

LightEmission

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Chapter 1

lactLightEmission - simulate artificial light sources

1.1 Introduction

The data format used with the IACT package for CORSIKA can also be used to simulate 'artificial' light sources rather than Cherenkov light emitted by extensive air showers. Like with CORSIKA, you set up detector fiducial spheres in some observation level plane (usually just one, and due to the '-DIACT_NO_GRID' definition in rather than above that plane), start a simulation 'run', with a single or many events in it. Instead of the air showers, light is produced by user-defined light sources, with given positions, alignment, angular and spectral emission characteristics, and emitting a requested number of photons in bunches of a given number of photons per bunch.

Many parameters in CORSIKA run header and events headers make little sense with these simulations, although those used to control subsequent telescope or camera simulations work just like for air showers. Thus output from these simulations of artificial light sources can be processed by sim_telarray just like CORSIKA air shower simulations. Light sources may illuminate an entire telescope, or an array of telescopes, as with shower simulations. Or they may illuminate a camera directly or, in a dual-mirror telescope, after only reflection on the secondary mirror (see BYPASS_OPTICS parameter for sim_telarray).

The typical program needs a

- [Run](#) class constructor and output file assignment.
- Setting up one or more light sources.
- A loop over a given number of events, each with a NextEvent() call for the run object, in order to write out any photon bunches from the preceding event and initialize the next on, followed by an Emit() call for each of the defined light sources (or a subset of those light sources if they alternate from event to event).
- Output of photon bunches from the final event and clean-up will be performed by the [Run](#) class destructor.

A very simple program might look like:

```
#include "lactLightEmission.hh"
int main()
{
    int iev, nev=10;
    Run run(1);
    SimpleLightSource ls;
    for ( iev=0; iev<nev; iev++ )
    {
        run.NextEvent();
        ls.Emit(run,1000.,1.,0.);
    }
    return 0;
}
```

1.2 Light sources types

There are a number of types of light sources available, from simple, point-like sources with isotropic emission, to extended and/or moving sources with user-defined emission characteristics. These include in particular:

- [SimpleLightSource](#) : Isotropic emission at fixed wavelength and very simple time profile (top-hat or Gaussian), characterized by its Full-Width-at-Half-of-Maximum (FWHM) pulse duration.
- [BackgroundLightSource](#) : A diffuse light source emitting for a given time period, isotropic within a given angular range, and with a user-defined spectral characteristic which defaults to the known night-sky background spectrum.
- [LightSource](#) : A more generic kind of light source with user-defined spectral, angular, and temporal characteristics (which can be a keyword like 'Gauss' followed by a colon and a value, or a table from which the corresponding distribution gets loaded) which can have a linear extension along its primary direction and can move along that with some finite speed. This may emulate, for example, an LED flasher, a muon emitting Cherenkov light, or a laser beam scattered in the atmosphere.

Chapter 2

Class Index

2.1 Class List

Here are the classes, structs, unions and interfaces with brief descriptions:

BackgroundLightSource	A light source class with diffuse emission characteristics	7
DirectionVect	A normalized direction vector	8
LightSource	A much more generic type of light source than the SimpleLightSource	10
Line	A line is defined by a point on the line and the direction vector along the line	13
photon_bunch	14
PixLed	15
PixPos	15
Plane	A plane in 3-D space and intersections	15
Run	Run class takes care of global initialization and winding-down/cleanup as well as advancing from one event to the next	17
SimpleLightSource	Simple light source: monochromatic, isotropic, pulse shape is Gaussian or top hat. In addition: point source, stationary	20
SpaceVect	A 3-D spatial with some arithmetic operations	21
TelPos	Telescope position as in the CORSIKA inputs card TELESCOPE, includes radius of sphere	23

Chapter 3

File Index

3.1 File List

Here is a list of all documented files with brief descriptions:

fake-muon.cc	Program to generate photon data similar to single muons in CORSIKA, using the CORSIKA/IACT interface	25
ff-1m.cc	Flat-fielding light source simulation for single-reflector cameras, using the CORSIKA/IACT interface, simulating light emitted from a flasher supposed to be in the center of the dish. The resulting photon bunches can be simulated by sim_telarray with the "Bypass_Optics=2" (or 1 if truly 1M) configuration to bypass reflection on the primary mirror	26
ff-gct.cc	Flat-fielding light source simulation for the GCT (CHEC) cameras, using the CORSIKA/IACT interface, simulating light emitted from four flashers supposed to be in the corners of the camera, here in a coordinate system relative to the center of the secondary mirror. The resulting photon bunches can be simulated by sim_telarray with the "Bypass_Optics=1" configuration to bypass reflection on the primary mirror but include reflection on the secondary mirror	27
fpls.cc	Focal plane lightsource example program for artificial light sources using the CORSIKA/IACT interface	28
lactLightEmission.cc	Simulate an artificial light source instead of running CORSIKA and write the resulting photon bunches in the same format as with shower simulation going through the IACT/atmo interface	28
lactLightEmission.hh	Class definitions and common inline code for the lactLightEmission package	32
ile_version.hh	This is the lactLightEmission package version 2023-01-31	34
ls-beam.cc	Simulate light from a laser beam crossing the field of view of a telescope, some of the laser light getting scattered along the beam. The modelling of the laser beam and its attenuation along the path is very simplified, and for performance reasons an isotropic emission angle distribution is to be preferred for now. The resulting photon bunches can be simulated by sim_telarray with the (default) "Bypass_Optics=0" configuration, not bypassing any optics raytracing	35
ls-geo-test.cc	A few tests for the LightEmission geometry classes	36
ls-test.cc	Test program for artificial light sources using the CORSIKA/IACT interface	37
ls-test2.cc	Test program for artificial light sources using the CORSIKA/IACT interface	37

nsbls.cc	Nightsky background lightsource example program for artificial light sources using the CORSIKA/IACT interface	38
octo.cc	Superluminal octocopter example program for artificial light sources using the CORSIKA/IACT interface	38
pixled.cc	A program to set an assumed LED in front of a specified subset of pixels in a camera as defined for sim_telarray. The main purpose is to test that trigger definitions work as intended	39
xyls.cc	Test program for artificial light sources using the CORSIKA/IACT interface. This program is not intended to be of practical use	41
xyzls.cc	Lightsource example program for artificial light sources using the CORSIKA/IACT interface . .	42

Chapter 4

Class Documentation

4.1 BackgroundLightSource Class Reference

A light source class with diffuse emission characteristics.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- **BackgroundLightSource** (const char *spec_fname=0, double pwid=0., double pdel=0., double vc=0., const [DirectionVect](#) &direct=[DirectionVect](#)(0., 0.,-1.))
- void **SetMaxViewcone** (double mvc)
- void **SetSpectrum** (const char *fname)
- void **SetEmissionPulseWidth** (double pw)
- void **SetDelay** (double t)
- int **Emit** (const [Run](#) &run, double photons, double bsize, double when)

Emit the requested number of photons in bunches of the requested size at the requested time. Depending on the distance between light source and detector fiducial sphere(s) and the solid angle actually subtended, the number of bunches that need to be written might be lower than expected from looking at photons and bsize alone.

4.1.1 Detailed Description

A light source class with diffuse emission characteristics.

The background (mainly nightsky background) light source is not a localized emitter but corresponds to a uniform flux of isotropic background, actually generated at the fiducial sphere. If there is more than one fiducial sphere the requested number of photons are split up to be distributed according to sphere cross section areas among the sphere. For efficiency reasons the emission is not across the whole hemisphere but within a 'viewing cone' up to a given limiting angle from a principal direction (which gets recorded as direction of the 'primary'). The spectrum can be specified as for the generic [LightSource](#) class but the default spectrum is the Benn&Ellison NSB spectrum. The light pulse time distribution can only be a flat-top distribution, the default being instantaneous for compatibility with other light source classes.

4.1.2 Member Function Documentation

4.1.2.1 Emit()

```
int BackgroundLightSource::Emit (
    const Run & run,
    double photons,
    double bsize,
    double when )
```

Emit the requested number of photons in bunches of the requested size at the requested time. Depending on the distance between light source and detector fiducial sphere(s) and the solid angle actually subtended, the number of bunches that need to be written might be lower than expected from looking at photons and bsize alone.

Parameters

<i>run</i>	Run class
<i>photons</i>	Number of photons that would be emitted into 4 pi steradian (isotropic).
<i>bsize</i>	Intended bunch size. A conservative approach would use a value of 1.0. Any value below one will be replaced by 1.0.
<i>when</i>	Nominal emission time, with individual photon bunches obtaining a random time value to reproduce the configured pulse shape, on average. Unit: nanoseconds.

Returns

0 (OK)

The documentation for this class was generated from the following files:

- [lactLightEmission.hh](#)
- [lactLightEmission.cc](#)

4.2 DirectionVect Class Reference

A normalized direction vector.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- [DirectionVect](#) ()
 - Note that length must be normalized, cannot be (0,0,0).*
- **DirectionVect** (double cxp, double cyp, double czp)
- **DirectionVect** (double alt, double az=0.)
- **DirectionVect** (const [SpaceVect](#) &sv)
- [SpaceVect operator*](#) (double d) const
 - A given distance along the direction defines a space vector.*
- [SpaceVect SV](#) () const
 - Reverse conversion without multiplication.*
- double [Altitude](#) () const
 - Altitude (or Elevation) angle above x/y plane, in radians.*

- double [Polar](#) () const
Angle from pole (or zenith), in radians.
- double [Azimuth](#) () const
The azimuthal angle in the x/y plane, counted in the mathematical (counter-clockwise!) sense, in radians.
- double [AltitudeDeg](#) () const
The Altitude angle converted into degrees.
- double [PolarDeg](#) () const
The polar (or zenith) angle converted into degrees.
- double [AzimuthDeg](#) () const
The azimuthal angle converted into degrees.
- double [AngleTo](#) (const [DirectionVect](#) &d2) const
Angle between two direction vectors (or via typecase between a direction vector and a space vector), in radians.
- double [AngleToDeg](#) (const [DirectionVect](#) &d2) const
Angle between to vectors converted into degrees.
- double [SProd](#) (const [DirectionVect](#) &dv2) const
Scalar (inner) product of two direction vectors.
- double **operator*** (const [DirectionVect](#) &dv2) const
- double [SProd](#) (const [SpaceVect](#) &sv2) const
Scalar (inner) product with a space vector, without typecast.
- [SpaceVect](#) [VProd](#) (const [SpaceVect](#) &sv2) const
Vector (outer) product with a space vector is a space vector (not normalized).
- [SpaceVect](#) [VProd](#) (const [DirectionVect](#) &dv2) const
Vector (outer) product with a direction vector is still not normalized.
- [DirectionVect](#) [NVProd](#) (const [DirectionVect](#) &dv2) const
Normalized vector (outer) product of two direction vectors.
- bool [IsParallelTo](#) (const [DirectionVect](#) &dv2) const
Better check if two vectors are (anti-) parallel before asking for the normalized vector product.
- bool [AlmostParallelTo](#) (const [DirectionVect](#) &dv2) const
Perhaps two direction vectors are (anti-) parallel within possible rounding errors.
- [DirectionVect](#) & [Rotate_z](#) (double phi)
Rotate counter-clockwise around given axis (active rotation)
- [DirectionVect](#) & [Rotate_x](#) (double phi)
- [DirectionVect](#) & [Rotate_y](#) (double phi)
- [DirectionVect](#) & [Rotate](#) (double alpha, double beta, double gamma)
Active rotation with Euler angles in standard x convention (sequence: z,x',z'')
- [DirectionVect](#) & [Scatter](#) (double dtheta, double dphi)
- ostream & [Print](#) (ostream &os) const
Show human-readable version of the class.

Public Attributes

- double **cx**
 - double **cy**
 - double **cz**
- Leave the direction cosines public for efficient access.*

4.2.1 Detailed Description

A normalized direction vector.

The [DirectionVect](#) class is, like [SpaceVect](#), just a 3-tuple of x/y/z coordinates but, in contrast to [SpaceVect](#), meant as a normalized(!) direction vector and not a position or difference of positions (which is again a [SpaceVect](#)). Scaling a [DirectionVect](#) with a length gives a [SpaceVect](#).

4.2.2 Member Function Documentation

4.2.2.1 Scatter()

```
DirectionVect& DirectionVect::Scatter (
    double dtheta,
    double dphi ) [inline]
```

New direction after scatter by angle dtheta between 0 and pi and (typically randomly chosen) angle dphi around the initial direction.

The documentation for this class was generated from the following file:

- [lactLightEmission.hh](#)

4.3 LightSource Class Reference

A much more generic type of light source than the [SimpleLightSource](#).

```
#include <IactLightEmission.hh>
```

Public Member Functions

- [LightSource](#) ()
Default light source constructor.
- [LightSource](#) (const [SpaceVect](#) &ls_pos, const char *spec_fname=0, const char *pulse_fname=0, const char *angdist_fname=0, const [DirectionVect](#) &direct=[DirectionVect](#)(0., 0.,-1.))
- [LightSource](#) (const [SpaceVect](#) &ls_pos, const string &spec_fname, const string &pulse_fname="Simple:0", const string &angdist_fname="isotropic", const [DirectionVect](#) &direct=[DirectionVect](#)(0., 0.,-1.))
- [LightSource](#) (const [SpaceVect](#) &ls_pos, const [LightSource](#) &ls1)
Light source constructor copying an existing source into another place.
- const [SpaceVect](#) & [GetPos](#) () const
Get the (starting) position of a light source.
- double [GetX](#) () const
- double [GetY](#) () const
- double [GetZ](#) () const
- void [SetSpectrum](#) (const char *fname)
Set the spectrum of a light source.
- void [SetSpectrum](#) (const string &fname)
- void [SetEmissionPulse](#) (const char *fname)
Set the pulse shape parameters for a light source.
- void [SetEmissionPulse](#) (const string &fname)
- void [SetAngularDistribution](#) (const char *fname)
Set the angular distribution parameters for a light source.
- void [SetAngularDistribution](#) (const string &fname)
- void [SetPosition](#) (const [SpaceVect](#) &npos)
Set the position of a light source.

- void [SetDirection](#) (const [DirectionVect](#) &dir)
Set the orientation of a light source.
- void [SetDirectionTo](#) (const [SpaceVect](#) &topos)
Make a light source point from its current location towards another location.
- void [SetLength](#) (double l)
Set the length of a light source.
- void [SetSpeed](#) (double b)
Set the propagation speed along the length of a light source, in units of c.
- void [SetDelay](#) (double t)
Set the time delay after which the light source should fire.
- void [SetName](#) (const string &n)
Set the name of a light source for later identification.
- void [SetName](#) (int n)
- double [GetLength](#) () const
- double [GetSpeed](#) () const
- double [GetDelay](#) () const
- const string & [GetName](#) () const
- int [Emit](#) (const [Run](#) &run, double photons, double bsize, double when)
Initiate actual emission of light pulse from this one light source.

4.3.1 Detailed Description

A much more generic type of light source than the [SimpleLightSource](#).

A fairly generic light source with a tabulated spectrum, a tabulated pulse shape, tabulated emission angle profile symmetric around a given pointing direction. If the emission angle distribution is missing, it is reduced to an isotropically emitting light source. Predefined angular distribution names not requiring to read any table are " \leftrightarrow Isotropic", "Gauss:<rms>", "Rayleigh", "Cone:<angle>", and "FilledCone:<angle>" as well as "Parallel", with all angles given in degrees, all with respect to the given direction vector (vertically downwards if missing). If the light source has a non-zero length and velocity (in units of the vacuum speed of light), it is handled as a moving source, in the given direction.

If the pulse shape is missing, it is reduced to instantaneous emission. Predefined pulse shapes not requiring to read any table are "Simple:<width>[,<mean>]" (for top-hat shape of given width [ns], like for a [SimpleLightSource](#)) or "Gauss:<rms>" (σ =r.m.s. value in [ns]) and others. Unfortunately, with versions prior to 2020-12-01 the "Gauss:..." value was interpreted as a FWHM (as in [SimpleLightSource](#)), in contrast to documentation. Therefore, more explicit variants have been added: "Gauss-rms:<rms>" (r.m.s. value in [ns]) and "Gauss-fwhm:<fwhm>" (giving the FWHM duration of a Gaussian pulse profile [ns], corresponding to a [SimpleLightSource](#) with the 'gs' (Gauss) flag set to true).

Optional <mean> time of emission (default 0) also in [ns].

If the spectrum is missing, a wavelength of 400 nm is assumed. If the spectrum string represents a number it is interpreted as a fixed wavelength in [nm]. Another predefined distribution is "Cherenkov:<from_wl>:<to_wl>", resulting in a $1/\lambda^2$ distribution from from_wl to to_wl (both in [nm]). In all other cases files are opened to read in the corresponding distributions.

Remaining restrictions are that the light source is still point-like or on a line, with a possible propagation of the point of emission along that line. There is no correlation of angular characteristics or spectrum with time of emission.

4.3.2 Constructor & Destructor Documentation

4.3.2.1 LightSource() [1/2]

```
LightSource::LightSource (
    const SpaceVect & ls_pos,
    const char * spec_fname = 0,
    const char * pulse_fname = 0,
    const char * angdist_fname = 0,
    const DirectionVect & direct = DirectionVect(0.,0.,-1.) ) [inline]
```

Light source constructor that can pass most, but not all, parameters along, with text parameters as C strings. Some less frequently used parameters may have to be set after creation.

4.3.2.2 LightSource() [2/2]

```
LightSource::LightSource (
    const SpaceVect & ls_pos,
    const string & spec_fname,
    const string & pulse_fname = "Simple:0",
    const string & angdist_fname = "isotropic",
    const DirectionVect & direct = DirectionVect(0.,0.,-1.) ) [inline]
```

Light source constructor that can pass most, but not all, parameters along, with text parameters as C++ strings. Some less frequently used parameters may have to be set after creation.

4.3.3 Member Function Documentation

4.3.3.1 Emit()

```
int LightSource::Emit (
    const Run & run,
    double photons,
    double bsize,
    double when )
```

Initiate actual emission of light pulse from this one light source.

Emit the requested number of photons in bunches of the requested size at the requested time. Depending on the distance between light source and detector fiducial sphere(s) and the solid angle actually subtended, as well as the angular distribution chosen, the number of bunches that need to be written might be lower than expected from looking at photons and bsize alone.

This more generic form of a light source allows for arbitrary angular emission shapes. Instead of being a point source, it can have a length (from the starting point towards a given direction). The actual emission point can move along that length with a given speed - which can cover scattered light from a narrow laser beam as well as Cherenkov or fluorescence emission from a charged particle.

Parameters

<i>run</i>	Run class
<i>photons</i>	Number of photons that would be emitted into the accessible solid angle (isotropic or in any configured angular distribution).
<i>bsize</i>	Intended bunch size. A conservative approach would use a value of 1.0. Any value below 1.0 will be replaced by 1.0.
<i>when</i>	Nominal emission time, with individual photon bunches obtaining a random time value to reproduce the configured pulse shape, on average. Unit: nanoseconds.

Returns

0 (OK)

The documentation for this class was generated from the following files:

- [lactLightEmission.hh](#)
- [lactLightEmission.cc](#)

4.4 Line Class Reference

A line is defined by a point on the line and the direction vector along the line.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- **Line** (const [SpaceVect](#) &ssv, const [DirectionVect](#) &sdv)
- const [SpaceVect](#) & **Position** () const
Retrieve configured point on line.
- const [DirectionVect](#) & **Direction** () const
- double **DistanceTo** (const [SpaceVect](#) &p) const
Shortest distance of a point in space from the line.
- double **DistanceFrom** (const [SpaceVect](#) &p) const
- double **Distance** (const [SpaceVect](#) &p) const
- [SpaceVect](#) **DistanceVectorFrom** (const [SpaceVect](#) &p) const
Vector representing the shortest line from a point to the given line.
- [SpaceVect](#) **DistanceVectorTo** (const [SpaceVect](#) &p) const
- double **DistanceTo** (const [Line](#) &l2) const
Shortest distance between two lines. Equivalent to `DistanceVectorTo().Length()` but faster.
- double **DistanceFrom** (const [Line](#) &l2) const
Same as `DistanceTo()` as the direction does not matter.
- double **Distance** (const [Line](#) &l2) const
Same as `DistanceTo()` and `DistanceFrom()`.
- ostream & **Print** (ostream &os) const
Show human-readable version of the class.

Public Attributes

- [SpaceVect](#) sv
Supporting vector on line.
- [DirectionVect](#) dv
Direction vector along line.

4.4.1 Detailed Description

A line is defined by a point on the line and the direction vector along the line.

The line is meant as infinite in either direction, although the defined point and direction can be retrieved. Any distance or distance vector is with respect to the both-sided infinite line.

4.4.2 Member Function Documentation

4.4.2.1 DistanceVectorTo()

```
SpaceVect Line::DistanceVectorTo (
    const SpaceVect & p ) const [inline]
```

Vector representing the shortest line (v) from the given line (starting at p-v) to a point (p). That is the opposite of [DistanceVectorFrom\(\)](#).

The documentation for this class was generated from the following file:

- [lactLightEmission.hh](#)

4.5 photon_bunch Struct Reference

Public Attributes

- double [photons](#)
Size of photon bunch.
- double [weight](#)
Weight = 1.0 (no thinning)
- double **xem**
- double **yem**
- double [zem](#)
Emission position.
- double **xol**
- double **yol**
- double [tol](#)
Position arriving at "observation level" (below telescope)
- double **dx**
- double **dy**
- double [dt](#)
Shift in ray propagation.
- double **u**
- double **v**
- double [w](#)
All direction cosines (not necessarily normalized)
- double [ctime](#)
Time of emission.
- double [lambda](#)
Wavelength of photons in bunch.

The documentation for this struct was generated from the following file:

- [lactLightEmission.cc](#)

4.6 PixLed Struct Reference

Public Member Functions

- **PixLed** (int ipix, bool rq=false, const string &p="", const string &s="", double d=0.)
- **PixLed** (int ipix, bool rq, const string &p, const string &s, double d, const vector< int > slave_ids)

Public Attributes

- int **id**
- bool **required**
- string **pulse**
- string **spectrum**
- double **delay**
- vector< int > **slaves**

The documentation for this struct was generated from the following file:

- [pixled.cc](#)

4.7 PixPos Struct Reference

Public Member Functions

- **PixPos** (int pid, double px, double py)

Public Attributes

- int **id**
- double **x**
- double **y**

The documentation for this struct was generated from the following file:

- [pixled.cc](#)

4.8 Plane Class Reference

A plane in 3-D space and intersections.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- [Plane](#) ()
Default plane is in x/y.
- [Plane](#) (const [DirectionVect](#) &nvs, double ds)
Plane defined by normal vector and distance to origin.
- [Plane](#) (const [SpaceVect](#) &sv, const [DirectionVect](#) &v1, const [DirectionVect](#) &v2)
Plane defined by support point on plane and two (non-parallel) direction vectors spanning the plane.
- [Plane](#) (const [SpaceVect](#) &sv, const [DirectionVect](#) &n)
Plane defined by support point on plane and the normal vector.
- const [DirectionVect](#) & [Direction](#) () const
Direction of normal vector, e.g. for calculating an angle with a line or such.
- double [DistanceTo](#) (const [SpaceVect](#) &p) const
Distance of a point in space from the plane.
- double [DistanceFrom](#) (const [SpaceVect](#) &p) const
- double [Distance](#) (const [SpaceVect](#) &p) const
- bool [HasIntersection](#) (const [Line](#) &l) const
Line is not parallel to plane and not in plane, thus must have an intersection.
- [SpaceVect](#) [Intersection](#) (const [Line](#) &l) const
Actual intersection, returns (0,0,0) if not possible.
- ostream & [Print](#) (ostream &os) const
Show human-readable version of the class.

Public Attributes

- [DirectionVect](#) nv
Normal vector to plane.
- double d

4.8.1 Detailed Description

A plane in 3-D space and intersections.

The plane is internally defined by its normal direction vector plus a constant, allowing for easy calculation of the distance of any points from the plane.

4.8.2 Member Data Documentation

4.8.2.1 d

```
double Plane::d
```

In plane definition: $nv[0]*x + nv[1]*y + nv[2]*z + d = 0$ for any point (x,y,z) being on the plane, within numerical precision.

The documentation for this class was generated from the following file:

- [lactLightEmission.hh](#)

4.9 Run Class Reference

`Run` class takes care of global initialization and winding-down/cleanup as well as advancing from one event to the next.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- `Run` (int rn=1, double h=2000e2, int atm=1, double r=12.5e2)
Run class constructor.
- `Run` (int rn, double h, const char *atmprof, double r=12.5e2)
Run class constructor.
- `Run` (int rn, double h, const string &atmprof, double r=12.5e2)
Run class constructor.
- void `SetOutput` (const char *fn)
We need a file name for the output.
- void `SetOutput` (const string &fn)
- void `SetDirection` (double stheta, double sph)
- void `SetDirection` (const `DirectionVect` &d)
- void `SetPrimary` (int prmid)
Report a primary particle type as requested (where applicable)
- void `SetPrimary` (const string &prnmam)
- void `SetEnergy` (double E_gev)
For some applications we may even report a useful energy for the primary.
- void `AddTelescope` (const `TelPos` &tp)
If the first telescope should not be at (0,0,0) or if you need more than one:
- void `AddHistory` (const char *txt)
Simplified function for adding history lines:
- void `AddHistory` (const string &txt)
- void `AddCmdLineHistory` (int argc, char **argv)
- void `SetCore` (double xcm, double ycm)
Report given core position in event header.
- void `SetScat` (int n, double r, double y)
Set random array scattering.
- void `SetReflDx` (double n)
Set a fixed index of refraction rather than deriving it from atmospheric profile.
- void `SetBunchSize` (double s)
Report maximum bunch size according to application; has no practical consequences.
- void `NextEvent` ()
Start of next event in run class.
- `size_t NumTel` () const
Return number of configured "telescopes".
- const `TelPos` & `Telescope` (size_t itel) const
Return telescope position.
- double `Area` () const
Sum of telescope sphere fiducial cross section areas.
- double `Area` (size_t itel) const
Individual telescope cross section.
- double `ObsLev` () const

- Return configured observation level [cm] a.s.l.*

 - int `Atmosphere` () const
Return configured atmospheric profile number.
 - int `RunNumber` () const
Return configured run number.
 - int `EventNumber` () const
Return current event number.
 - double `Rmax` () const
Return radius containing all fiducial spheres.
 - int `Primary` () const
 - int `Energy` () const
 - double `ReflDx` (double dz) const
Index of refraction at given altitude difference above observation level.
 - double `ReflDz` () const
Index of refraction at observation level.
 - double `Density` (double dz) const
Density [g/cm³] at given altitude difference above observation level.
 - double `Density` () const
Density [g/cm³] at observation level.
 - double `AirLightSpeed` () const
Propagation speed of light in air at observation level.
 - double `AirLightSpeed` (double dz) const
Propagation speed of light in air above observation level.
 - double `MeanAirLightSpeed` (double dz) const
Approximate mean propagation speed of light in air from start to observation level.

4.9.1 Detailed Description

`Run` class takes care of global initialization and winding-down/cleanup as well as advancing from one event to the next.

4.9.2 Constructor & Destructor Documentation

4.9.2.1 Run() [1/3]

```
Run::Run (
    int rn = 1,
    double h = 2000e2,
    int atm = 1,
    double r = 12.5e2 )
```

`Run` class constructor.

Parameters

<code>rn</code>	<code>Run</code> number.
<code>h</code>	Height of observation level a.s.l. [cm].
<code>atm</code>	Atmospheric profile table number -> atmprofd.dat.
<code>r</code>	Default radius of telescope fiducial sphere [cm].

4.9.2.2 Run() [2/3]

```
Run::Run (
    int rn,
    double h,
    const char * atmprof,
    double r = 12.5e2 )
```

[Run](#) class constructor.

Parameters

<i>rn</i>	Run number.
<i>h</i>	Height of observation level a.s.l. [cm].
<i>atmprof</i>	Atmospheric profile table given by name.
<i>r</i>	Default radius of telescope fiducial sphere [cm].

4.9.2.3 Run() [3/3]

```
Run::Run (
    int rn,
    double h,
    const string & atmprof,
    double r = 12.5e2 )
```

[Run](#) class constructor.

Parameters

<i>rn</i>	Run number.
<i>h</i>	Height of observation level a.s.l. [cm].
<i>atmprof</i>	Atmospheric profile table given by name.
<i>r</i>	Default radius of telescope fiducial sphere [cm].

4.9.3 Member Function Documentation**4.9.3.1 NextEvent()**

```
void Run::NextEvent ( )
```

Start of next event in run class.

Starting the next event includes emitting any header material data blocks for the first event and it includes wrapping up the previous event, including writing its photon bunches, before emitting the header of the new event.

4.9.3.2 SetCore()

```
void Run::SetCore (
    double xcm,
    double ycm )
```

Report given core position in event header.

Set a 'core position' info in the event header. It does not shift any 'telescope' positions or light source positions but changes event header data only in order to enhance consistency between fake muons (or other Cherenkov emitting light sources) and actual CORSIKA muon simulations.

Parameters

<i>xcm</i>	Shower core or track impact position x value at observation level [cm].
<i>ycm</i>	Shower core or track impact position y value at observation level [cm].

4.9.3.3 SetDirection()

```
void Run::SetDirection (
    double stheta,
    double sphl ) [inline]
```

Set the 'shower direction' to report, keeping in mind that the CORSIKA azimuth reports as 0 when coming from South, $\pi/2$ from East, ...

The documentation for this class was generated from the following files:

- [lactLightEmission.hh](#)
- [lactLightEmission.cc](#)

4.10 SimpleLightSource Class Reference

Simple light source: monochromatic, isotropic, pulse shape is Gaussian or top hat. In addition: point source, stationary.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- **SimpleLightSource** (const [SpaceVect](#) &ls_pos, double ls_wlen=400., double ls_fwhm=0.0, bool gs=true)
- void **SetPosition** (const [SpaceVect](#) &ls_pos)
- void **SetWavelength** (double ls_wlen)
- void **SetPulseFwhm** (double ls_fwhm)
- void **SetGaussianShape** (bool ls_gs)
- const [SpaceVect](#) & **Position** () const
- double **Wavelength** () const
- double **PulseFwhm** () const
- bool **HasGaussianPulse** () const
- int **Emit** (const [Run](#) &run, double photons, double bsize, double when)

Emit the requested number of photons in bunches of the requested size at the requested time. Depending on the distance between light source and detector fiducial sphere(s) and the solid angle actually subtended, the number of bunches that need to be written might be lower than expected from looking at photons and bsize alone.

4.10.1 Detailed Description

Simple light source: monochromatic, isotropic, pulse shape is Gaussian or top hat. In addition: point source, stationary.

4.10.2 Member Function Documentation

4.10.2.1 Emit()

```
int SimpleLightSource::Emit (
    const Run & run,
    double photons,
    double bsize,
    double when )
```

Emit the requested number of photons in bunches of the requested size at the requested time. Depending on the distance between light source and detector fiducial sphere(s) and the solid angle actually subtended, the number of bunches that need to be written might be lower than expected from looking at photons and bsize alone.

Parameters

<i>run</i>	Run class
<i>photons</i>	Number of photons that would be emitted into 4 pi steradian (isotropic).
<i>bsize</i>	Intended bunch size. A conservative approach would use a value of 1.0. Any value below one will be replaced by 1.0.
<i>when</i>	Nominal emission time, with individual photon bunches obtaining a random time value to reproduce the configured pulse shape, on average. Unit: nanoseconds.

Returns

0 (OK)

The documentation for this class was generated from the following files:

- [lactLightEmission.hh](#)
- [lactLightEmission.cc](#)

4.11 SpaceVect Class Reference

A 3-D spatial with some arithmetic operations.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- **SpaceVect** (double xp, double yp, double zp)
- **SpaceVect operator+** (const **SpaceVect** &sv2) const
- **SpaceVect operator-** (const **SpaceVect** &sv2) const
- **SpaceVect operator*** (double f) const
- **SpaceVect & operator+=** (const **SpaceVect** &sv2)
Translation by adding a second vector.
- **SpaceVect & operator-=** (const **SpaceVect** &sv2)
Translation by subtracting a second vector.
- **SpaceVect & operator*=** (double f)
Scaling the vector with a factor.
- bool **IsOrigin** () const
Is this a zero length vector (origin) ?
- double **Length** () const
Length of space vector (= distance to origin)
- double **LengthSqr** () const
Square of length of space vector.
- double **DistanceTo** (const **SpaceVect** &sv2) const
Distance as a length. v1.Distance(v2) is equivalent to (v2-v1).Length() but faster.
- double **DistanceFrom** (const **SpaceVect** &p) const
DistanceTo(), DistanceFrom(), and Distance() are the same thing.
- double **Distance** (const **SpaceVect** &p) const
- **SpaceVect & Rotate_z** (double phi)
Rotate counter-clockwise around given axis (active rotation); modifies the vector.
- **SpaceVect & Rotate_x** (double phi)
- **SpaceVect & Rotate_y** (double phi)
- **SpaceVect & Rotate** (double alpha, double beta, double gamma)
Active rotation with Euler angles in standard x convention (sequence: z,x',z'')
- double **SProd** (const **SpaceVect** &sv2) const
Scalar (inner) product of two vectors.
- double **operator*** (const **SpaceVect** &sv2) const
*v1*v2 denotes a scalar product. Use v1.SProd(v2) if this is confusing.*
- **SpaceVect VProd** (const **SpaceVect** &sv2) const
Vector (outer) product of two vectors.
- **SpaceVect operator/** (const **SpaceVect** &sv2) const
v1/v2 denotes a vector product. Use v1.VProd(v2) if this is confusing.
- double **TProd** (const **SpaceVect** &sv2, const **SpaceVect** &sv3) const
Triple product of three vectors. Returns the volume enclosed.
- ostream & **Print** (ostream &os) const
Readable output, including type and units of variable.

Public Attributes

- double **x**
 - double **y**
 - double **z**
- North, West, Altitude a.s.l. [cm], public for efficienc access.*

4.11.1 Detailed Description

A 3-D spatial with some arithmetic operations.

The [SpaceVect](#) class is just a 3-tuple of x/y/z coordinates. The class also covers translations, scaling, rotation as well as scalar, vector, and triple products.

The documentation for this class was generated from the following file:

- [lactLightEmission.hh](#)

4.12 TelPos Class Reference

Telescope position as in the CORSIKA inputs card TELESCOPE, includes radius of sphere.

```
#include <IactLightEmission.hh>
```

Public Member Functions

- **TelPos** (double rtel)
- **TelPos** (double xt看, double yt看, double zt看, double rtel=0.)
- **TelPos** (const [SpaceVect](#) &xyz, double rtel=0.)
- [SpaceVect Position](#) () const
Get sphere center position.
- double [Radius](#) () const
Get sphere radius.
- double [Area](#) () const
Get sphere cross section area.
- ostream & [Print](#) (ostream &os) const
Show human-readable version of the class.

Public Attributes

- double [x](#)
X coordinate of the fiducial sphere center.
- double [y](#)
Y coordinate of the fiducial sphere center.
- double [z](#)
Z coordinate of the fiducial sphere center.
- double [r](#)
Radius of the fiducial sphere.
- double [A](#)
*Cross section area of sphere ($\pi*r^2$).*

4.12.1 Detailed Description

Telescope position as in the CORSIKA inputs card TELESCOPE, includes radius of sphere.

The documentation for this class was generated from the following file:

- [lactLightEmission.hh](#)

Chapter 5

File Documentation

5.1 fake-muon.cc File Reference

Program to generate photon data similar to single muons in CORSIKA, using the CORSIKA/IACT interface.

```
#include "IactLightEmission.hh"
```

Functions

- void **syntax** (int argc, char **argv, int iarg)
- int **main** (int argc, char **argv)

Main program.

5.1.1 Detailed Description

Program to generate photon data similar to single muons in CORSIKA, using the CORSIKA/IACT interface.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o fake-muon \  
fake-muon.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \  
fileopen.c warning.c straux.c atmo.c sampling.c \  
rndm2.c sim_absorb.c -lm
```

Run as:

```
./fake-muon [ options see --help ]
```

Process the resulting 'fake-muon.iact.gz' through the usual telescope simulation, e.g.

```
/bin/rm -f fake-muon.simtel.gz; zcat fake-muon.iact.gz | \  
./sim_telarray -c cfg/CTA/CTA-ULTRA6-MST-NectarCam.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \  
-C Altitude=2150 -C iobuf_maximum=1000000000 \  
-C maximum_telescopes=1 -C trigger_telescopes=1 \  
-C atmospheric_transmission=atm_trans_2150_1_10_0_0_2150.dat \  
-C telescope_theta=0 -C telescope_phi=0 -C power_law=2.68 \  
-C output_file=fake-muon.simtel.gz -
```

and plot the resulting camera signal trace images with

```
read_cta -p fake-muon.ps --plot-without-reco fake-muon.simtel.gz
```

Author

Konrad Bernloehr

5.1.2 Function Documentation

5.1.2.1 main()

```
int main (
    int argc,
    char ** argv )
```

Main program.

< Index of refraction

5.2 ff-1m.cc File Reference

Flat-fielding light source simulation for single-reflector cameras, using the CORSIKA/IACT interface, simulating light emitted from a flasher supposed to be in the center of the dish. The resulting photon bunches can be simulated by `sim_telarray` with the "Bypass_Optics=2" (or 1 if truly 1M) configuration to bypass reflection on the primary mirror.

```
#include "IactLightEmission.hh"
```

Functions

- void **syntax** (int argc, char **argv, int iarg=0, const string &msg="")
- int **main** (int argc, char **argv)

5.2.1 Detailed Description

Flat-fielding light source simulation for single-reflector cameras, using the CORSIKA/IACT interface, simulating light emitted from a flasher supposed to be in the center of the dish. The resulting photon bunches can be simulated by `sim_telarray` with the "Bypass_Optics=2" (or 1 if truly 1M) configuration to bypass reflection on the primary mirror.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o ff-1m \
    ff-1m.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
    fileopen.c warning.c straux.c atmo.c sampling.c \
    rndm2.c sim_absorb.c -lm
```

Run as:

```
ff-1m [ options ]
```

See "ff-1m --help" for a list of known options and default values.

Process the resulting 'ff-gct.dat.gz' through the usual telescope simulation, partially bypassing ray-tracing in the telescope, e.g.

```
/bin/rm ff-1m.simtel.gz; zcat ff-1m.dat.gz | \
./sim_telarray -c cfg/CTA/CTA-ULTRA6-MST-FlashCam.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
-C Bypass_Optics=1 \
-C Altitude=2100 -C iobuf_maximum=1000000000 -C maximum_telescopes=1 \
-C atmospheric_transmission=atm_trans_2100_1_10_0_0_2100.dat \
-C telescope_theta=0 -C telescope_phi=0 -C power_law=2.5 \
-C output_file=ff-1m.simtel.gz
```

and plot the resulting camera signal trace images with

```
read_cta -p ff-1m.ps ff-1m.simtel.gz
```

or

```
read_cta -u -r 4 --tailcuts 100000,200000 -p ff-1m.ps ff-1m.simtel.gz
```

if you want to avoid the overlay of significant pixels and Hillas ellipse.

Author

Konrad Bernloehr

5.3 ff-gct.cc File Reference

Flat-fielding light source simulation for the GCT (CHEC) cameras, using the CORSIKA/IACT interface, simulating light emitted from four flashers supposed to be in the corners of the camera, here in a coordinate system relative to the center of the secondary mirror. The resulting photon bunches can be simulated by `sim_telarray` with the "Bypass_Optics=1" configuration to bypass reflection on the primary mirror but include reflection on the secondary mirror.

```
#include "IactLightEmission.hh"
```

Functions

- void **syntax** (int argc, char **argv, int iarg=0, const string &msg="")
- int **main** (int argc, char **argv)

5.3.1 Detailed Description

Flat-fielding light source simulation for the GCT (CHEC) cameras, using the CORSIKA/IACT interface, simulating light emitted from four flashers supposed to be in the corners of the camera, here in a coordinate system relative to the center of the secondary mirror. The resulting photon bunches can be simulated by `sim_telarray` with the "Bypass_Optics=1" configuration to bypass reflection on the primary mirror but include reflection on the secondary mirror.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o ff-gct \
  ff-gct.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
  fileopen.c warning.c straux.c atmo.c sampling.c \
  rndm2.c sim_absorb.c -lm
```

Run as:

```
ff-gct [ options ]
```

See "ff-gct --help" for a list of known options and default values.

Process the resulting 'ff-gct.dat.gz' through the usual telescope simulation, partially bypassing ray-tracing in the telescope, e.g.

```
/bin/rm ff-gct.simtel.gz; zcat ff-gct.dat.gz | \
  ./sim_telarray -c cfg/CTA/CTA-ULTRA6-SST-GCT-S.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
  -C Bypass_Optics=1 \
  -C Altitude=2100 -C iobuf_maximum=1000000000 -C maximum_telescopes=1 \
  -C atmospheric_transmission=atm_trans_2100_1_10_0_2100.dat \
  -C telescope_theta=0 -C telescope_phi=0 -C power_law=2.5 \
  -C FADC_BINS=128 -C DISC_BINS=128 \
  -C output_file=ff-gct.simtel.gz
```

and plot the resulting camera signal trace images with

```
read_cta -p ff-gct.ps ff-gct.simtel.gz
```

or

```
read_cta -u -r 4 --tailcuts 100000,200000 -p ff-gct.ps ff-gct.simtel.gz
```

if you want to avoid the overlay of significant pixels and Hillas ellipse.

Author

Konrad Bernloehr

5.4 fpls.cc File Reference

Focal plane lightsource example program for artificial light sources using the CORSIKA/IACT interface.

```
#include "IactLightEmission.hh"
```

Functions

- `int main ()`

5.4.1 Detailed Description

Focal plane lightsource example program for artificial light sources using the CORSIKA/IACT interface.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o fpls \
  fpls.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
  fileopen.c warning.c straux.c atmo.c sampling.c \
  rndm2.c sim_absorb.c -lm
```

Run as:

```
./fpls
```

Process the resulting 'fpls.dat.gz' through the usual telescope simulation, bypassing ray-tracing in the telescope, e.g.

```
/bin/rm fpls.simtel.gz; zcat fpls.dat.gz | \
  ./sim_telarray -c cfg/CTA/CTA-ULTRA6-MST-NectarCam.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
  -C Bypass_Optics=2 \
  -C Altitude=2100 -C iobuf_maximum=1000000000 -C maximum_telescopes=1 \
  -C atmospheric_transmission=atm_trans_2100_1_10_0_0_2100.dat \
  -C telescope_theta=0 -C telescope_phi=0 -C power_law=2.68 \
  -C FADC_BINS=128 -C DISC_BINS=128 \
  -C output_file=fpls.simtel.gz
```

and plot the resulting camera signal trace images with

```
read_cta -p fpls.ps fpls.simtel.gz
```

The enlarged FADC_BINS and DISC_BINS is only necessary where the rotation period is basically as large as the readout period. For single-reflector telescopes it makes no difference of Bypass_Optics=2 or Bypass_Optics=1.

Author

Konrad Bernloehr

5.5 IactLightEmission.cc File Reference

Simulate an artificial light source instead of running CORSIKA and write the resulting photon bunches in the same format as with shower simulation going through the IACT/atmo interface.

```
#include "IactLightEmission.hh"
#include "rpolator.h"
#include <time.h>
#include <sys/types.h>
#include <sys/time.h>
#include <unistd.h>
#include <ctype.h>
#include <stdexcept>
#include "fileopen.h"
#include "rndm2.h"
```

Classes

- struct [photon_bunch](#)

Macros

- #define **UNUSED(x)** UNUSED_ ## x

Typedefs

- typedef struct [photon_bunch](#) **Bunch**

Functions

- void **rmmard_** (double *a, int *b, int *UNUSED(c))
- double **rhof_** (double *height)
- double **thick_** (double *height)
- double **heigh_** (double *thick)
- void **random_initialize** ()
- void **strip_comments** (char *line)
- int **read_table_v** (const char *fname, vector< double > &col1, vector< double > &col2)
- int **read_table_v** (const string &fname, vector< double > &col1, vector< double > &col2)
- double **rpol_v** (const vector< double > &x, const vector< double > &y, double xp)
- int **make_random_ipol_table_v** (const vector< double > &x, const vector< double > &y, vector< double > &xr, vector< double > &yr, double xlow_user=0., double xhigh_user=0.)
Make fast lookup table for rather any number distribution.
- int **ls_verbose** ()
- int **ls_set_verbose** (int v)
- double **random_from_ipol_table_v** (const vector< double > &xr, const vector< double > &yr, double randnr)
Fast random number table lookup.
- void **SaveCommandLine** (int argc, char **argv)
- vector< string > **GetCommandLine** ()
- string **GetCommandLineString** ()
- [DirectionVect](#) **IsotropicRandomEmission** ()
- [DirectionVect](#) **IsotropicRandomEmissionUp** ()
- [DirectionVect](#) **IsotropicRandomEmissionDown** ()
- [DirectionVect](#) **IsotropicRandomEmission** (double maxang, const [DirectionVect](#) ¢ral_dir)
- [DirectionVect](#) **RandomEmission** (double angle, const [DirectionVect](#) ¢ral_dir, double phi_from=0., double phi_to=2.*M_PI)
- [DirectionVect](#) **InvertForRandomEmission** (const [DirectionVect](#) ¢ral_dir, const [DirectionVect](#) &emit_dir)
- double **refidx** (double alt)
- double **strtof** (const string &str)
STL string equivalent to C library atof.
- string **strip_whitespace** (const string &text)
- vector< double > **dbl_explode** (const string &text, const string &delim)
Given a text string and an optional list of separators, explode it into a vector of doubles.

5.5.1 Detailed Description

Simulate an artificial light source instead of running CORSIKA and write the resulting photon bunches in the same format as with shower simulation going through the IACT/atmo interface.

The resulting data file will be in the same format as CORSIKA IACT output and can thus be processed with `sim_telarray` or any other program handling the same type of data.

Author

Konrad Bernloehr

5.5.2 Function Documentation

5.5.2.1 `dbl_explode()`

```
vector<double> dbl_explode (
    const string & text,
    const string & delim )
```

Given a text string and an optional list of separators, explode it into a vector of doubles.

Simplified replacement for `stdtools` function.

5.5.2.2 `ls_set_verbose()`

```
int ls_set_verbose (
    int v )
```

Set verbose flag from code rather than environment variable.

5.5.2.3 `ls_verbose()`

```
int ls_verbose ( )
```

Verbose output from [LightSource](#) code or not.

5.5.2.4 `make_random_ipol_table_v()`

```
int make_random_ipol_table_v (
    const vector< double > & x,
    const vector< double > & y,
    vector< double > & xr,
    vector< double > & yr,
    double xlow_user = 0.,
    double xhigh_user = 0. )
```

Make fast lookup table for rather any number distribution.

The following procedures can be used for relatively fast production of random numbers according to a given table of (not to be normalized) probability values. Tails of a distribution should be sufficiently well reproduced within the limits imposed a) by the chosen generator of flat random numbers (really double, i.e. 48-bit abscissa, or effectively float, i.e. 24-bit abscissa) and b) by the assumption that the probability is constant within each bin of the input table.

The bin limits are calculated with the assumption of equidistant bins where the abscissa values of the input table give the middle of each bin. This must hold at least for the two lowest and the two highest bins.

Parameters

<i>x</i>	X coordinates of tabulated 'curve'
<i>y</i>	Y coordinates of tabulated 'curve'
<i>xr</i>	Vector of returned X values
<i>yr</i>	Vector of returned Y values (normalized to [0:1])

Returns

0 (o.k.), -1 (error)

5.5.2.5 `random_from_ipol_table_v()`

```
double random_from_ipol_table_v (
    const vector< double > & xr,
    const vector< double > & yr,
    double randnr )
```

Fast random number table lookup.

Get a random number from a lookup table as created by `make_random_ipol_table()` for a tabulated probability distribution.

Parameters

<i>xr</i>	Table as returned from <code>make_random_ipol_table()</code>
<i>yr</i>	Table as returned from <code>make_random_ipol_table()</code>
<i>randnr</i>	A uniform pseudorandom number in [0:1]

Returns

random number according to table

5.5.2.6 `random_initialize()`

```
void random_initialize ( )
```

Initialize our random number generator using system-provided sources of randomness.

5.5.2.7 `rpol_v()`

```
double rpol_v (
    const vector< double > & x,
    const vector< double > & y,
    double xp )
```

Interface from a pair of C++ vectors to the old C-based interpolation.

5.6 lactLightEmission.hh File Reference

Class definitions and common inline code for the lactLightEmission package.

```
#include "initial.h"
#include "io_basic.h"
#include "mc_tel.h"
#include "iact.h"
#include "atmo.h"
#include <vector>
#include <string>
#include "rndm2.h"
#include <iostream>
#include "fileopen.h"
```

Classes

- class [SpaceVect](#)
A 3-D spatial with some arithmetic operations.
- class [DirectionVect](#)
A normalized direction vector.
- class [Line](#)
A line is defined by a point on the line and the direction vector along the line.
- class [Plane](#)
A plane in 3-D space and intersections.
- class [TelPos](#)
Telescope position as in the CORSIKA inputs card TELESCOPE, includes radius of sphere.
- class [Run](#)
Run class takes care of global initialization and winding-down/cleanup as well as advancing from one event to the next.
- class [SimpleLightSource](#)
Simple light source: monochromatic, isotropic, pulse shape is Gaussian or top hat. In addition: point source, stationary.
- class [BackgroundLightSource](#)
A light source class with diffuse emission characteristics.
- class [LightSource](#)
A much more generic type of light source than the [SimpleLightSource](#).

Functions

- ostream & [operator<<](#) (ostream &os, const [SpaceVect](#) &t)
Non-member output operator for [SpaceVect](#).
- double [SProd](#) (const [SpaceVect](#) &sv1, const [SpaceVect](#) &sv2)
Non-member scalar product of two [SpaceVect](#).
- [SpaceVect](#) [VProd](#) (const [SpaceVect](#) &sv1, const [SpaceVect](#) &sv2)
Non-member vector product of two [SpaceVect](#).
- double [TProd](#) (const [SpaceVect](#) &sv1, const [SpaceVect](#) &sv2, const [SpaceVect](#) &sv3)
Non-member triple product of three [SpaceVect](#).
- [SpaceVect](#) [operator*](#) (double d, const [SpaceVect](#) &sv)
Non-member scaling of a [SpaceVect](#).

- ostream & **operator<<** (ostream &os, const [DirectionVect](#) &t)
Non-member print function.
- [SpaceVect](#) **operator*** (double d, const [DirectionVect](#) &dv)
Multiplication the other way round is mapped to member function.
- double **SProd** (const [DirectionVect](#) &dv1, const [DirectionVect](#) &dv2)
Non-member functions for scalar product of two direction vectors, basically the cosine of the angle between them.
- double **SProd** (const [SpaceVect](#) &sv1, const [DirectionVect](#) &dv2)
The scalar product between a [SpaceVect](#) and a [DirectionVect](#) includes the [SpaceVect](#), in addition to the angle cosine.
- double **SProd** (const [DirectionVect](#) &dv1, const [SpaceVect](#) &sv2)
- [SpaceVect](#) **VProd** (const [DirectionVect](#) &dv1, const [DirectionVect](#) &dv2)
- [SpaceVect](#) **VProd** (const [SpaceVect](#) &sv1, const [DirectionVect](#) &dv2)
Non-member function for the vector product between a [SpaceVect](#) and a [DirectionVect](#).
- [SpaceVect](#) **VProd** (const [DirectionVect](#) &dv1, const [SpaceVect](#) &sv2)
- ostream & **operator<<** (ostream &os, const [Line](#) &l)
Non-member print function.
- ostream & **operator<<** (ostream &os, const [Plane](#) &pl)
Non-member print function.
- double **Distance** (const [SpaceVect](#) &sv1, const [SpaceVect](#) &sv2)
Distance between to points (non-member function).
- double **Distance** (const [SpaceVect](#) &sv, const [Line](#) &l)
Shortest distance between point and line (non-member function).
- double **Distance** (const [Line](#) &l, const [SpaceVect](#) &sv)
Shortest distance between point and line, alternate order (non-member function).
- double **Distance** (const [SpaceVect](#) &sv, const [Plane](#) &a)
Shortest distance between point and plane (non-member function).
- double **Distance** (const [Plane](#) &a, const [SpaceVect](#) &sv)
Shortest distance between point and plane, alternate order (non-member function).
- double **Distance** (const [Line](#) &l1, const [Line](#) &l2)
Shortest distance between two lines (non-member function).
- ostream & **operator<<** (ostream &os, const [TelPos](#) &t)
Non-member print function.
- int **ls_verbose** ()
- int **ls_set_verbose** (int v)
- void **SaveCommandLine** (int argc, char **argv)
- vector< string > **GetCommandLine** ()
- string **GetCommandLineString** ()
- [DirectionVect](#) **IsotropicRandomEmission** ()
- [DirectionVect](#) **IsotropicRandomEmissionUp** ()
- [DirectionVect](#) **IsotropicRandomEmissionDown** ()
- [DirectionVect](#) **IsotropicRandomEmission** (double maxang, const [DirectionVect](#) ¢ral_dir)
- [DirectionVect](#) **RandomEmission** (double angle, const [DirectionVect](#) ¢ral_dir)
- vector< double > **dbl_explode** (const string &text, const string &delim=",")
Simplified replacement for stdtools function.
- ostream & **operator<<** (ostream &os, const vector< double > &t)

5.6.1 Detailed Description

Class definitions and common inline code for the lactLightEmission package.

Author

Konrad Bernloehr

5.6.2 Function Documentation

5.6.2.1 `dbl_explode()`

```
vector<double> dbl_explode (
    const string & text,
    const string & delim )
```

Simplified replacement for `stdtools` function.

Simplified replacement for `stdtools` function.

5.6.2.2 `ls_set_verbose()`

```
int ls_set_verbose (
    int v )
```

Set verbose flag from code rather than environment variable.

5.6.2.3 `ls_verbose()`

```
int ls_verbose ( )
```

Verbose output from [LightSource](#) code or not.

5.6.2.4 `VProd()`

```
SpaceVect VProd (
    const DirectionVect & dv1,
    const DirectionVect & dv2 ) [inline]
```

Non-member function for the vector product between two [DirectionVect](#). Since the result is not normalized, it is returned as a [SpaceVect](#).

5.7 `ile_version.hh` File Reference

This is the `lactLightEmission` package version 2023-01-31.

Macros

- `#define ILE_VERSION "2023-01-31"`

5.7.1 Detailed Description

This is the lactLightEmission package version 2023-01-31.

Version

2023-01-31

5.8 ls-beam.cc File Reference

Simulate light from a laser beam crossing the field of view of a telescope, some of the laser light getting scattered along the beam. The modelling of the laser beam and its attenuation along the path is very simplified, and for performance reasons an isotropic emission angle distribution is to be preferred for now. The resulting photon bunches can be simulated by `sim_telarray` with the (default) "Bypass_Optics=0" configuration, not bypassing any optics raytracing.

```
#include "IactLightEmission.hh"
#include "fileopen.h"
```

Functions

- void **syntax** (int argc, char **argv, int iarg=0, const string &msg="")
- int [solve_quadratic_equation](#) (double a, double b, double c, double *x1, double *x2)
- int **main** (int argc, char **argv)

5.8.1 Detailed Description

Simulate light from a laser beam crossing the field of view of a telescope, some of the laser light getting scattered along the beam. The modelling of the laser beam and its attenuation along the path is very simplified, and for performance reasons an isotropic emission angle distribution is to be preferred for now. The resulting photon bunches can be simulated by `sim_telarray` with the (default) "Bypass_Optics=0" configuration, not bypassing any optics raytracing.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o ls-beam \
  ls-beam.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
  fileopen.c warning.c straux.c atmo.c sampling.c \
  rndm2.c sim_absorb.c -lm
```

Run as:

```
ls-beam [ options ]
```

See "ls-beam --help" for a list of known options and default values.

Example use:

```
ls-beam --laser-position 1000,0,0 --laser-theta 60 --laser-phi 180 \
  --telescope-theta 30 --telescope-phi 0 --fov 15
```

Process the resulting 'ls-beam.dat.gz' through the usual telescope simulation, partially bypassing ray-tracing in the telescope, e.g.

```
/bin/rm ls-beam.simtel.gz; zcat ls-beam.iact.gz | \
  ./sim_telarray -c cfg/CTA/CTA-ULTRA6-MST-FlashCam.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
  -C Altitude=2150 -C iobuf_maximum=1000000000 -C maximum_telescopes=1 \
  -C atmospheric_transmission=atm_trans_2150_1_10_0_0_2150.dat \
  -C telescope_theta=30 -C telescope_phi=0 -C power_law=2.5 \
  -C output_file=ls-beam.simtel.gz
```

and plot the resulting camera signal trace images with

```
read_cta -p ls-beam.ps ls-beam.simtel.gz
```

or

```
read_cta -u -r 4 --tailcuts 100000,200000 -p ls-beam.ps ls-beam.simtel.gz
```

if you want to avoid the overlay of significant pixels and Hillas ellipse.

Author

Konrad Bernloehr

5.8.2 Function Documentation**5.8.2.1 solve_quadratic_equation()**

```
int solve_quadratic_equation (
    double a,
    double b,
    double c,
    double * x1,
    double * x2 )
```

Solve a quadratic equation $a*x**2 + b*x + c = 0$

Returns

0: OK, we have solutions, -1: no solution

5.9 Is-geo-test.cc File Reference

A few tests for the LightEmission geometry classes.

```
#include "IactLightEmission.hh"
```

Functions

- int **main** ()

5.9.1 Detailed Description

A few tests for the LightEmission geometry classes.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o ls-geo-test \
ls-geo-test.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
fileopen.c warning.c straux.c atmo.c sampling.c \
rndm2.c sim_absorb.c -lm
```

Run as:

```
./ls-geo-test
```

Author

Konrad Bernloehr

5.10 Is-test.cc File Reference

Test program for artificial light sources using the CORSIKA/IACT interface.

```
#include "IactLightEmission.hh"
```

Functions

- `int main ()`

5.10.1 Detailed Description

Test program for artificial light sources using the CORSIKA/IACT interface.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o ls-test \  
ls-test.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \  
fileopen.c warning.c straux.c atmo.c sampling.c \  
rndm2.c sim_absorb.c -lm
```

Run as:

```
./ls-test
```

Author

Konrad Bernloehr

5.11 Is-test2.cc File Reference

Test program for artificial light sources using the CORSIKA/IACT interface.

```
#include "IactLightEmission.hh"
```

Functions

- `int main ()`

5.11.1 Detailed Description

Test program for artificial light sources using the CORSIKA/IACT interface.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o ls-test2 \  
ls-test2.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \  
fileopen.c warning.c straux.c atmo.c sampling.c \  
rndm2.c sim_absorb.c -lm
```

Run as:

```
./ls-test2
```

Author

Konrad Bernloehr

5.12 nsbls.cc File Reference

Nightsky background lightsource example program for artificial light sources using the CORSIKA/IACT interface.

```
#include "IactLightEmission.hh"
```

Functions

- void **syntax** (int argc, char **argv, int iarg)
- int **main** (int argc, char **argv)

5.12.1 Detailed Description

Nightsky background lightsource example program for artificial light sources using the CORSIKA/IACT interface.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o nsbls \
  nsbls.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
  fileopen.c warning.c straux.c atmo.c sampling.c \
  rndm2.c sim_absorb.c -lm
```

Run as:

```
./nsbls
Process the resulting 'nsbls.dat.gz' through the usual telescope
simulation, bypassing ray-tracing in the telescope, e.g.
@code{.sh}
/bin/rm nsbls.simtel.gz; zcat nsbls.iact.gz | \
./sim_telarray -c cfg/CTA/CTA-ULTRA6-MST-NectarCam.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
-C Altitude=2150 -C iobuf_maximum=1000000000 -C maximum_telescopes=1 \
-C atmospheric_transmission=atm_trans_2150_1_10_0_0_2150.dat \
-C telescope_theta=0 -C telescope_phi=0 \
-C output_file=nsbls.simtel.gz
```

and plot the resulting map of registered true photo-electrons with

```
read_cta -p nsbls.ps \
  --plot-with-true-pe \
  --plot-with-pixel-pe \
  --plot-with-sum-only \
  --plot-without-reco \
  nsbls.simtel.gz
```

Author

Konrad Bernloehr

5.13 octo.cc File Reference

Superluminal octocopter example program for artificial light sources using the CORSIKA/IACT interface.

```
#include "IactLightEmission.hh"
```

Functions

- int **main** ()

5.13.1 Detailed Description

Superluminal octocopter example program for artificial light sources using the CORSIKA/IACT interface.

This program is for demonstration purposes only and has no practical use.

If not build by default, use the Makefile dependencies with

```
make octo
```

or compile it manually (will actually need paths to source files and for header files included, here assuming we are under the `sim_telarray` directory):

```
g++ -I. \
  -I../corsika -I../common -I../hessioxxx/include -I../stdtools \
  -DIACT_NO_GRID -DCORSIKA_VERSION=6999 -DWITH_STDTOOLS \
  -DWITH_RPOLATOR_SEPARATE -DWITH_MC_ATMPROF_SEPARATE \
  -Wall -Wextra -o octo \
  octo.cc IactLightEmission.cc \
  ../corsika/iact.c ../corsika/atmo.c ../corsika/sampling.c \
  ../../hessioxxx/src/io_simtels.c \
  ../../hessioxxx/src/eventio.c \
  ../../hessioxxx/src/fileopen.c \
  ../../hessioxxx/src/warning.c \
  ../../hessioxxx/src/straux.c \
  ../../hessioxxx/src/mc_atmprof.c \
  ../common/rndm2.c ../common/rpolator.c ../common/sim_absorb.c \
  -lm
```

Run as:

```
./octo
```

Process the resulting 'octo.iact.gz' through the usual telescope simulation, e.g.

```
../sim_telarray -c cfg/CTA/CTA-PROD6-SST.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
-C Altitude=2150 -C iobuf_maximum=1000000000 -C maximum_telescopes=1 \
-C atmospheric_transmission=atm_trans_2150_1_10_0_0_2150.dat \
-C telescope_theta=0 -C telescope_phi=0 -C power_law=2.68 \
-C output_file=octo.simtel.zst octo.iact.gz
```

and plot the resulting camera signal trace images.

Author

Konrad Bernloehr

5.14 pixled.cc File Reference

A program to set an assumed LED in front of a specified subset of pixels in a camera as defined for `sim_telarray`. The main purpose is to test that trigger definitions work as intended.

```
#include "IactLightEmission.hh"
#include <straux.h>
#include <list>
```

Classes

- struct [PixPos](#)
- struct [PixLed](#)

Functions

- string **clean_trigger_list** (const string &plist, bool keep_slaves)
- int **read_cam** (const string &fname, vector< PixPos > &cam, bool keep_slaves, bool read_triggers)
- int **get_simple_use** (const char *use, vector< PixLed > &pl)
- int **get_complete_use** (const char *use, vector< PixLed > &pl)
- int **get_trg_use_random** (vector< PixLed > &pl)
- int **get_trg_use** (size_t j, vector< PixLed > &pl)
- void **syntax** (int argc, char **argv, int iarg=0, const string &msg="")

Report syntax errors and program usage.
- void **emit_all_events** (vector< LightSource > &vls, size_t nreq, const vector< size_t > &selected, const vector< size_t > &available, Run &run, double nphot_per_pos, double bunch_size, double dt, vector< LightSource > &vls, const vector< size_t > &nsl, const vector< size_t > fsl)

Recursive function to produce as many events as possible pixel combinations.
- int **main** (int argc, char **argv)

The main program.

Variables

- vector< string > **trglist**

5.14.1 Detailed Description

A program to set an assumed LED in front of a specified subset of pixels in a camera as defined for sim_telarray. The main purpose is to test that trigger definitions work as intended.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o pixled \
  pixled.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
  fileopen.c warning.c straux.c atmo.c sampling.c \
  rndm2.c sim_absorb.c -lm
```

Run as:

```
./pixled [ --camera-file <name> ] [ --use <pixel-list > ]
  [ --require <npix> ] [ --random <npix> ] [ --events <nev> ] \
  ... [ -o <output_file> ]
```

Note that the name of the camera definition file and the pixel list are necessary. If the number of required or random pixels is missing, all specified pixels will get fired. Process the resulting output file through the usual telescope simulation, bypassing ray-tracing in the telescope, e.g.

```
/bin/rm pixled.simtel.gz; zcat pixled.dat.gz | \
  ./sim_telarray -c cfg/CTA/CTA-PROD4-LST.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
  -C Bypass_Optics=2 \
  -C Altitude=2150 -C iobuf_maximum=1000000000 -C maximum_telescopes=1 \
  -C atmospheric_transmission=atm_trans_2150_1_10_0_0_2150.dat \
  -C telescope_theta=0 -C telescope_phi=0 -C power_law=2.68 \
  -C output_file=pixled.simtel.gz
```

and plot the resulting camera signal trace images with

```
read_cta -p pixled.ps pixled.simtel.gz
```

Author

Konrad Bernloehr

5.14.2 Function Documentation

5.14.2.1 get_complete_use()

```
int get_complete_use (
    const char * use,
    vector< PixLed > & pl )
```

Complete list of pixels to illuminate and how to illuminate can only be obtained from a separate file (one line per pixel).

5.14.2.2 get_simple_use()

```
int get_simple_use (
    const char * use,
    vector< PixLed > & pl )
```

Get a simple list of pixels to use, that is illuminate by light. The simple list is provided on the command line and its only pixel-specific option is to say that a pixel is always required.

5.14.2.3 read_cam()

```
int read_cam (
    const string & fname,
    vector< PixPos > & cam,
    bool keep_slaves,
    bool read_triggers )
```

Read a camera definition file as it gets used by sim_telarray, extracting the x/y projected positions of each pixel center.

5.15 xyIs.cc File Reference

Test program for artificial light sources using the CORSIKA/IACt interface. This program is not intended to be of practical use.

```
#include "IactLightEmission.hh"
```

Functions

- int main (int argc, char **argv)

5.15.1 Detailed Description

Test program for artificial light sources using the CORSIKA/IACt interface. This program is not intended to be of practical use.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACt_NO_GRID -Wall -Wextra -o xyIs \
    xyIs.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
    fileopen.c warning.c straux.c atmo.c sampling.c \
    rndm2.c sim_absorb.c -lm
```

Run as:

```
./xyIs [ -r rcam ] [ -x xpos ] [ -y ypos ] [ -h height ] [ -o fname ]
```

Author

Konrad Bernloehr

5.16 xyzls.cc File Reference

Lightsource example program for artificial light sources using the CORSIKA/IACT interface.

```
#include "IactLightEmission.hh"
#include "straux.h"
```

Functions

- int **read_telpos_file** (const char *fname, bool corsika_fmt, vector< [TelPos](#) > &tpv)
- int **main** (int argc, char **argv)

5.16.1 Detailed Description

Lightsource example program for artificial light sources using the CORSIKA/IACT interface.

Compile with (will actually need paths to source files):

```
g++ -DCTA_PROD3 -DIACT_NO_GRID -Wall -Wextra -o xyzls \
xyzls.cc IactLightEmission.cc iact.c io_simtel.c eventio.c \
fileopen.c warning.c straux.c atmo.c sampling.c \
rndm2.c sim_absorb.c -lm
```

Run as:

```
./xyzls
```

Process the resulting 'xyzls.dat.gz' through the usual telescope simulation, bypassing ray-tracing in the telescope, e.g.

```
/bin/rm xyzls.simtel.gz; zcat xyzls.dat.gz | \
./sim_telarray -c cfg/CTA/CTA-ULTRA6-MST-NectarCam.cfg -DNUM_TELESCOPES=1 -Icfg/CTA \
[ -C Bypass_Optics=2 ] \
-C Altitude=2150 -C iobuf_maximum=1000000000 \
-C maximum_telescopes=1 -C trigger_telescopes=1 \
-C atmospheric_transmission=atm_trans_2150_1_10_0_0_2150.dat \
-C telescope_theta=0 -C telescope_phi=0 -C power_law=2.68 \
-C output_file=xyzls.simtel.gz
```

and plot the resulting camera signal trace images with

```
read_cta -p xyzls.ps xyzls.simtel.gz
```

For single-reflector telescopes it makes no difference if Bypass_Optics=2 or Bypass_Optics=1.

Author

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