Quo vadis, neutrino flavor models

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The flavor problem



origin of neutrino mass?

why 3 generations?

The flavor problem



origin of neutrino mass?

why 3 generations?

The flavor problem

Mass hierarchy?

Mixing hierarchy?

Is the flavor problem the evidence of some underlying symmetry?

origin of neutrino mass?

why 3 generations?



triplet irreducible representation of a flavor group?

Lepton mixing: status

M, proceeding NOW12

large solar large atmospheric small reactor



Is the atmospheric angle maximal?

Reactor angle not zero!

Is the atmospheric angle maximal?



With abelian flavor symmetry difficult to get maximal mixings

Strong indications in favor of non abelian flavor symmetry everything is possible: there are no particular indication on the flavor symmetry...

No

maximal mixing & mu-tau symmetry

$$\sin\theta_{23} = \frac{1}{\sqrt{2}}$$
$$\sin\theta_{13} = 0$$



exchange symmetry In the neutrino sector



the second second

Babu, Ma, Valle 02 Grimus, Lavoura, 04 Altarelli, Feruglio 05 Caravaglios, M, 05

 $\mu \leftrightarrow \tau$

exchange symmetry broken in the charged sector

.....

maximal mixing & mu-tau symmetry



charged sector

maximal atmospheric mixing origin of mu-tau neutrino symmetry



Non-abelian symmetry Symmetry of tetrahedron Triplet representation Finite subgroup of SU(3)

$\mu \leftrightarrow \tau$

Broken in the charged lepton sector: but **Z3** preserved

 $u_{\mathrm{u}} \leftrightarrow \nu_{\tau}$

In the neutrino sector

maximal atmospheric mixing and reactor angle

$$\sin heta_{_{13}}\simeq\lambda_{_c}
eq 0$$

M, Patel, Peinado 11 Boucenna, M, Tortola, Valle 11 King, 12 King, Luhn 12

We expect deviations of the $\mathcal{V}_{\mu} \leftrightarrow \mathcal{V}_{\tau}$

exchange symmetry in the neutrino sector

What if the atmospheric angle is not maximal?

Anarchy: neutrino mass matrix has random entries

 $M = m \begin{pmatrix} 0(1) & 0(1) & 0(1) \\ 0(1) & 0(1) & 0(1) \\ 0(1) & 0(1) & 0(1) \end{pmatrix}$

Hall, Murayama, Weiner 99

implies mixing and ratio of masses of order one in rough agreement with data discrete non abelian symmetries so far are not excluded

connection between atmospheric and reactor angles, i.g.

$$\sin\theta_{23} = \frac{1}{\sqrt{2}} + \sin\theta_{13}\cos\delta$$

mixing sum rule

Ballet, King, Luhn, Pascoli, Schmidt 13

Discrete symmetries & Mass sum rules

$$lpha m_{1} + eta m_{2} = m_{3} \ rac{lpha}{m_{1}} + rac{eta}{m_{2}} = rac{1}{m_{3}} \ lpha \sqrt{m_{1}} + eta \sqrt{m_{2}} = rac{1}{m_{3}} \ rac{lpha}{\sqrt{m_{1}}} + rac{eta}{\sqrt{m_{2}}} = rac{1}{\sqrt{m_{3}}} \ rac{lpha}{\sqrt{m_{1}}} + rac{eta}{\sqrt{m_{2}}} = rac{1}{\sqrt{m_{3}}} \ rac{1}{\sqrt{m_{1}}} + rac{eta}{\sqrt{m_{2}}} = rac{1}{\sqrt{m_{3}}} \ rac{1}{\sqrt{m_{3}}} \ rac{1}{\sqrt{m_{3}}} = rac{1}{\sqrt{m_{3}}} \ rac{1}{\sqrt{m_{3}}} \ rac{1}{\sqrt{m_{3}}} \ rac{1}{\sqrt{m_{3}}} = rac{1}{\sqrt{m_{3}}} \ rac$$

Altarelli, Feruglio, Hagedorn 08 Hirsch, M, Valle 08 Bazzocchi, M, Merlo 09 Altarelli, Meloni 09 Barry, Rodejohann 10/11 Dorame, Meloni, M, Peinado, Valle11

See talk of Merle

Mass sum rules & Onubb

King, Merle, M, Shimizu, Tanimoto 14

See talk of Merle



Flavor symmetries: multi-Higgs vs flavon

 $y_{ijk} \overline{L}_i l_{Rj} H_k = \frac{1}{\Lambda} y_{ijk} \overline{L}_i l_{Rj} H \varphi_k$

Flavor symmetries: multi-Higgs vs flavon

 $y_{ijk} \overline{L}_i l_{Rj} H_k = \frac{1}{\Lambda} y_{ijk} \overline{L}_i l_{Rj} H \varphi_k$



No signal at LHC,...

Flavor symmetries: multi-Higgs vs flavon

 $y_{_{i\,ik}}L_{_i}l_{_{R_j}}H_{_k}$

 $\frac{1}{\Lambda}y_{ijk}\overline{L}_{i}l_{Rj}H\varphi_{k}$

New Higgs at the weak scale FCNC, LFV,... $H \rightarrow ee, H \rightarrow \mu\mu,...$

Flavor symmetry gives relations between branching ratios

See talk of Vicente

Higgs decay and flavor symmetry, i.g.



Based on A4: it is distinguishable and falsifiable

See also: triality, Ma PRD (10) Higgs decay in S3, Bhattacharyya et al PRD (12)

$$R_K = \frac{\text{BR}(B \to K\mu^+\mu^-)}{\text{BR}(B \to Ke^+e^-)} = 0.745^{+0.090}_{-0.074} \pm 0.036$$

Only 2,6 sigma

LHCb collaboration PRL (14)

R = 1 in the SM



$$R_K = \frac{\text{BR}(B \to K\mu^+\mu^-)}{\text{BR}(B \to Ke^+e^-)} = 0.745^{+0.090}_{-0.074} \pm 0.036$$

Only 2,6 sigma

LHCb collaboration PRL (14)

• Minimal Flavor Violation:

Alonso et al 1505.05164 Lee, Tandeon 1505.04692

Celis et al1505.03079 Crivellin et al 1504.07928

Neutrino oscillation

• 7'

Boucenna et al 1503.07099

Leptoquark

Hiller, Schmaltz JHEP (15)



$$\mathcal{L} = -\lambda_{q\ell} \,\Delta \left(\bar{q} P_L \ell \right)$$

$$\mathcal{L} = -\lambda_{q\ell} \,\Delta^* \left(\bar{q}\ell \right)$$

SU(2) doublet

SU(2) triplet

Non-Abelian flavor symmetry

$$\lambda \equiv \begin{pmatrix} \lambda_{de} & \lambda_{d\mu} & \lambda_{d\tau} \\ \lambda_{se} & \lambda_{s\mu} & \lambda_{s\tau} \\ \lambda_{be} & \lambda_{b\mu} & \lambda_{b\tau} \end{pmatrix} \sim \lambda_0 \begin{pmatrix} \rho_d \kappa & \rho_d & \rho_d \\ \rho \kappa & \rho & \rho \\ \kappa & 1 & 1 \end{pmatrix}$$

De Medeiros, Hiller 1503.01084

$$\begin{aligned} \mathcal{B}(B \to K \mu^{\pm} e^{\mp}) &\simeq 3 \cdot 10^{-8} \,\kappa^2 \left(\frac{1 - R_K}{0.23}\right)^2 \,, \\ \mathcal{B}(B \to K e^{\pm} \tau^{\mp}) &\simeq 2 \cdot 10^{-8} \,\kappa^2 \left(\frac{1 - R_K}{0.23}\right)^2 \,, \\ \mathcal{B}(B \to K \mu^{\pm} \tau^{\mp}) &\simeq 2 \cdot 10^{-8} \left(\frac{1 - R_K}{0.23}\right)^2 \,, \end{aligned}$$

$$\begin{split} \mathcal{B}(\mu \to e\gamma) &\simeq 2 \cdot 10^{-12} \, \frac{\kappa^2}{\rho^2} \left(\frac{1 - R_K}{0.23} \right)^2 \,, \\ \mathcal{B}(\tau \to e\gamma) &\simeq 4 \cdot 10^{-14} \, \frac{\kappa^2}{\rho^2} \left(\frac{1 - R_K}{0.23} \right)^2 \,, \\ \mathcal{B}(\tau \to \mu\gamma) &\simeq 3 \cdot 10^{-14} \, \frac{1}{\rho^2} \left(\frac{1 - R_K}{0.23} \right)^2 \,, \\ \mathcal{B}(\tau \to \mu\eta) &\simeq 4 \cdot 10^{-11} \, \rho^2 \left(\frac{1 - R_K}{0.23} \right)^2 \,. \end{split}$$

Flavor symmetries: from neutrino to quarks

While large neutrino mixing seems to give indication of non-abelian family symmetry

Typically the neutrino flavor symmetries do not naturally fit the quark sector

Non abelian symmetries must be introduced to fit also the quark sector



Flavor symmetries: from neutrino to quarks







Summary

The future direction of flavor model building will strongly depend on experimental results:

atmospheric mixing, leptonic CP violation, mass hierarchy,.....

If the atmospheric angle is VERY close to be maximal, evidence of a neutrino mu-tau symmetry

If the atmospheric angle is NOT maximal, even anarchy is possible

BUT

flavor symmetries imply testable mixing & mass sum rules

Summary

In the multi-Higgs case, hopefully LHC will give us some indication

We can also have indications from Flavor Physics like B decays

From the theoretical point of view, a possible criteria to proceed is to link very different topics and problems like dark matter and neutrino and so on....

Tri-bi-maximal mixing: theoretical ansatz



maximal atmospheric Angle