text{ and } \chi_{L,R} \rightarrow -\chi_{L,R}$$

Dark Matter

- ▶ Dirac DM, SM singlet-like: $\chi = \chi_R + \chi_L$
- ▶ Coupling to the new gauge boson:

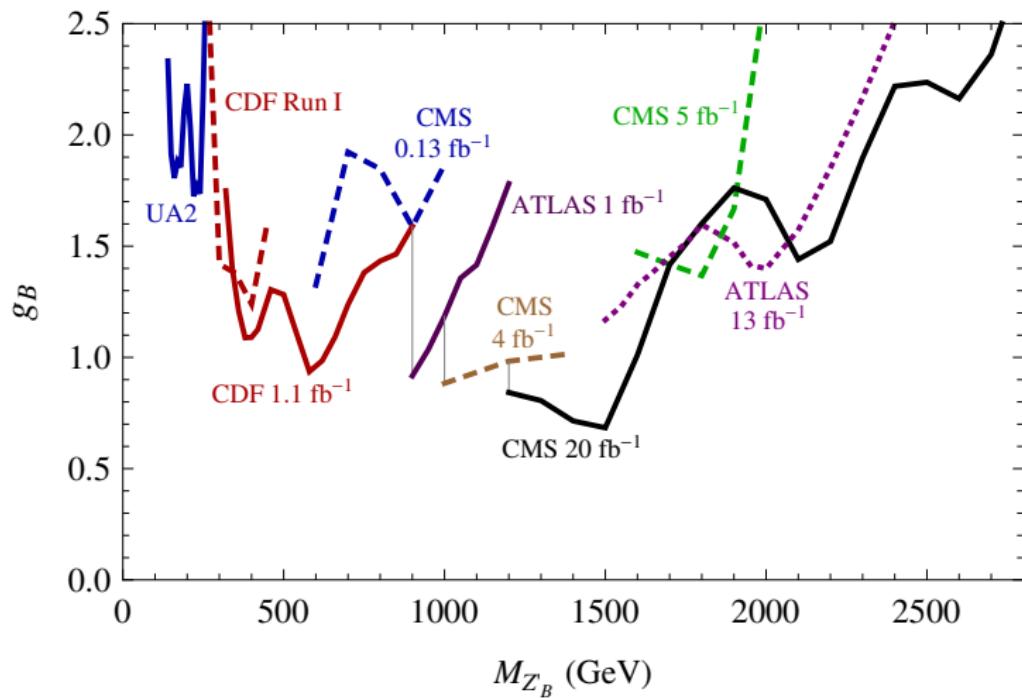
$$\mathcal{L} \supset g_B \bar{\chi} \gamma_\mu Z_B^\mu (B_2 P_L + B_1 P_R) \chi$$

- ▶ Model has only four free parameters:

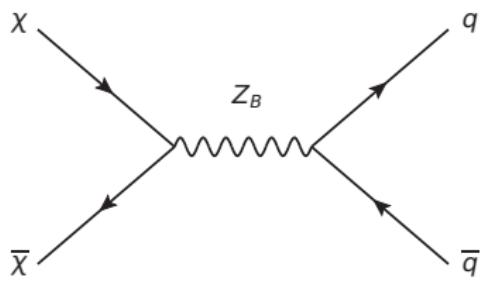
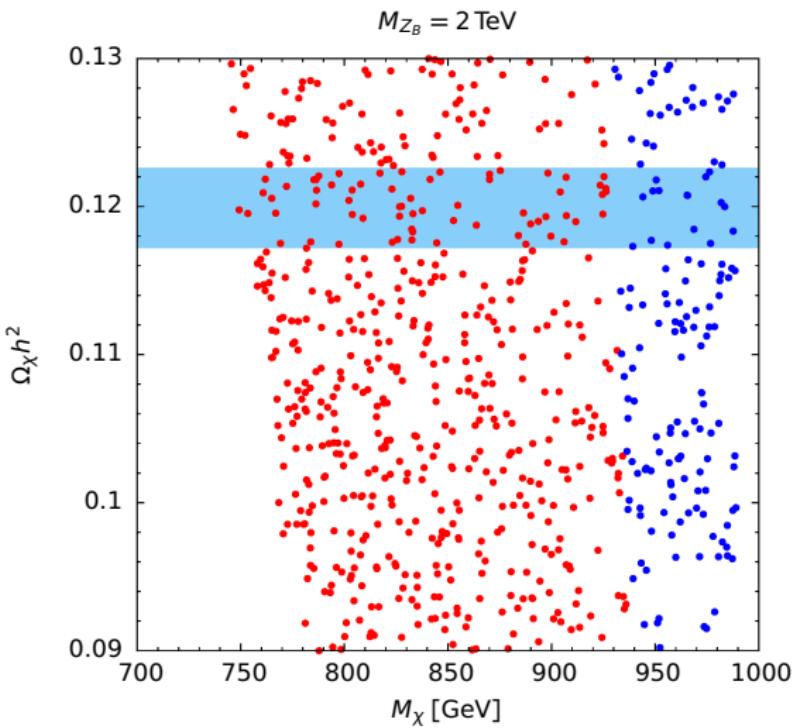
$$M_\chi, M_{Z_B}, g_B, \text{ and } B_1 + B_2$$

⇒ fully testable by combining LHC, DM direct detection,
DM relic density.

New Gauge Boson at the LHC



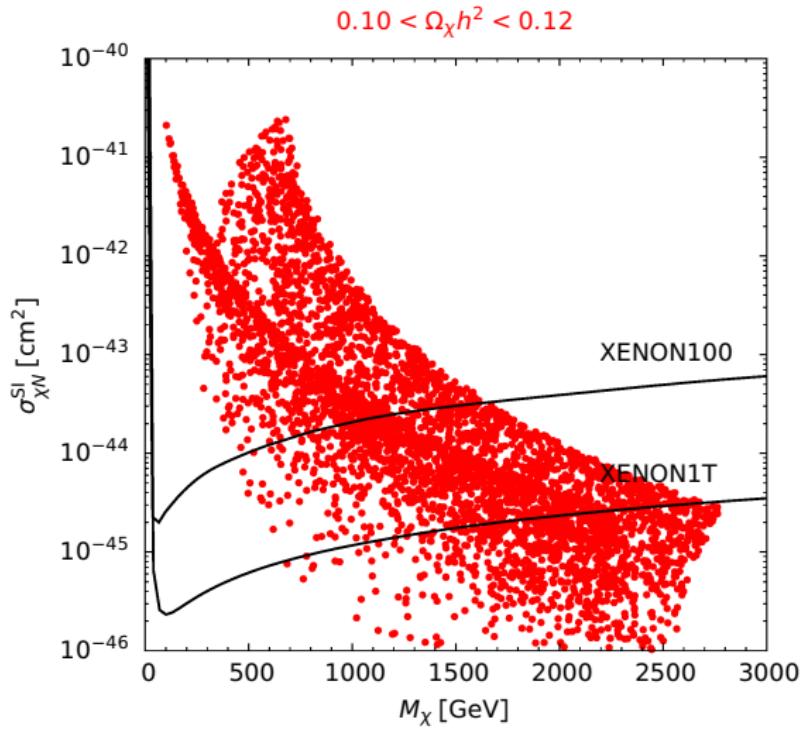
Dark Matter Relic Density



► Planck: $\Omega_{\text{DM}} h^2 = 0.1199 \pm 0.0027$

$g_B \in [0.10, 0.25]$
 $g_B \in [0.25, 0.50]$

Dark Matter Direct Detection



Left-right symmetric model

SM fields

$$G_{LR} = SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$$

$$Q_L \sim (\mathbf{2}, \mathbf{1}, 1/3)$$

$$Q_R \sim (\mathbf{1}, \mathbf{2}, 1/3)$$

$$\ell_L \sim (\mathbf{2}, \mathbf{1}, -1)$$

$$\ell_R \sim (\mathbf{1}, \mathbf{2}, -1)$$

- ▶ Connects neutrino masses and spontaneous parity violation.
- ▶ Standard version uses hybrid version of type I and type II seesaw mechanism for neutrino masses.

Pati, Salam, PRD **10** (1974) 275, Mohapatra, Pati, PRD **11** (1975) 2558, Senjanovic, Mohapatra, PRD **12** (1975) 1502

$$\Rightarrow SU(2)_L \otimes SU(2)_R \otimes U(1)_B \otimes U(1)_L$$

He, Rajpoot, PRD **41** (1990) 1636

Anomaly Cancellation

- ▶ Anomalies that need to be cancelled:

$$\mathcal{A}_1 \left(SU(2)_L^2 \otimes U(1)_B \right) = 3/2$$

$$\mathcal{A}_2 \left(SU(2)_L^2 \otimes U(1)_L \right) = 3/2$$

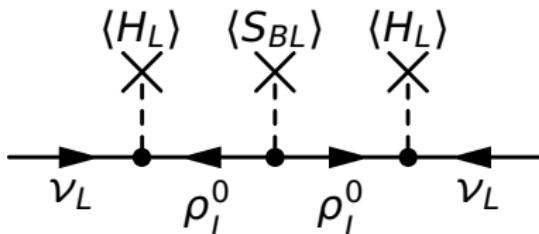
$$\mathcal{A}_3 \left(SU(2)_R^2 \otimes U(1)_B \right) = -3/2$$

$$\mathcal{A}_4 \left(SU(2)_R^2 \otimes U(1)_L \right) = -3/2$$

- ▶ Simplest solution: type III seesaw fields

$\rho_L \sim (\mathbf{3}, \mathbf{1}, -3/4, -3/4)$ and

$\rho_R \sim (\mathbf{1}, \mathbf{3}, -3/4, -3/4)$



M. Duerr, P. Fileviez Pérez, M. Lindner, arXiv:1306.0568 [hep-ph]

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Simple extension of the SM:

- ▶ B and L are gauge symmetries broken at a low scale
- ▶ no proton decay \Rightarrow no need for a desert
- ▶ Neutrino masses
- ▶ Fermionic DM candidate
- ▶ Testable at the LHC

Summary

Simple extension of the SM:

- ▶ B and L are gauge symmetries broken at a low scale
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Thank you!

Backup slides

Neutrino parameters

- Pontecorvo–Maki–Nakagawa–Sakata mixing matrix

$$\begin{aligned}s_{ij} &= \sin \theta_{ij} \\c_{ij} &= \cos \theta_{ij} \\&\alpha, \beta: \text{Majorana phases} \\&\delta: \text{Dirac phase}\end{aligned}$$

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \text{diag} \left(1, e^{i\frac{\alpha}{2}}, e^{i(\frac{\beta}{2} + \delta)} \right)$$

- Oscillation parameters ($\Delta m_{ij}^2 = m_i^2 - m_j^2$)

$\sin^2 \theta_{12}$	$\sin^2 \theta_{23}$	$\sin^2 \theta_{13}$
$0.302^{+0.013}_{-0.012}$	$0.413^{+0.037}_{-0.025} \oplus 0.594^{+0.021}_{-0.022}$	$0.0227^{+0.0023}_{-0.0024}$
$\delta_{CP}/^\circ$	$\Delta m_{21}^2 [10^{-5} \text{ eV}^2]$	$\Delta m_{31}^2 [10^{-3} \text{ eV}^2]$ (NO) $\Delta m_{32}^2 [10^{-3} \text{ eV}^2]$ (IO)
300^{+66}_{-138}	$7.50^{+0.18}_{-0.19}$	$+2.473^{+0.070}_{-0.067}$ $-2.427^{+0.042}_{-0.065}$

M. C. Gonzalez-Garcia et al., arXiv:1209.3023 [hep-ph]

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