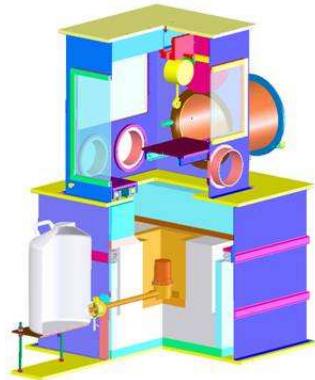


High Sensitive Gamma-Spectrometers of GERDA for Material Screening: Part 2



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XIV-th International School on `Particles and Cosmology'
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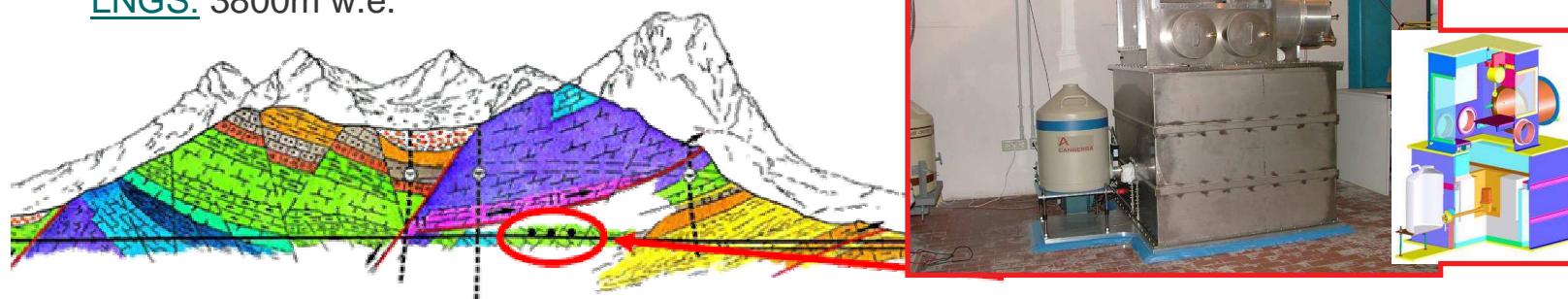


Outline

- § basic design considerations
- § setup of the CORRADO spectrometer (15m w.e.)
- § setup of the GeMPI III spectrometer (3800m w.e.)
- § comparison of achieved background-reduction

GeMPI III:

LNGS: 3800m w.e.



Basic Design Considerations

The central question:

How to decrease the lowest detectable specific activity:

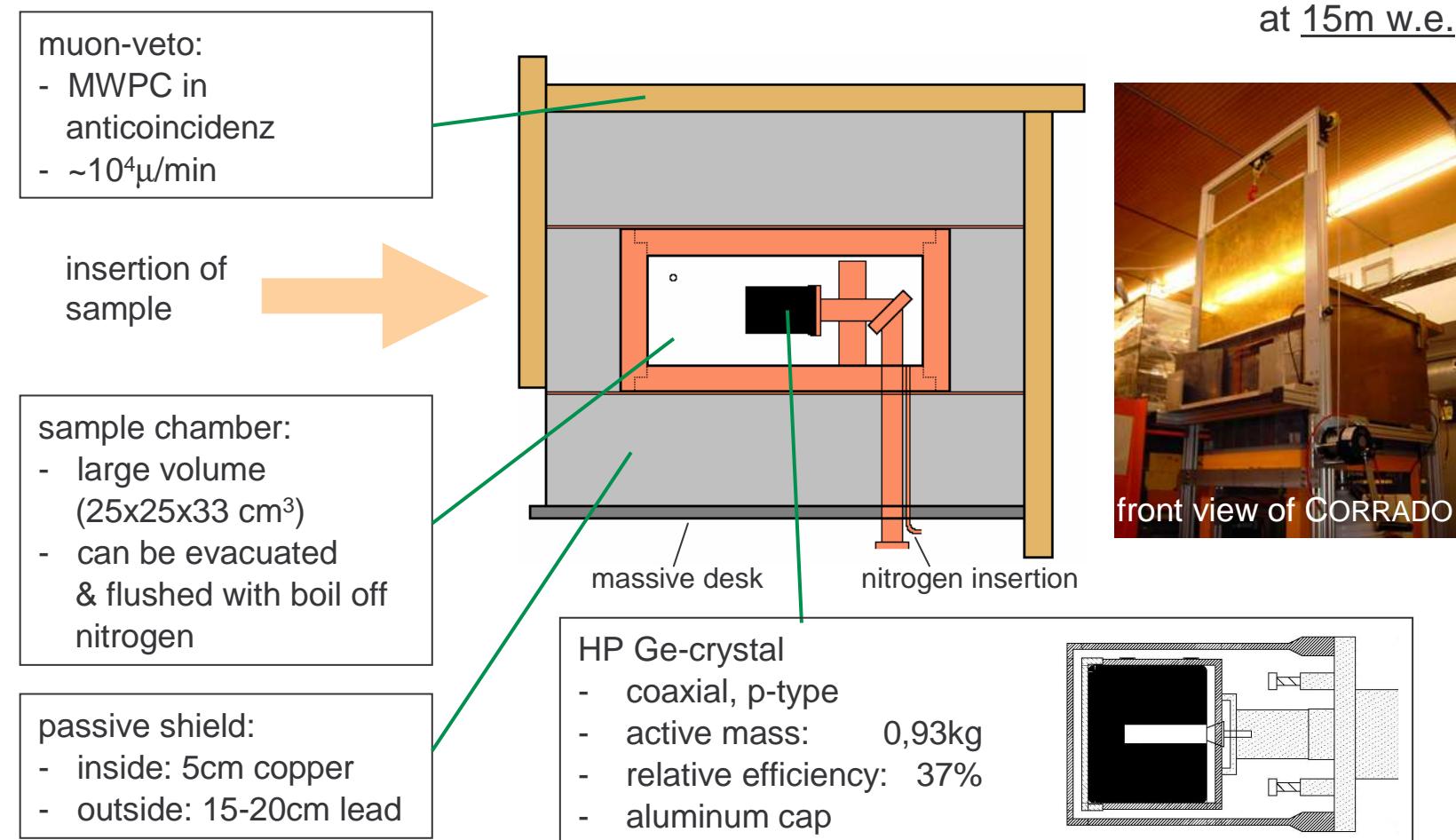
$$A_{spec} \approx \frac{\sqrt{B}}{M \cdot \epsilon \cdot t}$$

- § optimize signal count rate (i.e. $M \cdot \epsilon$ is maximal)
 - use large Ge-crystal
 - use high sample masses in efficient geometry
i.e. choose optimal sample chamber dimensions
- § low background

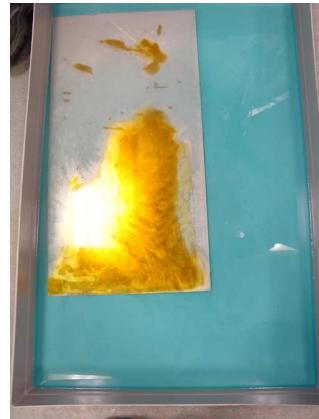
Background Reduction Techniques

- § external background (environmental radioactivity, radon & progenies, neutrons from fission and (α ,n)-reactions)
 - passive shielding (Pb, Cu, polyethylene)
 - air tightness, N₂-flushing
- § internal background in detector & shielding material
 - strict material selection (iterative process)
 - minimize cosmogenic production in detector & shielding copper during production
 - surface cleaning, machining
- § cosmic rays
 - go underground
 - active muon veto (in shallow depth)

Design of CORRADO



Assembly of CORRADO



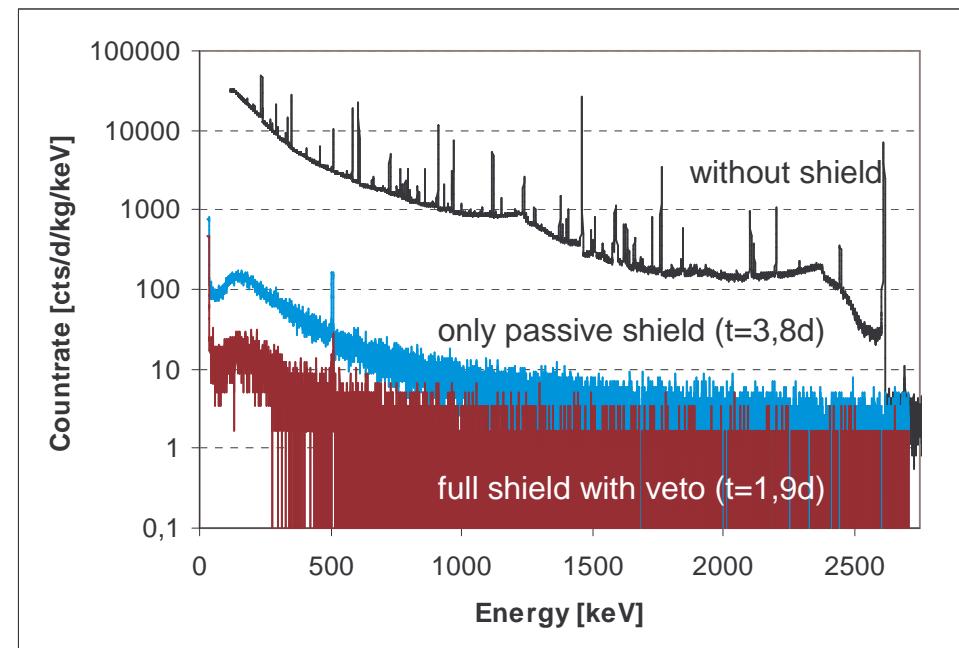
Background Reduction of CORRADO

- based on first preliminary background spectrum:
- **passive shield**: reduction of environmental & airborne radioactivity by factor ~ 100
- **active shield**: suppression of muon induced background by factor ~ 10 (88%)
- still no background lines observable

count rate [cts/day]

40-2700 keV		5600
609 keV	(^{214}Bi)	<13
1461 keV	(^{40}K)	<2.3
2615 keV	(^{208}Tl)	<2.9

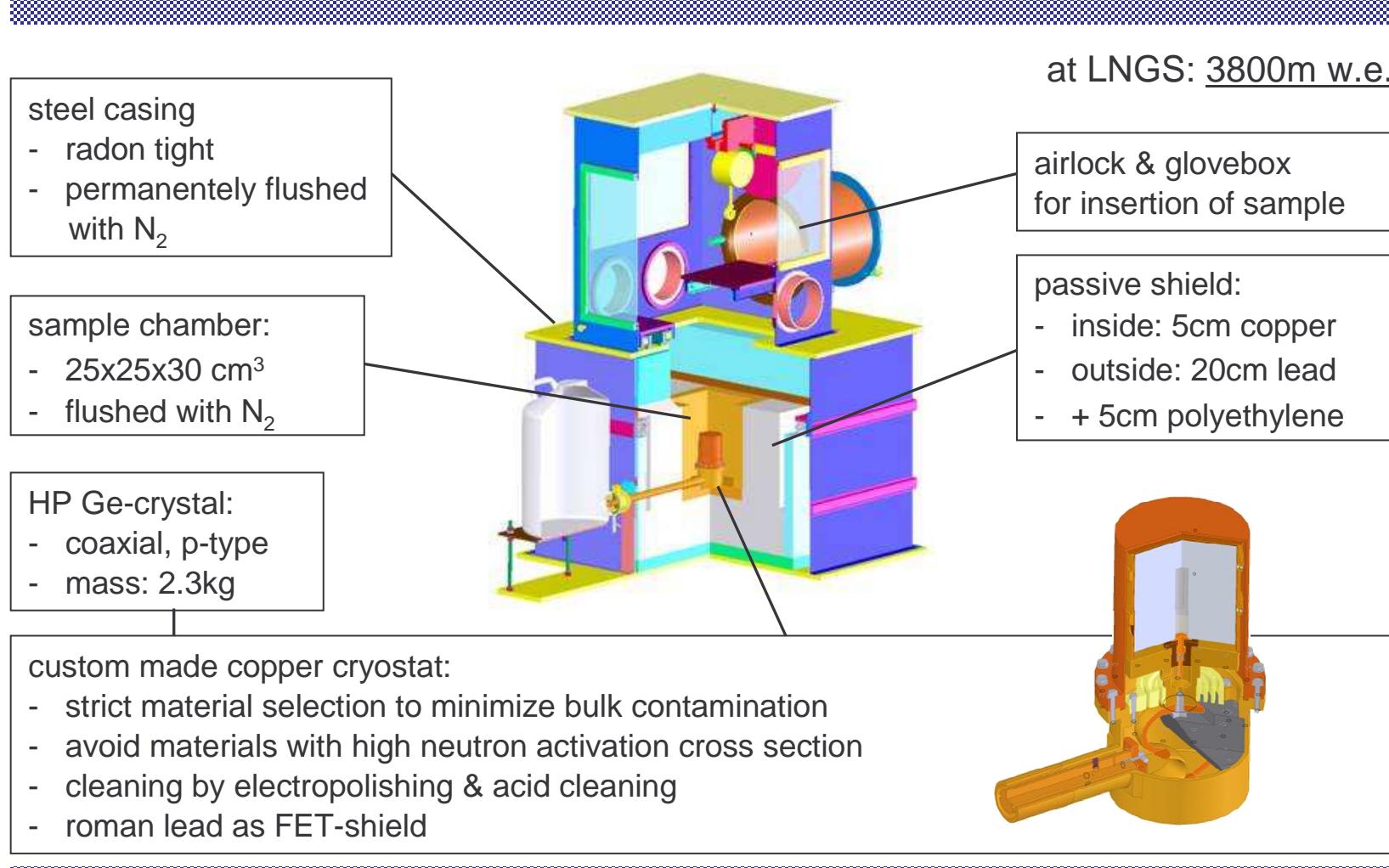
- coming up: long term background measurement



expected sensitivity:

~1 mBq/kg

Design of GeMPI III

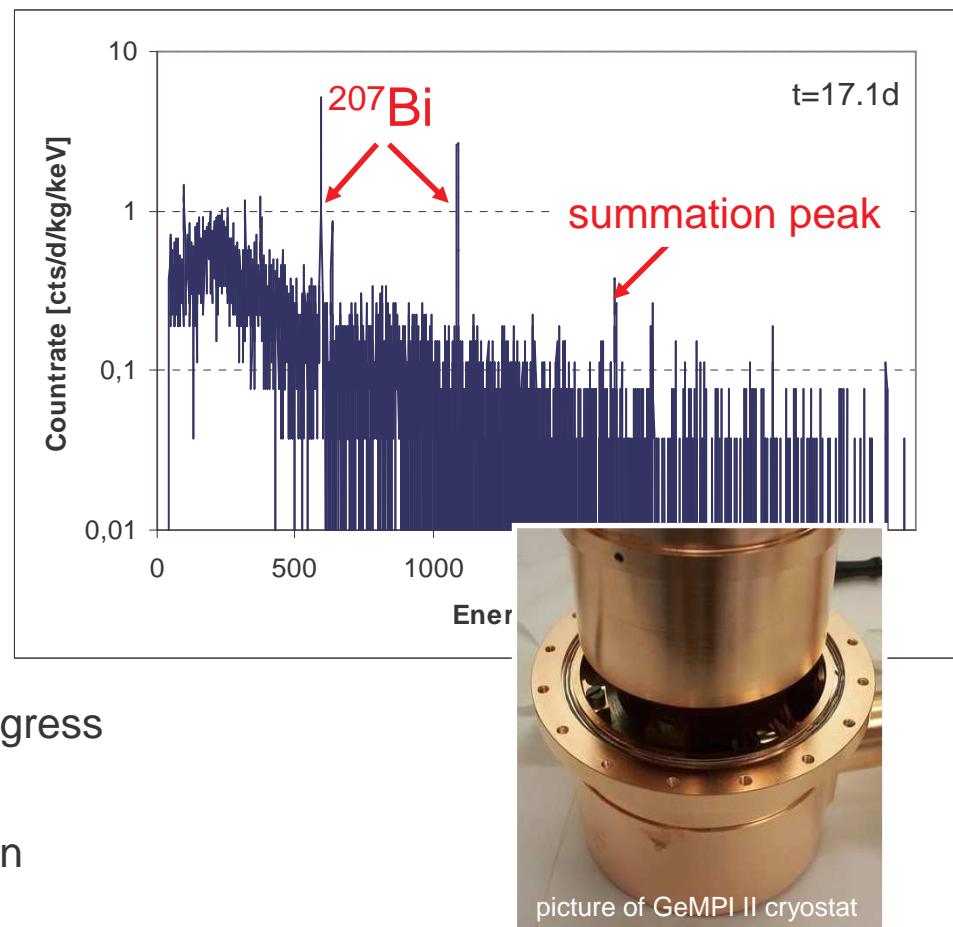


Assembly of GeMPI III

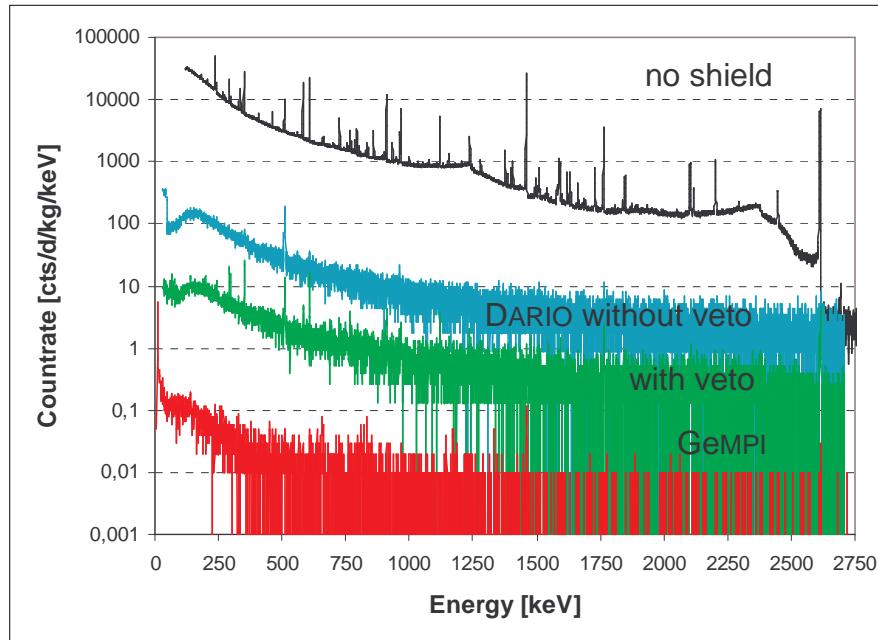


First Spectra of GeMPI III

- ^{222}Rn (^{214}Bi) lines from air in sample chamber visible (due to provisional tightness)
- high **^{207}Bi contamination!**
570keV: $(37,8 \pm 1,6)$ cts/d
1064keV: $(21,2 \pm 1,2)$ cts/d
sum: $(3,0 \pm 0,5)$ cts/d
- from line ratio: contamination is inside of shielding
- suspecting the lead: possible production via (p,n)-reaction on ^{207}Pb
- screening of candidates in progress
- from summation peak & MC:
distance & direction information
(in progress)



Comparison of Background Reduction



longlive isotopes	Energy	GeMPI	DARIO
^{226}Ra	352keV 609keV	<24 <25	4790 ± 590 4070 ± 460
^{228}Th	583keV 2615keV	<21 18 ± 5	1440 ± 330 4520 ± 280
^{40}K	1461keV	86 ± 12	302 ± 161
^{60}Co	1173keV 1332keV	43 ± 10 35 ± 8	<375 <348
total	100- 2730	6840 ± 110	1320000 ± 4400

achieved sensitivity: GeMPI ~10 $\mu\text{Bq}/\text{kg}$
DARIO/CORRADO ~1 mBq/kg

Conclusion

- § new high sensitive Ge-spectrometers are being build in order to provide sufficient screening capacity for GERDA.
- § with a shielding system appropriate to the depth of the detector site sensitivities of
 - ~1mBq/kg @ 15m w.e.
 - ~10 μ Bq/kg @ 3800m w.e.can be achieved.
- § a good knowledge of the detector's background, it's geometry, as well as large sample masses at long measurement times are required.