

Observations of gamma-ray binaries with VERITAS

PSR J1023+0038, HESS J0632+057 & LS I +61 303



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for the VERITAS Collaboration



Alliance for Astroparticle Physics

- array of four 12 m Imaging Atmospheric Cherenkov Telescopes located in southern Arizona
- energy range: 85 GeV to >30 TeV
- field of view of 3.5°
- angular resolution $\sim 0.08^\circ$
- point source sensitivity (5σ detection): 1% Crab in < 25 h (10% in 25 min)

The VERITAS Binary Program

	type	D (kpc)	orbital period [d]	GeV/TeV detection	type of observation	reference (VERITAS)
LS I +61 303	Be+neutron star? BH?	1.6	26.5	✓/✓	regular since 2006 (10-30 h/season)	ApJ 2008, 2009, 2011, 2013
HESS J0632+057	B0pe + ??	1.5	315	x/✓	regular since 2006 (10-30 h/season)	ApJ 2009, 2014
LS 5039	06.5V+neutron star? BH?	2.5	3.9	✓/✓	(~10 h/season)	-
Cygnus X-1	O9.7Iab + BH	2.2	5.6	(x/✓)	ToO (X-rays/LAT)	-
Cygnus X-3	Wolf Rayet + BH?	7	0.2	(✓)/x	ToO (X-rays/LAT)	ApJ 2013
1A0535+262	Be/pulsar binary	2	111	x/x	ToO (triggered by Swift XRT)	ApJ 733, 96 (2011)
V407 Cygni	Nova in a symbiotic binary	2.7		✓/x	ToO (triggered by Fermi)	ApJ 754, 77 (2012)
Be/X-ray Binary discover program	Be-XRB	-	-	-	filler program	-
BAT flaring hard X-ray objects	SGRs+XRBs	-	-	-	ToO	-
Millisecond pulsar binaries	MSPB	-	-	-	regular (10-15 h/season)	-
Magnetars	SGRs+AXPs	-	-	-	ToO (GRB pipeline)	Proc of ICRC 2009

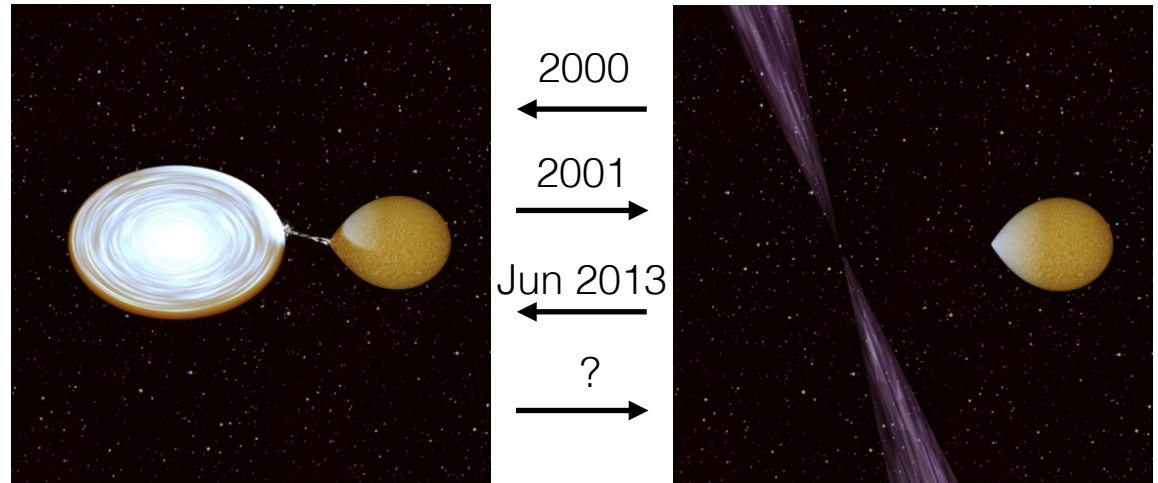


PSR J1023+0038: A new type of gamma-ray binary?

PSR J1023+0038: 1.69 ms spin period, 4.8 hr orbital period at 1.3 kpc

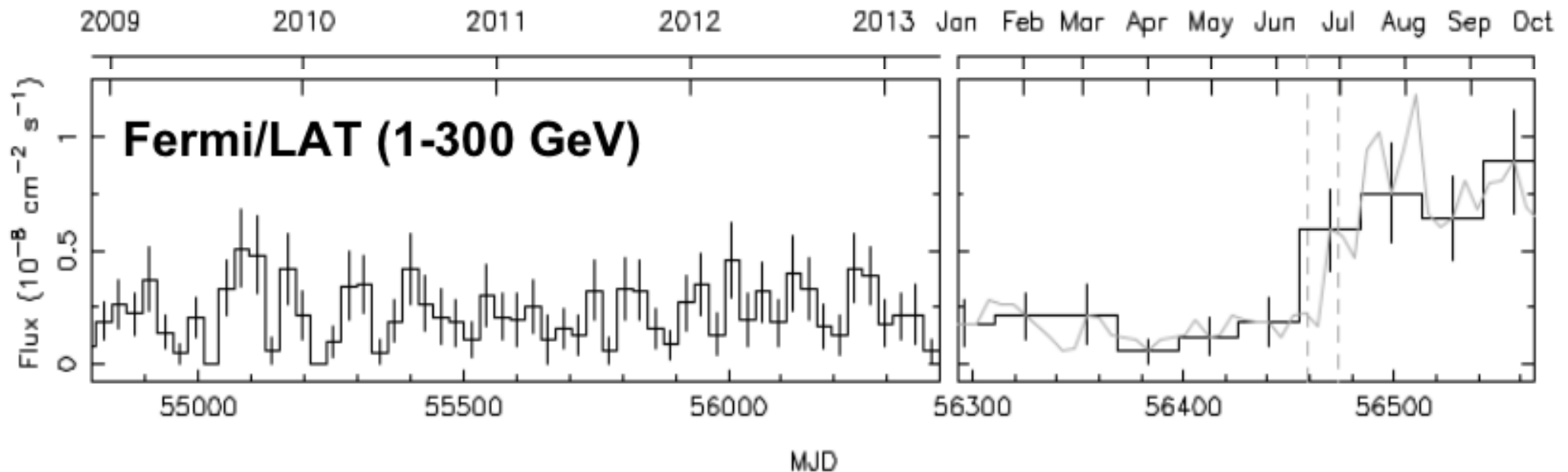
The June 2013 state change:

- Radio pulsations disappear
- an accretion disk appears in optical
- GeV gamma-rays flux increases by a factor >5



LMXB

Radio msPulsar



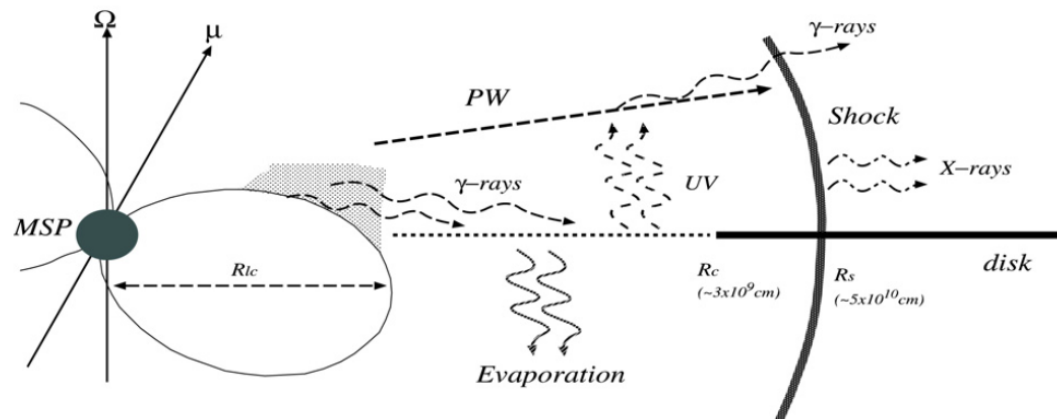
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VERITAS observations in the Radio msPulsar phase

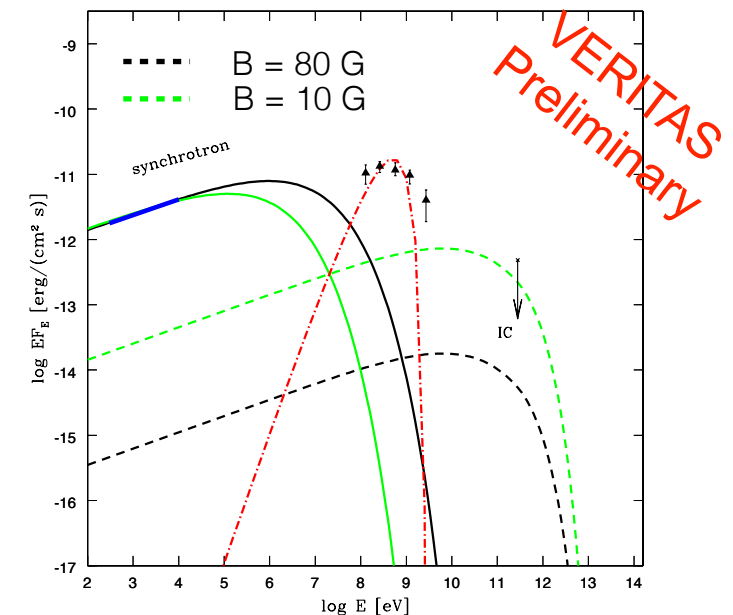
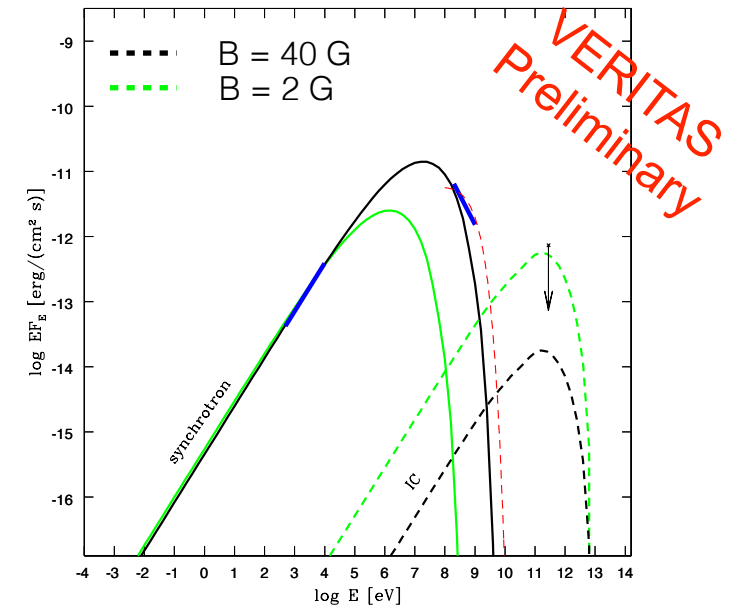
- 20 h of observations between Dec 2010 - Feb 2011
- UL (95% , $E > 280$ GeV): $F < 1.1 \times 10^{-8}$ $\text{cm}^{-2}\text{s}^{-1}$

VERITAS observations in the LMXB phase

- 10 h of observations during Dec 2013
- UL (95% , $E =$): $F < \text{erg cm}^{-2}\text{s}^{-1}$

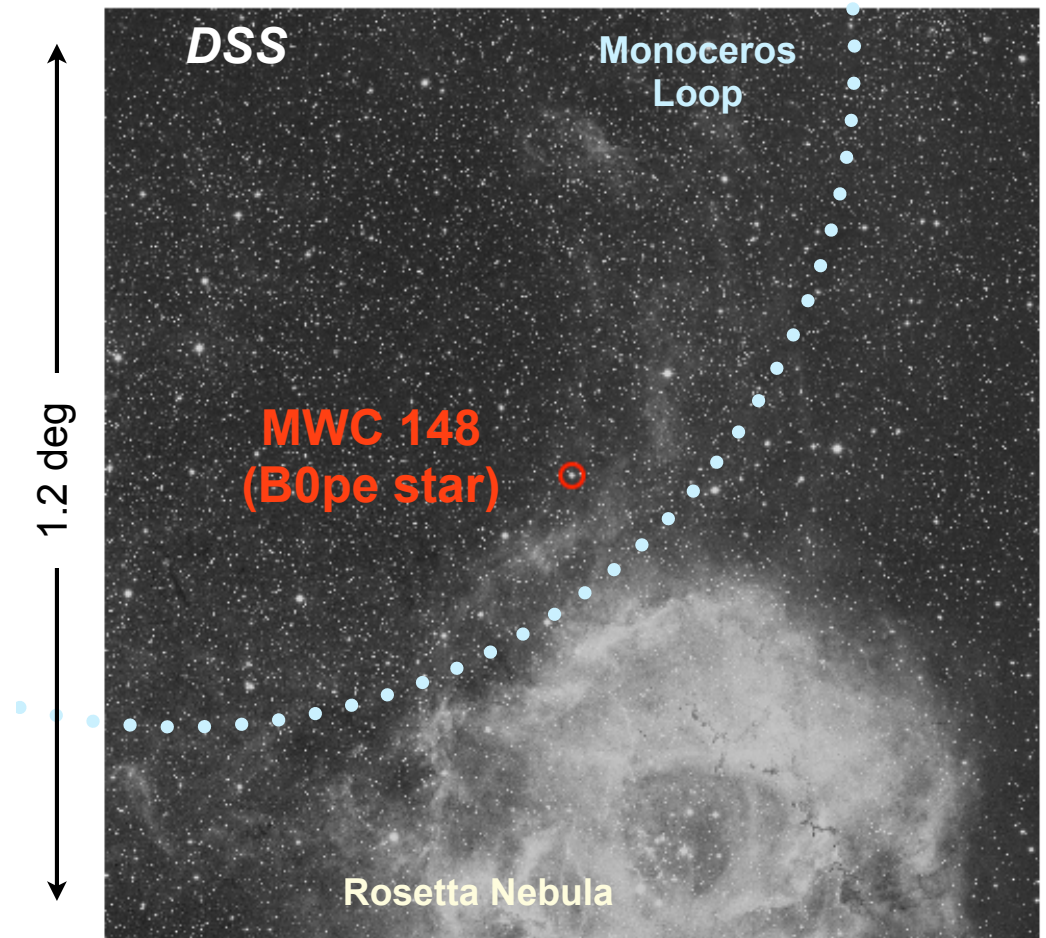


Takata et al 2014

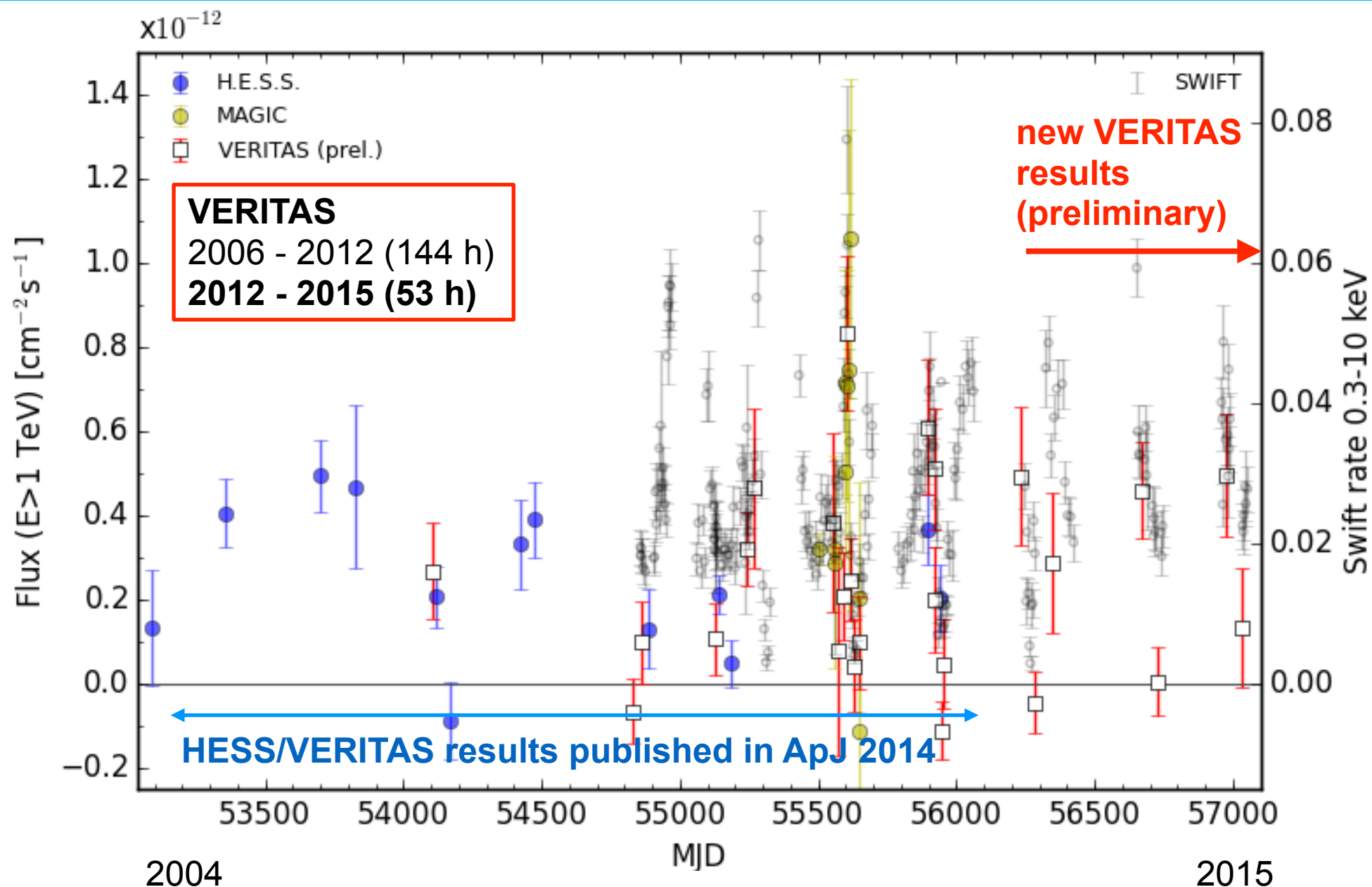


HESS J0632+057

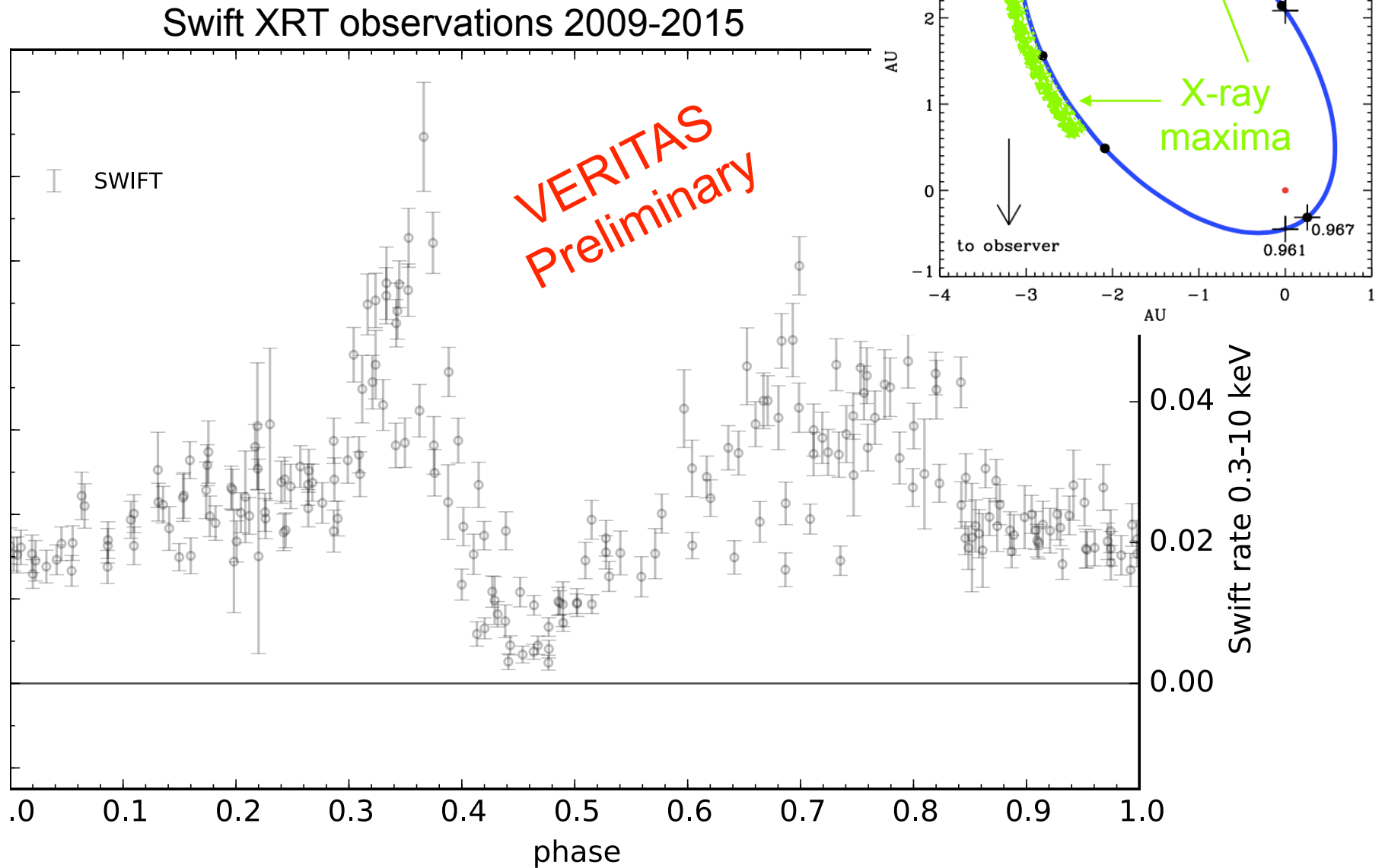
- Be star (MWC 148; $16 M_{\odot}$) + compact object at 1.6 kpc
- Period ~ 315 days
- Discovered in Gamma-rays
 - serendipitous discovery by HESS ($\sim 3\%$ Crab Nebula flux)
 - variability detected by VERITAS
 - observations by HESS, MAGIC and VERITAS (2004-2015)
- binary nature shown with Swift XRT observations
- orbital parameters determined by X-ray and optical observations



HESS J0632+057/VER J0633+057: X-ray and TeV



HESS J0632+057: X-ray light curve



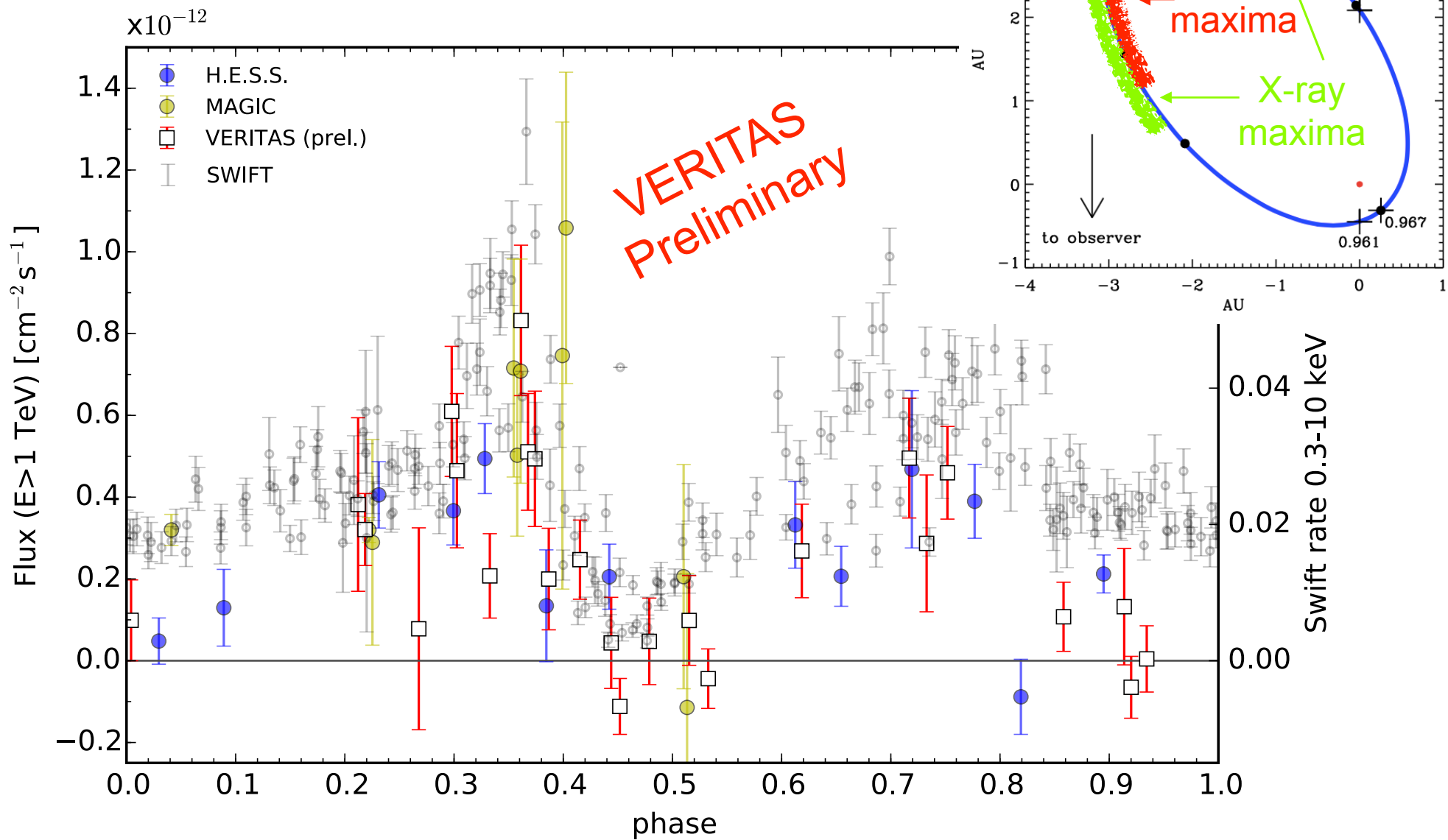
(315 d orbital phase)

Orbital parameters from Bongiorno
et al (2011), Casares et al 2012
(Fig from Dubus 2013),
Aliu et al 2014



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HESS J0632+057: TeV light curve



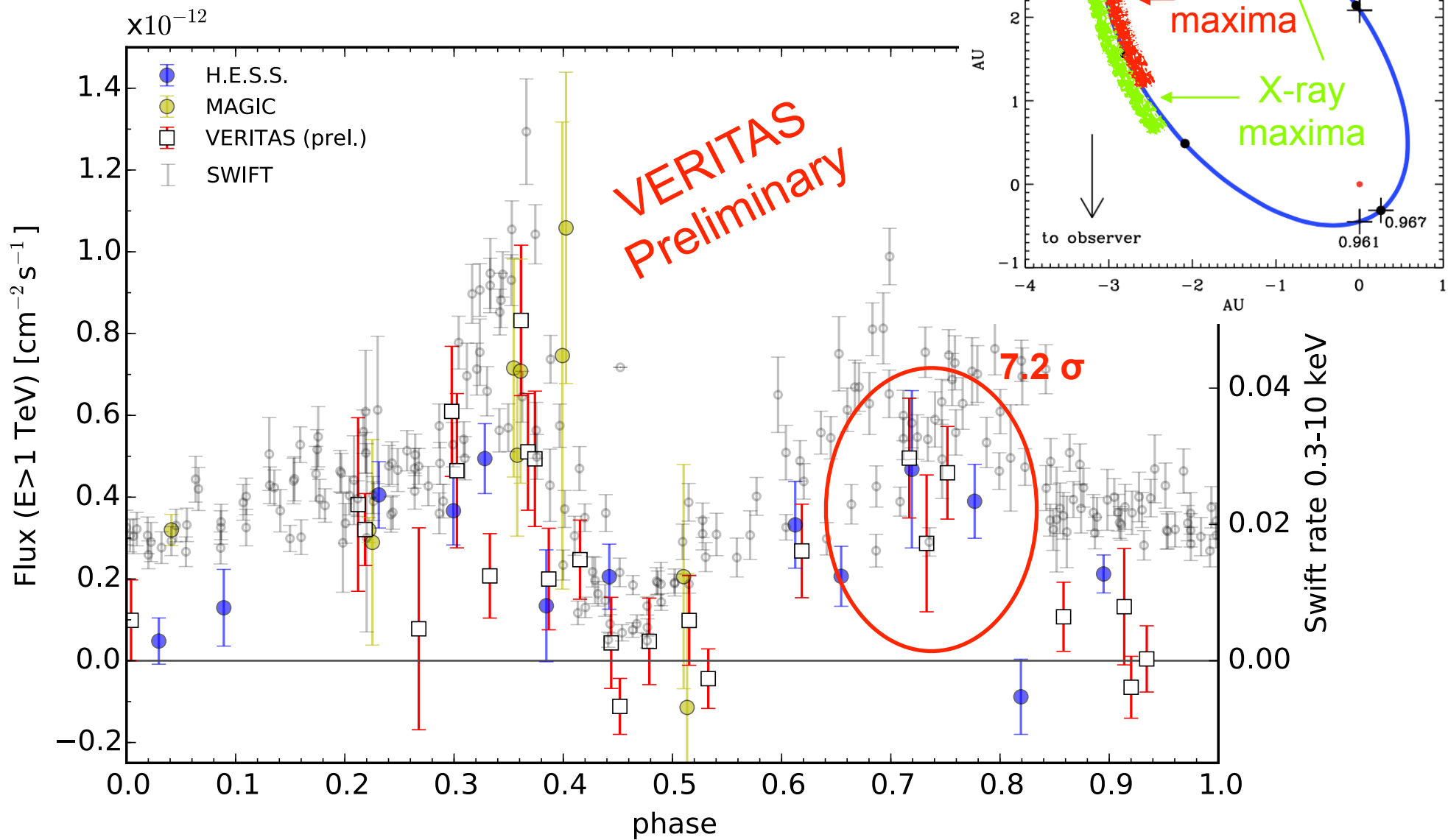
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HESS J0632+057: TeV light curve

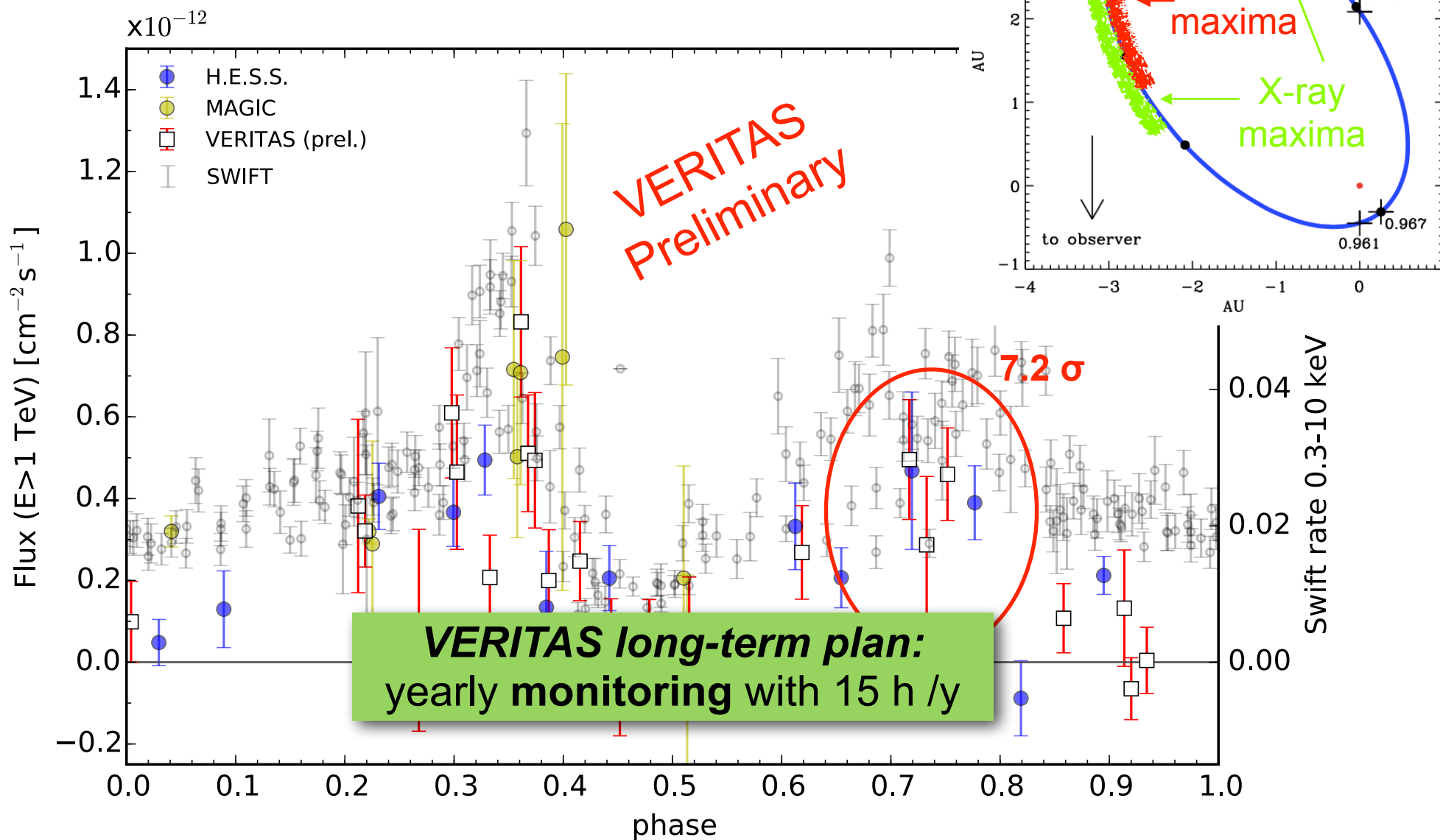


(315 d orbital phase)

Orbital parameters from Bongiorno et al (2011), Casares et al 2012 (Fig from Dubus 2013), Aliu et al 2014



HESS J0632+057: TeV light curve



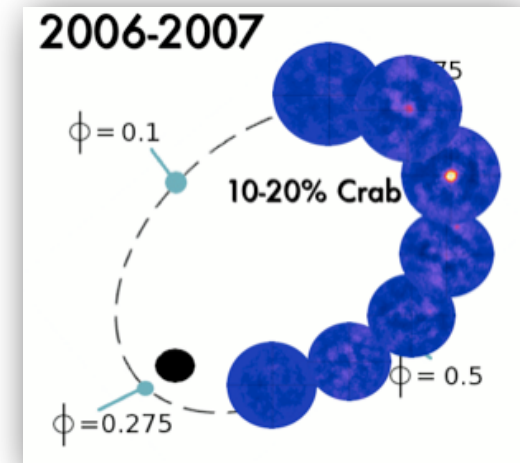
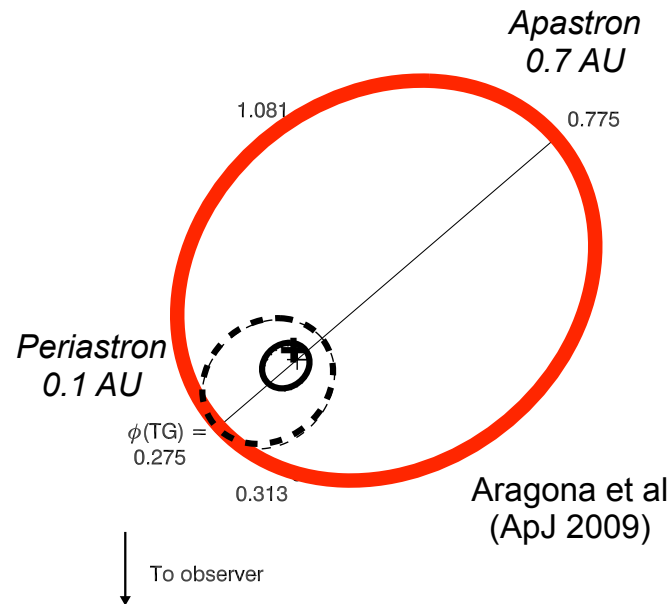
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Orbital parameters from Bongiorno et al (2011), Casares et al 2012 (Fig from Dubus 2013), Aliu et al 2014



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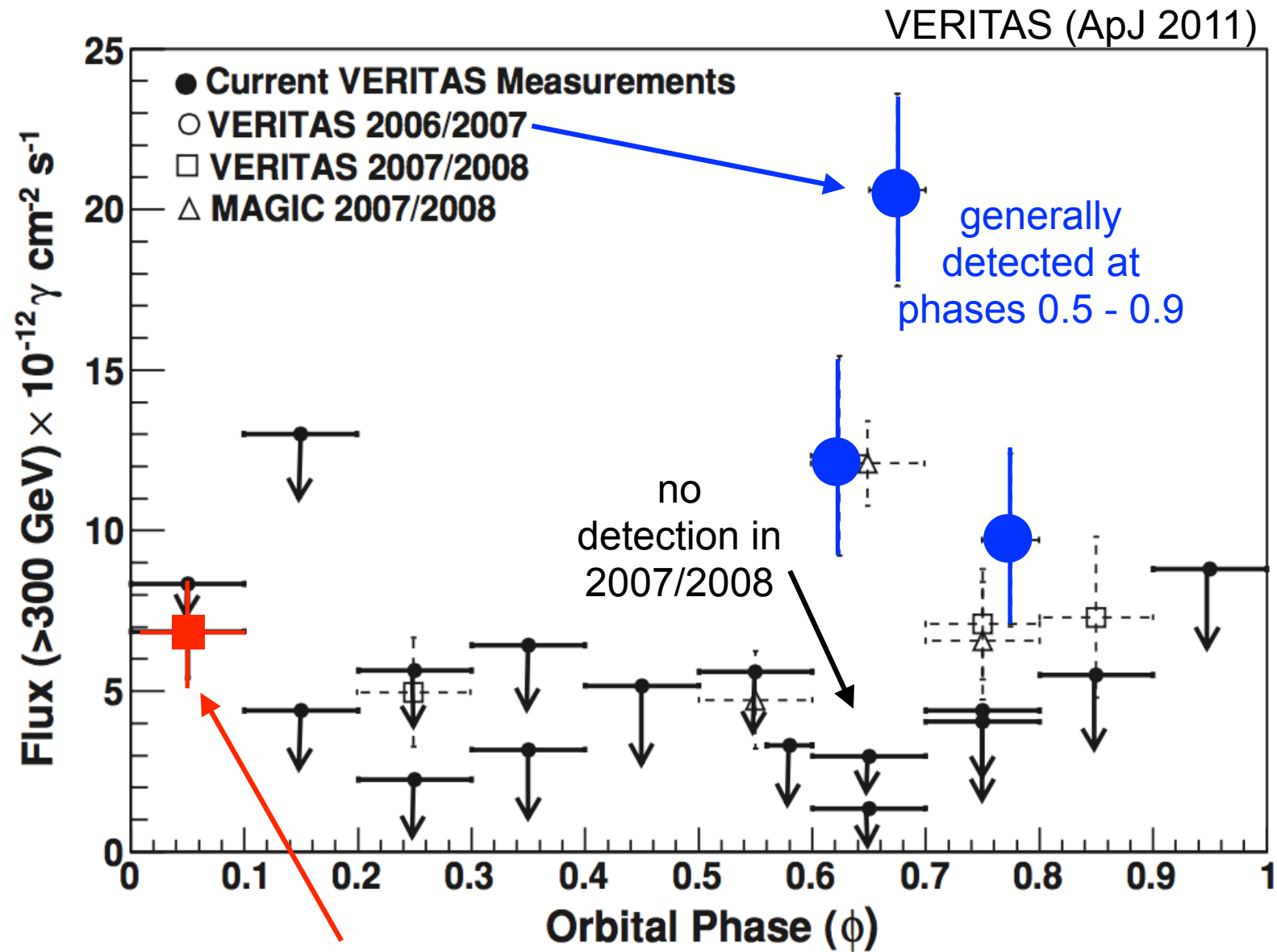
- > Be star + compact object at 1.6 kpc
- > 26.5 day orbit; unknown inclination
- > 1667 day super-orbital modulation



- high X-ray activity
orbital, super-orbital variability & magnetar-like flares (e.g. Smith et al 2009)
- extended radio emission throughout orbit (Dhwawan 2006)
- GeV flux modulated by orbital and super-orbital period (LAT 2013)
- indications for X-ray / VHE correlation during one orbit (MAGIC 2009),
not confirmed by using multiple orbits (VERITAS 2011)

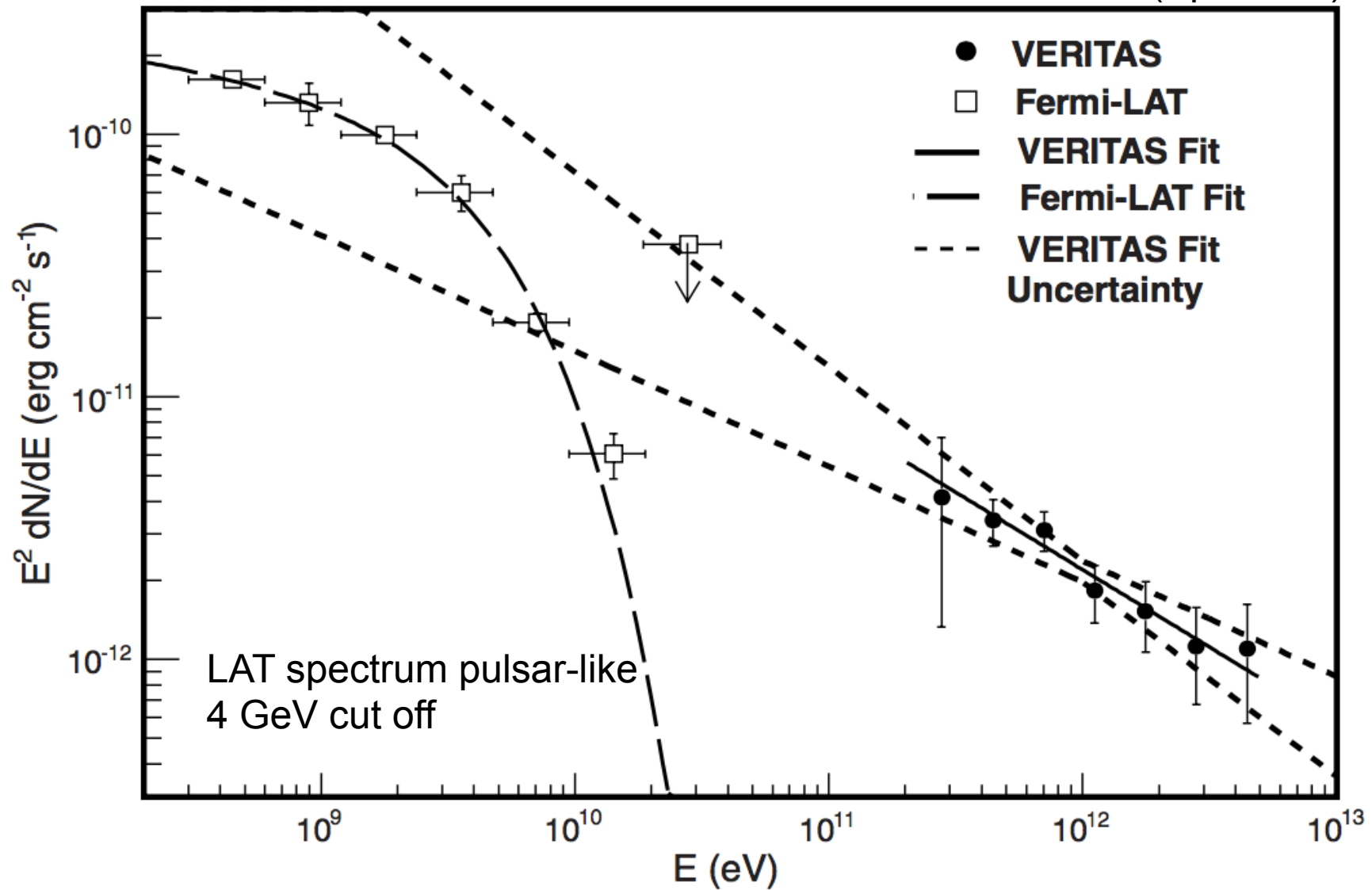
>180 h of VERITAS observations (2006-2015)

LS I +61 303 in gamma rays

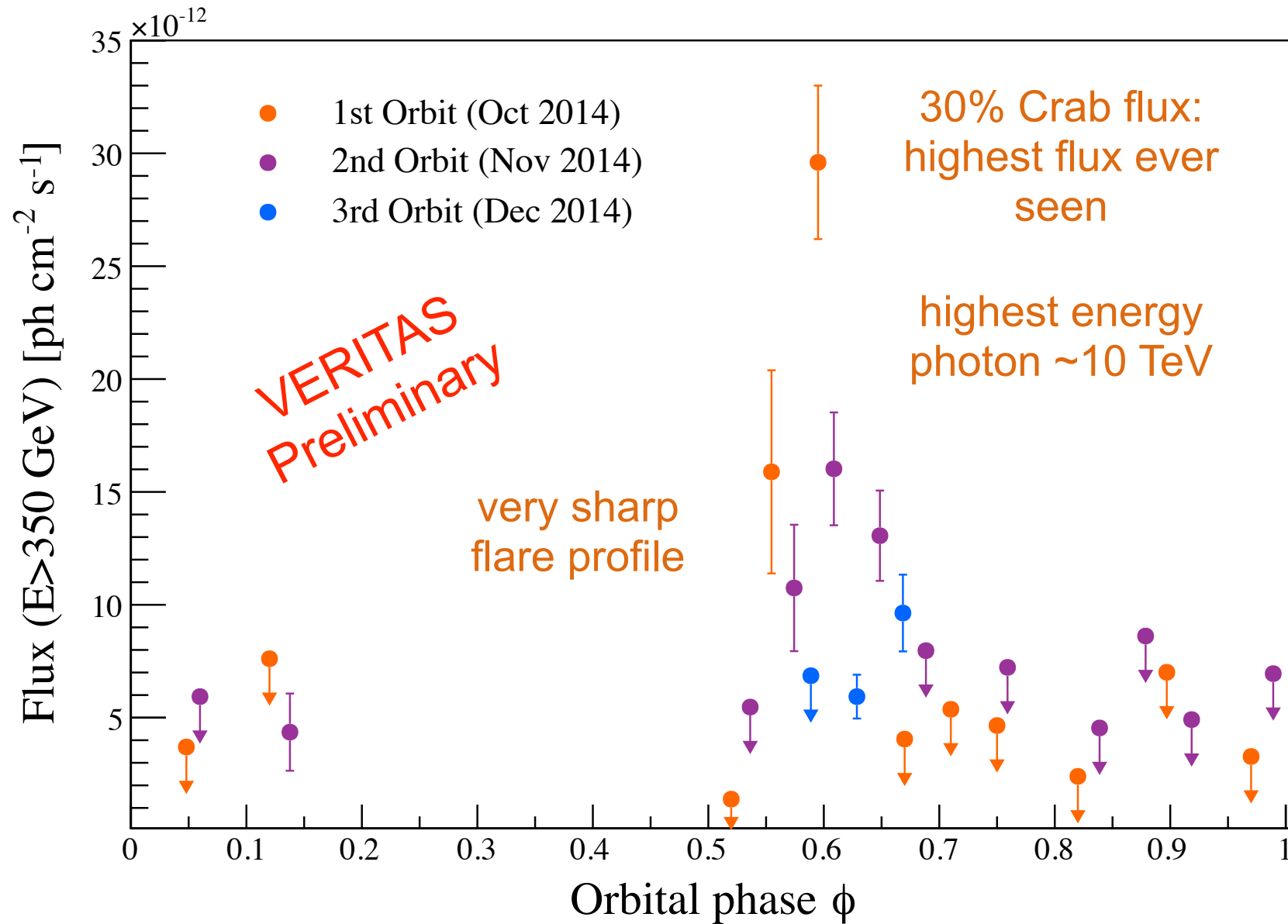


LS I +61 303 in gamma rays

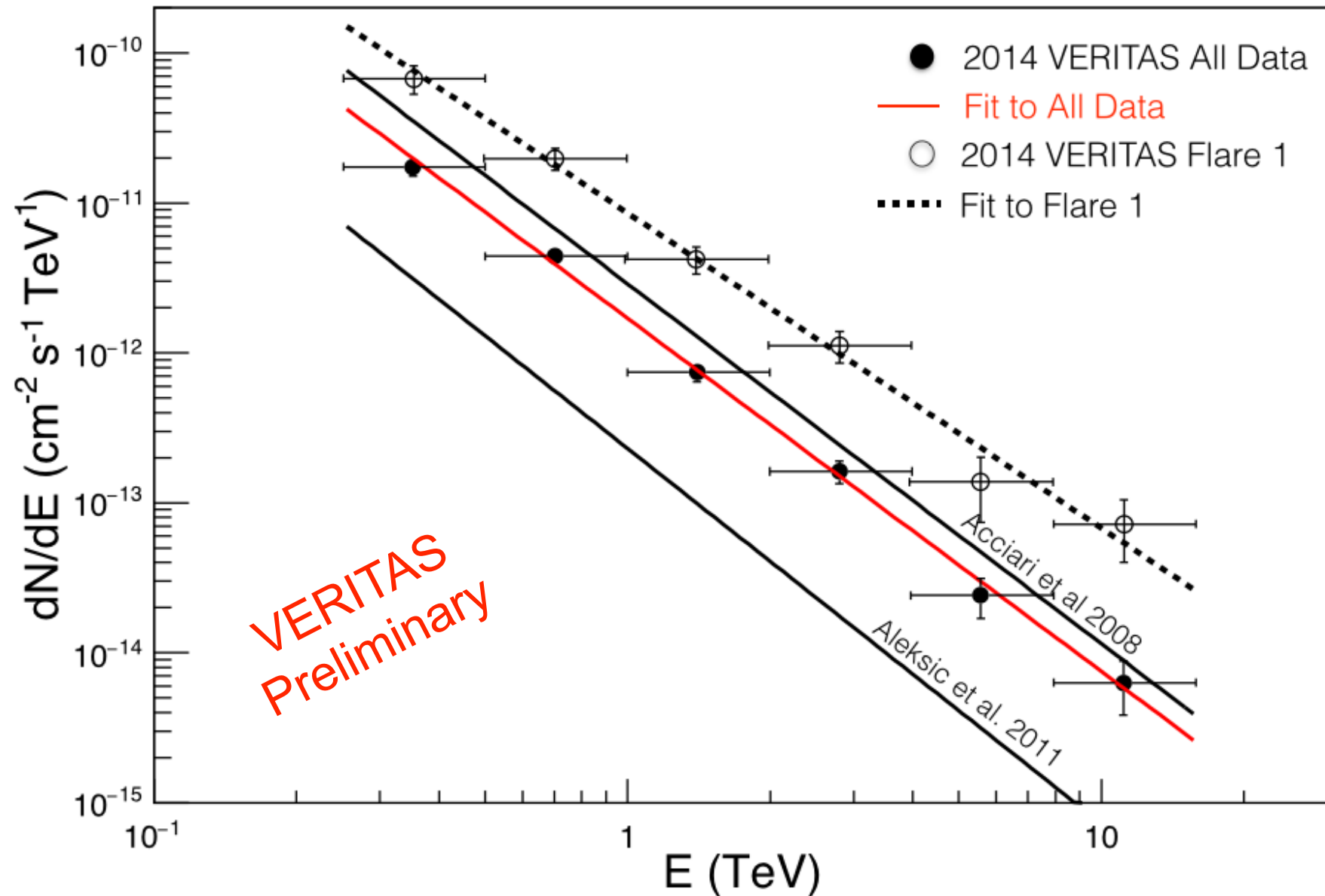
VERITAS (ApJ 2013)



Historically bright flares of LS I +61 303 in 2014

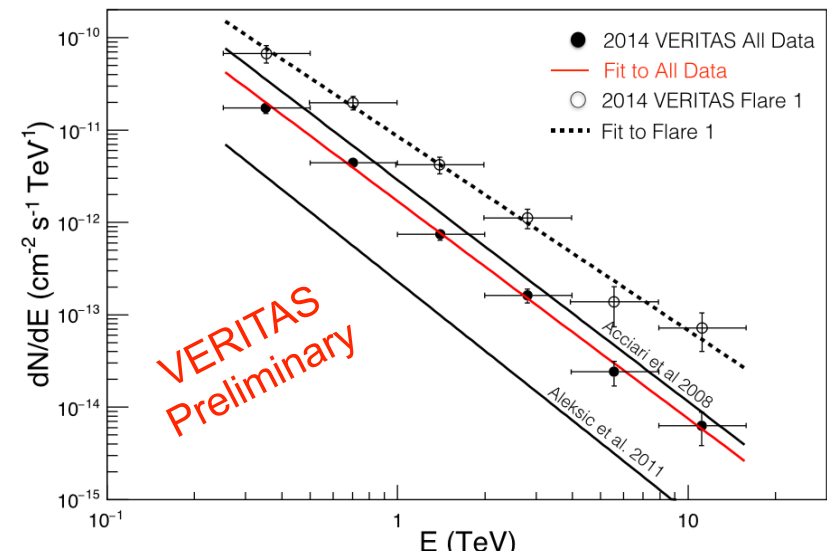
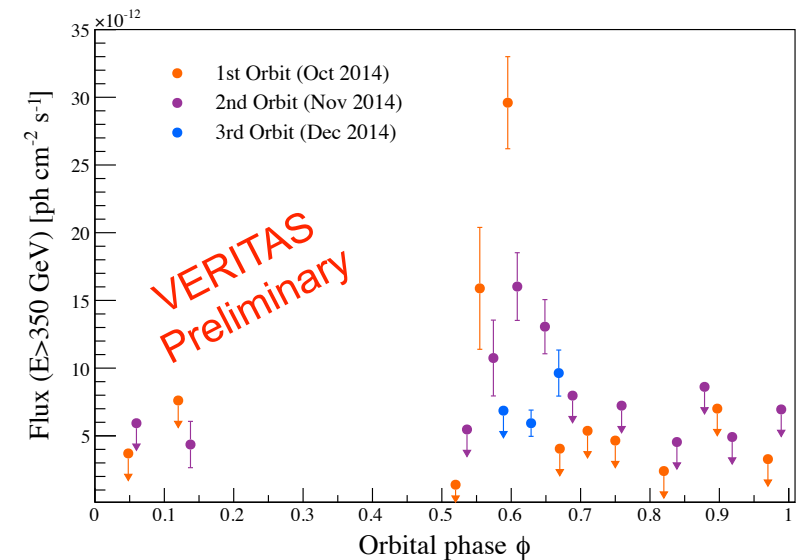


Historically bright flares of LS I +61 303 in 2014



LS I +61 303 in 2014 and future plans

- Atel #6785, publication in preparation
- time-scale for variability <2 days
- multi-wavelength analysis (LAT, Swift XRT, Radio (AMI & RATAN), Ritter Observatory)
- preliminary results of XRT and LAT analysis show no evidence for similar high emission
- plans for future observations
 - yearly monitoring (12 h/y) is part of the VERITAS long-term observing plan
 - coordinated observations with MAGIC (started in Nov 2014)



Summary of VERITAS Binary observations

- > observations of PSR J1023+0038 in LMXB and MSP phase
- > long-term monitoring of HESS J0632+057
 - clear detection of second peak (phases 0.7-0.8) with 7.2σ
- > complex variability in LS I +61 303
 - bright flare of LS I +61 303 in Autumn 2014 with day-scale variability
- > VERITAS binary discovery program
 - systematic search for gamma-ray emission from O/Be-X-ray binary systems
 - filler observing program

