Estimate of the <sup>42</sup>Ar content in the Earth's atmosphere

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# Decay scheme of <sup>42</sup>Ar $^{42}$ Ar (33 yr, $Q_{\beta} = 600 \text{ keV})$ $^{42}$ K (12.36 h, Q<sub> $\beta$ </sub> = 3.52 MeV) $E_{v} = 1.524 \text{ MeV}$ 1.922 MeV, 2.423 MeV..

### Decay scheme of <sup>42</sup>Ar

#### Bgd. in LAr: example <sup>42</sup>Ar

<sup>42</sup>Ar / <sup>nat</sup>Ar = 3·10<sup>-21</sup> (30 μBq/kg)

[Barabash et al., LAr-TPC @ LNGS]



#### References to the theme

 A.S.Barabash, V.N.Kornoukhov, V.E.Jants NIM A385 (1997) 530-534

 A.J.Peurrung, T.W.Bowyer, R.A.Craig, P.L.Reeder
NIM A396 (1997) 425-426

# <sup>42</sup>Ar production modes

• A two-fold neutron capture process starting with the <sup>40</sup>Ar isotope:

 $^{40}Ar(n,\gamma)^{41}Ar,$  $^{41}Ar(n,\gamma)^{42}Ar$ 

• Spallation reactions from the cosmic ray nucleon component on <sup>40</sup>Ar:

#### $^{40}Ar(\alpha,2p)^{42}Ar$

 Interactions of cosmic ray muons with the nuclei of the atmosphere and rocks near the surface of the Earth negligible.

#### Two-fold neutron capture

 $^{40}Ar(n,\gamma)^{41}Ar$  $^{41}Ar(n,\gamma)^{42}Ar$ 

 $t_{1/2}(41) = 1.83 h$  high flux of neutrons

The sources of neutrons in the atmosphere:

- from cosmic rays (<  $10^{-42}$  of  $^{42}$ Ar /  $^{nat}$ Ar);
- from nuclear bomb tests (!)

# Bomb testing

- 1<sup>st</sup> stage: 1945-1957
- 2d stage: 1961-1962

#### Total energy release ~ 400 Mtn

- 1 explosion  $\implies$  60 (90?) Mtn (30.10.1961)
- 3 explosions ➡ 30 Mtn
- 1 explosion  $\implies$  25 Mtn

### <sup>42</sup>Ar yield in bomb testing

N(42)=  $(1/8\pi) \times \sigma(40) \times \sigma(41) \times n(40) \times (\phi \cdot t) \times dV$ 

N(42)=  $(1/8\pi) \times \sigma(40) \times \sigma(41) \times n(40) \times N^2 \times P^2 \times \int_{x}^{few \ km} \frac{\int_{x}^{few \ km} \exp(-2 \ r/L)/(1/r^2) \ dr}{r_{min}}$ 

- L fast neutron attenuation length,  $L = 2 \cdot 10^4$  cm
- $\sigma(40) = 0.65 \cdot 10^{-24} \text{ cm}^2$
- $\sigma(41) = 0.5 \cdot 10^{-24} \text{ cm}^2$
- $n(40) = 2.5 \cdot 10^{17} \text{ cm}^{-3}$

Neutrons production Types of weapons:

1. nuclear bomb (U/Pu fission);  $N_n \sim 2.25 \cdot 10^{26}$  neutrons/Mtn

- 2. thermonuclear bomb:
  - based on liquid deuterium;
  - based on chemical compound of D and <sup>6</sup>L

 $N_n \sim 2.25 \cdot 10^{26}$  neutrons/Mtn

## <sup>42</sup>Ar yield calculation

few km  
N(42)=1.64 · 10<sup>20</sup> · P<sup>2</sup> · m · 
$$\int \exp(-2 r/L)/(1/r^2) dr$$
  
 $r_{min}$ 

Where  $r_{min}$  – radius of scattering debris  $r_{min} = f(P, Mtn)$  1st stage of the explosion,  $t < 10^{-7}$  sec

• Energy release ~  $4.2 \cdot 10^{22}$  erg/Mtn:

radiation;

shock vawe.

- $t \sim (n*10) \cdot 10^{6} \circ C$
- moderation of fast neutrons (on the light elements of TNT + the unspent nuclear charge + the matter of bomb shell)

# 2d stage of the explosion, $t \le 10^{-5}$ sec;

Expansion of the scattering debris ≡
≡ the speed of shock vawe

$$r = (E/\rho)^{0.2} \cdot t^{0.4}$$



 $r = 13 \text{ m for } E \approx 1 \text{ Mtn}$ 

### 3d stage of the explosion

 Partly moderated neutrons start to escape from the scattering bomb shell and interact with the nuclei of the air

• 60% of neutron flux  $30 \text{ eV} \le E \le 1 \text{ MeV}$ • 40% of neutron flux  $1 \text{ MeV} \le E \le 14 \text{ MeV}$ 

# Estimation of <sup>42</sup>Ar production $N \sim 3.10^{20}$ atoms $(3 \cdot 10^{-22} \text{ atoms of } {}^{42}\text{Ar} / {}^{\text{nat}}\text{Ar atoms})$ Overestimation due to the assumptions: 1. All neutrons are thermal ones $\implies k \ge 10$ ; 2. An air density is 2 times lower at heights of a few kilometers $\longrightarrow$ $k \sim 2$

$$N \sim 1.5 \cdot 10^{19}$$
 atoms



#### <sup>42</sup>Ar after bomb testing

 $\leq 1.5 \cdot 10^{-23}$  atoms of  $^{42}$ Ar / <sup>nat</sup>Ar atoms

taking into account that nuclear testing in the Earth's atmosphere was finished in 1962

 $\leq 6 \cdot 10^{-24}$  atoms of  $^{42}$ Ar /  $^{nat}$ Ar atoms

# $^{40}$ Ar( $\alpha$ ,2p) $^{42}$ Ar (1)

 $\alpha$ -particles ~ 14% of proton flux of cosmic ray  $\lambda \cdot N(42) = \sigma \cdot \phi \cdot n(40) \cdot S$ 

Where

- $\sigma(\alpha, 2p) = 2.5 \text{ mb}$  (? uncertainty k = 2)
- $\phi(\alpha \text{ particles}) = 0.1 \text{ m}^{-2} \text{ sec}^{-1} (200 \text{ MeV} 1 \text{ GeV})$
- $S = 4 \cdot \pi \cdot R^2 = 5.15 \cdot 10^{14} \text{ m}^2 (\text{R} \text{the Earth' radius})$

The range of α particles in the air: R(E) = 2.9 g·cm<sup>-2</sup> · (E/200)<sup>1.8</sup> = 30 g· cm<sup>-2</sup> then n(40) = 4.316 · 10<sup>21</sup> at ·cm<sup>-2</sup> N(40) = 5.15·10<sup>14</sup>\*10<sup>4</sup>\*4.316·10<sup>21</sup>= 2.2\*10<sup>40</sup>

# $^{40}Ar(\alpha, 2p)^{42}Ar$ (2)

 $\lambda \cdot N(42) = 2.5 \text{ mb* } 0.1 \text{ m}^{-2} \text{ sec}^{-1} \text{ }^*2.2 \text{ }^*10^{40} \text{ at} = 6 \text{ }^*10^{12} \text{ sec}^{-1}$ 

 $N(42) = 9 \cdot 10^{21} atoms$ 

 $9 \cdot 10^{21} / 10^{42} = 9 \cdot 10^{-21} \sim 10^{-20} \text{ at } {}^{42}\text{Ar/at Ar}$ 

#### CONCLUSION (1)

1. Bomb testing:  $\leq 6.10^{-24} \text{ atoms } {}^{42}\text{Ar} / {}^{\text{nat}}\text{Ar}$ 

2. Spallation reactions from the cosmic ray nucleon component: <sup>40</sup>Ar(α,2p)<sup>42</sup>Ar
≤ 9 ⋅ 10<sup>-21</sup> atoms <sup>42</sup>Ar / <sup>nat</sup>Ar

3. From experiment (A.Barabash, TAUP2001):  $\leq 3.10^{-21}$  atoms  $^{42}$ Ar /  $^{nat}$ Ar

#### CONCLUSION (2)

## If $\leq 3.10^{-21}$ atoms <sup>42</sup>Ar/Ar $\leq 26 \mu Bq/kg$ $\sim 2.2 decay/day$

# How to measure the <sup>42</sup>Ar content?

