


# The Fermilab Short-Baseline Neutrino (SBN) Program



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# The SBN Program and Heidelberg



- **Built up over time in different phases**
- **Utilized by many different users**
- **Long history of physics**



# Where I hope we ultimately differ

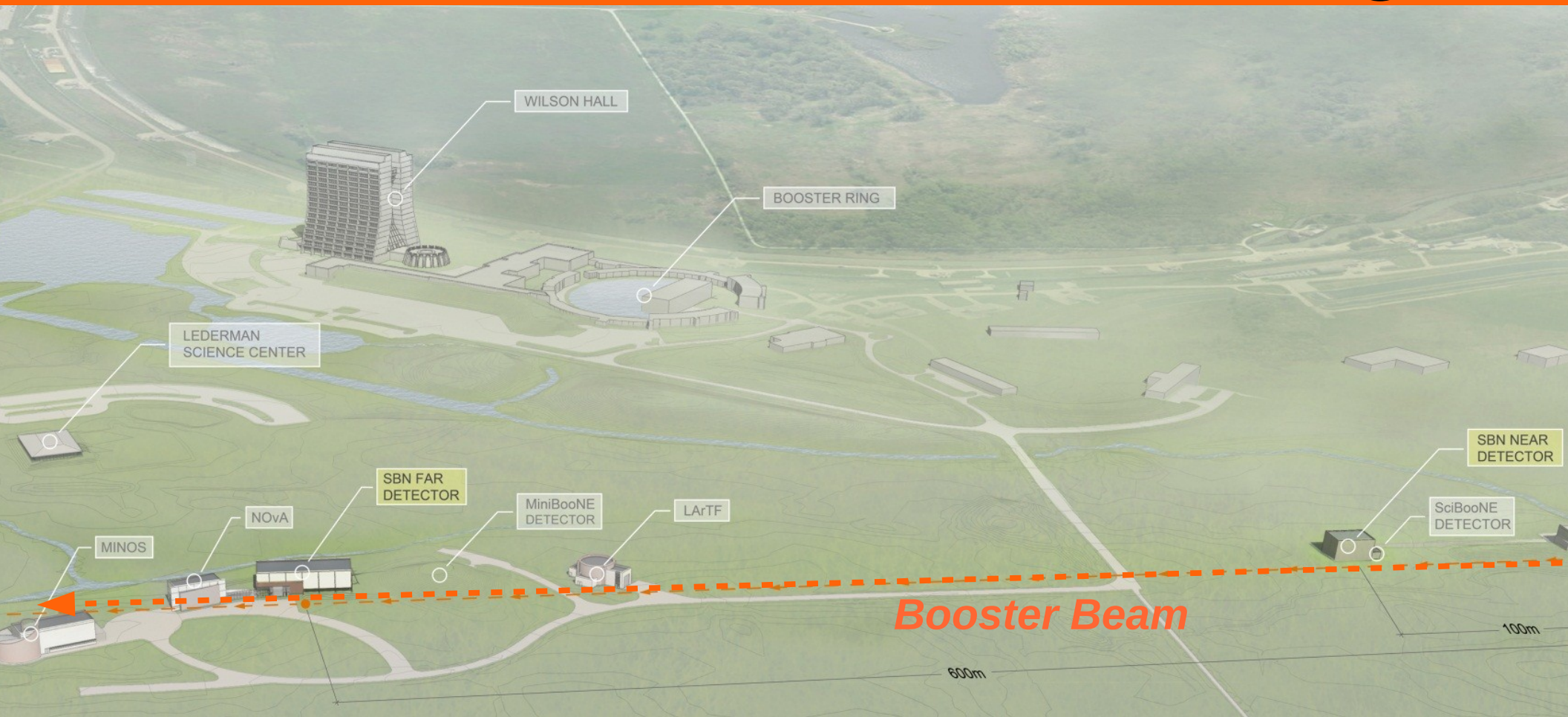


- Hopefully SBN will be destroyed fewer times by the French
- Hopefully SBN will be struck by lightning fewer times
- Hopefully more neutrino interactions will be seen at the SBN





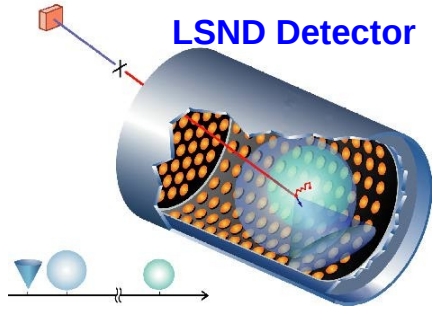
# The Short-Baseline Neutrino Program



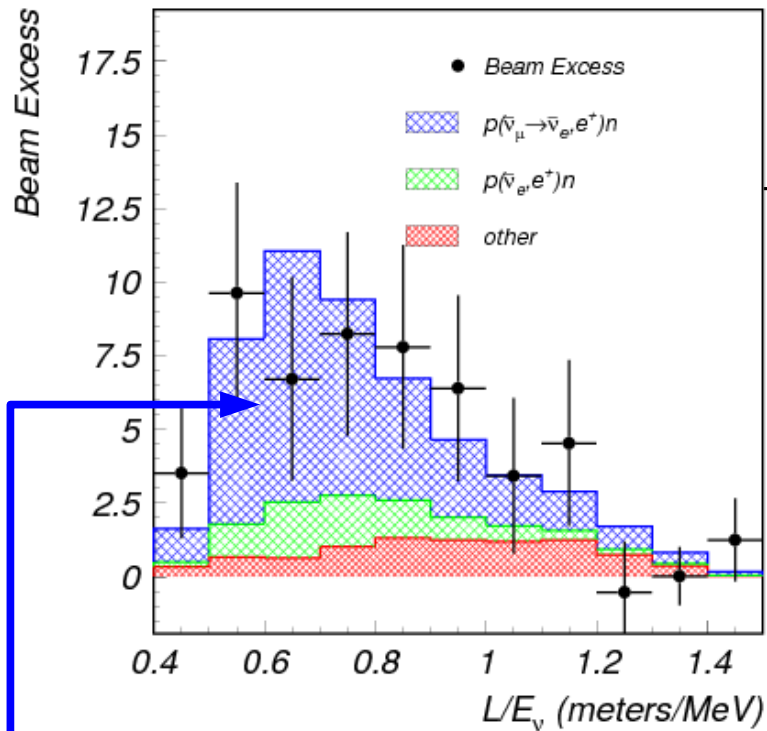
The story of the Short-Baseline Neutrino Program can best be understood through the history of the physics that we've been following



# Hints of new physics?

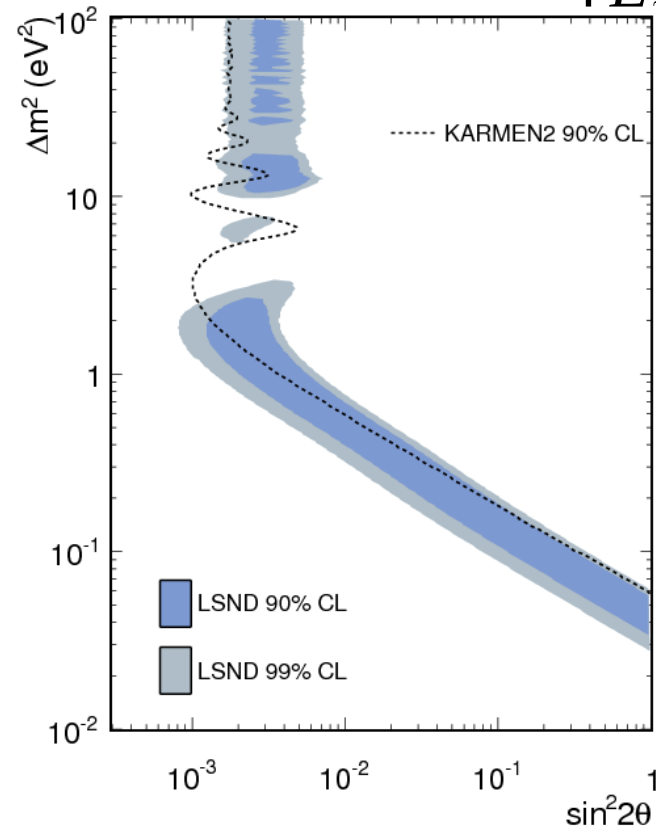


- LSND ran in a low energy  $\bar{\nu}_\mu$  beam from a decay-at-rest source
- Detected a  $3.8\sigma$  excess in the appearance of  $\bar{\nu}_e$ 
  - The result was interpreted from within the neutrino oscillation model as an additional mixing



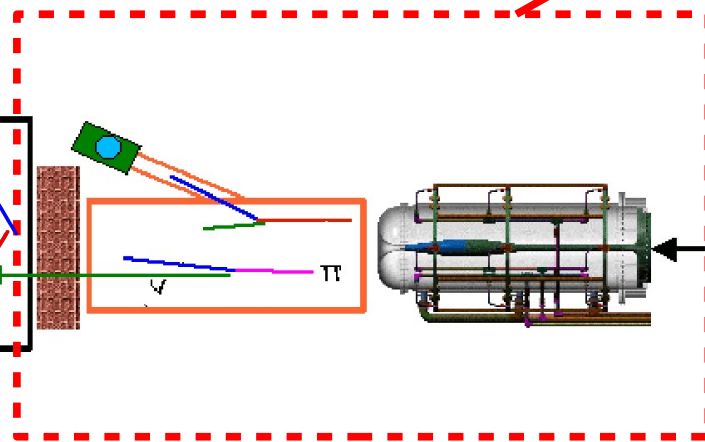
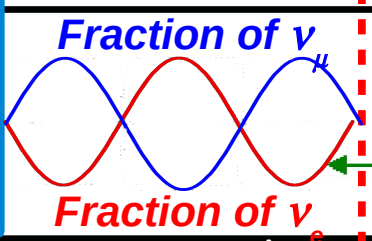
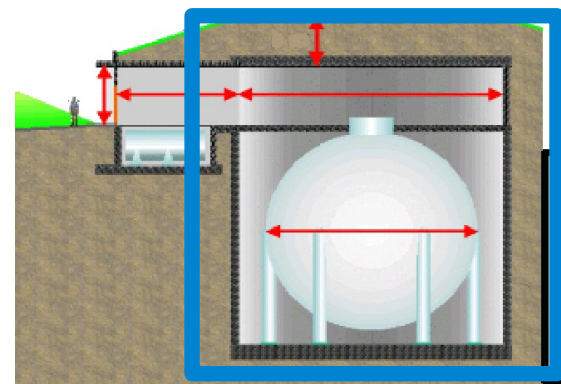
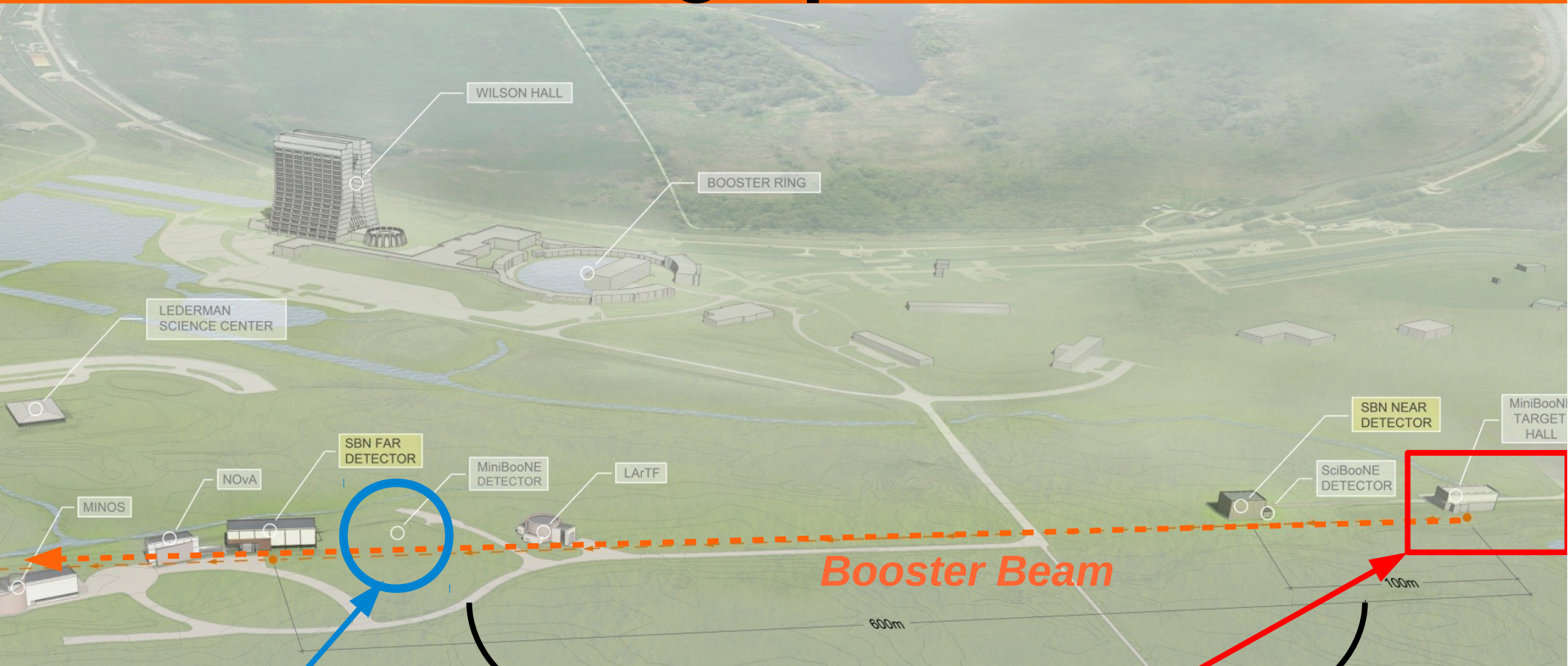
**Could this excess be evidence for sterile neutrinos?**

$$P(\nu_\beta \rightarrow \nu_\alpha) = \sin^2(2\theta) \sin^2\left(\frac{\Delta m^2 L}{4E_\nu}\right)$$



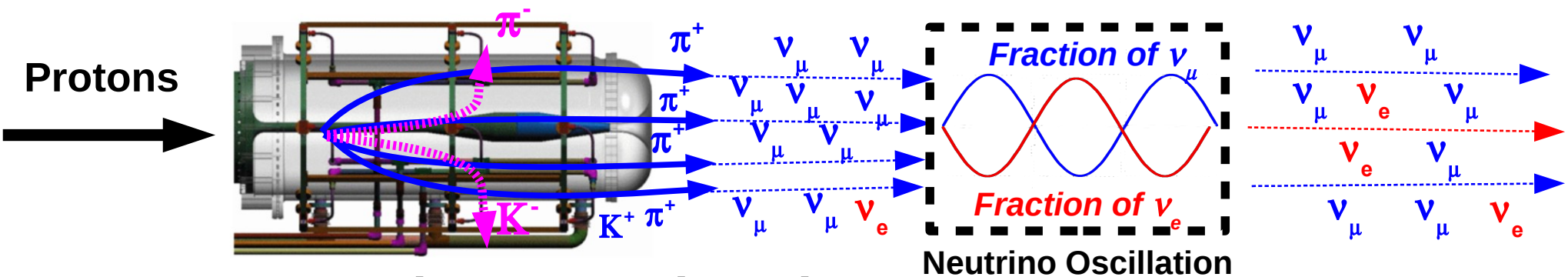


# Following up on LSND



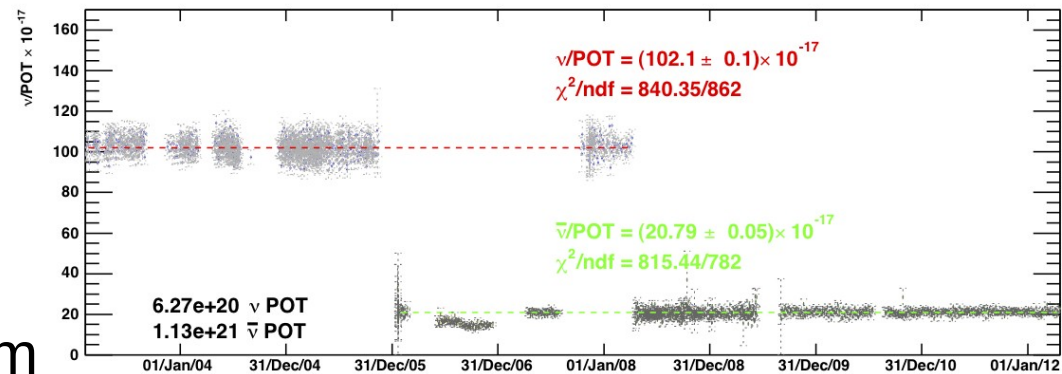


# Booster Neutrino Beam



- **Booster Neutrino Beam (BNB)** has been operating for a decade!

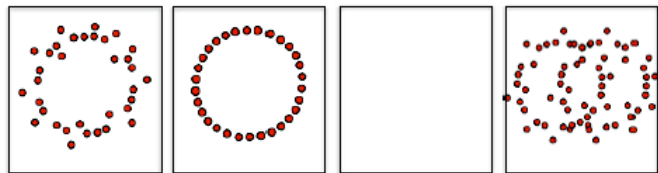
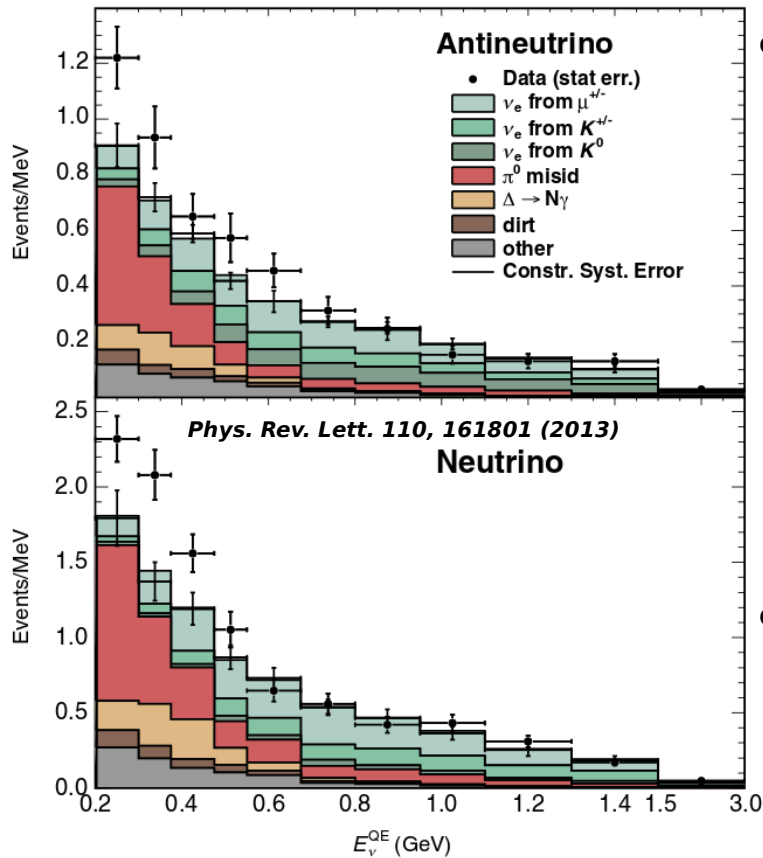
- A very well understood and characterized beam
- Low ( $< 0.5\%$ ) contamination from intrinsic  $\nu_e$



- **Neutrino beam created from 8 GeV protons colliding on a beryllium target and having sign selected pions focused by a magnetic horn**



# More hints of new physics?



Electron,  
Photon

Muon

Proton

$\pi^0 \rightarrow \gamma + \gamma$

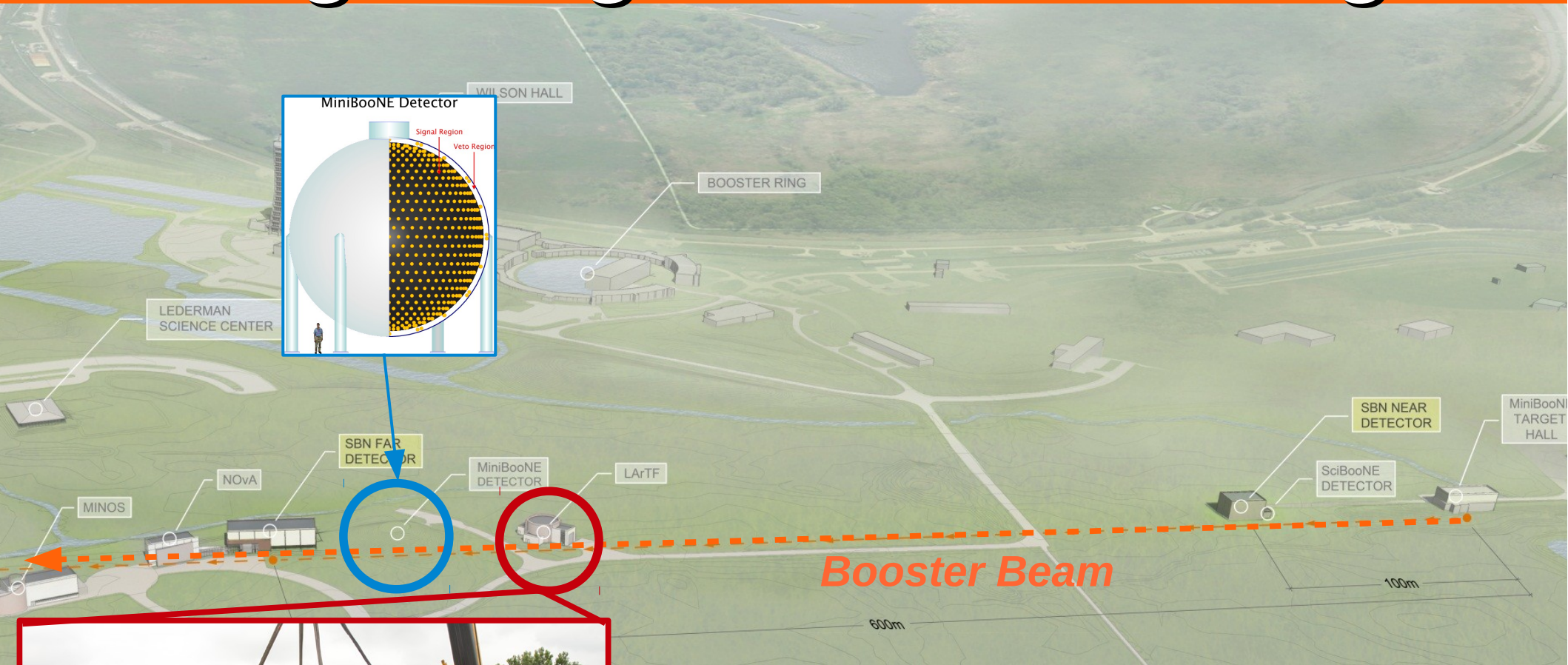
- MiniBooNE ran at a similar L/E (utilizing decay in flight neutrino source) and sees a low energy excess in  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  and  $\nu_\mu \rightarrow \nu_e$  appearance search
  - Excess can also be interpreted as an additional mixing
- However, MiniBooNE (Cherenkov detector) has a difficult time determining the composition of the excess

- Electron like?
- Photon like?

What you would like is an experiment that **sees the same beam** as MiniBooNE, at (nearly) the **same distance** as MiniBooNE but with superior **electron/photon separation ability**



# The beginning of the SBN Program

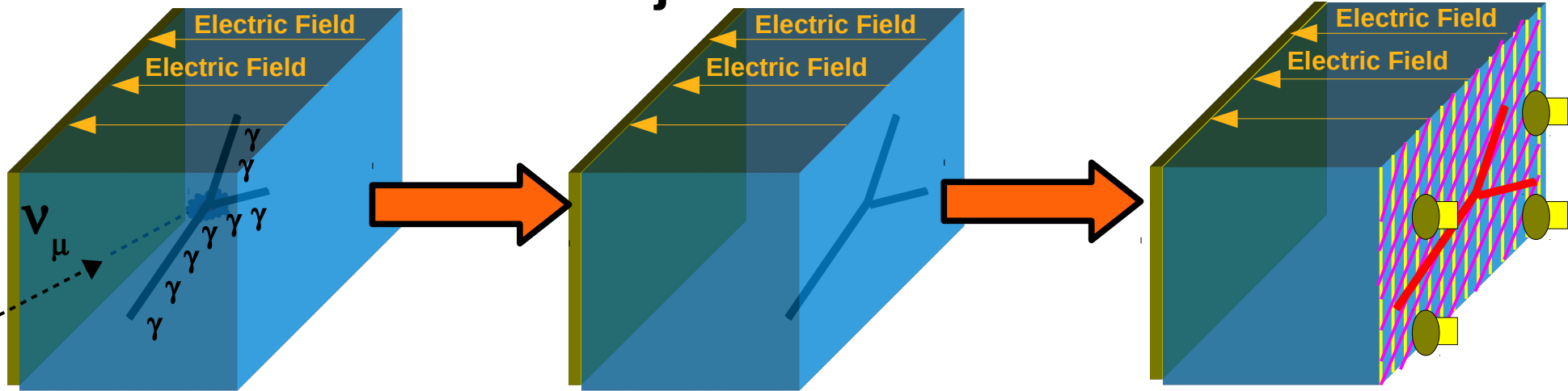


**MicroBooNE is the first LArTPC detector on the short-baseline and kicks off the SBN program**



# LArTPC's

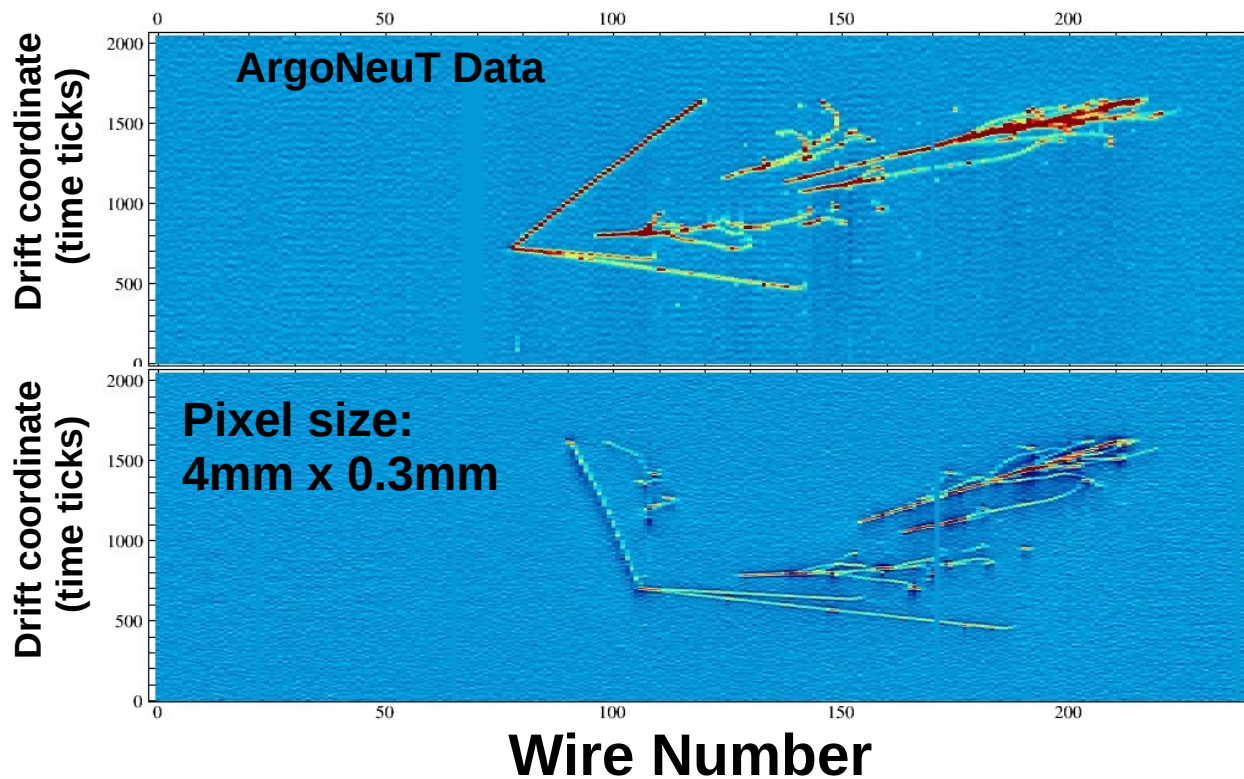
## Time Projection Chamber



Neutrino interaction in LAr produces ionization and scintillation light

Drift the ionization charge in a uniform electric field

Read out charge and light produced using precision wires and PMT's

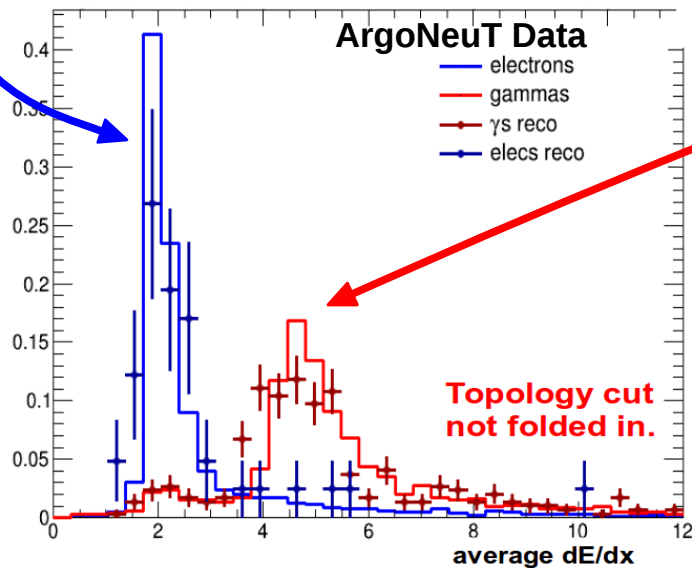
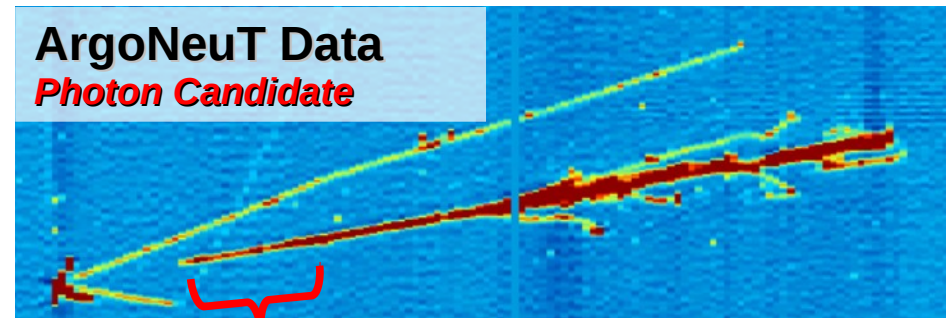
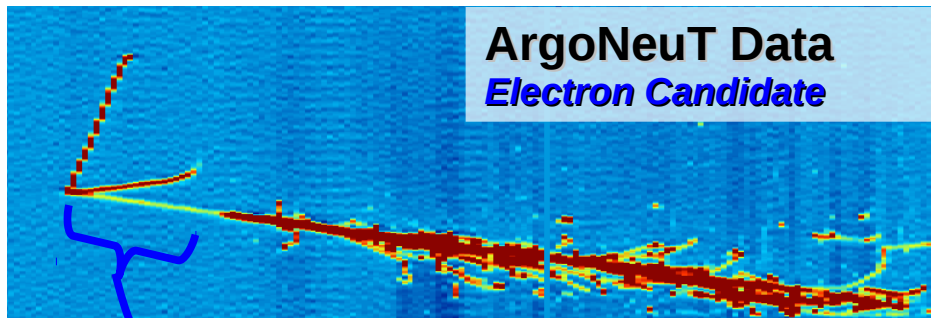


LArTPC's offer incredible fine grain tracking along with electron/photon separation



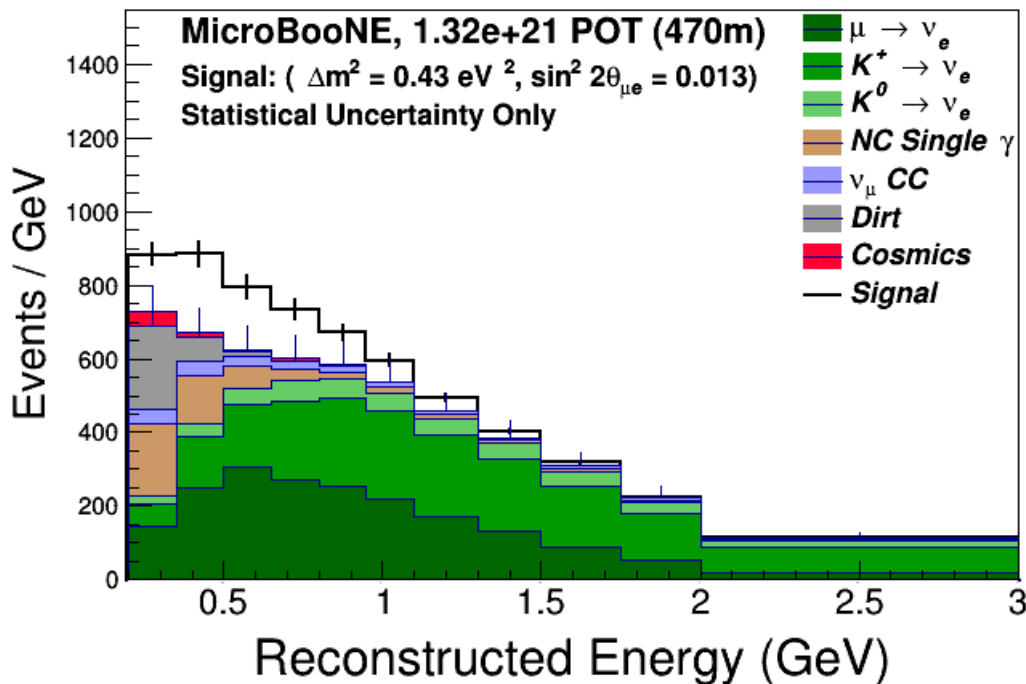
# MicroBooNE

- MicroBooNE will utilize the electron / photon discrimination power of LArTPC's to determine if the MiniBooNE excess is **electron like** (from  $\nu_e$  appearance) or **photon like** (unaccounted for background)



By analyzing the topology and the dE/dX of the electromagnetic shower, disentangling the MiniBooNE low energy excess becomes possible

# MicroBooNE



- MicroBooNE is the largest LArTPC ever built in the U.S.

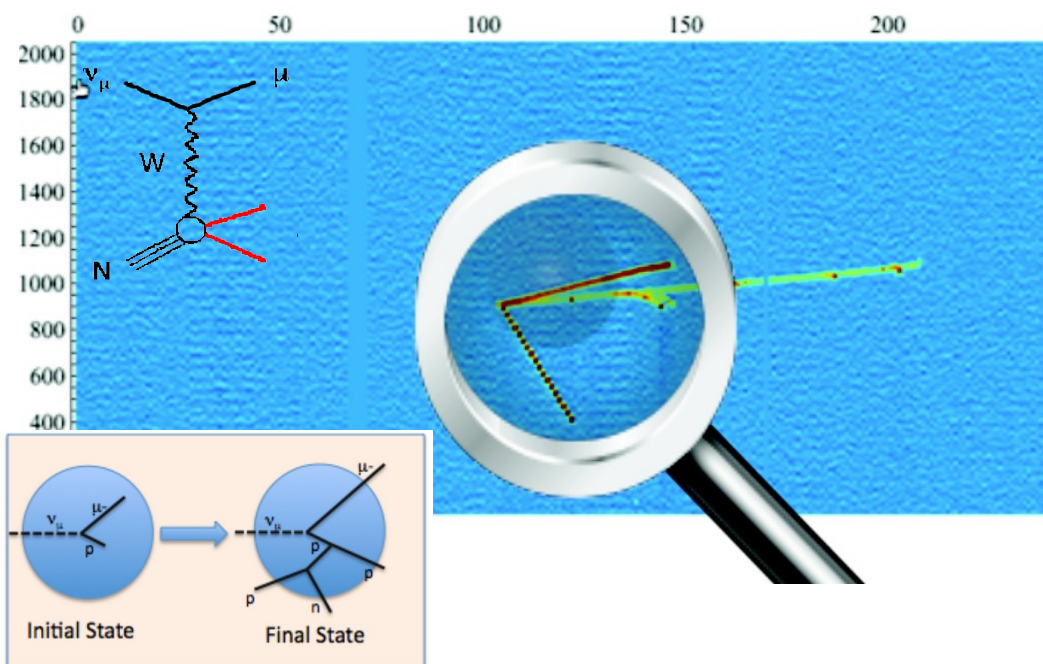
– 89 Tons of active mass



*about this big*

- MicroBooNE also has a rich physics program planned

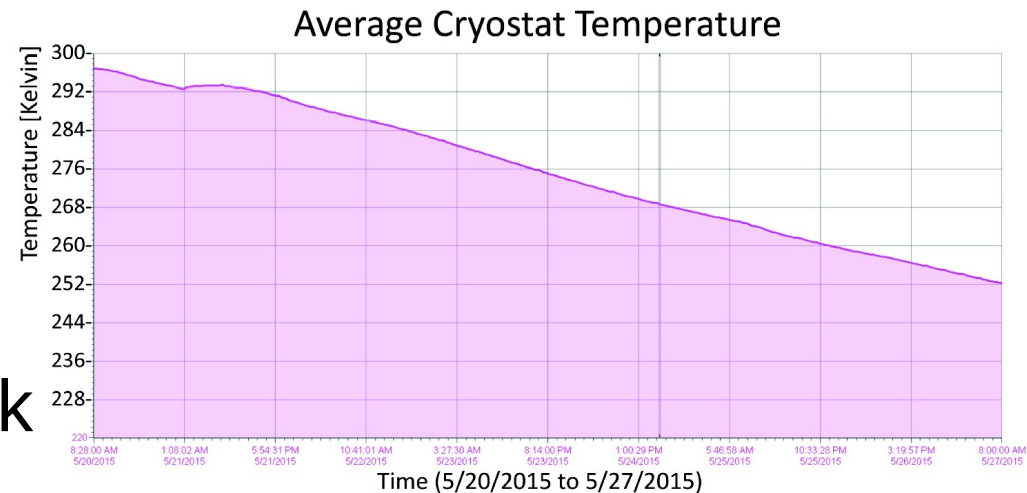
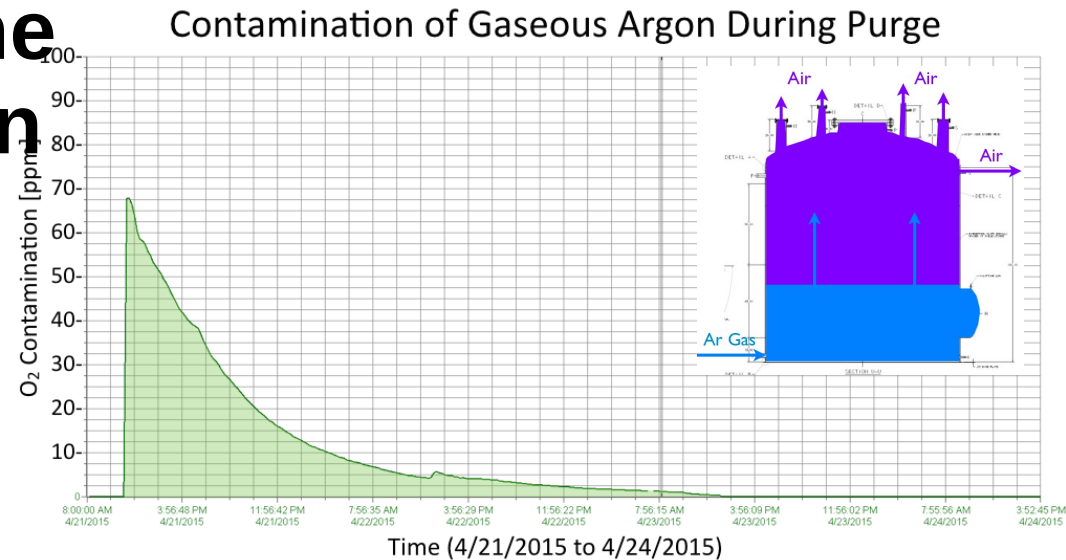
- Determining the nature of the MiniBooNE low energy excess
- Neutrino cross-sections
- Studying nuclear final state interactions
- Exploring the capabilities for LArTPC to look at astroparticle and exotic phenomenon



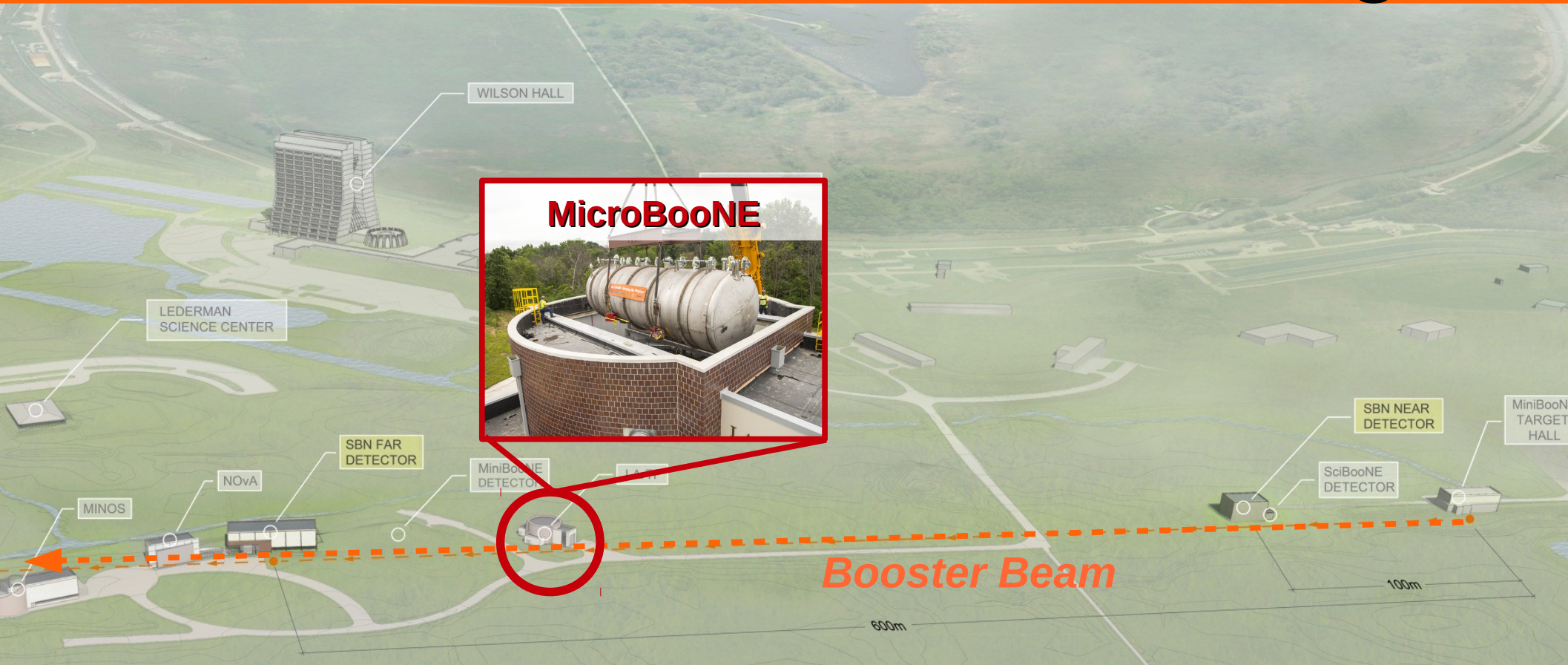


# MicroBooNE: Current Status

- MicroBooNE is fully installed and starting the O<sub>2</sub> purge and cool down process
- 24 hour commissioning shifts have begun
  - Exercising our full readout and monitoring systems
- We expect to be full of LAr and taking data this summer
  - Neutrino data run to start when the beam comes back this fall



# The Short-Baseline Neutrino Program



**What do I need to add to the existing program (top notch neutrino beam + world class neutrino detectors) to make a definitive search eV scale for sterile neutrinos?**

- Normalization of the unoscillated neutrino beam (**Near detector**)
- High statistics in the appearance channel (**large mass far detector**)
- Look for complimentary muon disappearance (**near/far comparison**)

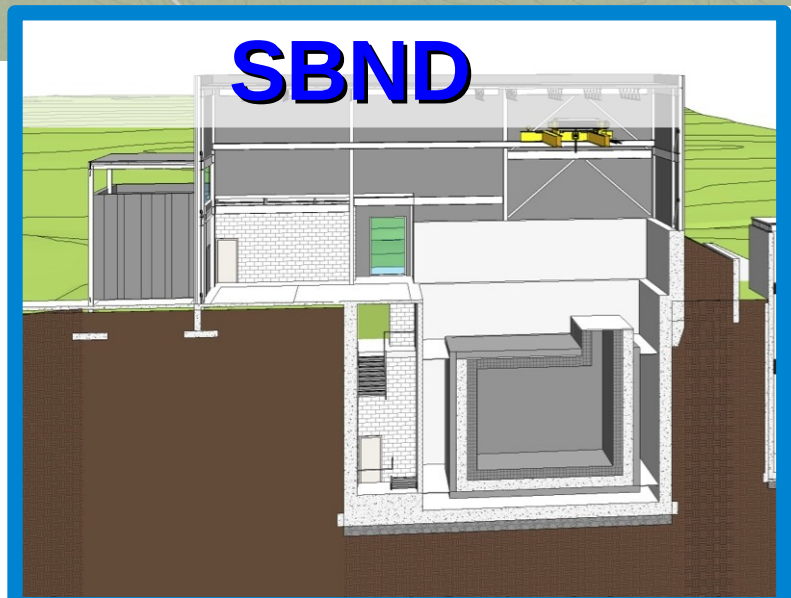
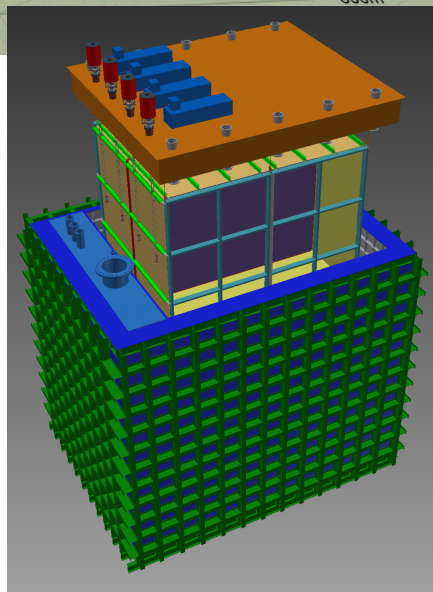


# The Short-Baseline Neutrino Program



**The Short-Baseline Near Detector (SBND) will be a 112 ton LArTPC located 110 meters from the target**

- Characterize the beam before oscillation
- Cancel many dominant systematic



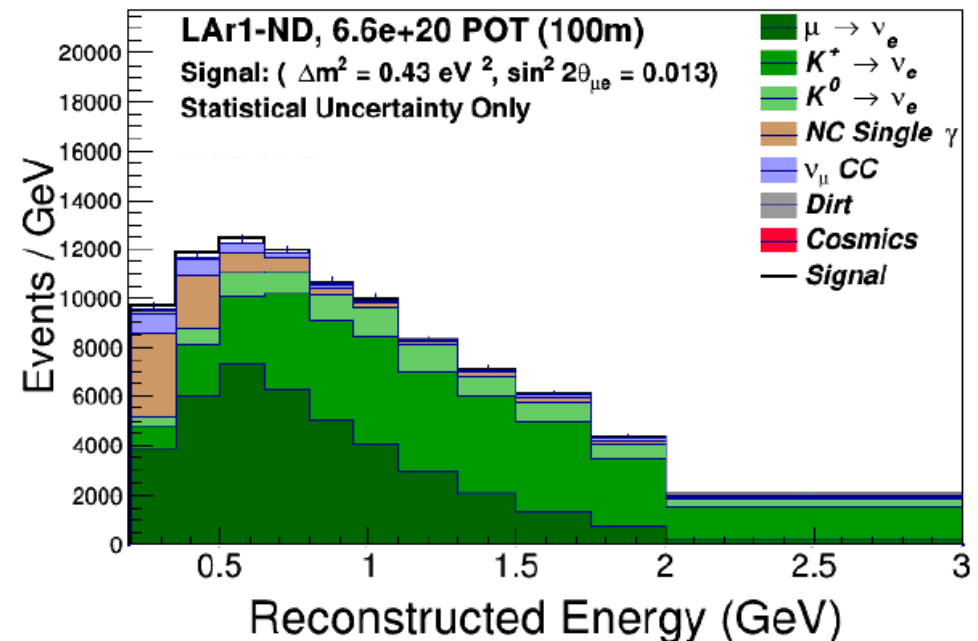


# Short Baseline Near Detector (SBND)

Process		No. Events
<i><math>\nu_\mu</math> Events (By Final State Topology)</i>		
CC Inclusive		5,212,690
CC 0 $\pi$	$\nu_\mu N \rightarrow \mu + Np$	3,551,830
	· $\nu_\mu N \rightarrow \mu + 0p$	793,153
	· $\nu_\mu N \rightarrow \mu + 1p$	2,027,830
	· $\nu_\mu N \rightarrow \mu + 2p$	359,496
	· $\nu_\mu N \rightarrow \mu + \geq 3p$	371,347
CC 1 $\pi^\pm$	$\nu_\mu N \rightarrow \mu + \text{nucleons} + 1\pi^\pm$	1,161,610
CC $\geq 2\pi^\pm$	$\nu_\mu N \rightarrow \mu + \text{nucleons} + \geq 2\pi^\pm$	97,929
CC $\geq 1\pi^0$	$\nu_\mu N \rightarrow \mu + \text{nucleons} + \geq 1\pi^0$	497,963
NC Inclusive		1,988,110
NC 0 $\pi$	$\nu_\mu N \rightarrow \text{nucleons}$	1,371,070
NC 1 $\pi^\pm$	$\nu_\mu N \rightarrow \text{nucleons} + 1\pi^\pm$	260,924
NC $\geq 2\pi^\pm$	$\nu_\mu N \rightarrow \text{nucleons} + \geq 2\pi^\pm$	31,940
NC $\geq 1\pi^0$	$\nu_\mu N \rightarrow \text{nucleons} + \geq 1\pi^0$	358,443
<i><math>\nu_e</math> Events</i>		
CC Inclusive		36798
NC Inclusive		14351
<b>Total <math>\nu_\mu</math> and <math>\nu_e</math> Events</b>		<b>7,251,948</b>

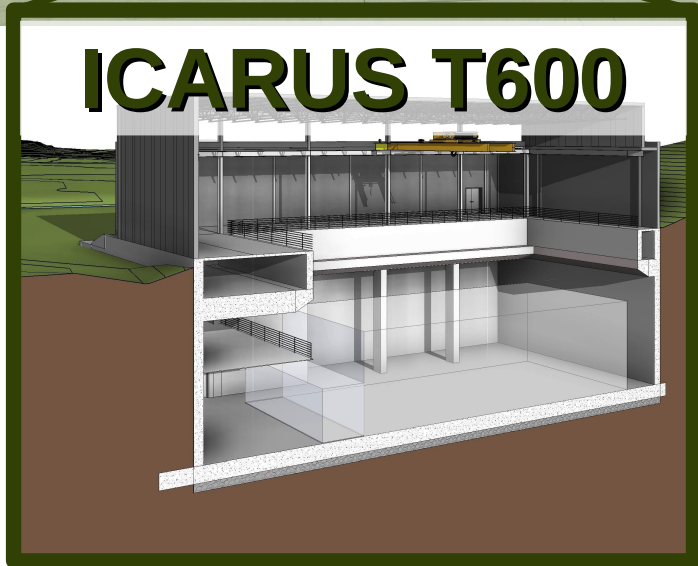
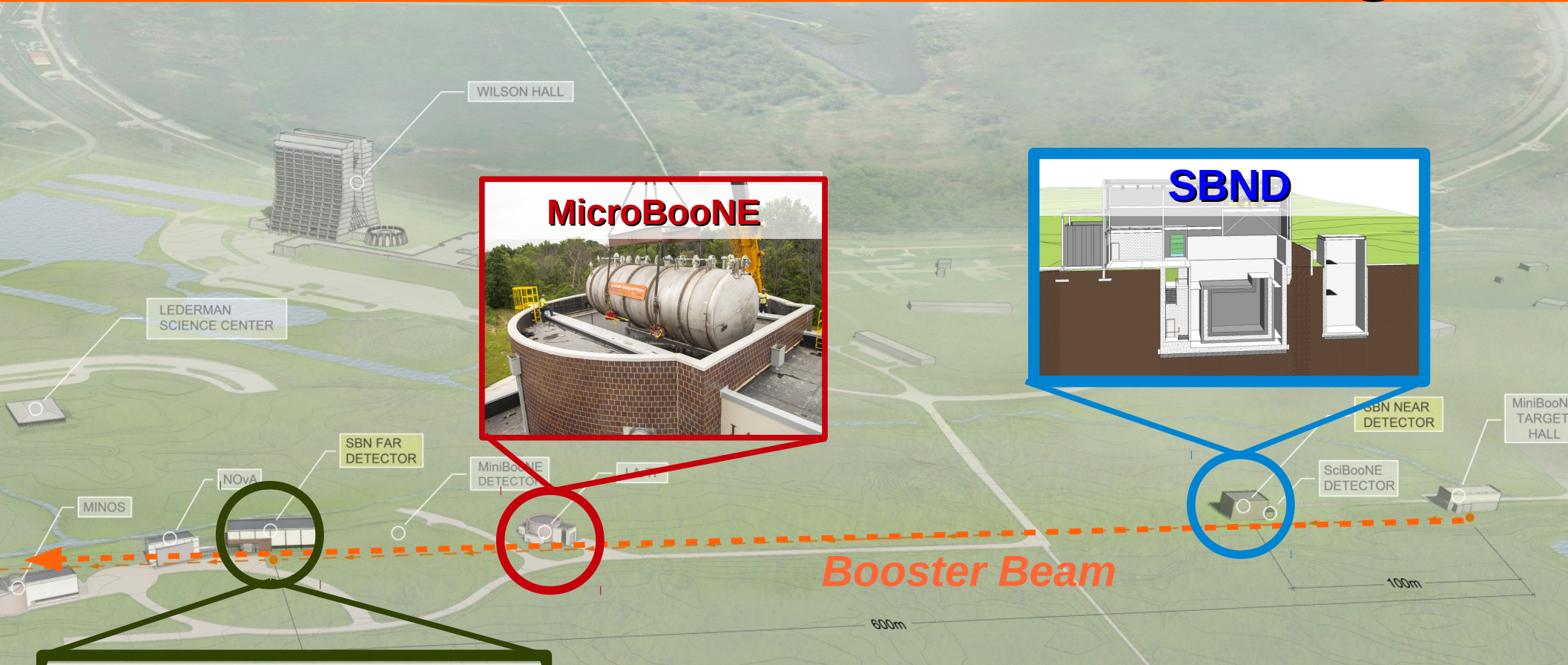
- **SBND will collect millions of neutrino interactions**

- High statistics, precision neutrino cross-sections measurements



- **Provides an un-oscillated spectrum for the electron neutrino appearance search**

# The Short-Baseline Neutrino Program



**The ICARUS detector is the largest LArTPC ever built**

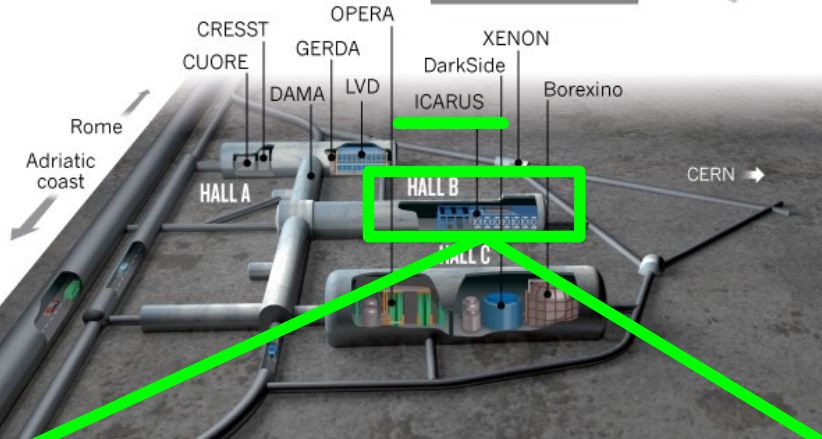
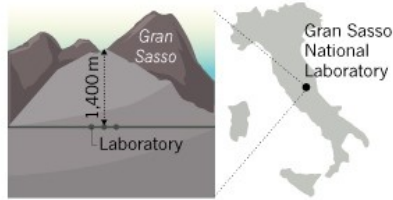
- Adding the large mass allows for precision oscillation search



# ICARUS T600

## THE A, B AND C OF GRAN SASSO

Experiments at the Gran Sasso National Laboratory are housed in and around three huge halls carved deep inside the mountain, where they are shielded from cosmic rays by 1,400 metres of rock.



- **ICARUS was the first large scale LArTPC to run in a neutrino beam line**

- Ran in the CNGS beam from CERN to Gran Sasso Laboratory from 2010 – 2013

- **After completing a successful neutrino run demonstrating the power of the LArTPC technology in an underground laboratory the detector has been moved from Gran Sasso to CERN**

**ICARUS Detector @ Gran Sasso**

# ICARUS T600

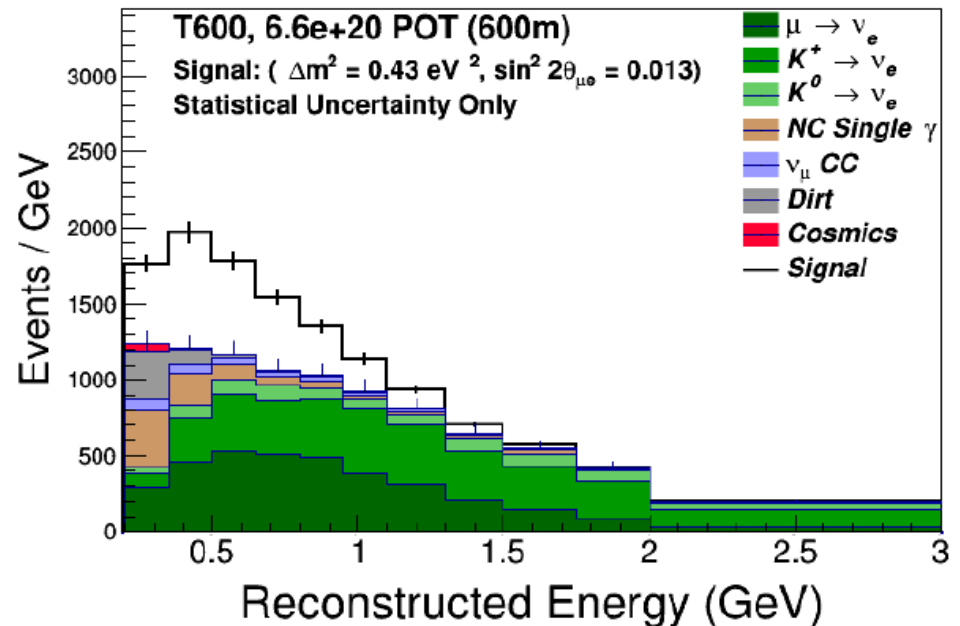


ICARUS TPC @ CERN

- The ICARUS detector is at CERN for refurbishment before it is shipped to Fermilab
  - The first module is expected to be finished in 2015
- This large mass detector will provide increased sensitivity to the electron neutrino appearance search



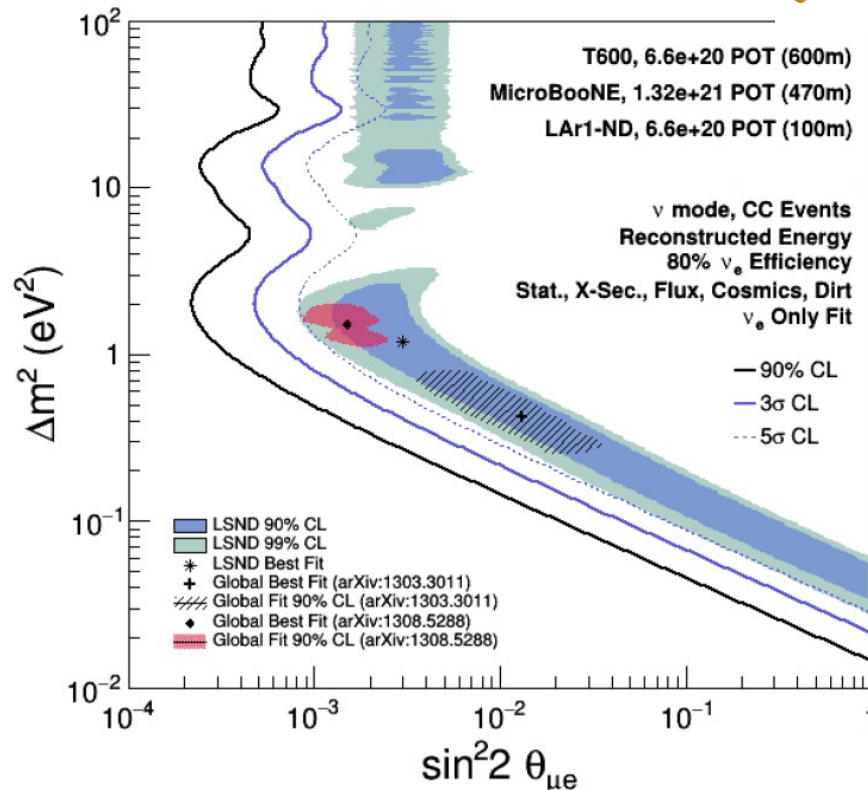
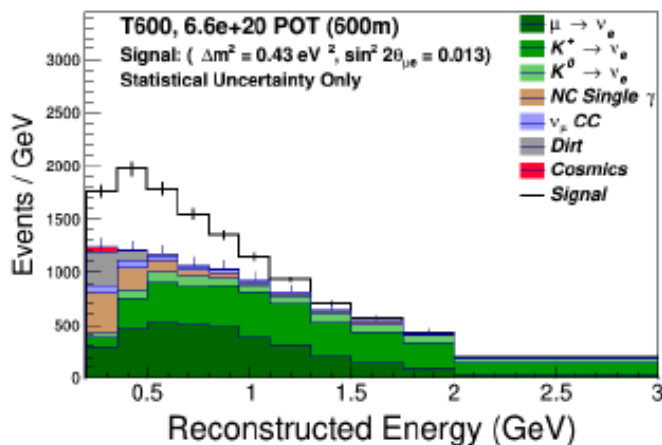
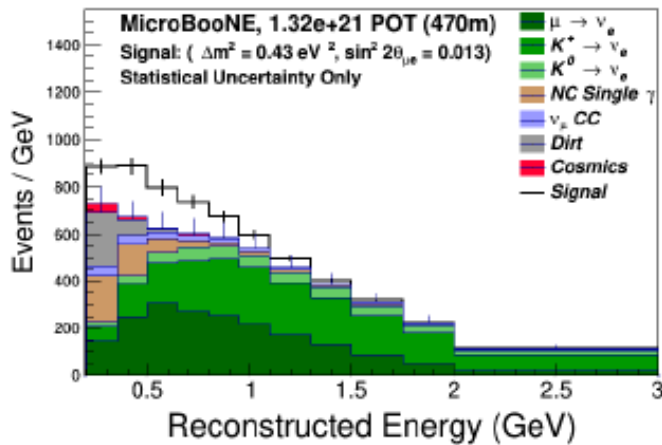
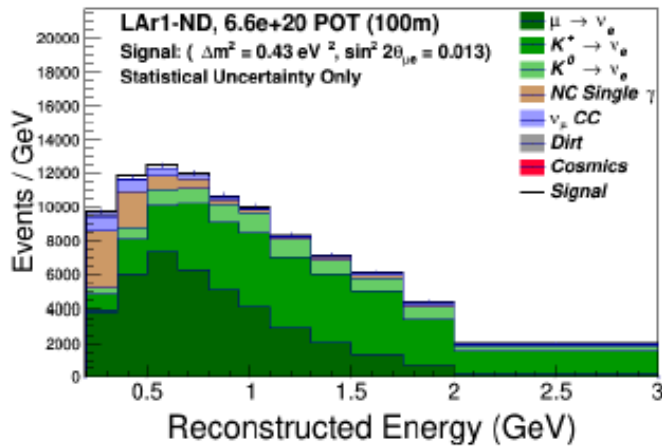
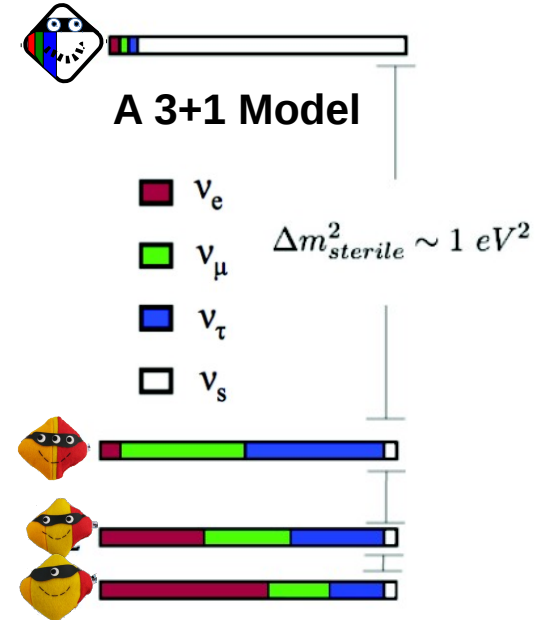
Far Detector Building Design



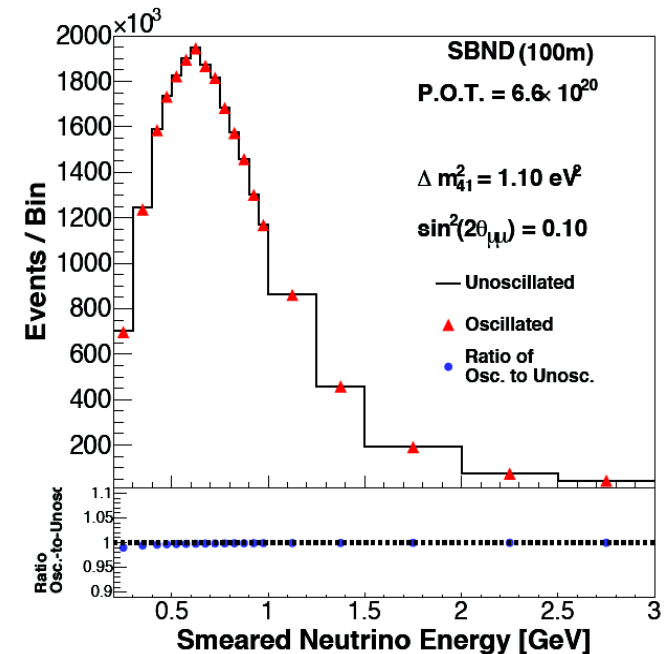
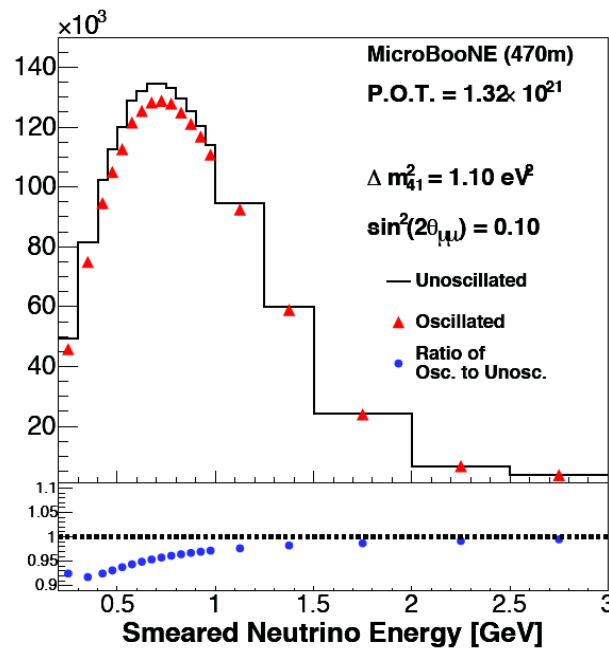
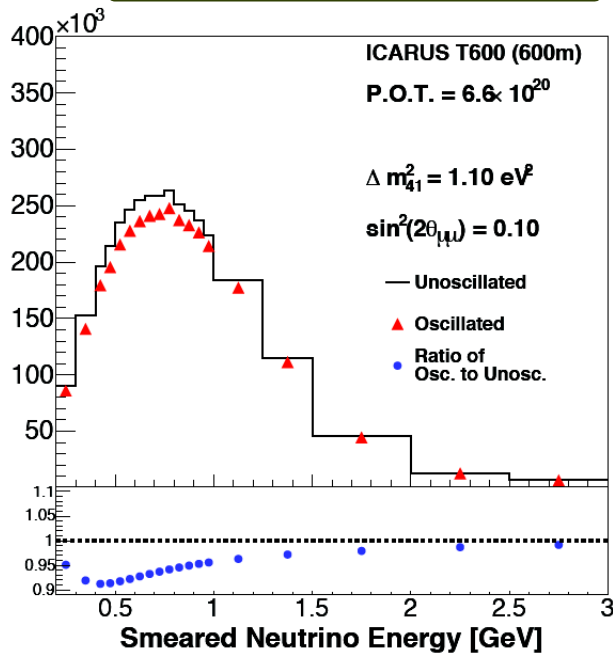
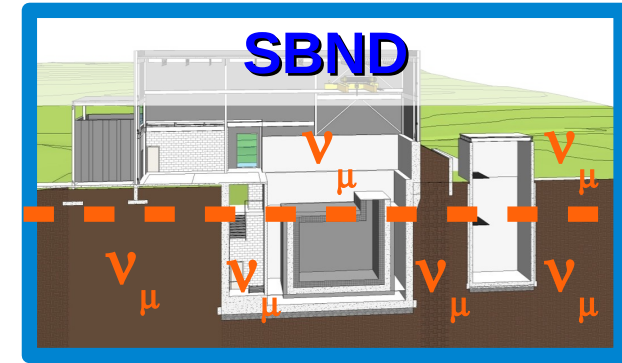
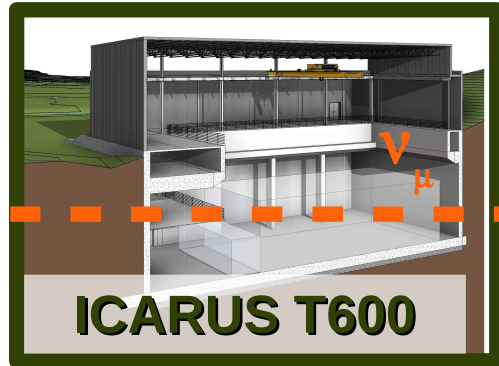


# The SBN Program

Utilizing three similar detectors at three different distances along the same neutrino beam allows for a **definitive measurement** of the allowed sterile neutrino parameter space



# The SBN Program



- The three detector configuration also allows you to search for the muon neutrino disappearance channel as well
  - Complimentary to the electron neutrino appearance search



# Conclusions

- **Fermilab stands at the dawn of the next generation of precision neutrino experiments**
- **The MicroBooNE experiment will begin to take data this summer**
  - This turns the key on the launch of the short-baseline experiment at Fermilab
- **Ground breaking on the buildings for the near and far detector will occur this summer**
  - Planning and design work on the near detector is moving ahead at full speed
  - The refurbishment of the ICARUS detector is ongoing at CERN and is expected to be complete in 2016

# Thank you for your attention

## The ICARUS-WA104 Collaboration

M. Antonello<sup>16</sup>, B. Baibussinov<sup>31</sup>, V. Bellini<sup>3</sup>, P. Benetti<sup>32</sup>, S. Bertolucci<sup>6</sup>, H. Bilokon<sup>15</sup>, F. Boffelli<sup>32</sup>, M. Bonesini<sup>17</sup>, J. Bremer<sup>6</sup>, E. Calligarich<sup>32</sup>, S. Centro<sup>31</sup>, A.G. Cocco<sup>19</sup>, A. Dermenev<sup>20</sup>, A. Falcone<sup>32</sup>, C. Farnese<sup>31</sup>, A. Fava<sup>31</sup>, A. Ferrari<sup>6</sup>, D. Gibin<sup>31</sup>, S. Gninenko<sup>20</sup>, N. Golubev<sup>20</sup>, A. Guglielmi<sup>31</sup>, A. Ivashkin<sup>20</sup>, M. Kirsanov<sup>20</sup>, J. Kisiel<sup>38</sup>, U. Kose<sup>6</sup>, F. Mammoliti<sup>5</sup>, G. Mannocchi<sup>15</sup>, A. Menegolli<sup>32</sup>, G. Meng<sup>31</sup>, D. Mladenov<sup>6</sup>, C. Montanari<sup>32</sup>, M. Nessi<sup>6</sup>, M. Nicoletto<sup>31</sup>, F. Noto<sup>6</sup>, P. Picchi<sup>15</sup>, F. Pietropaolo<sup>31</sup>, P. Płoński<sup>42</sup>, R. Potenza<sup>3</sup>, A. Rappoldi<sup>32</sup>, G. L. Raselli<sup>32</sup>, M. Rossella<sup>32</sup>, C. Rubbia<sup>\*,6,11,16</sup>, P. Sala<sup>18</sup>, A. Scaramelli<sup>18</sup>, J. Sobczyk<sup>44</sup>, M. Spanu<sup>32</sup>, D. Stefan<sup>18</sup>, R. Sulej<sup>43</sup>, C.M. Sutura<sup>5</sup>, M. Torti<sup>32</sup>, F. Tortorici<sup>5</sup>, F. Varanini<sup>31</sup>, S. Ventura<sup>31</sup>, C. Vignoli<sup>16</sup>, T. Wachala<sup>12</sup>, and A. Zani<sup>32</sup>

## The SBND Collaboration

C. Adams<sup>45</sup>, C. Andreopoulos<sup>23</sup>, A. Ankowski<sup>41</sup>, J. Asaadi<sup>40</sup>, L. Bagby<sup>10</sup>, B. Baller<sup>10</sup>, N. Barros<sup>33</sup>, M. Bass<sup>30</sup>, S. Bertolucci<sup>6</sup>, M. Bishai<sup>3</sup>, A. Bitadze<sup>25</sup>, J. Bremer<sup>6</sup>, L. Bugel<sup>26</sup>, L. Camilleri<sup>9</sup>, F. Cavanna<sup>\*,10</sup>, H. Chen<sup>3</sup>, C. Chi<sup>9</sup>, E. Church<sup>10</sup>, D. Cianci<sup>7</sup>, G. Collin<sup>26</sup>, J.M. Conrad<sup>26</sup>, G. De Geronimo<sup>3</sup>, R. Dharmapalan<sup>1</sup>, Z. Djurcic<sup>1</sup>, A. Ereditato<sup>2</sup>, J. Esquivel<sup>40</sup>, J. Evans<sup>25</sup>, B.T. Fleming<sup>45</sup>, W.M. Foreman<sup>7</sup>, J. Freestone<sup>25</sup>, T. Gamble<sup>37</sup>, G. Garvey<sup>24</sup>, V. Genty<sup>9</sup>, D. Göldi<sup>2</sup>, H. Greenlee<sup>10</sup>, R. Guenette<sup>30</sup>, A. Hackenburg<sup>45</sup>, R. Hänni<sup>2</sup>, J. Ho<sup>7</sup>, J. Howell<sup>10</sup>, C. James<sup>10</sup>, C.M. Jen<sup>41</sup>, B.J.P. Jones<sup>26</sup>, L.M. Kalousis<sup>41</sup>, G. Karagiorgi<sup>25</sup>, W. Ketchum<sup>24</sup>, J. Klein<sup>33</sup>, J. Klinger<sup>37</sup>, U. Kose<sup>6</sup>, I. Kreslo<sup>2</sup>, V.A. Kudryavtsev<sup>37</sup>, D. Lissauer<sup>3</sup>, P. Livesly<sup>22</sup>, W.C. Louis<sup>24</sup>, M. Lu<sup>□thi</sup><sup>2</sup>, C. Mariani<sup>41</sup>, K. Mavrokoridis<sup>23</sup>, N. McCauley<sup>23</sup>, N. McConkey<sup>37</sup>, I. Mercer<sup>22</sup>, T. Miao<sup>10</sup>, G.B. Mills<sup>24</sup>, D. Mladenov<sup>6</sup>, D. Montanari<sup>10</sup>, J. Moon<sup>26</sup>, Z. Moss<sup>26</sup>, S. Mufson<sup>14</sup>, M. Nessi<sup>6</sup>, B. Norris<sup>10</sup>, F. Noto<sup>6</sup>, J. Nowak<sup>22</sup>, S. Pal<sup>37</sup>, O. Palamara<sup>\*,b,10</sup>, J. Pater<sup>25</sup>, Z. Pavlovic<sup>10</sup>, J. Perkin<sup>37</sup>, G. Pulliam<sup>40</sup>, X. Qian<sup>3</sup>, L. Qiuguang<sup>24</sup>, V. Radeka<sup>3</sup>, R. Rameika<sup>10</sup>, P.N. Ratoff<sup>22</sup>, M. Richardson<sup>37</sup>, C. Rudolf von Rohr<sup>2</sup>, D.W. Schmitz<sup>\*,7</sup>, M.H. Shaevitz<sup>9</sup>, B. Sippach<sup>9</sup>, M. Soderberg<sup>40</sup>, S. Söldner-Rembold<sup>25</sup>, J. Spitz<sup>26</sup>, N. Spooner<sup>37</sup>, T. Strauss<sup>2</sup>, A.M. Szelc<sup>25,45</sup>, C.E. Taylor<sup>24</sup>, K. Terao<sup>9</sup>, M. Thiesse<sup>37</sup>, L. Thompson<sup>37</sup>, M. Thomson<sup>4</sup>, C. Thorn<sup>3</sup>, M. Toups<sup>26</sup>, C. Touramanis<sup>23</sup>, R.G. Van De Water<sup>24</sup>, M. Weber<sup>2</sup>, D. Whittington<sup>14</sup>, T. Wongjirad<sup>26</sup>, B. Yu<sup>3</sup>, G.P. Zeller<sup>10</sup>, and J. Zennaro<sup>7</sup>

## The MicroBooNE Collaboration

R. Acciarri<sup>10</sup>, C. Adams<sup>45</sup>, R. An<sup>13</sup>, A. Ankowski<sup>41</sup>, J. Asaadi<sup>40</sup>, L. Bagby<sup>10</sup>, B. Baller<sup>10</sup>, G. Barr<sup>30</sup>, M. Bass<sup>30</sup>, M. Bishai<sup>3</sup>, A. Blake<sup>4</sup>, T. Bolton<sup>21</sup>, C. Bromberg<sup>27</sup>, L. Bugel<sup>26</sup>, L. Camilleri<sup>9</sup>, D. Caratelli<sup>9</sup>, B. Carls<sup>10</sup>, F. Cavanna<sup>\*,10</sup>, H. Chen<sup>3</sup>, E. Church<sup>10</sup>, G.H. Collin<sup>26</sup>, J.M. Conrad<sup>26</sup>, M. Convery<sup>39</sup>, S. Dytman<sup>34</sup>, B. Eberly<sup>39</sup>, A. Ereditato<sup>2</sup>, J. Esquivel<sup>40</sup>, B.T. Fleming<sup>\*,45</sup>, W.M. Foreman<sup>7</sup>, V. Genty<sup>9</sup>, D. Göldi<sup>2</sup>, S. Gollapinni<sup>21</sup>, M. Graham<sup>39</sup>, E. Gramellini<sup>45</sup>, H. Greenlee<sup>10</sup>, R. Grosso<sup>8</sup>, R. Guenette<sup>30</sup>, A. Hackenburg<sup>45</sup>, O. Hen<sup>26</sup>, J. Hewes<sup>25</sup>, J. Ho<sup>7</sup>, G. Horton-Smith<sup>21</sup>, C. James<sup>10</sup>, C.M. Jen<sup>41</sup>, R.A. Johnson<sup>8</sup>, B.J.P. Jones<sup>26</sup>, J. Joshi<sup>3</sup>, H. Jostlein<sup>10</sup>, D. Kaleko<sup>9</sup>, L. Kalousis<sup>41</sup>, G. Karagiorgi<sup>25</sup>, W. Ketchum<sup>24</sup>, B. Kirby<sup>3</sup>, M. Kirby<sup>10</sup>, T. Kobilarcik<sup>10</sup>, I. Kreslo<sup>2</sup>, Y. Li<sup>3</sup>, B. Littlejohn<sup>13</sup>, D. Lissauer<sup>3</sup>, S. Lockwitz<sup>10</sup>, W.C. Louis<sup>24</sup>, M. Lu<sup>□thi</sup><sup>2</sup>, B. Lundberg<sup>10</sup>, A. Marchionni<sup>10</sup>, C. Mariani<sup>41</sup>, J. Marshall<sup>4</sup>, K. McDonald<sup>35</sup>, V. Meddage<sup>21</sup>, T. Miceli<sup>28</sup>, G.B. Mills<sup>24</sup>, J. Moon<sup>26</sup>, M. Mooney<sup>3</sup>, M.H. Moulai<sup>26</sup>, R. Murrells<sup>25</sup>, D. Naples<sup>34</sup>, P. Nienaber<sup>36</sup>, O. Palamara<sup>\*,b,10</sup>, V. Paolone<sup>34</sup>, V. Papavassiliou<sup>28</sup>, S. Pate<sup>28</sup>, Z. Pavlovic<sup>10</sup>, S. Pordes<sup>10</sup>, G. Pulliam<sup>40</sup>, X. Qian<sup>3</sup>, J.L. Raaf<sup>10</sup>, V. Radeka<sup>3</sup>, R. Rameika<sup>10</sup>, B. Rebel<sup>10</sup>, L. Rochester<sup>39</sup>, C. Rudolf von Rohr<sup>2</sup>, B. Russell<sup>45</sup>, D.W. Schmitz<sup>7</sup>, A. Schukraft<sup>10</sup>, W. Seligman<sup>9</sup>, M. Shaevitz<sup>9</sup>, M. Soderberg<sup>40</sup>, J. Spitz<sup>26</sup>, J. St. John<sup>8</sup>, T. Strauss<sup>2</sup>, A.M. Szelc<sup>25,45</sup>, N. Tagg<sup>29</sup>, K. Terao<sup>9</sup>, M. Thomson<sup>4</sup>, C. Thorn<sup>3</sup>, M. Toups<sup>26</sup>, Y. Tsai<sup>39</sup>, T. Usher<sup>39</sup>, R. Van de Water<sup>24</sup>, M. Weber<sup>2</sup>, S. Wolbers<sup>10</sup>, T. Wongjirad<sup>26</sup>, K. Woodruff<sup>28</sup>, M. Xu<sup>13</sup>, T. Yang<sup>10</sup>, B. Yu<sup>3</sup>, G.P. Zeller<sup>\*,10</sup>, J. Zennaro<sup>7</sup>, and C. Zhang<sup>3</sup>

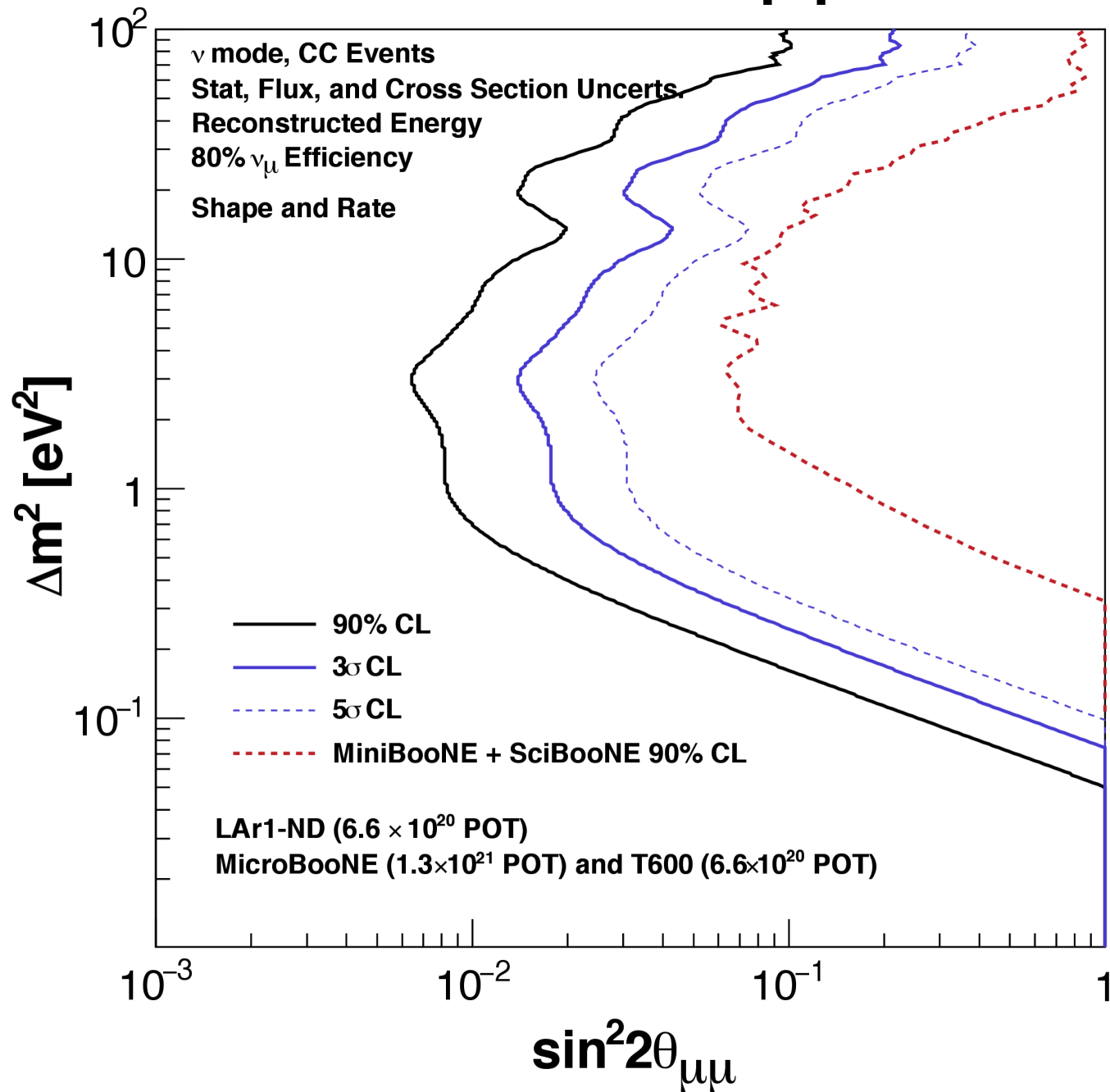
## Additional Fermilab Contributors

W. Badgett<sup>10</sup>, K. Biery<sup>10</sup>, S. Brice<sup>10</sup>, S. Dixon<sup>10</sup>, M. Geynisman<sup>10</sup>, C. Moore<sup>10</sup>, E. Snider<sup>10</sup>, and P. Wilson<sup>10</sup>



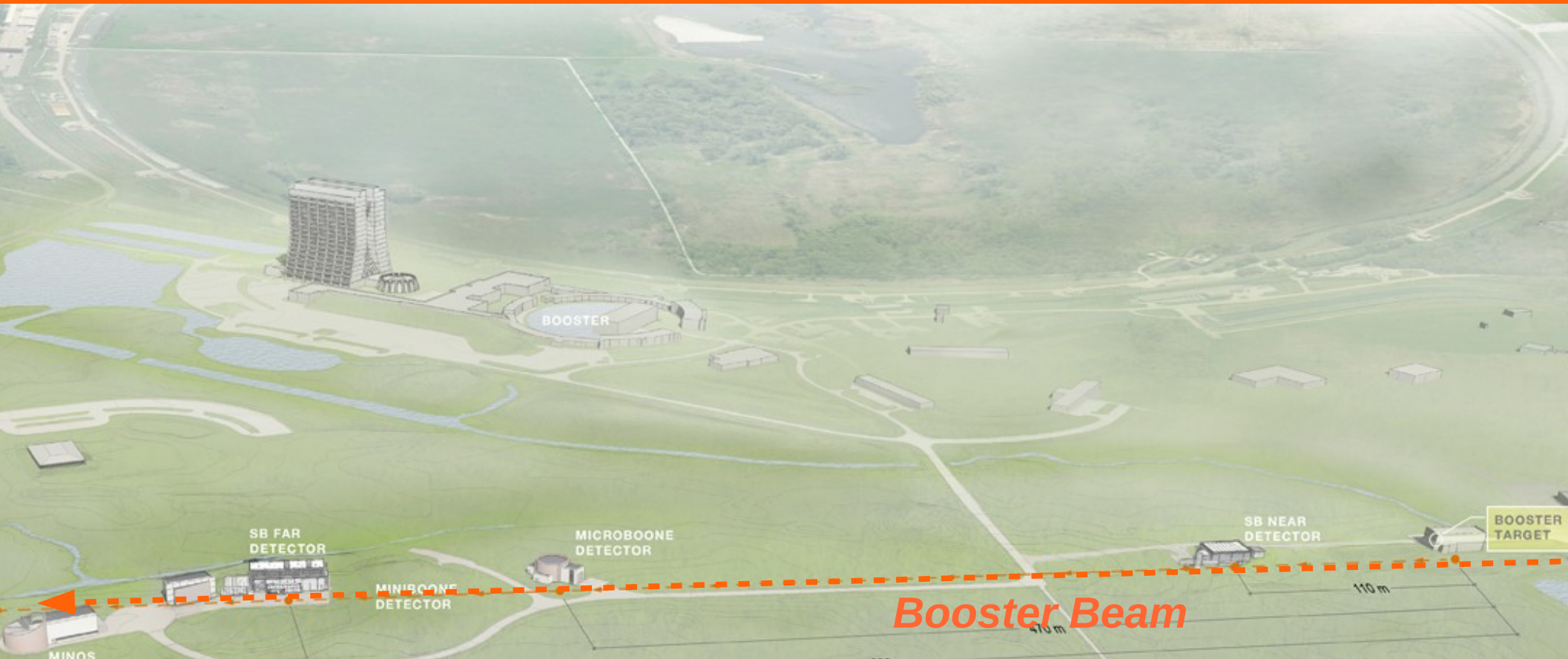
# Backups

# Muon Neutrino Disappearance

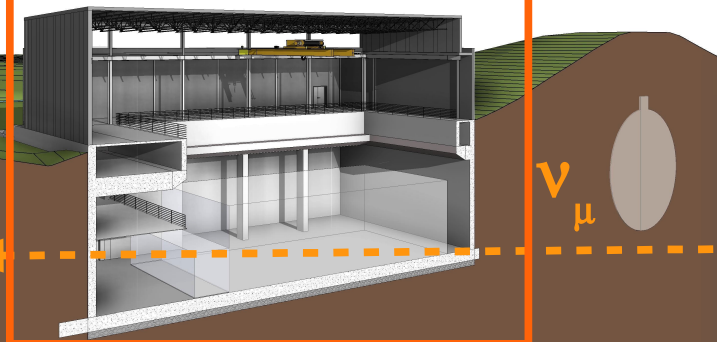




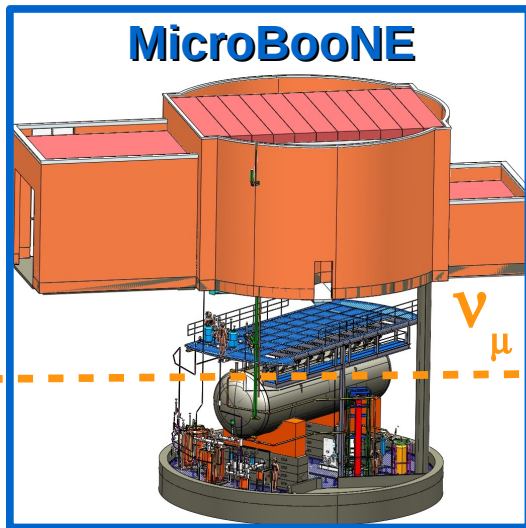
# The Short-Baseline Neutrino Program



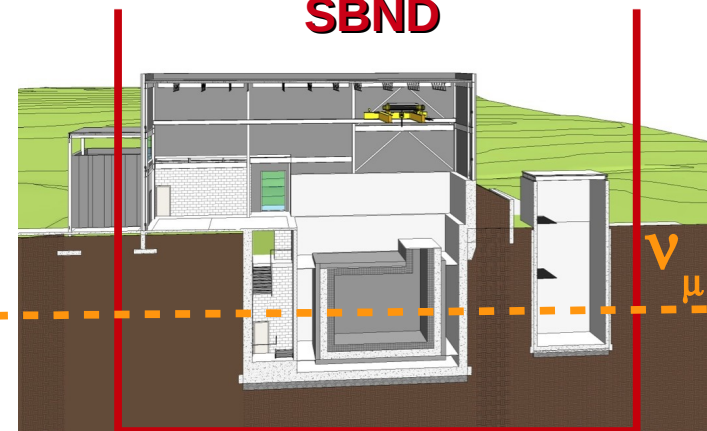
**ICARUS: T600**



**MicroBooNE**

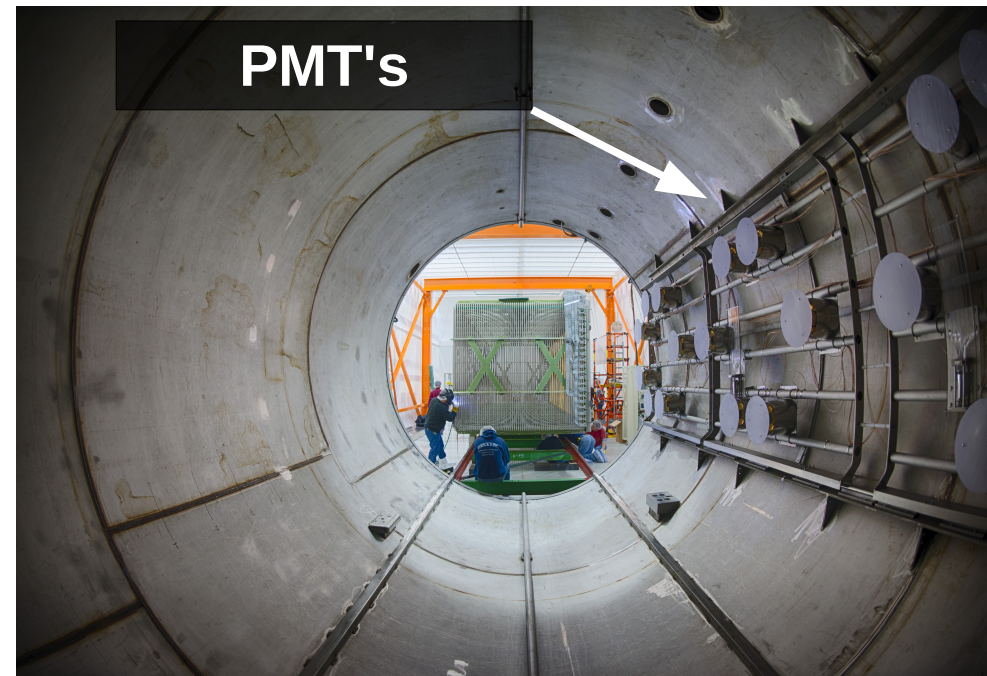
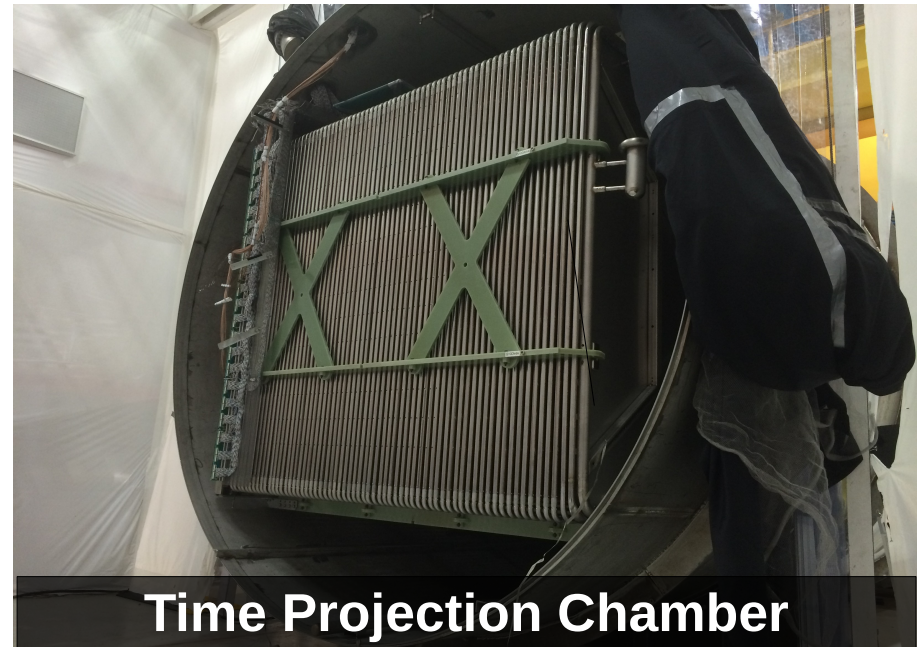


**SBND**



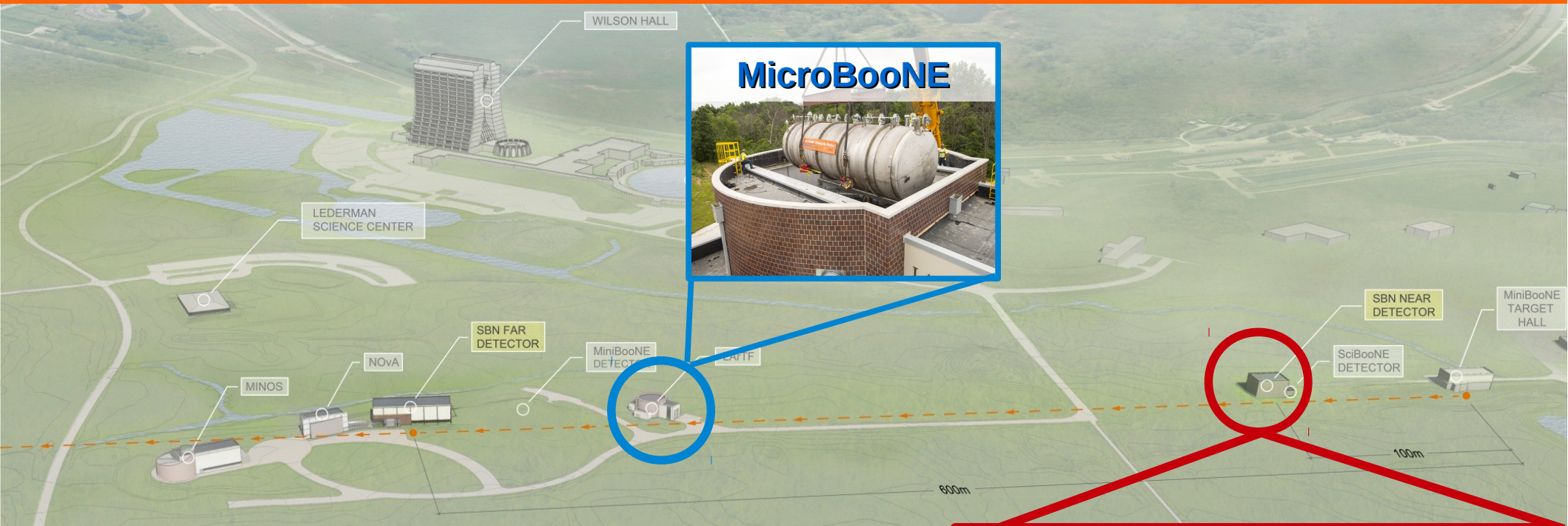
# MicroBooNE

- MicroBooNE is a 170 ton (total volume) LArTPC
- TPC Dimensions:
  - 10.3 m long x 2.3 m tall x 2.5 m wide (drift distance)
  - 89 ton active mass
- 8256 wire channels
  - 3456 Collection channels
    - Wires oriented w.r.t. the vertical
  - 4800 Induction channels
    - Wires oriented +/- 60°
- 32 8" cryogenic PMT's
  - Provides event  $t_0$  as well as cosmic ray removal
- UV Laser Calibration System

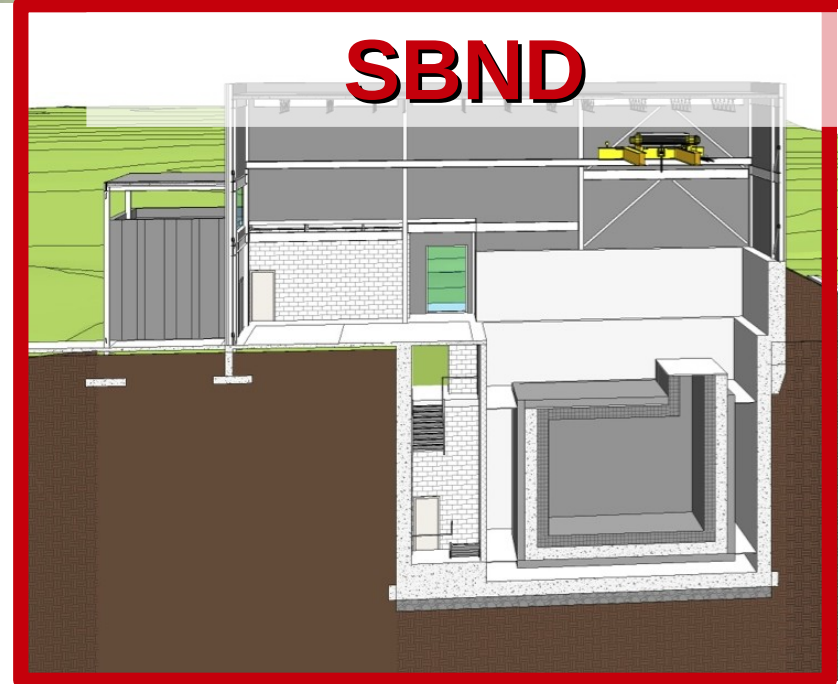
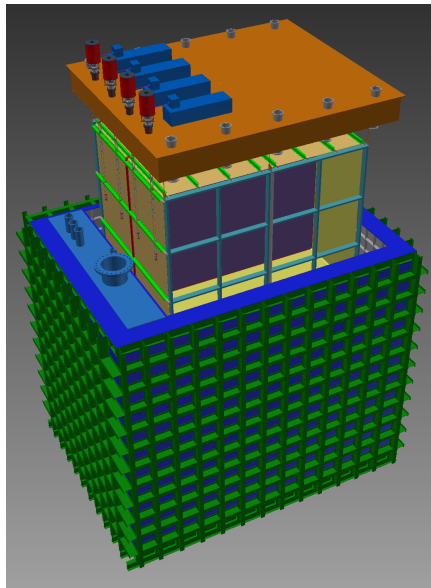




# Current Status of SBN

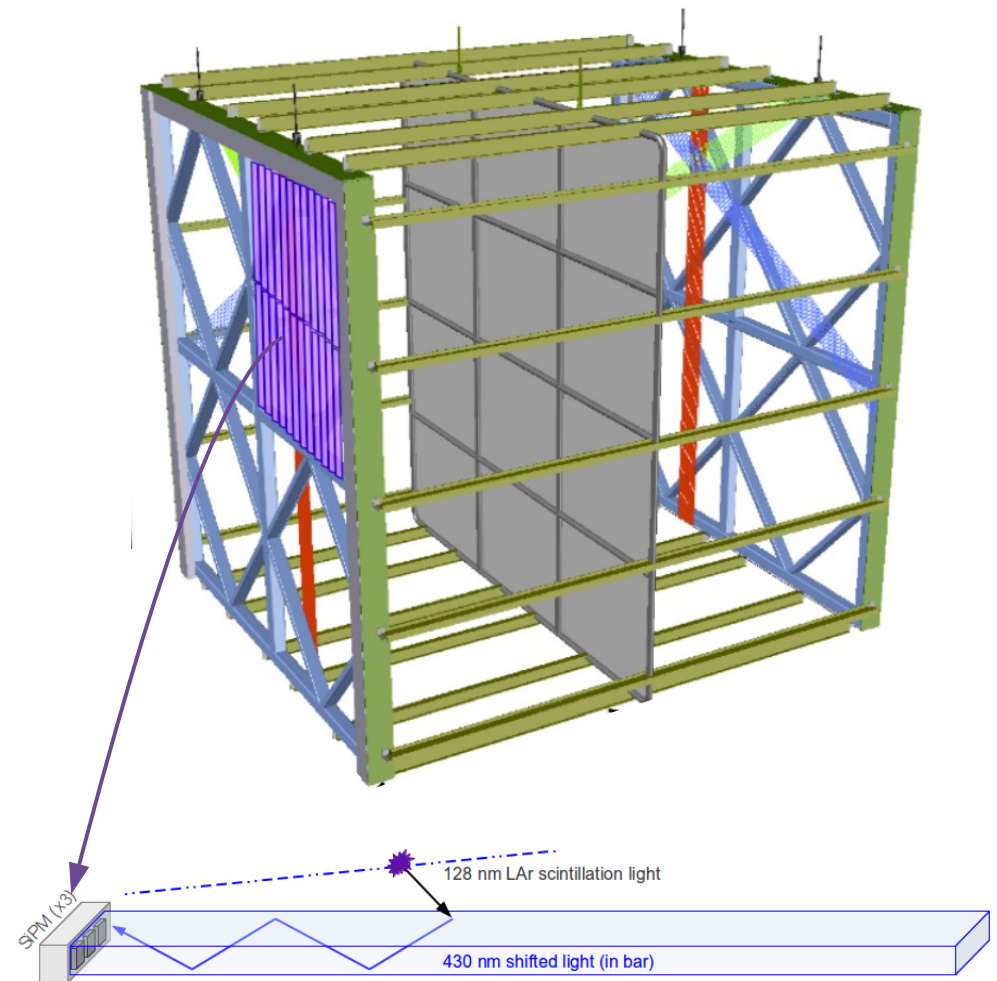


**SBND will serve as the near detector and**

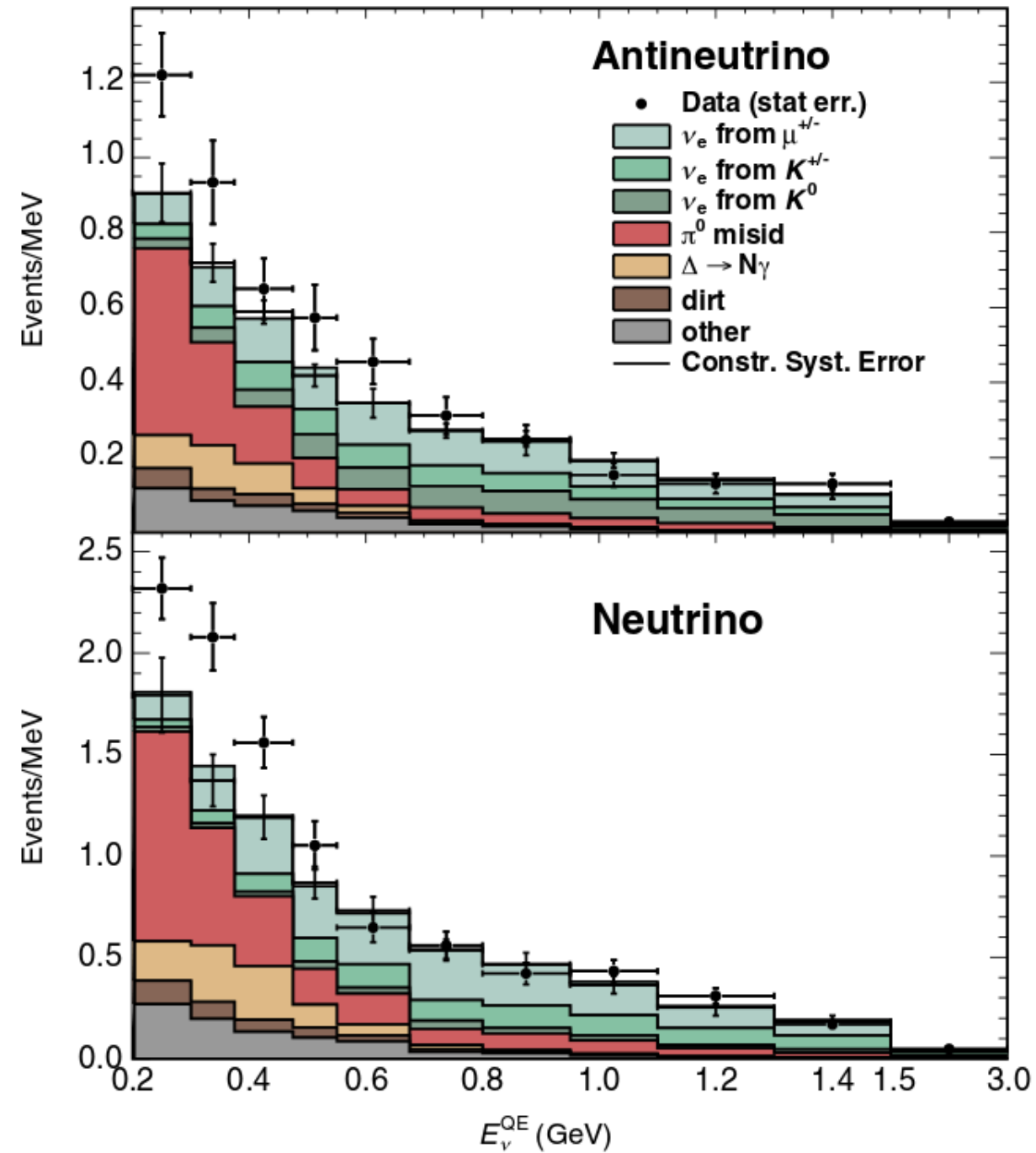


# Short Baseline Near Detector (SBND)

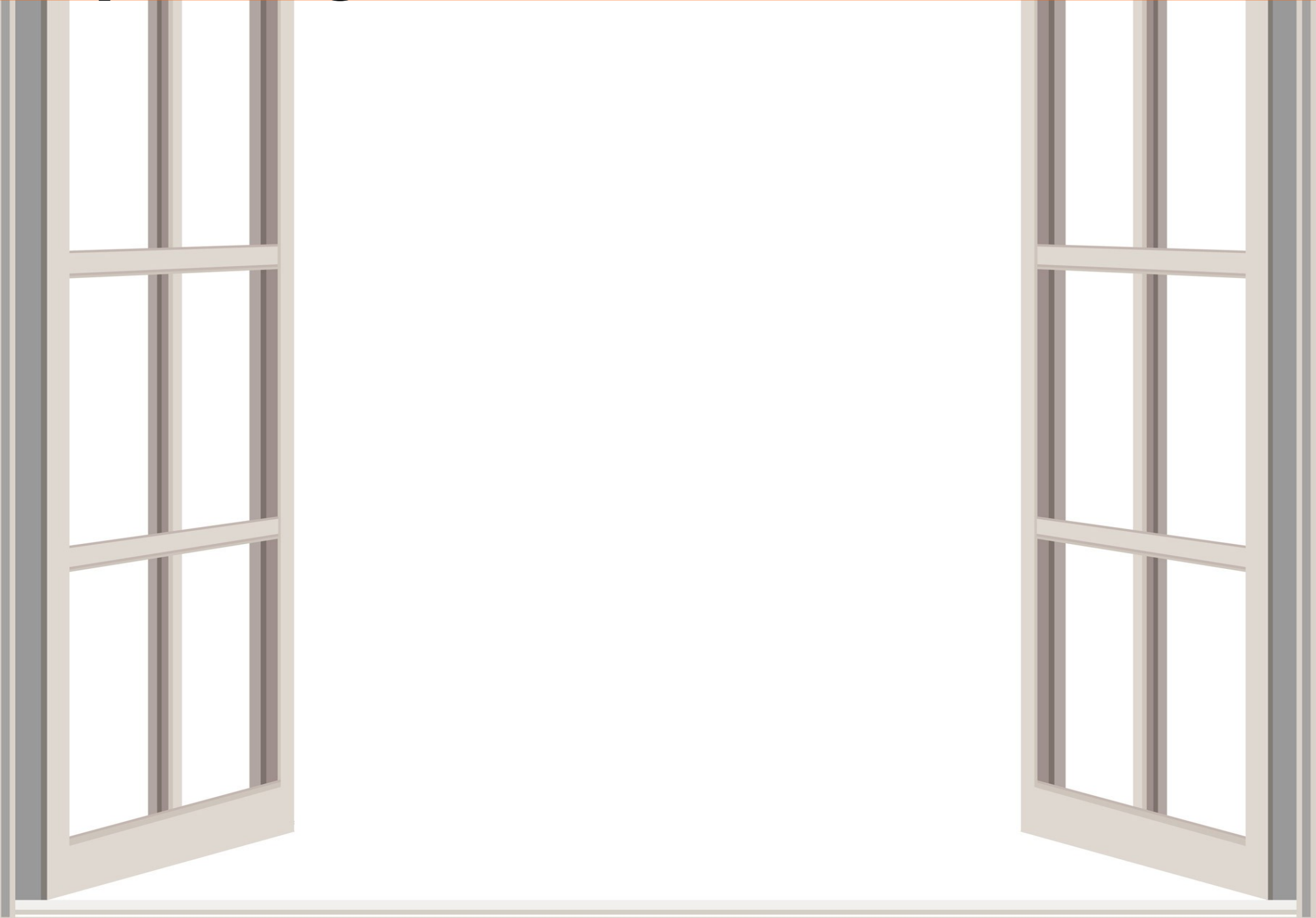
- SBND is a 210 ton (total volume) LArTPC
- TPC Dimensions:
  - 5 m long x 4 m tall x 4 m wide (drift distance)
  - 112 ton active mass
- 11,264 wire channels
  - ???? Collection channels
    - Wires oriented w.r.t. the vertical
  - ???? Induction channels
    - Wires oriented +/- 60°
- Finalizing the Light Detection System Desing
  - Provides event  $t_0$  as well as cosmic ray removal
- UV Laser Calibration System







# Opening windows on new areas





# Opening windows on new areas

