

Positioning the 1st COMET transport solenoid

Phillip Litchfield (COMET)

Charged lepton flavour violation

We already know that lepton flavour is not conserved

- Weak mixing mechanism & non-degenerate neutrino masses
- Neutrino (lack of) mass & charge means this is easiest to observe in neutrino oscillations, but can also lead to CLFV:



- The basic SM amplitudes can be related to the neutrino oscillation parameters, but requires some radiation to conserve energy & momentum.
- The μ e system is particularly simple because the radiated 'mass' must be neutral, and lighter than a muon.

Options for decaying muons

The most obvious candidate for the transition to radiate is a photon, and the branching ratio is:

$$\frac{\Gamma(\mu \to e\gamma)}{\Gamma(\mu \to e\nu\nu)} \propto \left| \sum_{i} \frac{m_i^2}{m_W^2} U_{\mu i}^* U_{ei} \right|^2 \sim O(10^{-54})$$

For a free muon, γ (or *ee*) is the only option...

...but in a muonic atom the radiation can be virtual The nucleus absorbs it, and recoils slightly.

- Because of the relatively large nuclear mass, the electron is effectively mono-energetic.
- Because the process does not require a 'real' photon, other diagrams are possible...



 $\mu \rightarrow e\gamma$ (see note)

Note: The γ can connect anywhere, not just in the loop

New physics

Similar processes exist in a wide variety of new physics scenarios.
Muon decay is at low energy, so reduce to effective operators:

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$\mu N \rightarrow eN$ and $\mu \rightarrow e\gamma$

$$\mathcal{L}_{\mu e} \sim \frac{1}{\Lambda^2} \left[\frac{1}{\kappa + 1} m_{\mu} \bar{\mu} \sigma_{\mu \nu} e \cdot F^{\mu \nu} + \frac{\kappa}{\kappa + 1} \bar{\mu} \gamma_{\mu} e \cdot \bar{q} \gamma_{\mu} q \right]$$

- New physics \rightarrow CLFV in rare muon decays.
- Energy scale <u>A</u> affects the rate of all such processes.
- Parameter *k* depends on the nature of the new physics

Both $\mu \rightarrow e\gamma$ and $\mu - e$ conversion are sensitive to dipole terms, but $\mu - e$ conv. is also sensitive to 4-femion terms.

- More sensitive to some models.
- (If signal seen) the comparison allows discrimination between models







Basics of a (modern) $\mu - e$ conversion experiment

Cartoon: Flip Tanedo http://www.particlebites.com/

Muon decays

Muons allowed stop in suitable target.

- Initially Aluminium, but other materials (Ti) under study.
- Conversion from 1s orbital: $\mu N \rightarrow eN$ gives a mono-energetic electron at 105MeV ($\approx m_{\mu} - B_{1s}^{\mu}$)

'Normal' decays are backgrounds

- Nuclear muon capture: $\mu N(Z) \rightarrow \nu N(Z-1)$
- Decay in Orbit [DIO]: $\mu N \rightarrow e v \overline{\nu} N$

For a free muon, cuts off at $\frac{1}{2}m_{\mu}$, but bound state has a small tail up to m_{μ}



UCL

μ

AI

Backgrounds

Three main background processes: Decay in orbit, as before Energy resolution!

- Decay in flight:
 Electrons from energetic free 10² muons can be boosted to 105MeV.
 - Use momentum selection in muon transport (see later)

Results from SINDRUM-II (BR <7 \times 10⁻¹³ @ 90%CL)



Beam backgrounds:

Significant number of prompt e^- and π^- produced by beam. Can eliminate this with timing *if* we have reliably beam-free time windows.

Beam-free windows?

Naïvely, this sounds easy, but...

- High intensity pulsed muon beams are uncommon \rightarrow new facilities
- Need $\tau_{\mu} \gg \sigma_{\text{Pulse}}$ so choose stopping targets with long lifetime \mapsto low *Z*, conflicts with high *A* preferred for coherent signal.

Aluminium (Z = 13, $\tau_{\mu} = 0.88 \mu s$) preferred



Really beam-free windows?

RCS

UCL

COMET

Main Ring 4/9 buckets filled

Synchrotrons have stable acceleration buckets

Even if you don't inject protons into them, stray protons can remain in stable acceleration.

The signal process is rare, so requirements on the *extinction* between pulses is very strict



Muon source

Main driver of sensitivity: Need lots of low energy muons!

- Use dedicated high-power *pulsed* proton beam lines (~8 GeV)
- Resonant slow extraction onto pion production target
- Collect *backward*-going pions with capture solenoid
- Can use gradient field to reflect low-energy forward pions. (Mu2e ►)
- Maximise field at target to give larger solid angle apperture (COMET)



- Pions decay to muons en-route to stopping target.
- Many neutrons produced, requires careful shielding. The curved transport line helps to eliminate direct line-of sight.

Muon transport







Muon transport is a curved solenoid

- Particles are channelled in spiral paths [solenoid], which naturally tend up/down [curvature] depending on momentum and charge
 - Gives charge sign and momentum selection, which can be enhanced by using a collimator.
- Use to eliminate high momentum muons, other particles.
- Eliminates line-of-sight from production target



▲ 3D printed model of Mu2e solenoids



- S-shape and off-centre collimators that can rotate for BG studies
- Stopping target is 17 × 0.2mm AI foils
- Target & detector surrounded by solenoid
 - Electron transport
 - Magnetic mirror





Electrons spiral from target to tracker and EM calorimeter

Mu2e Tracker





MU2e

- Tracker made from straw tubes to measure particle momentum
- Minimum radius of 380mm corresponds to momentum ~ 60MeV
 - 'Complete' tracks need mom > 90 MeV





Mu2e calorimeter

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Inorganic scintillator crystals
Provides fast response for trigger

Mu2e

• Energy measurement combined with momentum from tracker gives excellent μ/e discrimination

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▲ View from (newly furnished) COMET Counting Room

Two phases of COMET





In time-revered order: Phase II...





- Muon transport is $180^\circ \rightarrow$ larger dispersion.
- COMET transport coils use compensating dipole so selected tracks stay level.
- COMET uses a second curved solenoid as an **electron spectrometer**. This filters out 'low' momentum and +ive backgrounds
- Final detector is tracker / EM calorimeter (like Mu2e) but *full plane* – thanks to

spectrometer.

...and Phase I



Phase I has 2 goals:

- Investigate backgrounds for phase II
- Perform search at 100× sensitivity of SINDRUM-II

For Phase I measurement use a cylindrical drift chamber around the stopping target.

• Triggering by **auxillary hodoscopes** Also include prototypes/partial elements of Phase II detectors for development and characterising backgrounds at low current



COMET developments



COMET hall construction completed!



Magnets laid-out for new beam switchyard

Transport Solenoid





Corrective dipoles

▼Completed 90° muon transport arc (including octagonal return yoke)

Extinction test





2014: 7×10^{-13} 90% U.L. [SINDRUM-II] (since 2004) ~2017: 3×10^{-15} S.E.S. [COMET Phase-I] (~ 6mo) ~2021: 3×10^{-17} S.E.S. [COMET Phase-II & Mu2e]

beyond 2021: PRISIM /PRIME @J-PARC? [Goal 3×10^{-19}] Mu2e × ProjectX @FNAL? Charged lepton flavour violation is 'natural' in new physics scenarios.

- The (arguably) SM process is driven by neutrino mixing, and is hugely suppressed: O(10⁻⁵⁴)
- Complimentary approaches from $\mu \rightarrow e\gamma \mu \rightarrow eee$ and

 $\mu - e$ conversion

- Current limit is $O(10^{-12})$ from SINDRUM-II
- COMET & Mu2e can improve on this by 4 orders of magnitude in 5~10 years
- Information on *what* the NP is through similar experiments.





Production target

UCL

Phase-I baseline (unlikely to change): 60cm × 2cm dia. graphite (IG-43) target.

Higher Z is better for pion production, but **graphite** is a 'safer' choice:

- IG-43 is used for T2K target (FX, >200kW beam) so is known to be capable of handling our beam.
- Lower irradiation of target and shield make safer in case of replacement in Phase-II
- At Phase-I power, radiative cooling is suffic





Capture Solenoid



Comet needs *low energy* pions so collect from **back and sides** of target.



- **Gradient field** converts transverse momentum into longitudinal momentum.
- Effectively increases the solid angle aperture into the transport solenoid.



Cooling and shielding

A 5T solenoid is (unsurprisingly?) superconducting.And therefore cryogenically cooled...

But there is a high power beam hitting a target in the middle!

- Phase I: this heating is estimated up to 30W
- Phase II: heating can be 120W

[c.f. other sources ~15W]

Shielding is needed, for radiation and thermal heating.

- Copper and tungsten shield
- Cooled with water
- Will probably need upgrade for Phase II, gets very (radioactively) hot.

Non-trivial engineering challenge!



Phase-I Detector



Phase-I will have a dedicated detector for $\mu \rightarrow e$ conversion measurements.

Because of the charged particle tracks in the centre channel, a co-axial cylinder geometry is used. B_z(tesla)

└→ CyDet

The detector and capture target will sit within a 1T solenoid field.

Low momentum particles do not reach the detector



CyDet

The main part of the detector is a coaxial drift chamber

- Helium-based gas mixture to reduce multiple scattering.
 – Resolution ~ 200 keV
- *z* measurement by stereo layers
- Large inner radius to reduce DIO hit rate
 Dim: 150cm × 84cm_(outer) // 50cm_(inner)
- 19 concentric sense layers
- Triggering from hodoscopes at ends





Drift chamber progress



Electron track [~3Hz] **Proton tracks** [~7 in 500ns] identals

- Event display showing event projection
- Stringing wires and CR test of prototype section

Beam extinction



*currently prefer

3/9 buckets

filled

protons will 'leak' into empty buckets.

Signal process is rare so even a small leak is a major background

Extinction measurement



COMET design requires that we can achieve an extinction:

$$E = \frac{N_{Empty}}{N_{Filled}} < 10^{-9}$$

Extinction can be improved by increasing RF voltage, but this heats the cavities.

(And there is a limit...)



2012 test at 30 GeV demonstrated this is possible for RF > 120kV, But study time at 8 GeV not allocated until this year...