# Background characterization for the GERDA experiment



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### **Outline**

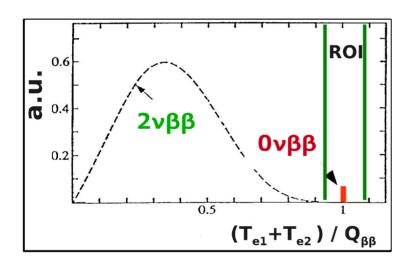
- Physics motivation and goals of the GERDA experiment
- Experimental setup and status of Phase-I data taking
- Analysis of Phase-I data:
  - $\rightarrow$  measurement of the T<sub>1/2</sub> of 2νββ decay of <sup>76</sup>Ge
  - background due to alpha-induced events
  - decomposition of the background spectrum
- Conclusions

### **Motivation**

GERDA experiment is searching for the neutrinoless double beta ( $0\nu\beta\beta$ ) decay of  $^{76}$ Ge.

### Neutrino accompanied double beta decay $(2\nu\beta\beta)$

$$(A, Z) \rightarrow (A, Z+2) + 2e^{-} + 2\overline{\nu}_{e}$$



- → SM process, observed
- $\rightarrow$  T<sub>1/2</sub>  $\sim$  (10<sup>19</sup> 10<sup>24</sup>) yr Rarest decay measured in lab
- → Experimental signature: continuous spectrum of the sum of electrons kinetic energies

### Neutrinoless double beta decay $(0v\beta\beta)$

$$(A, Z) \rightarrow (A, Z+2) + 2e^{-}$$

- → Non SM process:  $\Delta L = 2$
- $\rightarrow$  More rare than 2νββ
- $\rightarrow$  Experimental signature: peak at  $Q_{\beta\beta}$

### Goals

GERDA experiment is searching for the neutrinoless double beta ( $0\nu\beta\beta$ ) decay of  $^{76}$ Ge.

limit:  $T_{1/2}^{0\nu}(^{76}\text{Ge}) > 1.9 \times 10^{25} \text{ y (90\% C.L.)}$  from HdM Collaboration [Eur. Phys. J. A 12, 147154 (2001)]

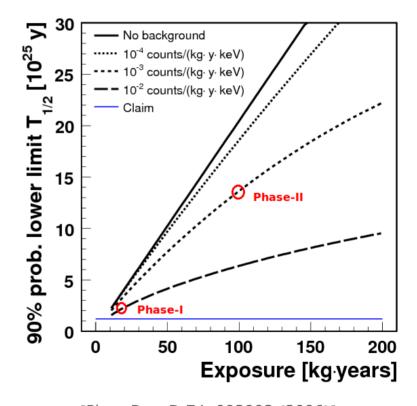
claim:  $T_{1/2}^{0\nu}(^{76}\text{Ge}) = 1.2 \times 10^{25} \text{ y}$  [Phys. Lett. B 586 (2004) 198-212]

**Phase-I:**  $T_{1/2} > 2 \times 10^{25} \text{ y} \rightarrow \text{test the claim}$ 

**Phase-II:**  $T_{1/2} > 10^{26} \text{ y}$ 

For a higher sensitivity on the  $T_{1/2}$ 

- → larger exposure
- $\rightarrow$  lower background rate around  $Q_{BB}$ 
  - ▶ Background characterization & suppression



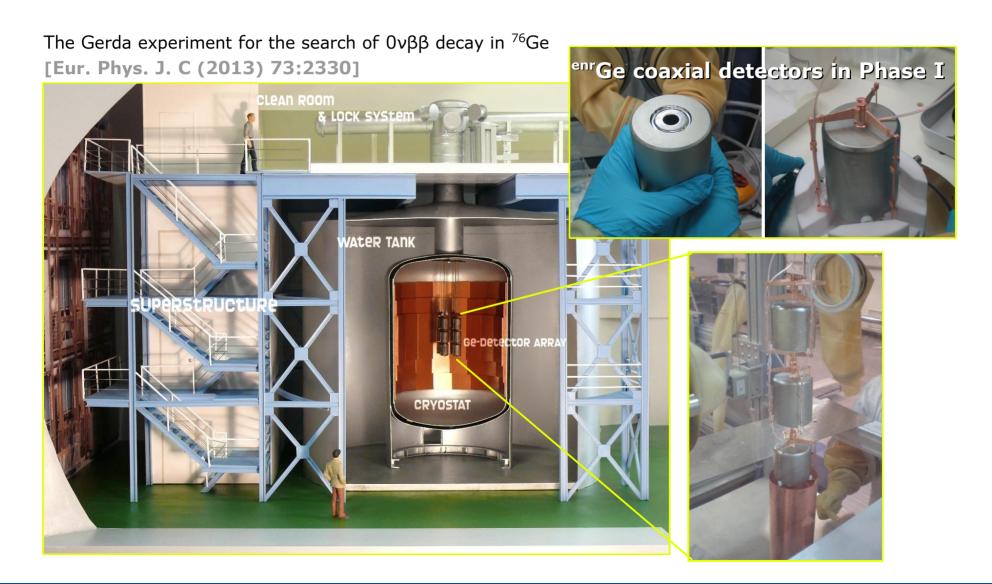
[Phys. Rev. D 74, 092003 (2006)]

### **Experimental setup**

### **GERDA** @ LNGS of INFN, Italy

- HPGe detectors directly submerged in LAr
- ♦ High-purity shields: LAr, H₂O

- ◆ Active muon veto: detection of Cherenkov radiation in water
- Minimal amount of screened material around the detectors



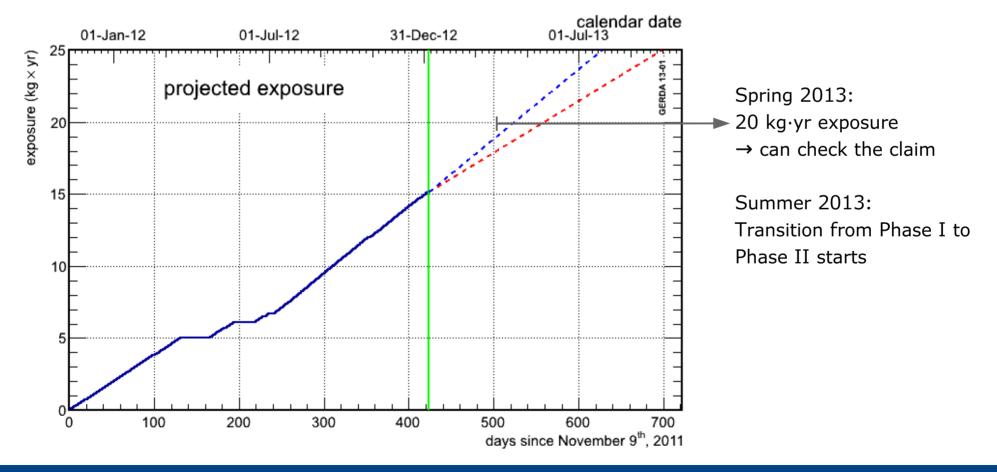
Phase-I started on 9 November 2011 with 3 natGe and 8 enrGe (from HdM and IGEX) coax detectors

2 <sup>nat</sup>Ge detectors removed, 5 <sup>enr</sup>Ge BEGe detectors deployed on July 2012

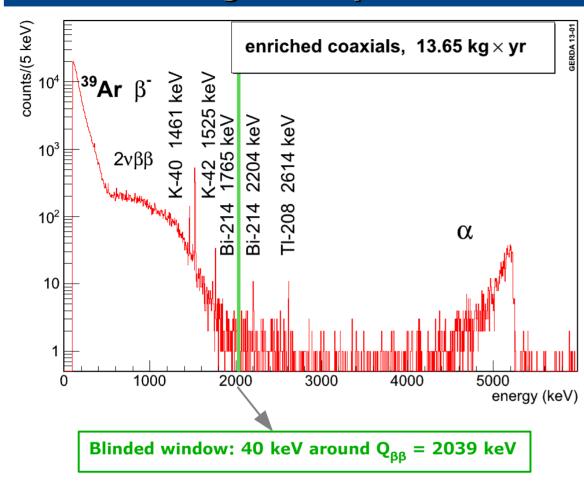
enrGe mass for physics analysis: 14.6 kg (coaxial) and 3.6 kg (BEGe)

**Energy resolution:** FWHM at  $Q_{\beta\beta}=4.5$  keV (mass weighted average for the <sup>enr</sup>Ge coax detectors)

**Stability:** energy resolution within  $\sim 0.5$  keV, energy scale within  $\sim 1$  keV.



### Phase-I background spectrum



9 November 20 11 - 5 January 2013

Live time: 340.94 days

Detectors: 6 enrGe coax (14.63 kg)

Total exposure: 13.65 kg  $\times$  yr

#### **Background components:**

No contribution at  $Q_{\beta\beta}$ 

 $^{39}$ Ar ( $Q_{\beta} = 565 \text{ keV}$ ),  $2\nu\beta\beta$ ,  $^{40}$ K,  $^{228}$ Ac

Contribution at  $Q_{\beta\beta}$ 

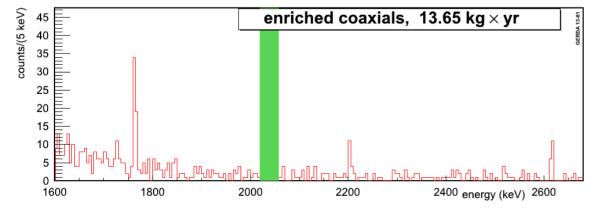
• 
$$^{42}$$
K ( $^{42}$ Ar)  $\rightarrow$  Q <sub>$\beta$</sub>  = 3.5 MeV, E <sub>$\gamma$</sub>  = 2.4 MeV

• 
$$^{214}$$
Bi ( $^{238}$ U) →  $Q_{\beta} = 3.3$  MeV,  $E_{\gamma} = 2.1, 2.2, 2.4$  MeV

• 208TI (232Th) 
$$\rightarrow$$
 E<sub>Y</sub> = 2.6 MeV

• **60Co** 
$$\to$$
 Q<sub>B</sub> = 2.8 MeV

• α-induced events (from isotopes in <sup>238</sup>U chain)



### Background index in $Q_{\beta\beta} \pm 100 \text{ keV}$

0.022 cts/(keV kg yr) for 13.6 kg·yr data excluding 1.30 kg·yr (period following detector substitutions in July):

0.017 cts/(keV kg yr) for 12.3 kg·yr

10x lower than previous experiments (HdM, IGEX)

# Phase I: Measurement of the $T_{1/2}$ of $2\nu\beta\beta$ of $^{76}$ Ge

enrGe coax detectors, total exposure: 5 kg·yr

2vββ spectrum clearly visible with the first 126 days data

#### Binned maximum likelihood fit

Fit window: 600 - 1800 keV

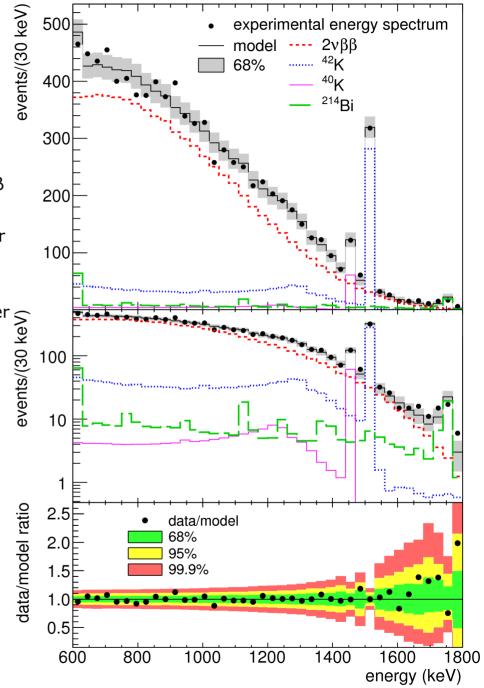
- $\rightarrow$  above <sup>39</sup>Ar end-point energy (Q<sub>B</sub> = 565 keV)
- $\rightarrow$  in E>1800 keV range 0.02% probability of 2vββ

**Background components:** <sup>40</sup>K and <sup>214</sup>Bi close source, <sup>42</sup>K in LAr

Fit Parameters: Active detector masses, enrichment fractions, background contributions,  $T_{1/2}$  common parameter background contributions,  $T_{1/2}$  backgro

$$T_{1/2}^{2\nu} = \left(1.84_{-0.08 \text{ fit } -0.06 \text{ syst}}^{+0.11}\right) \times 10^{21} \text{yr} = \left(1.84_{-0.10}^{+0.14}\right) \times 10^{21} \text{yr}$$

Item	Uncer	tainty on $T_{1/2}^{2\nu}$ (%)
Non-identified background components Energy spectra from $^{42}$ K, $^{40}$ K and $^{214}$ Bi Shape of the $2\nu\beta\beta$ decay spectrum	+5.3 ±2.1 ±1	
Subtotal fit model		+5.8 -2.3
Precision of the Monte Carlo geometry model Accuracy of the Monte Carlo tracking	±1 ±2	
Subtotal Monte Carlo		±2.2
Data acquisition and selection		±0.5
Grand total		+6.2 -3.3



# Phase I: Measurement of the $T_{1/2}$ of $2\nu\beta\beta$ of $^{76}Ge^{-1}$

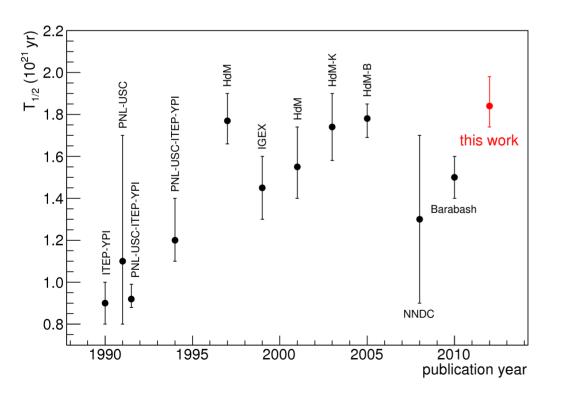
The GERDA collaboration [J. Phys. G 40 (2013) 035110]

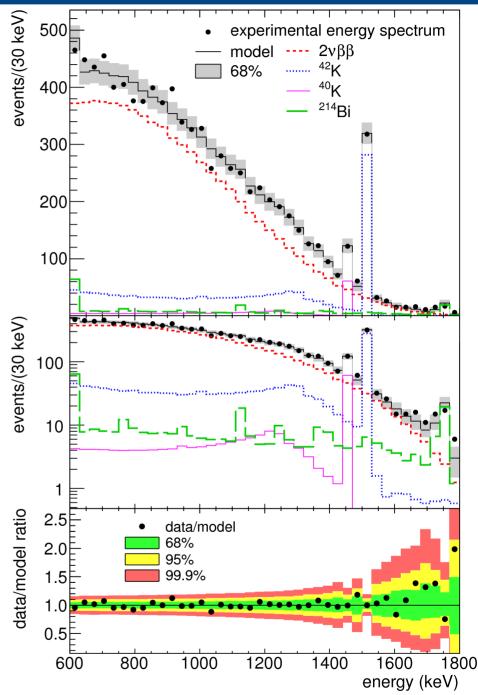
$$T_{1/2}^{2\nu} = (1.84_{-0.08 \text{ fit } -0.06 \text{ syst}}^{+0.11}) \times 10^{21} \text{yr} = (1.84_{-0.10}^{+0.14}) \times 10^{21} \text{yr}$$

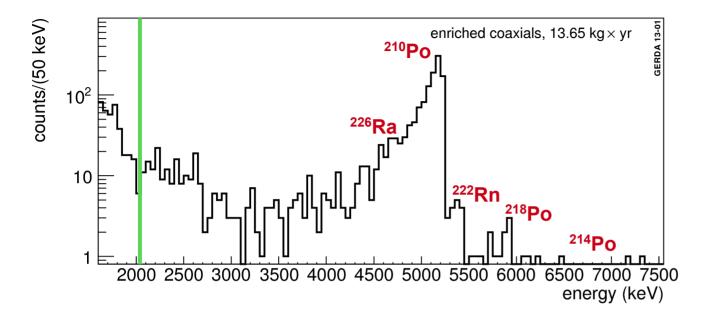
Signal-to-background ratio in the fit range 4:1

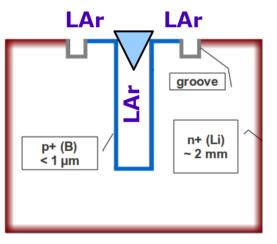
→ uncertainty comparable to previous measurements despite much smaller exposure

Good agreement with reanalysis of HdM data (HdM-K, HdM-B)



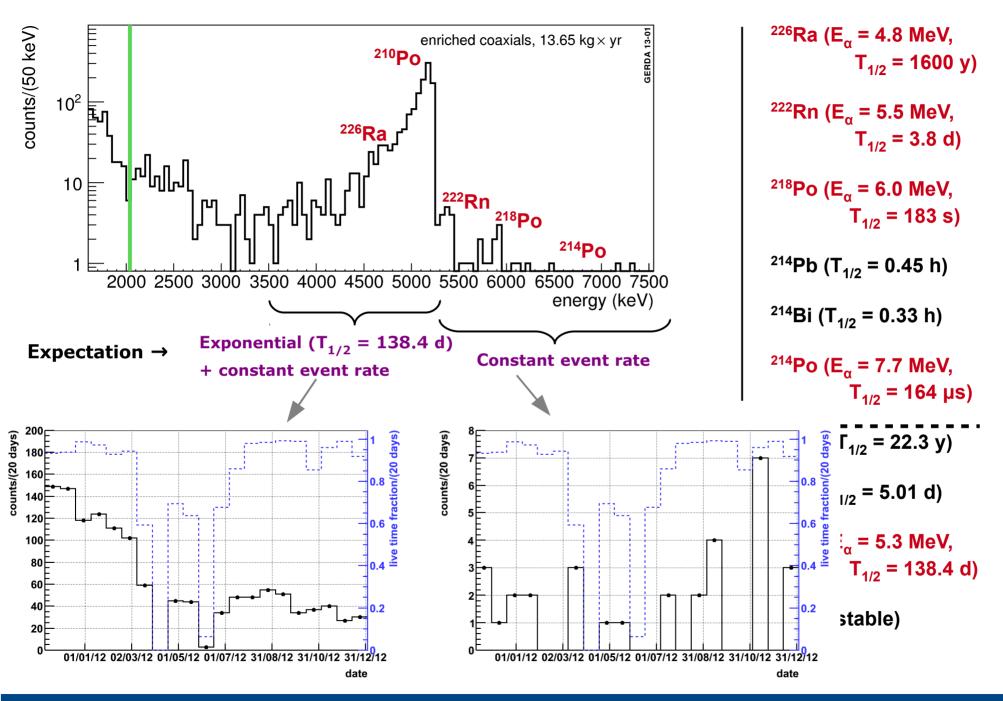


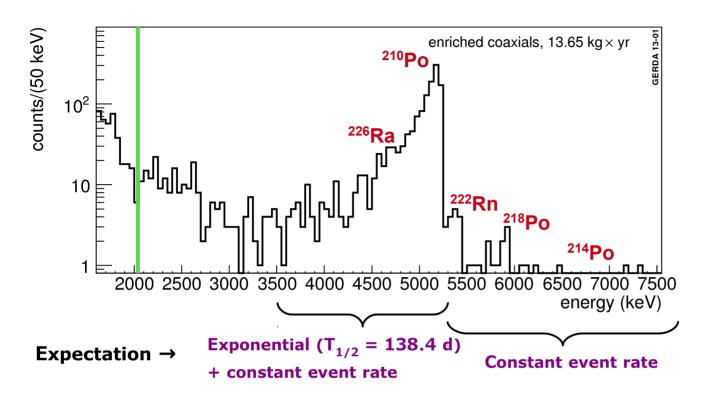




- Range of alphas with E ~ 4 MeV 9 MeV
- $\rightarrow$  14 µm 41 µm in Ge
- $\rightarrow$  34  $\mu$ m 113  $\mu$ m in LAr
- Possible origin of alpha-induced events:
  - → Separate <sup>226</sup>Ra and <sup>210</sup>Po contaminations on thin dead layer (thinDL) surfaces.

```
^{226}Ra (E<sub>g</sub> = 4.8 MeV,
               T_{1/2} = 1600 \text{ y}
^{222}Rn (E<sub>a</sub> = 5.5 MeV,
               T_{1/2} = 3.8 d
^{218}Po (E<sub>\alpha</sub> = 6.0 MeV,
              T_{1/2} = 183 s)
<sup>214</sup>Pb (T_{1/2} = 0.45 \text{ h})
^{214}Bi (T_{1/2} = 0.33 h)
<sup>214</sup>Po (E_{\alpha} = 7.7 \text{ MeV},
              T_{1/2} = 164 \mu s
<sup>210</sup>Pb (T_{1/2} = 22.3 \text{ y})
^{210}Bi (T_{1/2} = 5.01 d)
^{210}Po (E<sub>a</sub> = 5.3 MeV,
              T_{1/2} = 138.4 d
<sup>206</sup>Pb (stable)
```



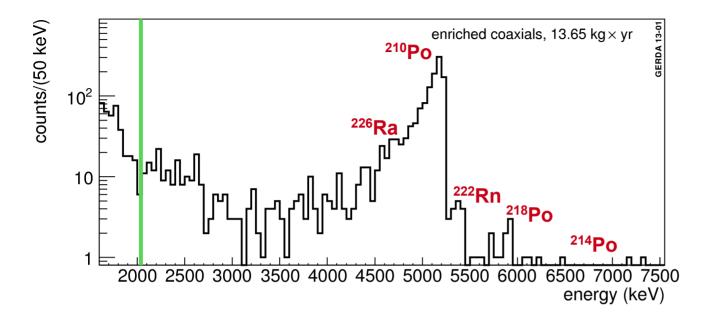


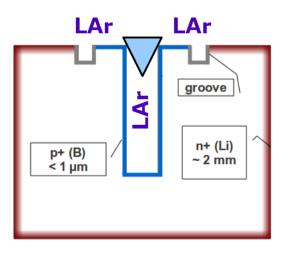
#### Results from fitting the event rate distributions (details in the backup):

	$rac{ ext{C}}{ ext{[cts/day]}}$	$egin{array}{c} \mathbf{N}_0 \ [\mathbf{cts/day}] \end{array}$	$egin{array}{c} \mathbf{T}_{1/2} \ [\mathbf{days}] \end{array}$	p-value
$(3.5{ m MeV}{<}{ m E}{<}5.3{ m MeV})$				
expo		$9.26 \pm 0.26$	$138.4 \pm 0.2$	0.11
$\exp o + const$	$0.57 \pm 0.16$	$7.91 \pm 0.44$	$138.4 \pm 0.2$	0.87
$(5.3{ m MeV}{<}{ m E}{<}7.5{ m MeV})$				
const	$0.09 \pm 0.02$			0.86

Given a strong prior probability on the half life parameter

226Ra (
$$E_{\alpha}$$
 = 4.8 MeV,  
 $T_{1/2}$  = 1600 y)  
222Rn ( $E_{\alpha}$  = 5.5 MeV,  
 $T_{1/2}$  = 3.8 d)  
218Po ( $E_{\alpha}$  = 6.0 MeV,  
 $T_{1/2}$  = 183 s)  
214Pb ( $T_{1/2}$  = 0.45 h)  
214Pi ( $T_{1/2}$  = 0.33 h)  
214Po ( $E_{\alpha}$  = 7.7 MeV,  
 $T_{1/2}$  = 164 μs)  
210Pb ( $T_{1/2}$  = 22.3 y)  
210Pb ( $T_{1/2}$  = 5.01 d)  
210Po ( $E_{\alpha}$  = 5.3 MeV,  
 $T_{1/2}$  = 138.4 d)





#### MC simulations to model the energy spectrum:

- 1) <sup>210</sup>Po on thinDL surface
- 2) <sup>226</sup>Ra & daughters on thinDL surface
- 3) <sup>222</sup>Rn & daughters in LAr close to thinDL surfaces

thinDL thickness: 300 ... 600 nm

$$^{226}$$
Ra (E $_{\alpha}$  = 4.8 MeV,  $T_{1/2}$  = 1600 y)

 $^{222}$ Rn (E $_{\alpha}$  = 5.5 MeV,  $T_{1/2}$  = 3.8 d)

 $^{218}$ Po (E $_{\alpha}$  = 6.0 MeV,  $T_{1/2}$  = 183 s)

 $^{214}$ Pb ( $T_{1/2}$  = 0.45 h)

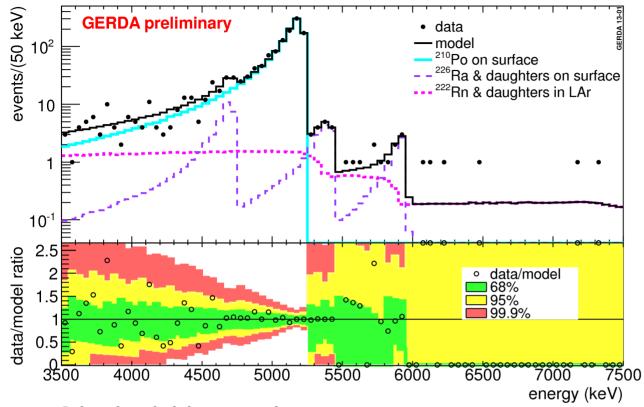
 $^{214}$ Bi ( $T_{1/2}$  = 0.33 h)

 $^{214}$ Po (E $_{\alpha}$  = 7.7 MeV,  $T_{1/2}$  = 164 μs)

 $^{210}$ Pb ( $T_{1/2}$  = 22.3 y)

 $^{210}$ Bi ( $T_{1/2}$  = 5.01 d)

 $^{210}$ Po (E $_{\alpha}$  = 5.3 MeV,  $T_{1/2}$  = 138.4 d)



enrGe coaxials, exposure: 13.65 kg·yr

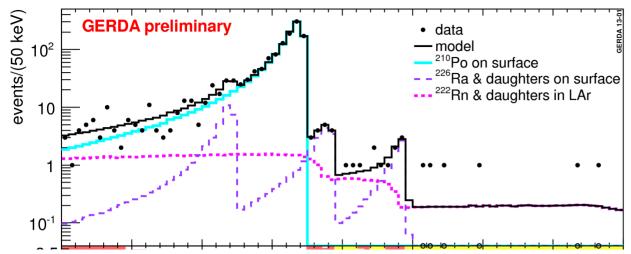
**Binned maximum likelihood fit Fit window:** 3500 - 7500 keV

**p-value:** 0.7

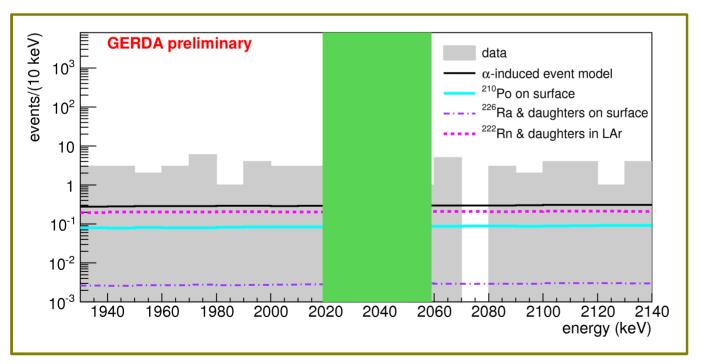
Note: the model explains the  $\alpha$ -induced events in <sup>nat</sup>Ge detectors as well (backup)

**Colored probability intervals:** 

[R. Aggarwal and A. Caldwell, Eur. Phys. J. Plus 127 24 (2012)]



Extrapolation of the alpha-induced event model to ROI (160 keV): (1939 - 2019) plus (2059 - 2139) keV

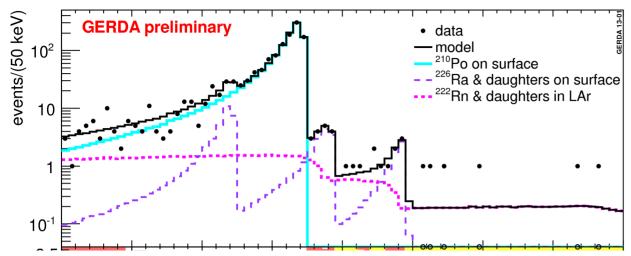


enrGe coaxials, exposure: 13.65 kg·yr

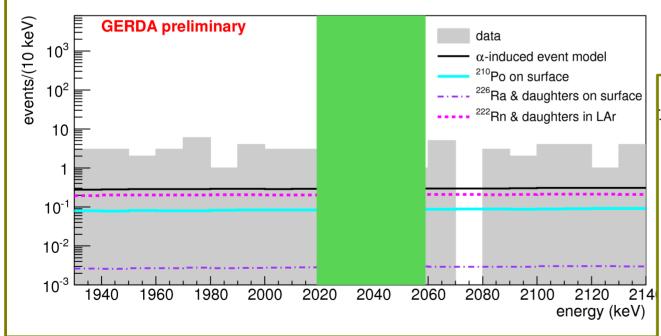
**Binned maximum likelihood fit Fit window:** 3500 - 7500 keV

**p-value:** 0.7

Note: the model explains the  $\alpha$ -induced events in <sup>nat</sup>Ge detectors as well (backup)



Extrapolation of the alpha-induced event model to ROI (160 keV): (1939 - 2019) plus (2059 - 2139) keV



enrGe coaxials, exposure: 13.65 kg·yr

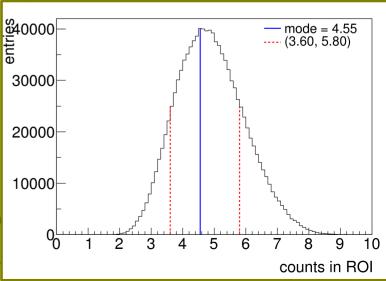
Binned maximum likelihood fit

**Fit window:** 3500 – 7500 keV

**p-value:** 0.7

Note: the model explains the  $\alpha$ -induced events in <sup>nat</sup>Ge detectors as well (backup)

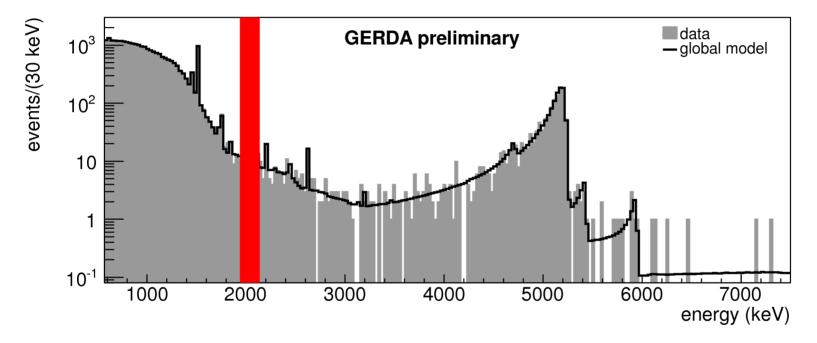
- → data: 49 events
- $\rightarrow$  alpha model: 4.55<sup>+1.25</sup><sub>-0.95</sub> events
- ~ 9% contribution from alphas



### Phase-I: decomposition of the background spectrum

Binned maximum posterior fit to the sum en Ge coax spectrum in (570 - 7500) keV window

- $\bullet$  fit window enlarged to include  $Q_{\beta\beta}$
- background components considered in the global fit:
  - → K-42, K-40, Bi-214, Ac-228 & Th228 (beta- / gamma-induced events) and alpha-induced event model
- p-value of the fit: 0.3



Work in progress: analysis with more background components, various source positions, different data sets is ongoing. systematics under investigation

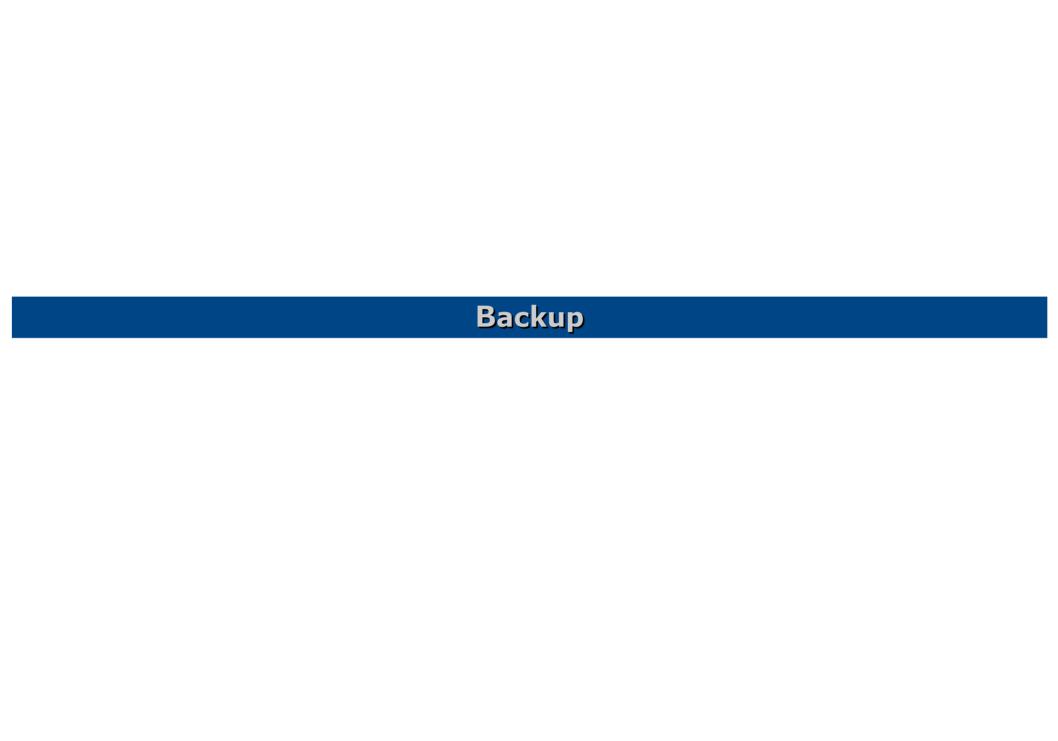
Preliminary: Dominant background contributions around  $Q_{\beta\beta}$   $\rightarrow$  K-42, Bi-214, Tl-208 and alphas

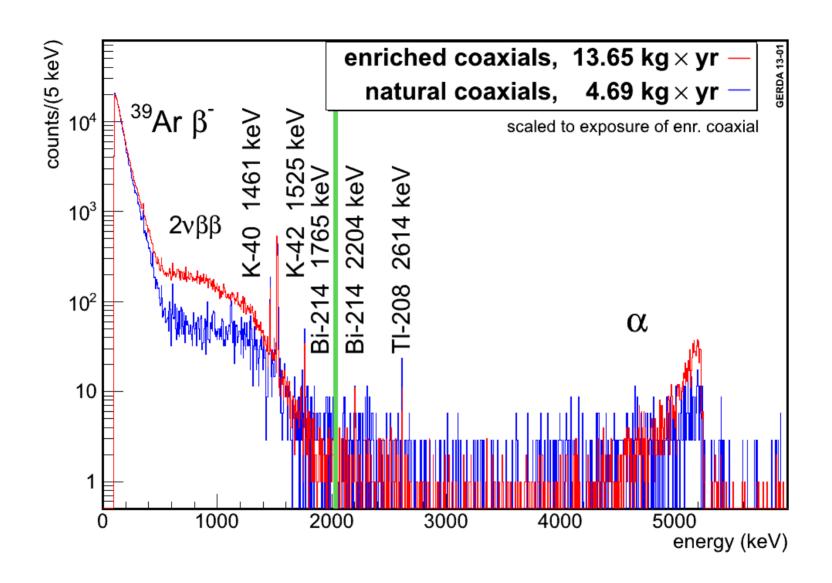
### **Conclusions**

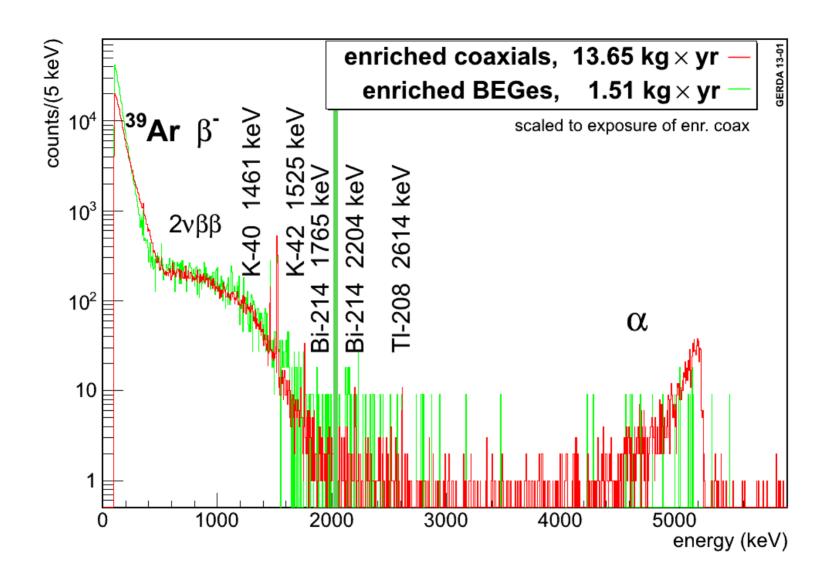
- GERDA Phase I started in Nov 2011 ongoing.
- Phase I background an order of magnitude lower than HdM, IGEX.
- Measurement of  $T_{1/2}$  of  $2\nu\beta\beta$  with the first 5 kg·yr (s:b = 4:1)

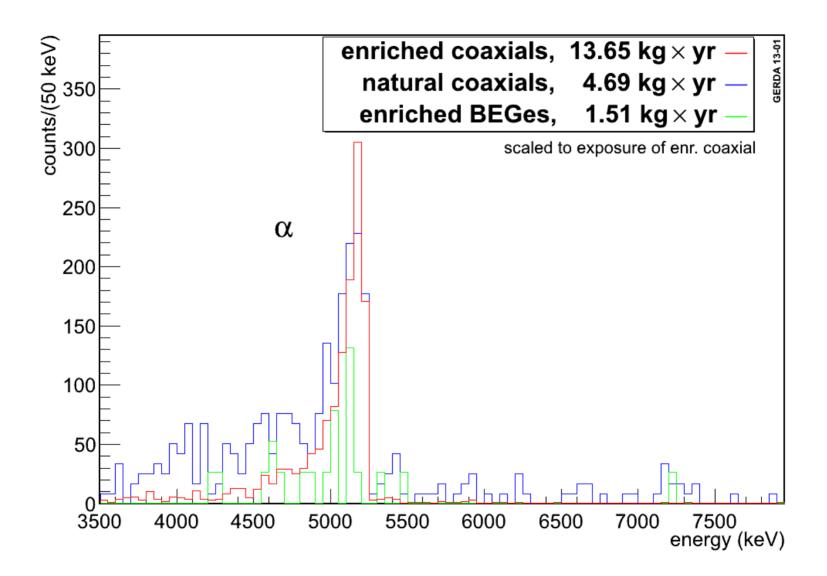
$$T_{1/2}^{2\nu} = \left(1.84_{-0.08 \text{ fit } -0.06 \text{ syst}}^{+0.11}\right) \times 10^{21} \text{yr} = \left(1.84_{-0.10}^{+0.14}\right) \times 10^{21} \text{yr}$$

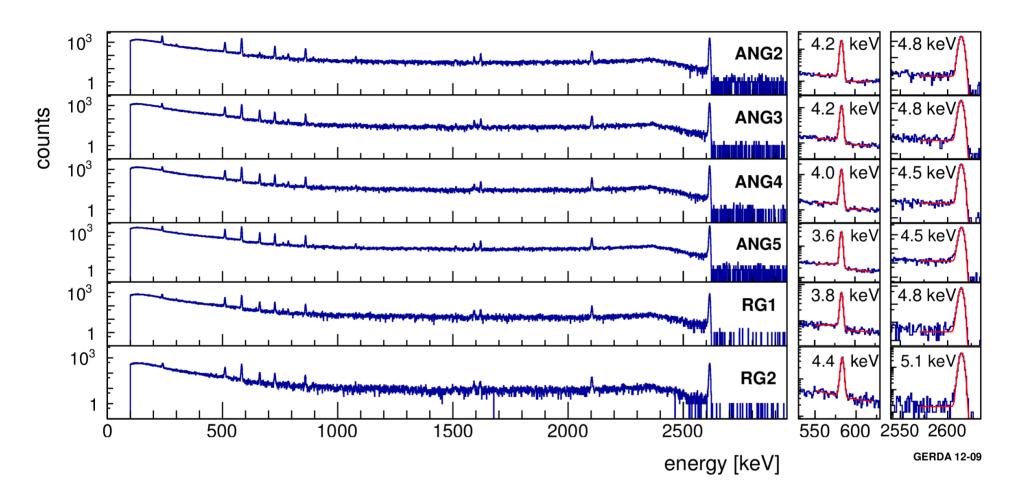
- $_{\mbox{\scriptsize o}}$  Blind analysis: 40 keV window around  $Q_{\beta\beta}=$  2039 keV is blinded.
  - → unblinding: after 20 kg·yr exposure collected (Spring 2013)
- Model the background energy spectrum before unblinding: Promising "preliminary" results!
  - for the upcoming  $0\nu\beta\beta$  analysis: expected number of background events, shape of the background spectrum around  $Q_{\beta\beta}$
  - → for Phase II: understand the background sources in Phase I & mitigate it further

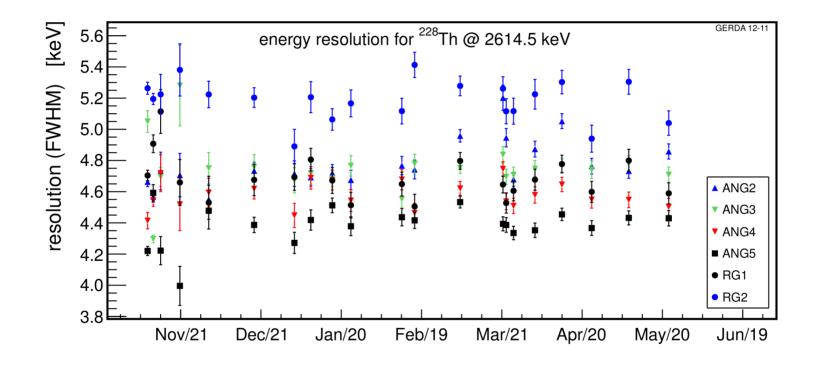


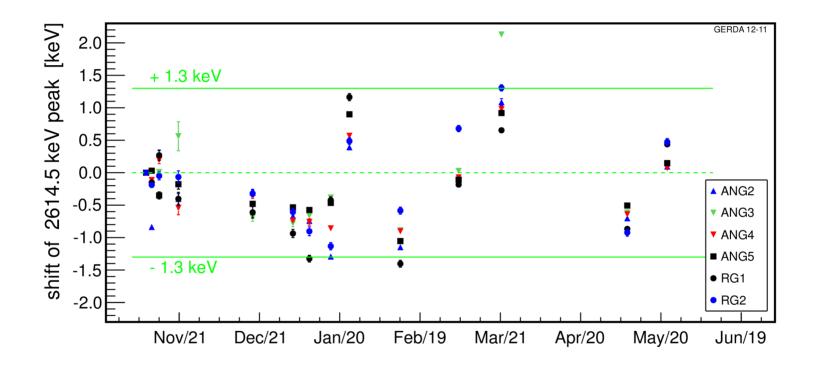












### Analysis of event rate distributions:

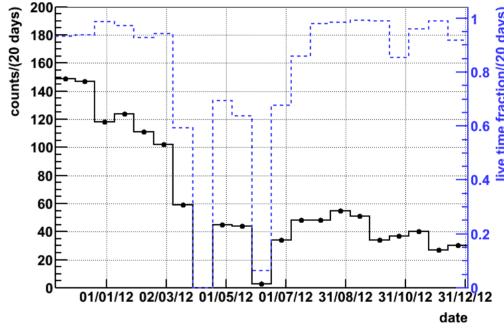
- Fit the distribution with an exponential function  $N(t) = N_0 \cdot e^{-ln2\,t/T_{1/2}}$
- Maximized quantity posterior probability:

$$P(\vec{\lambda}|\vec{n}) \propto P(\vec{n}|\vec{\lambda})P_0(\vec{\lambda})$$

- Set a prior on the half life parameter:  $P_0 (T_{1/2}) = Gaus(138.4 days, 0.2 days)$  half-life of <sup>210</sup>Po
- Likelihood:

$$P(\vec{n}|\vec{\lambda}) = \prod_{i} P(n_i|\lambda_i) = \prod_{i} \frac{e^{-\lambda_i} \lambda_i^{n_i}}{n_i!}$$

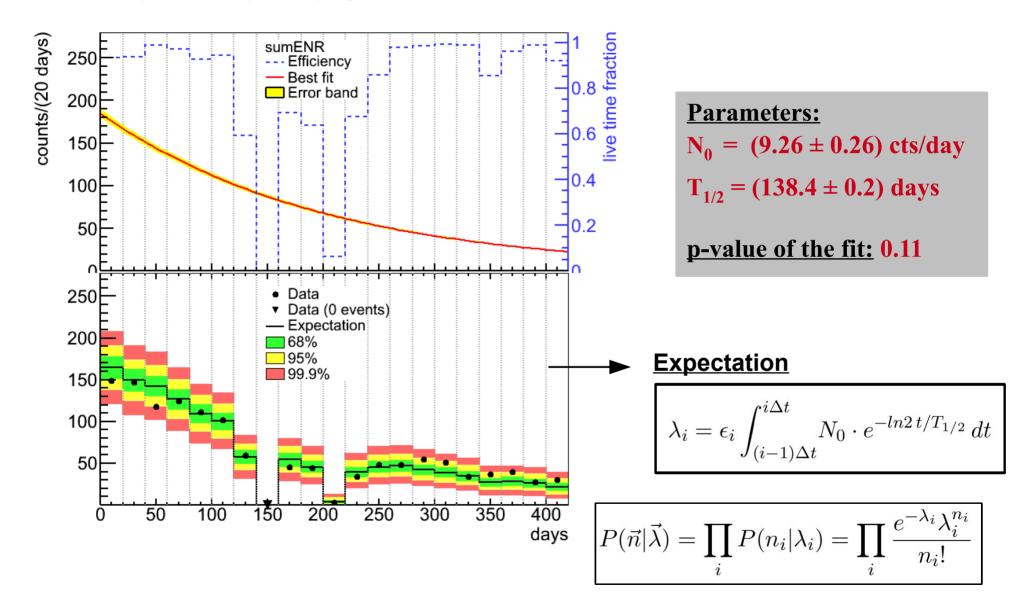
$$\begin{split} &n_i: \text{raw number of counts in i-th bin}\\ &(\text{not scaled, not corrected for live time fraction,}\\ &\lambda_i: \text{expectation in the i-th bin}\\ &\text{corrected with the live time fraction in that bin} \end{split}$$



$$\lambda_i = \epsilon_i \int_{(i-1)\Delta t}^{i\Delta t} N_0 \cdot e^{-\ln 2t/T_{1/2}} dt$$

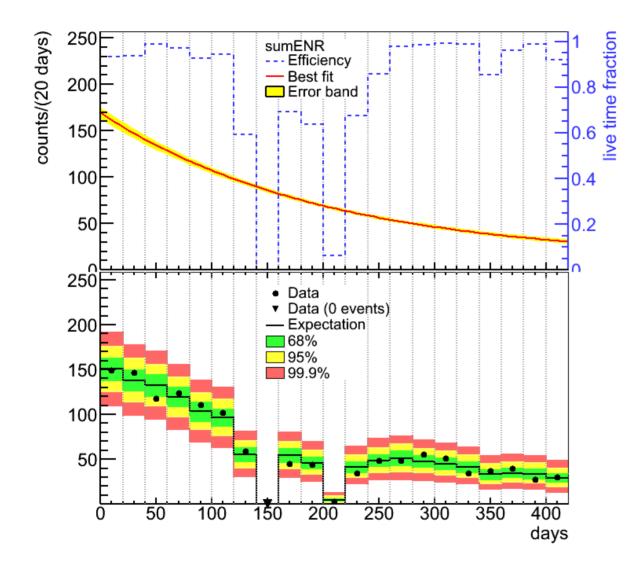
Event rate distribution of events with 3500 keV < E < 5300 keV in sum enrGe-coax

Model: exponentially decaying event rate



Event rate distribution of events with 3500 keV < E < 5300 keV in sum enrGe-coax

Model: exponential + constant rate



### **Parameters:**

$$C = (0.57 \pm 0.16) \text{ cts/day}$$

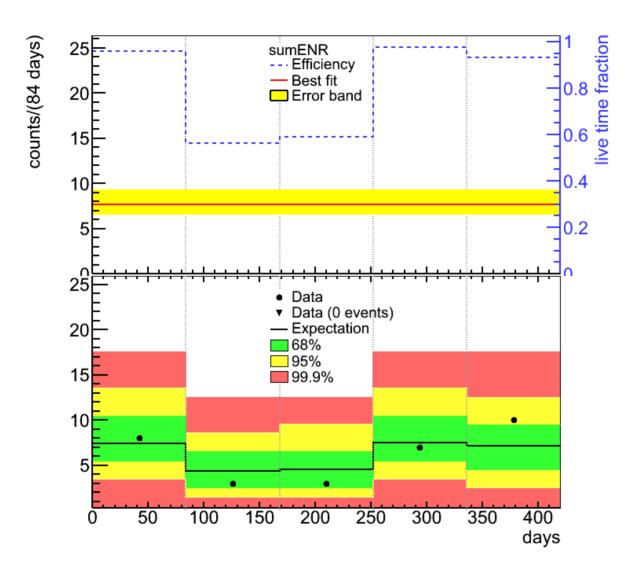
$$N_0 = (7.91 \pm 0.44) \text{ cts/day}$$

$$T_{1/2} = (138.4 \pm 0.2)$$
 days

p-value of the fit: 0.87

#### Event rate distribution of events with E > 5300 keV in sum enrGe-coax

Model: constant rate

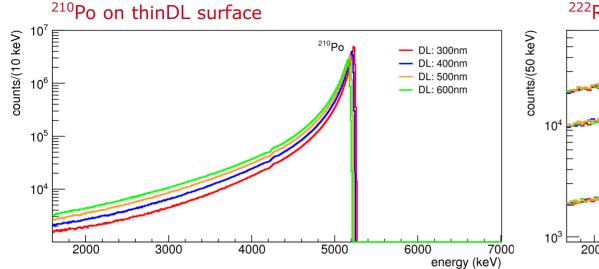


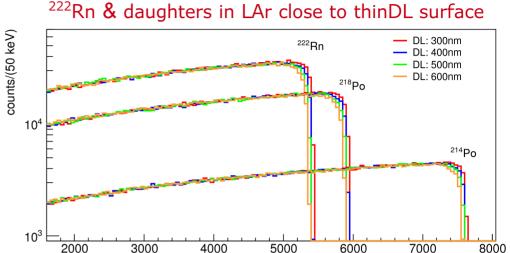
### **Parameters:**

$$C = (0.09 \pm 0.02) \text{ cts/day}$$

p-value of the fit: 0.86

#### Simulated energy spectrum of different model components





### Binned maximum posterior fit to the sum enrGe-coax spectrum in (3500 - 7500) keV window

Maximized the posterior probability using Markov Chain Monte Carlo in Bayesian Analysis Toolkit BAT:

[A. Caldwell et. al., Comput. Phys. Commun. 180, 2197 (2009)]

Posterior probability:  $P(\vec{\lambda}|\vec{n}) \propto P(\vec{n}|\vec{\lambda})P_0(\vec{\lambda})$ 

Likelihood:

$$P(\vec{n}|\vec{\lambda}) = \prod_{i} P(n_i|\lambda_i) = \prod_{i} \frac{e^{-\lambda_i} \lambda_i^{n_i}}{n_i!}$$

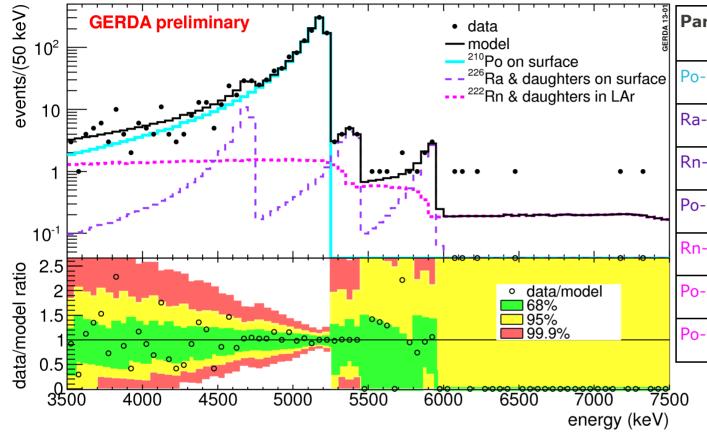
 $n_i$  number of observed,  $\lambda_i$  number of expected events in i-th bin

$$\lambda_i = \sum \lambda_{i,M} \quad o$$
 sum contribution of each model component M

$$\lambda_i = \sum \lambda_{i,M} o ext{sum contribution of each model component M}$$
 
$$\lambda_{i,M} = N_M \int_{\Delta E_i} f_M(E) dE$$
 
$$\int_{\text{scaling parameter for the component M}}$$

energy (keV)

### sumEnrCoax, all runs fit window: (3500 - 7500) keV p-value: 0.72



Parameters	Mode
	(smallest int.)
Po-210 sur	1205 (1160, 1240)
Ra-226 sur	39.5 (30.0, 55.0)
Rn-222 sur	20.5 (14.0, 28.0)
Po-218 sur	11.5 (7.0, 16.0)
Rn-222 in LAr	40.5 (26.0, 61.0)
Po-218 in LAr	22.5 (14.0, 38.0)
Po-214 in LAr	14.5 (9.0, 21.0)

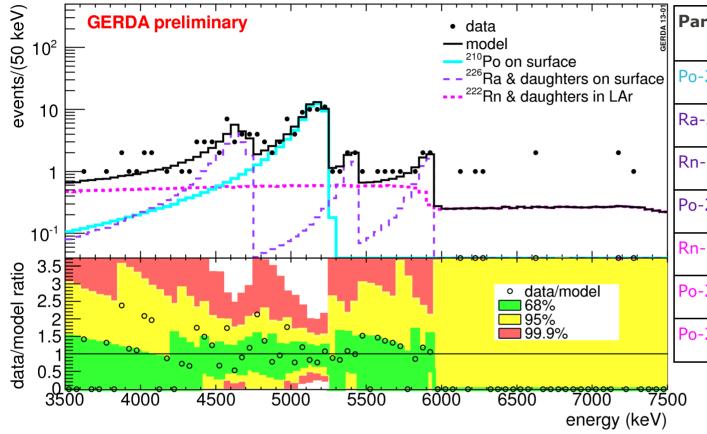
Ra-226	on	p+	surfa	ce:
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A = 81.3/340.96/24/3600

 $A = 3 \mu Bq$ 

	BI [10 <sup>-3</sup> cts/(keV kg yr)]
alpha model	2.85 (2.3, 3.4)
Po-210 on surface	0.67 (0.64, 0.70)
Ra-226 & daughters on surface	0.045 (0.03, 0.06)
Rn-222 & daughters in LAr	2.15 (1.6, 2.7)

### GTF112, golden data set fit window: (3500 - 7500) keV p-value: 0.89



Parameters	Mode (smallest int.)
Po-210 sur	55.0 (40, 70)
Ra-226 sur	18.5 (13, 26)
Rn-222 sur	10.5 (6, 15)
Po-218 sur	9.5 (6, 14)
Rn-222 in LAr	17.5 (9, 25)
Po-218 in LAr	15.5 (9, 24)
Po-214 in LAr	15.5 (10, 22)

Ra-226	on	p+ s	urface:
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A = 38/308.37/24/3600

 $A = 1.4 \mu Bq$ 

	BI [10 <sup>-3</sup> cts/(keV kg yr)]
alpha model	6.4 (5.1, 7.9)
Po-210 on surface	0.17 (0.13, 0.20)
Ra-226 & daughters on surface	0.16 (0.12, 0.20)
Rn-222 & daughters in LAr	6.1 (4.7, 7.5)