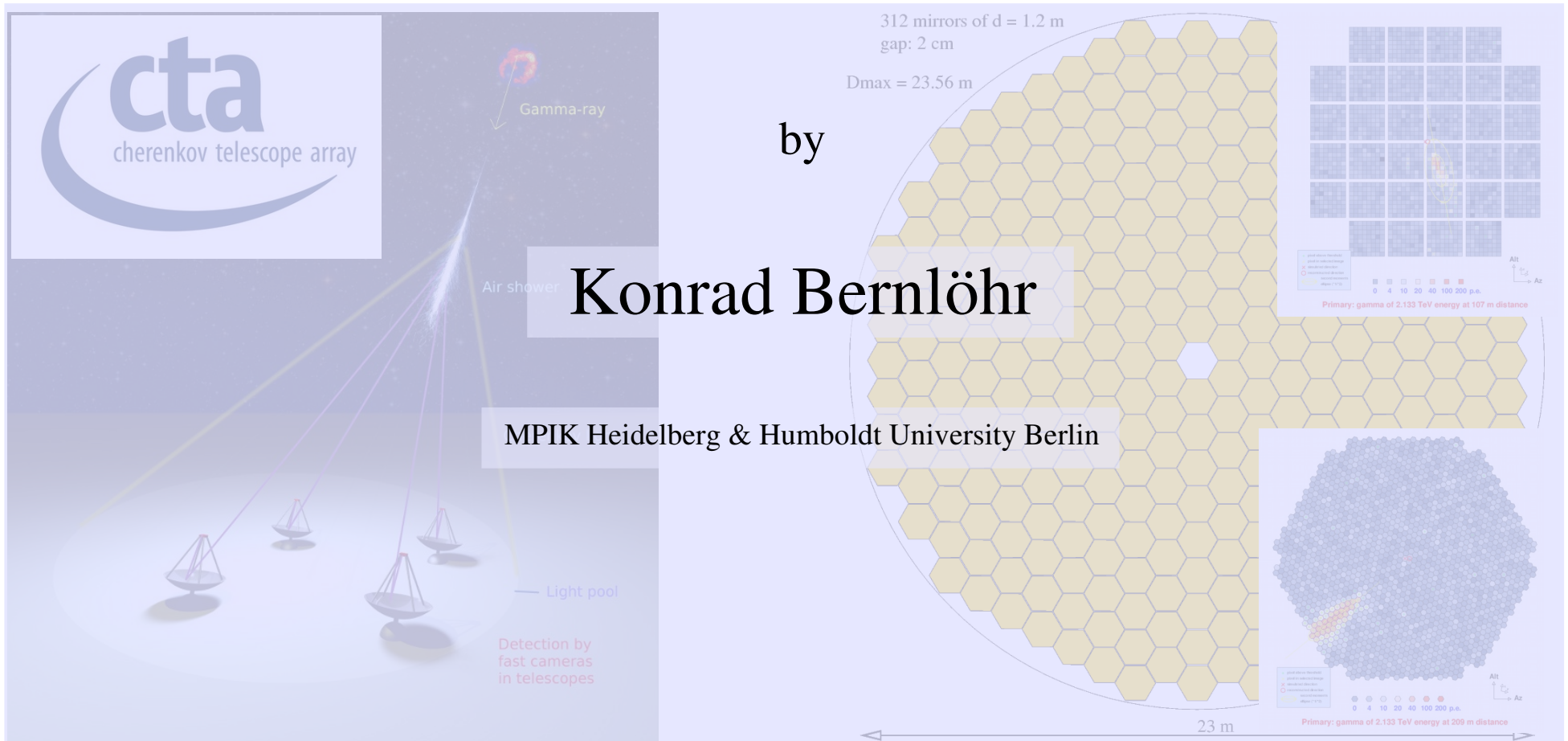


Tutorial 2: MC data reduction and analysis with 'read_cta' ('read_hess')



Goals of this tutorial

Show how to

- check what you have in your data,
- produce DSTs from raw data files
 - Note: a good integration of prod-2 traces is still missing!
- analysis procedures
- lookup tables
- output histograms for performance calculations
- calculating the sensitivity

read_hess

Syntax: /home/konrad32c/hess/hessio/bin/read_hess [options] [- | input_fname ...]

Options:

- p ps_filename (Write a PostScript file with camera images.)
- r level (Use 10/5 tail-cut image cleaning and redo reconstruction.)
 - level >= 1: show parameters from sim_hessarray.
 - level >= 2: redo shower reconstruction
 - level >= 3: redo image cleaning (and shower reconstruction with new image parameters)
 - level >= 4: redo amplitude summation
 - level >= 5: PostScript file includes original and new shower reconstruction.
- v (More verbose output)
- q (Much more quiet output)
- s (Show data explained)
- S (Show data explained, including raw data)
- ...

And many more options. Options in red not available with the reduced version read_hess_nr. Both full and reduced version are now distributed within CTA.

read_hess example

```
opts="-r 2 -u -q ${cuts} ...." # e.g.: --not-telescope 5,6,7,8
```

```
gamma_dst=" ../DST/gamma_${t1}_${t2}_-2.57.simhess-dst.gz"
```

```
proton_dst=" ../DST/proton_${t1}_${t2}_-2.70.simhess-dst.gz"
```

```
electron_dst=" ../DST/electron_${t1}_${t2}_-3.30.simhess-dst.gz"
```

```
read_hess ${opts} --powerlaw -2.57 --auto-lookup "${gamma_dst}"
```

```
read_hess ${opts} --powerlaw -2.57 --auto-lookup "${gamma_dst}"
```

```
read_hess ${opts} --powerlaw -2.57 --auto-lookup "${gamma_dst}" New!
```

```
read_hess ${opts} --powerlaw -2.57 "${gamma_dst}"
```

```
read_hess ${opts} --powerlaw -2.70 --theta-scale 6 "${proton_dst}"
```

```
read_hess ${opts} --powerlaw -3.30 --theta-scale 6 "${electron_dst}"  
Note: electrons at this point still assumed to follow power law
```

The read_hess output

Output from read_hess may include:

- Histograms (again in eventio/hessio format, can be converted to ROOT, HBOOK, eventio/iactio), some used as lookups for further analysis.
- Reconstructed/selected data (DSTs etc.)
- Text-mode n-tuples of MC/image/shower data.
- Postscript camera images.
- All the data explained in full detail.

Heidelberg CTA-MC Analysis 1

- Simple analysis (with updates since last meeting):
 - Cleaning: 4/7, 5/10, 8/12 p.e. tail-cuts and other scheme;
 - Image amplitude > 30 p.e., up to 200 p.e.
 - Image c.o.g. Radius + 0.5 image length $< 0.85 R_{\text{cam}}$ and variations of it (“**edge cut**”). Potential problem at high E .
 - Geometrical shower reconstruction (direction and core position) from Hillas parameters, using weighted mean of pair-wise intersections of image major axes.

This is certainly sub-optimal and can be improved!

Heidelberg CTA-MC Analysis 2

- Generating lookup tables width+length of gamma rays.
- Using the lookups, get mean reduced scaled width+length

$$mrscw = \frac{1}{N} \sum_{i=1, N} \frac{(w_i - \bar{w}(r_c, A, z))}{\sigma_w}$$

and cut on *mrscw* and *mrscL* for gammas and protons (“**shape cuts**”). Energy dependent cuts.

- Get angular resolution and apply point source selection (“**angle cut**”). Multiplicity-dependent cut.

Heidelberg CTA-MC Analysis 3

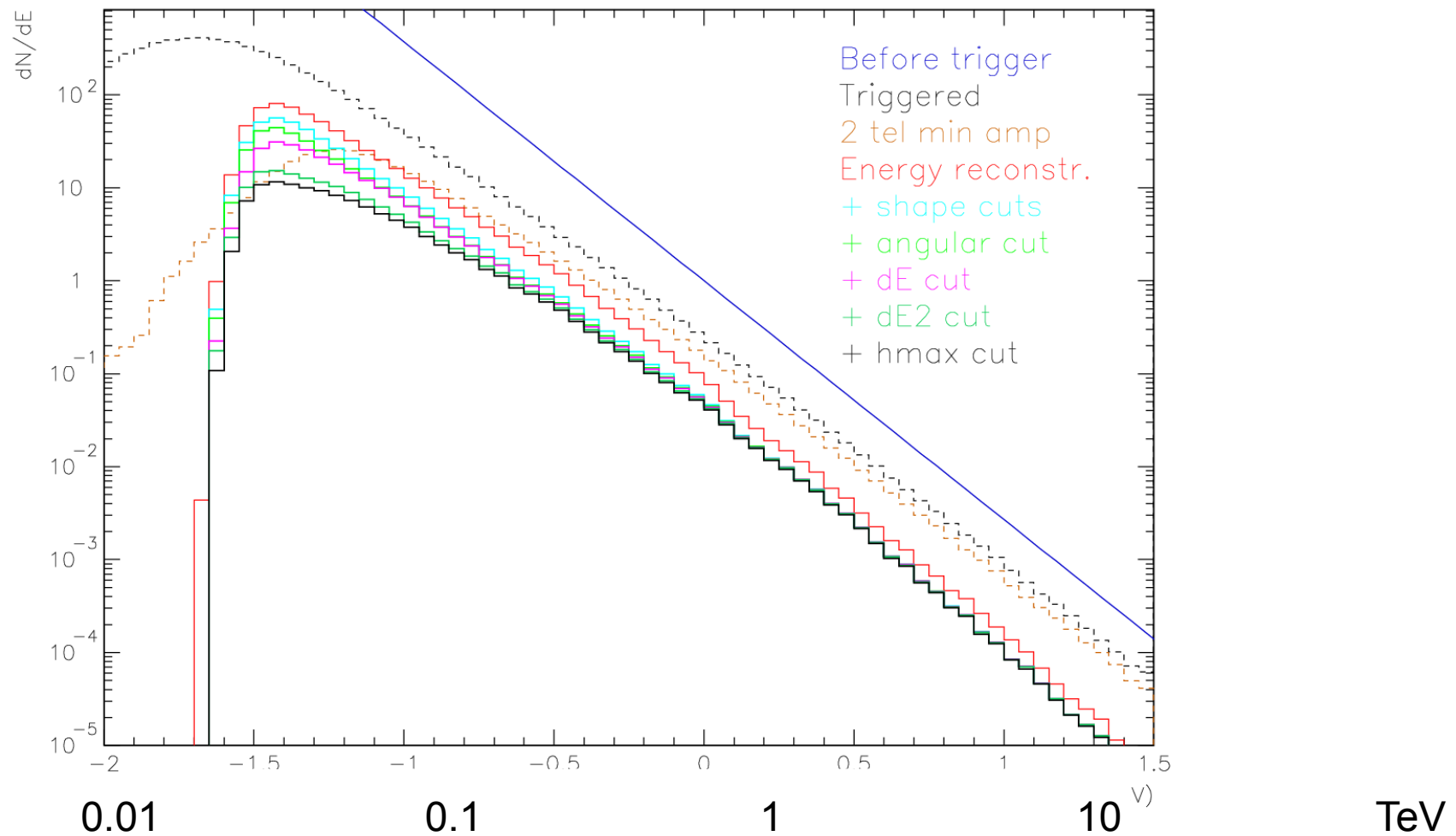
- Generating lookup tables for image amplitude / energy (I/E) as with width and length.
- Using that lookup an energy estimate and estimate of its fluctuation for each telescope is obtained.
- Get energy E , energy accuracy σ_E/E , and $\chi^2/\text{n.d.f.}$
- Discard showers with bad σ_E/E (“dE cut”).
Rejects also gammas at large (but energy dependent) R_c .
- Discard showers with bad $\chi^2/\text{n.d.f.}$ (“dE2 cut”).
- Calculate distance of shower maximum and discard showers inconsistent with gammas (“hmax cut”).

Heidelberg CTA-MC analysis 4

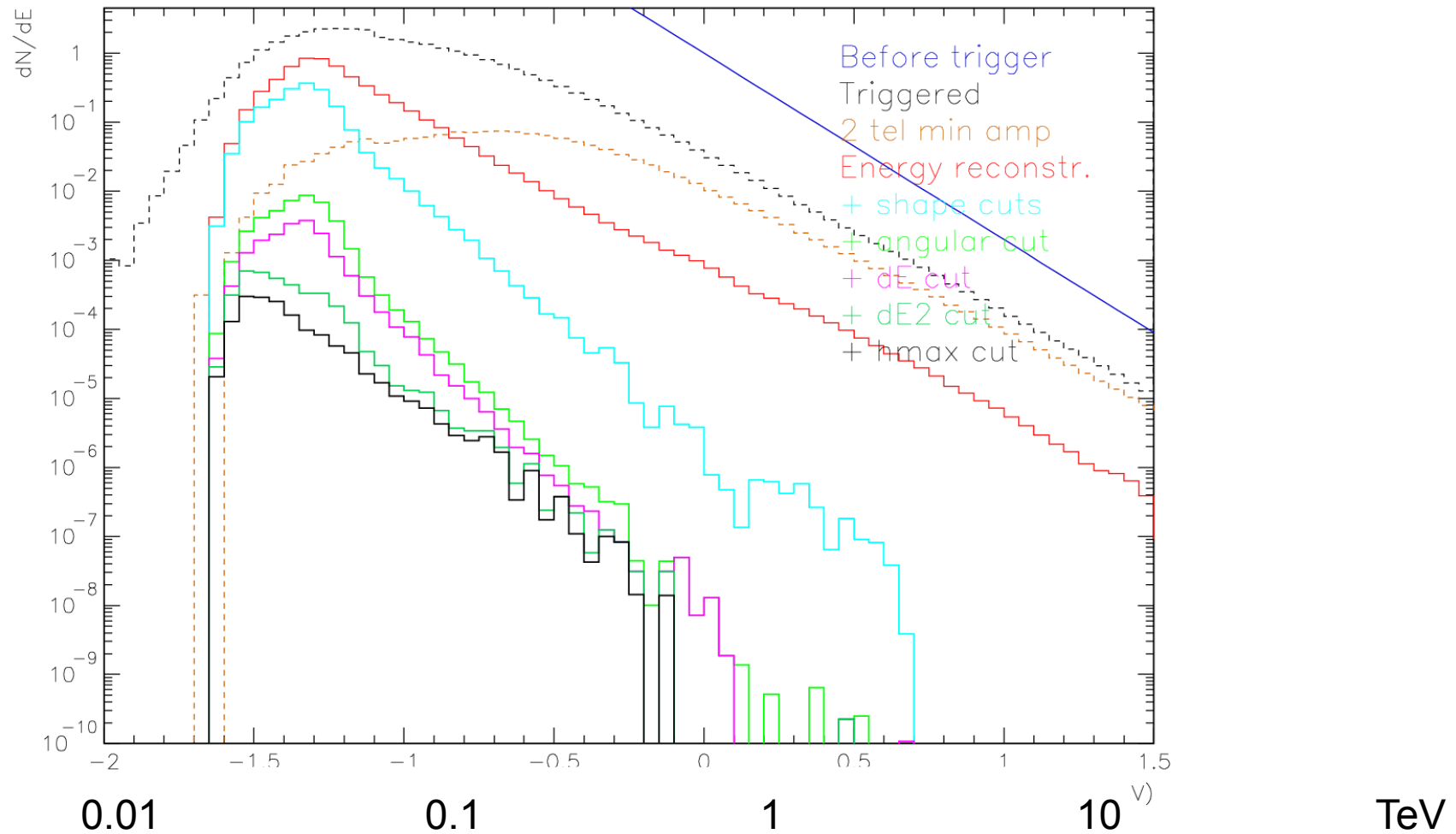
Also in the eventio/hessio format for some time now:

- Inter-telescope trigger time difference w.r.t. nominal shower plane.
- “Online” pixel pulse shape analysis (peak, rise, width). For pixel amplitudes above threshold.
- Pulse shape analysis summary at DST level:
peak time gradient (along major axis), r.m.s. residuals, mean pixel pulse widths (20%, 50%), mean rise time.

Cut efficiencies (gammas)



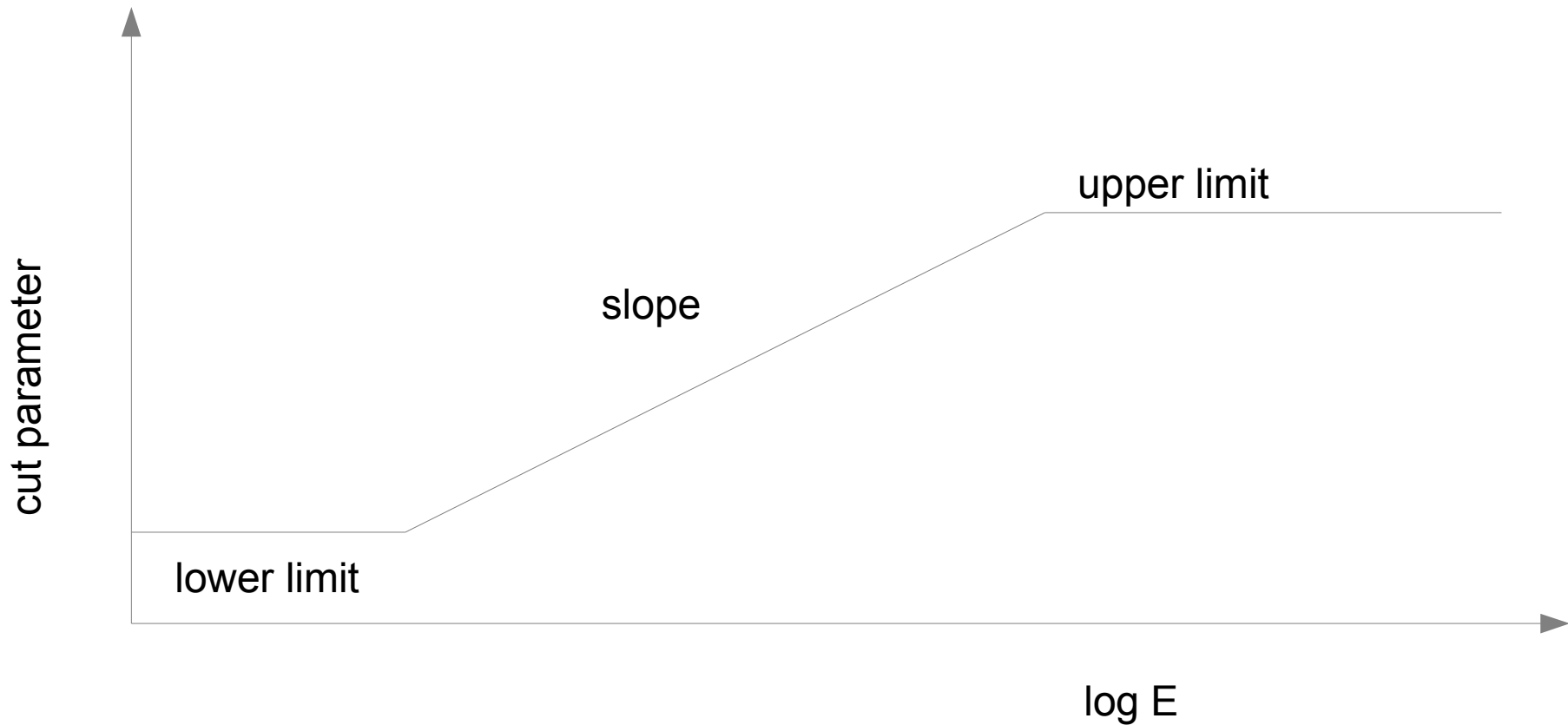
Cut efficiencies (protons)



Optimizing shape cuts etc.

- Background falls off rapidly at higher energies.
- Optimum shape cuts are, compared to fixed cuts:
 - more strict at low energies (better hadron rejection)
 - less strict at high energies (more gamma signal)
- Similar for additional cuts like
 - dE cut (accepting events with well-determined energy),
 - dE^2 cut (events with consistent energy estimates from individual telescopes), and
 - $hmax$ cuts (height of maximum as expected for gammas) .

Allowed dependence of cut parameters



Example read_hess parameters

Example (from initial prod-2 analysis, ~ from prod-1):

- `read_hess -r 2 -u -q --min-trg-tel 2 --min-tel 2 --min-pix 4 \
--min-amp 55 --tail-cuts 5,10 --shape-cuts -4.0,0.5,-2.0,1.2 \
--width-cut 0.5,0.3,0.2,2.0 --length-cut 1.2,0.6,1.2,2.0 \
--dE-cut 1.0,0.0,1.0,1.0 --dE2-cut 0.6,0.35,0.12,2.0 --hmax-cut 0.8 \
--only-telescopes 2,3,5,6,8,10,12,23-26,31-33,43-46,51-53,97-168 \
--type 1,1,5 --tail-cuts 7.655,15.310 --min-amp 81.550 --min-pix 4 \
--type 2,5,59 --tail-cuts 7.468,14.937 --min-amp 79.68 --min-pix 3 \
--type 3,97,168 --tail-cuts 1.991,3.983 --min-amp 24.91 --min-pix 3 \
--powerlaw -2.57 --auto-lookup -f ../gamma_dst.lis`

Energy-dependent cut values

- Energy-dependent cut parameter example:
dE2 is cutting on how consistent the energy estimates from the individual images are,
 - **--dE2-cut 0.6,0.35,0.12,2.0**
 - Lower limit of parameter: 0.6,
 - rising by 0.35 per decade in energy,
 - lower limit used for $\log_{10}(E/\text{TeV}) < 0.12$,
 - upper limit of $0.6+(2.0-0.12)*0.35$ used for $\log_{10}(E/\text{TeV}) > 2.0$

Telescope-type dependent parameters

Example: image cleaning tail-cuts:

- `--tail-cuts 5,10 \`
`--type 2,5,59 --tail-cuts 7.468,14.937`
- Default value: 5, 10 (p.e.), meant as nominal for 100 MHz NSB
 - applied if no type-specific value seen
 - used for file names
- Telescope type 2 is for telescopes with ID from 5 to 59
 - IDs not strictly necessary (auto grouping by optical parameters)
- For telescopes of type 2 the tail-cut levels are 7.468 and 14.937 p.e.
 - Higher tail-cut levels for telescopes with more NSB: $\propto \sqrt{\text{NSB}}$

Histogram files

- Generated in eventio-based format
- Conversion tools available into
 - old PAW (HBOOK), needs CERNLIB: hdata2hbook
 - ROOT histograms: hdata2root
- Some specialized tools directly working on them, like
 - gen_lookup (generating look-ups from accumulated hist.s)
- → Let's look into a histogram file