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MARE-1 in Milan: an update

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**fondazione
c a r i p l o**

Outline

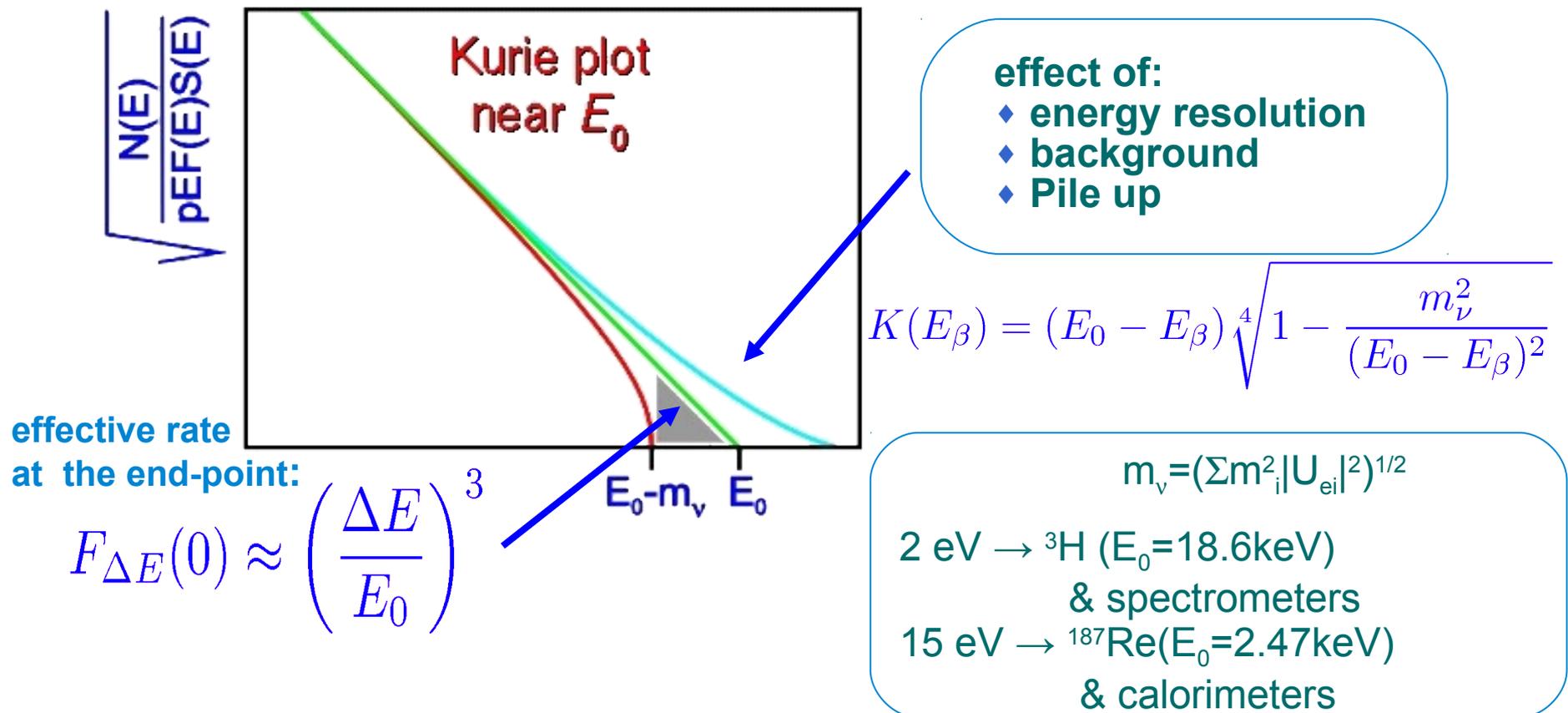
- Direct neutrino mass measurement
- The calorimetric approach
- MARE: a project for a new Re/ho experiment
- MARE-1 in Milano
- MKDs R&D @ Milano Bicocca

Direct neutrino mass measurement

neutrino oscillations evidence $\rightarrow m_\nu \neq 0$
 BUT oscillation experiments give only Δm^2

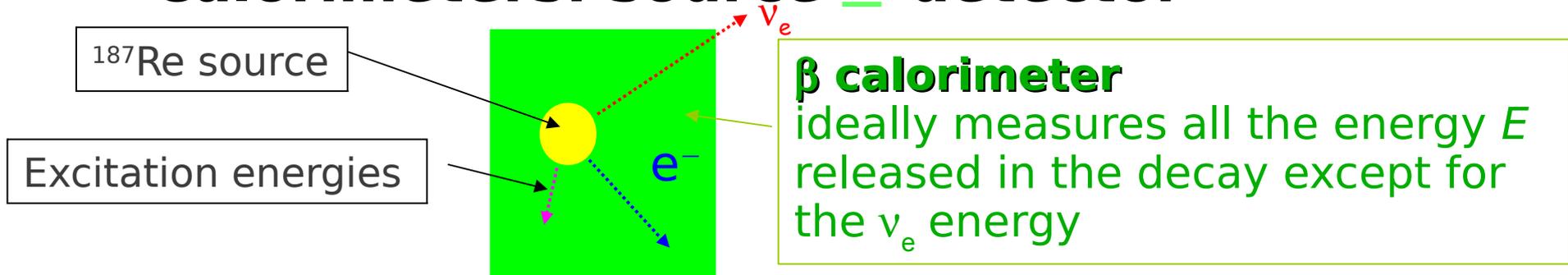


direct neutrino mass measurement



The calorimetric approach

- **Calorimeters: source \subseteq detector**

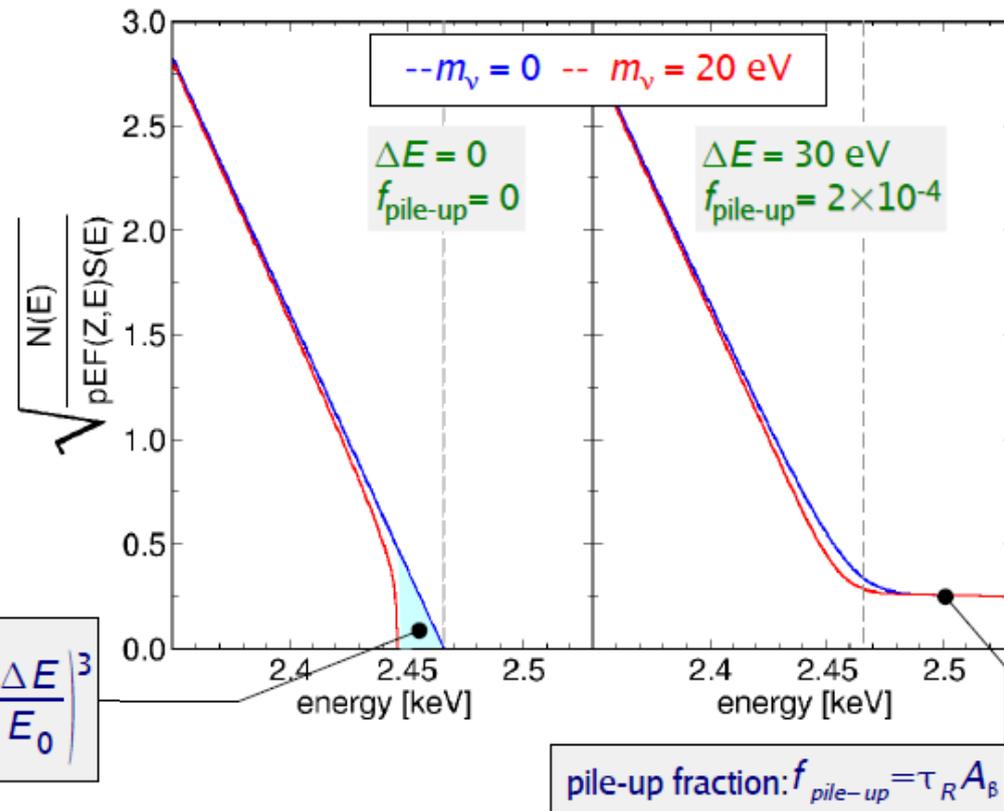


General experimental requirements:

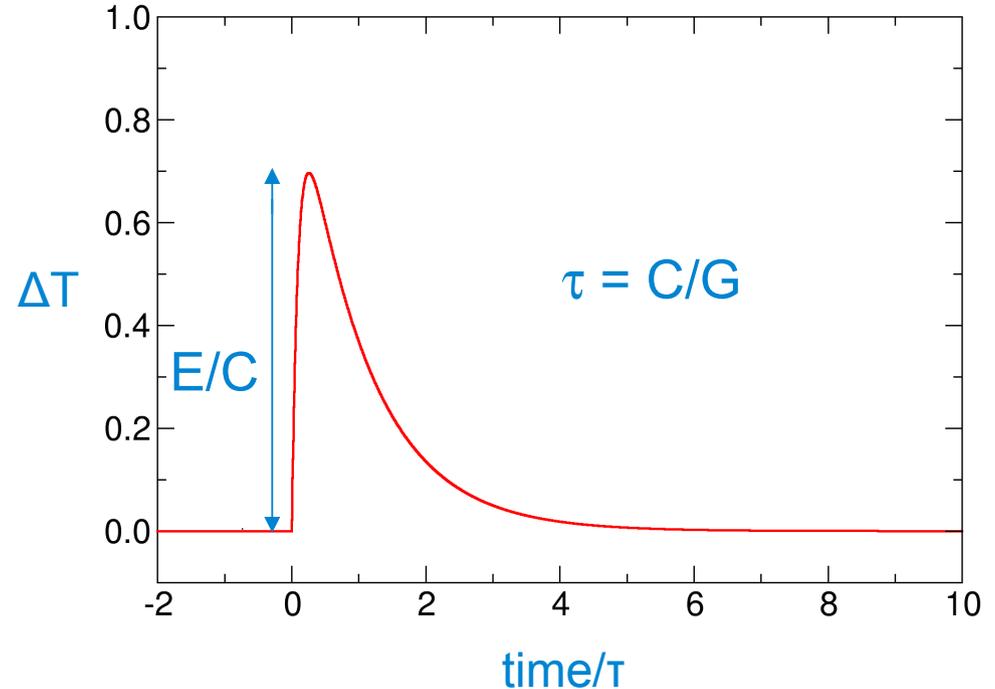
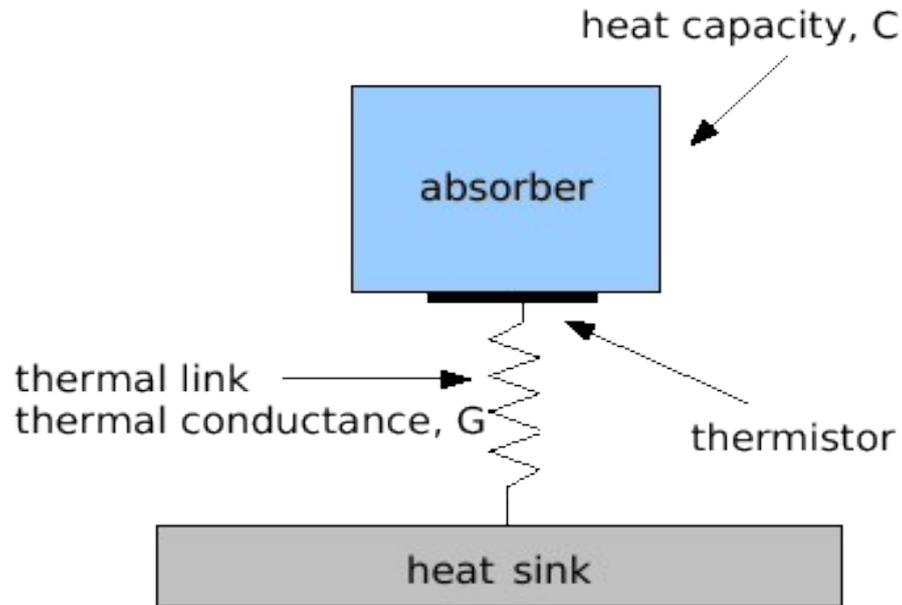
- High statistics at the beta spectrum end-point
- high energy resolution ΔE
- high Signal to Noise ratio
- small systematic effects

Calorimeters measure the **entire spectrum** at once:

- low E_0 β decaying isotopes for more statistics near the end-point
- best choice ^{187}Re :
 - $E_0 = 2.5 \text{ keV}$, $\tau^{1/2} = 4 \times 10^{10} \text{ y}$
- other option ^{163}Ho EC:
 - $E_0 \approx 2.6 \text{ keV}$, $\tau^{1/2} \approx 4600 \text{ y}$



Cryogenic detectors as calorimeters



Detection Principle:

- $\Delta T = E/C$ where C is the total thermal capacity
 - low C : $C \sim (T/\Theta_D)^3$ in superconductors below T_c & dielectric
 - low T (10 ÷ 100 mK)
- ultimate limit to energy resolution:
 - statistical fluctuation of internal energy $\Delta E = (k_B T^2 C)^{1/2}$
- detect all deposited energy, including short-lived excited states (100 μ s)
- achieve very good energy resolution in the keV range

MARE - A project for a new Re/Ho experiment

Goal: a sub-eV direct neutrino mass measurement complementary to the KATRIN experiment

MARE 1

- activities aiming at isotope/technique selection (^{187}Re or ^{163}Ho options)
- activities using medium sized arrays to improve ^{187}Re measurement understanding and possibly calorimetric m_ν limit
- detector and absorber coupling R&D activities



MARE 2

- very large experiment with a m_ν statistical sensitivity close to KATRIN but still improvable
- requires new improved detector technologies

MARE for sub-eV calorimetric m_ν measurement

MARE: Microcalorimeter Arrays for a Rhenium Experiment

Università di Genova e INFN Sez. di Genova, Italy
Univ. di Milano-Bicocca, Univ. dell'Insubria e INFN Sez. di Milano-Bicocca, Italy
Kirkhhof-Institute Physik, Universitat Heidelberg, Germany
University of Miami, Florida, USA
Wisconsin University, Madison, Wisconsin, USA
Universidade de Lisboa and ITN, Portugal
Università di Roma "La Sapienza" e INFN Sez. di Roma1, Italy
Goddard Space Flight Center, NASA, Maryland, USA
PTB, Berlin, Germany
FBK, Trento e INFN Sez. di Padova, Italy
NIST, Boulder, Colorado, USA
SISSA - Trieste, GSI Darmstad, JPL/Caltech, CNRS Grenoble, ...



<http://crio.mib.infn.it/wig/silicini/proposal/>

MARE 1

MARE-1: collection of activities aiming at isotope/technique selection

- o **^{187}Re** – high statistics measurement
 - o asses systematics
 - o test large arrays
 - o lower limit to few eV
- o **^{163}Ho** – high statistics measurement – R&D for ^{163}Ho production
 - o measure Q_{EC}
 - o study spectrum shape
 - o asses systematics

Different techniques:

- TES – Transition Edge Sensor
- MMC – Magnetic MicroCalorimeter
- MKID – Microwave Kinetic Inductance Detector

 • **multiplexed readout**
• **large arrays**

MARE-1 in Milan / Re option

MARE-1 in Milan: Milano/FBK/Wisconsin/NASA

- $m_{\nu_e} < 2 \text{ eV}/c^2$
- 10^{10} events - 300 sensors
- 8 arrays of Si:P thermistors with AgReO_4 absorbers
- energy resolution 25 eV @ 2.6 keV

The first phase is needed:

- because it's the only possible one with present technology
- To investigate systematics in thermal calorimeters

 **very important to cross-check spectrometer results**

MARE-1 detectors

- ^{187}Re β -decay

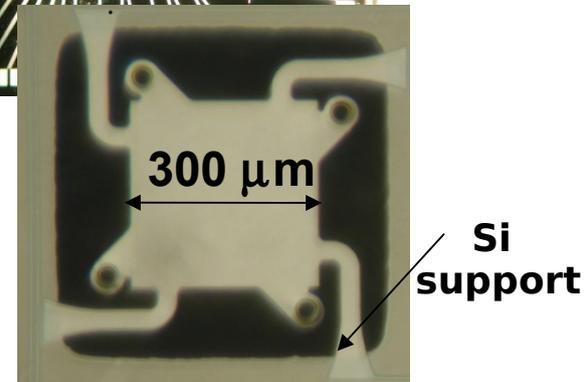
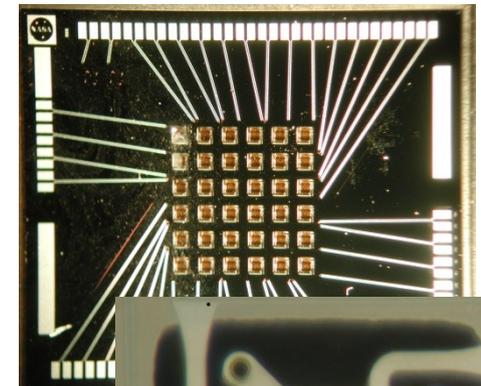
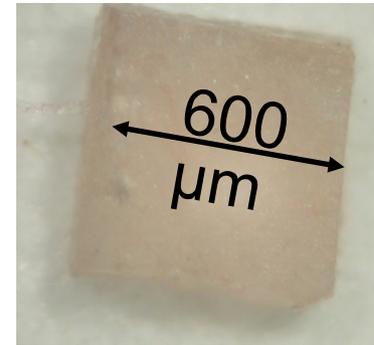
- $^{187}\text{Re} \rightarrow ^{187}\text{Os} + e^- + \nu_e$ $E_0=2.47$ keV
- i. a. 63% and $\tau=42.3$ Gy

- Single crystal of silver perrhenate (AgReO_4)

- mass ~ 500 μg per pixel ($A_\beta \sim 0.3$ decay/sec)
- regular shape ($600 \times 600 \times 250$ μm^3)
- low heat capacity due to Debye law

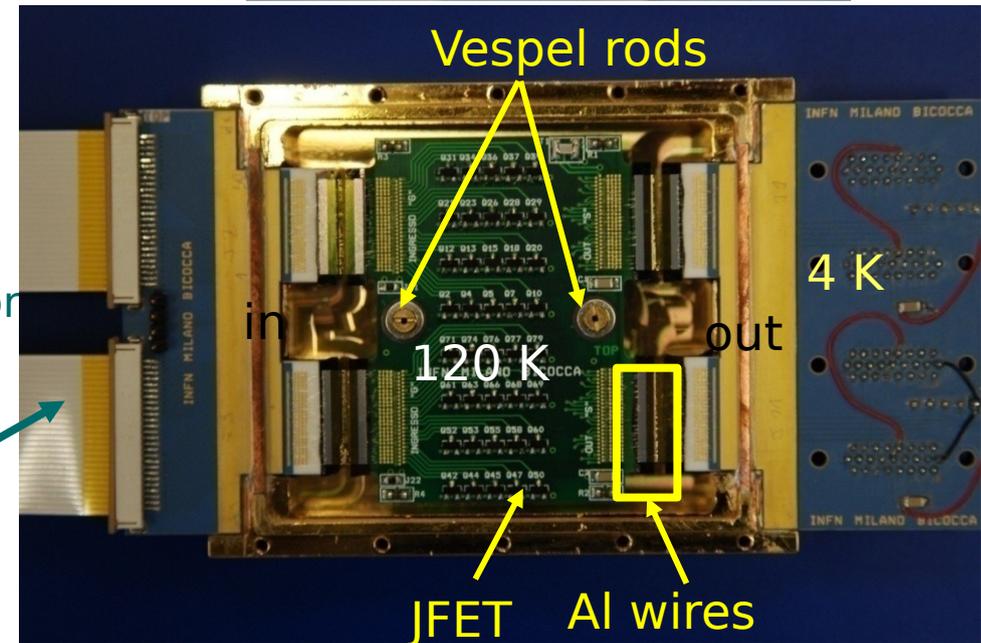
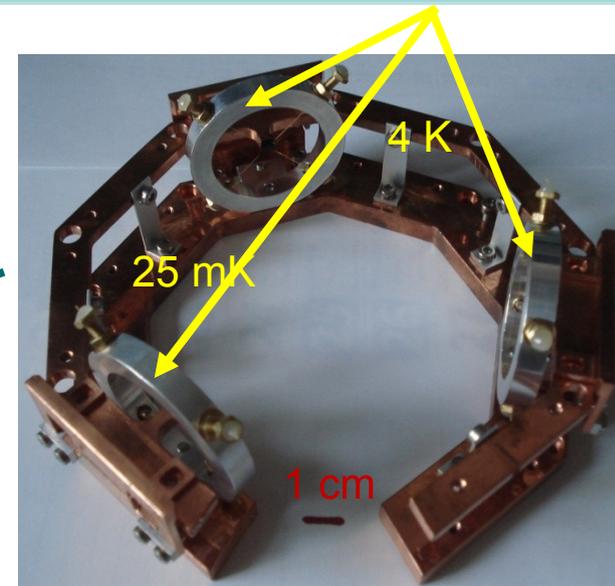
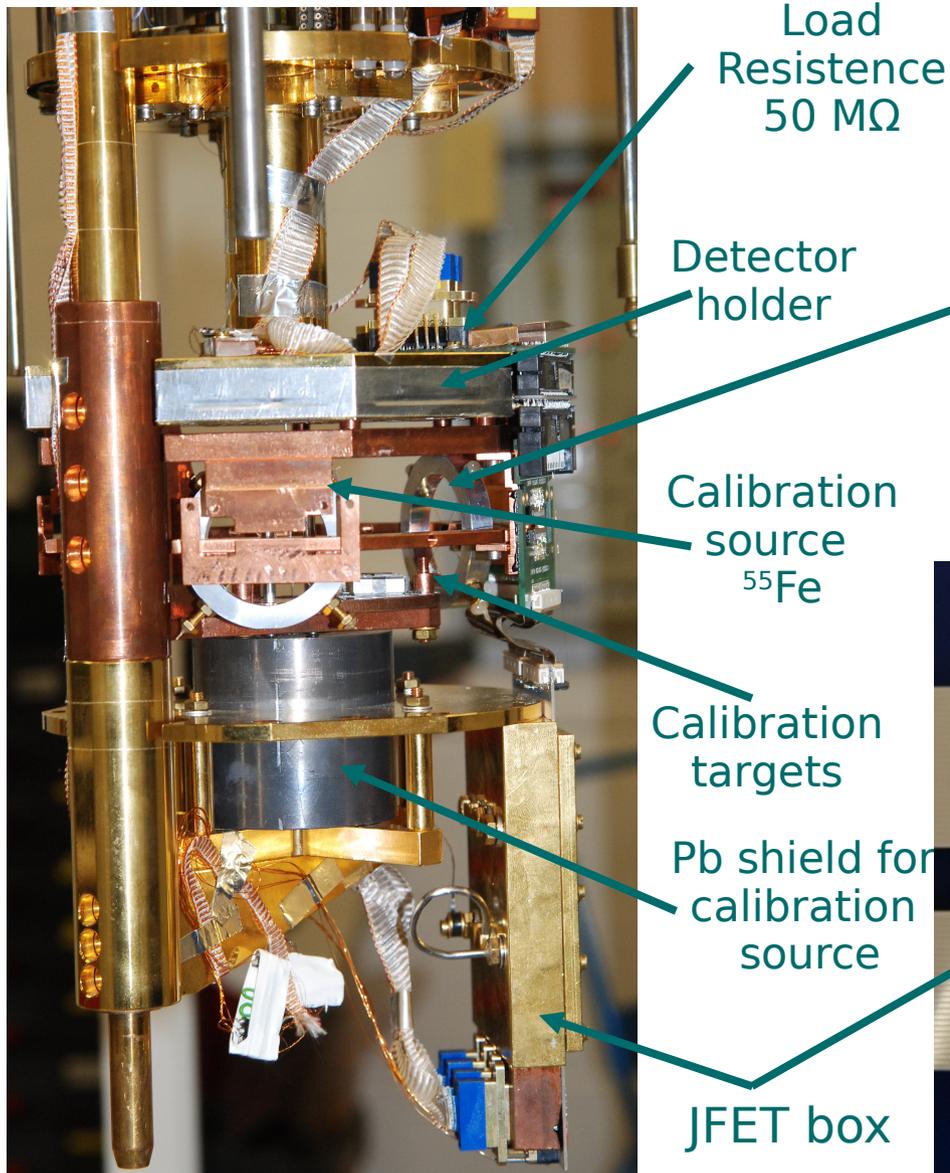
- 6x6 array of Si:P semiconductors (NASA-GSFC)

- pixel: $300 \times 300 \times 1.5$ μm^3
- high energy resolution
- developed for X-ray spectroscopy with HgTe absorber (ASTRO-E2)



Cryogenic set-up of MARE 1 @ Milano Bicocca

Kevlar crosses



MARE 1 @ Milano-Bicocca

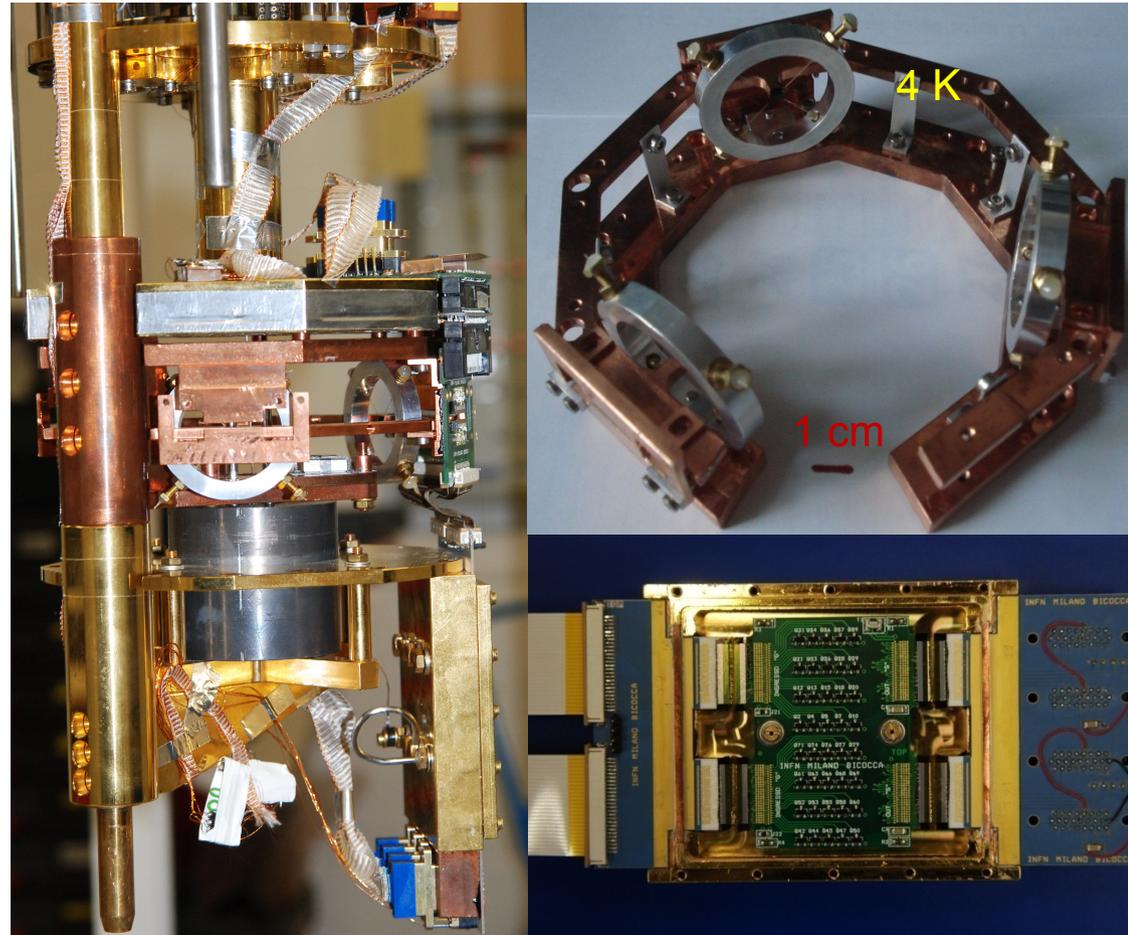
All the problems concerning the cryogenic set-up have been solved.



Thanks to the improvements added to the cryogenic set-up the detector target performances have been achieved.

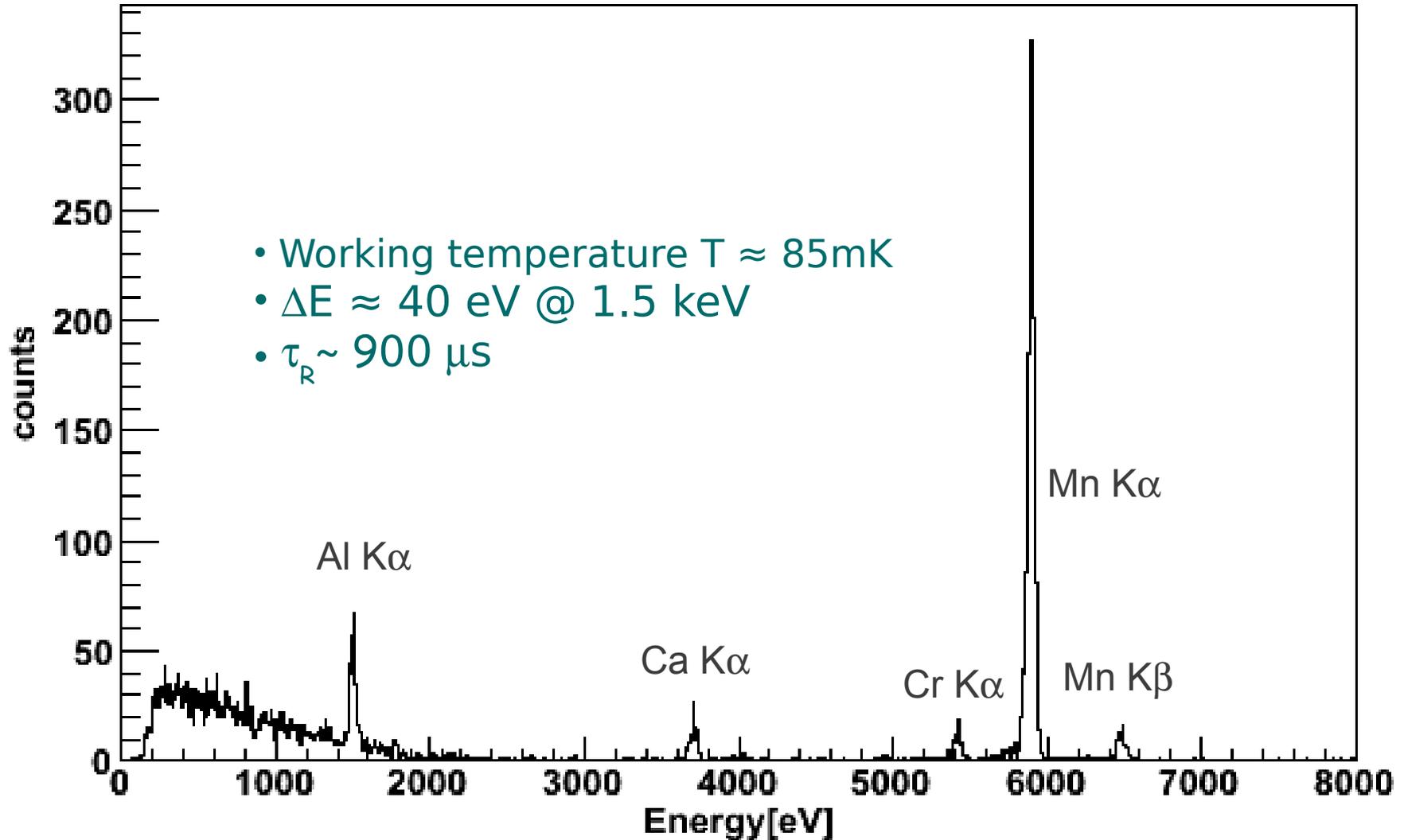


- First spectra acquired
- Completed assembly of the first array



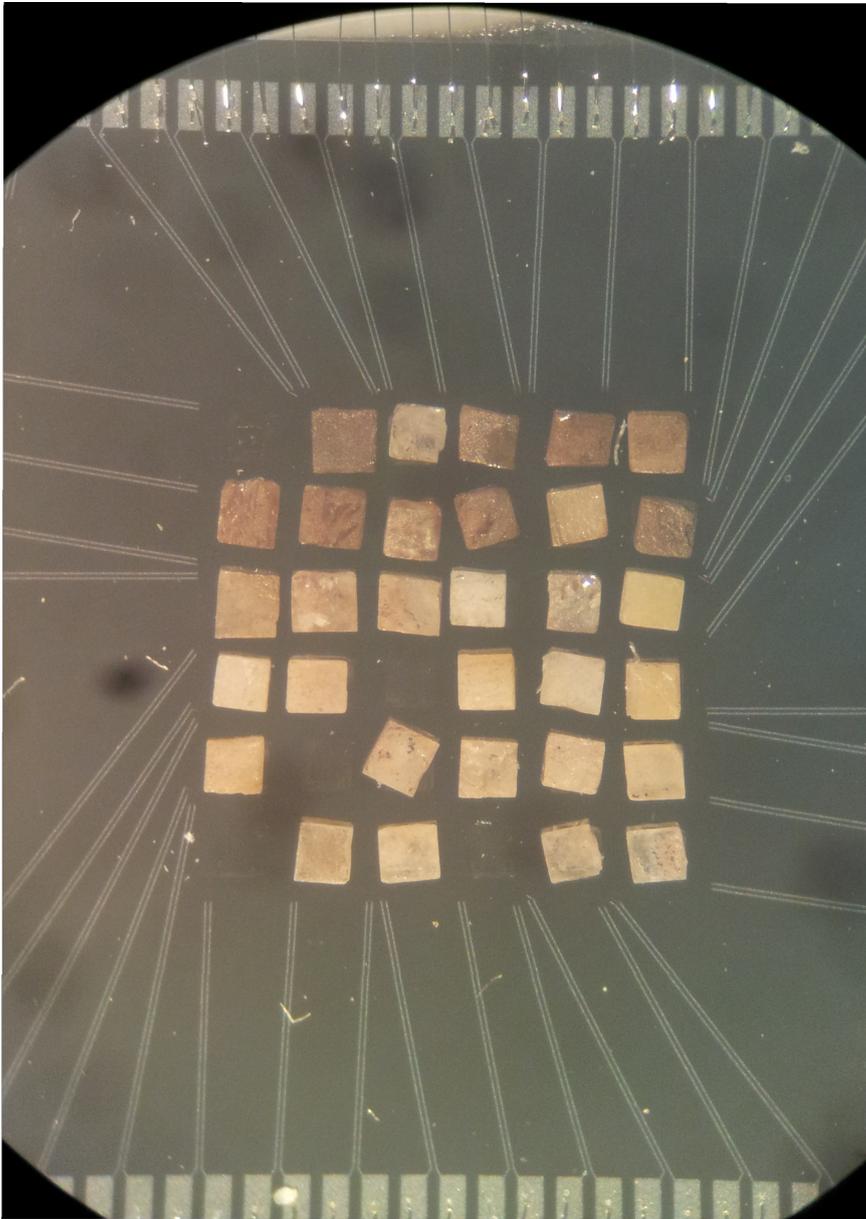
MARE 1 @ Milano-Bicocca

first spectrum acquired after the improvements added to MARE-1 cryogenic set-up



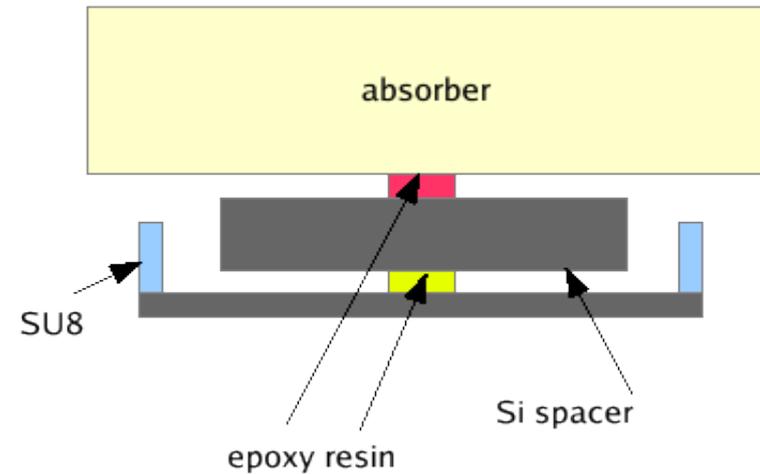
Measured 7 pixels: $\Delta E_{\text{ave}} \sim 30\text{eV @ } 1,5\text{keV}$

First array of MARE-1



Thermal coupling

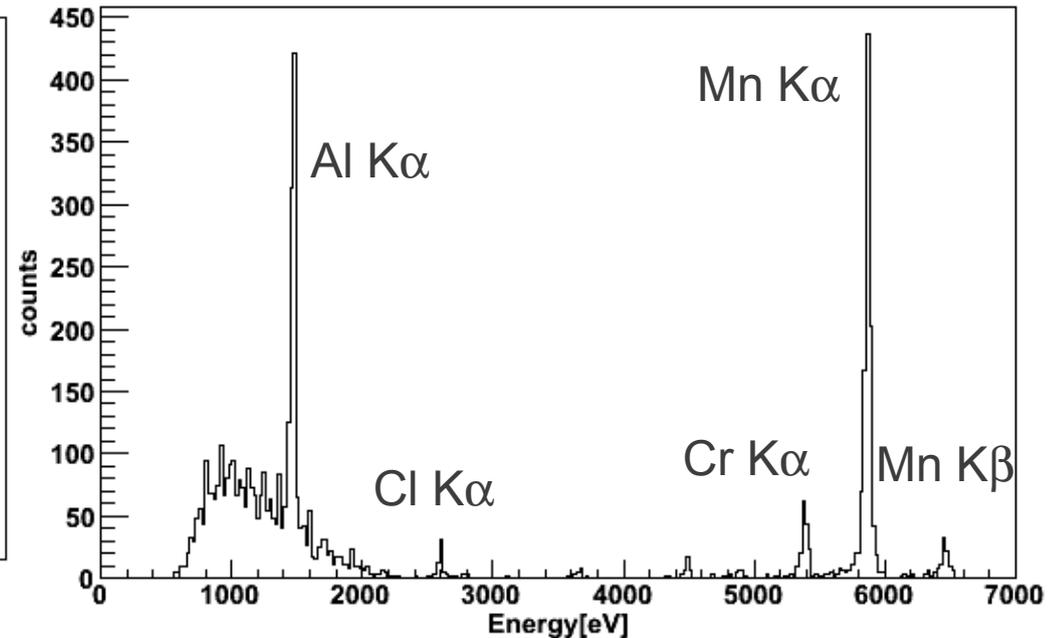
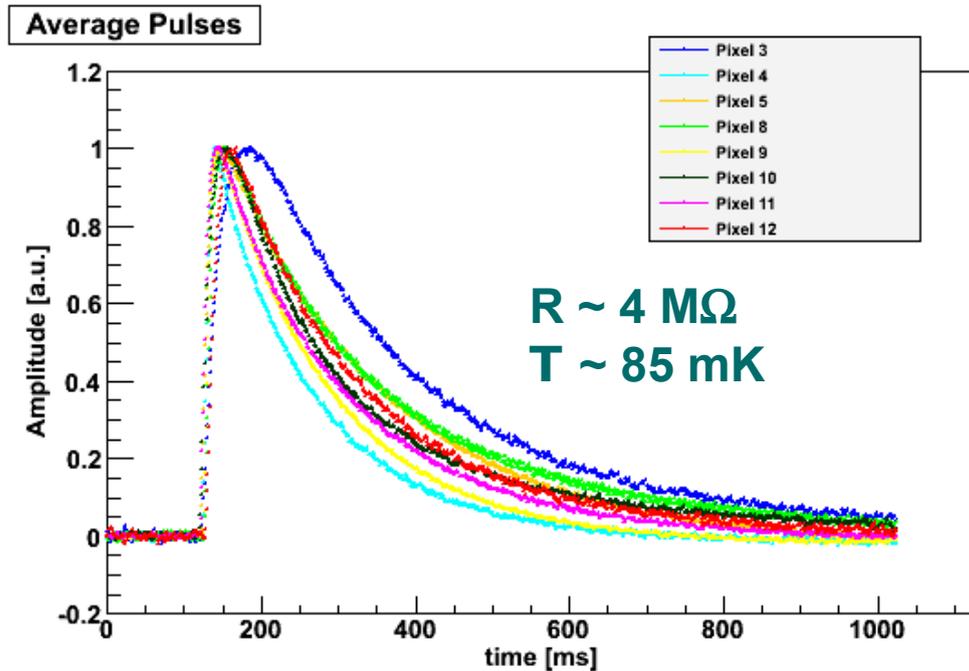
- **Araldit or ST1266:** thermistor/spacer
- **ST2850:** spacer/ AgReO_4



- 6 silicon spacers are attached with **Araldite Normal**
- 10 with **Araldite Rapid**
- 15 with **ST1266**

First array of MARE-1

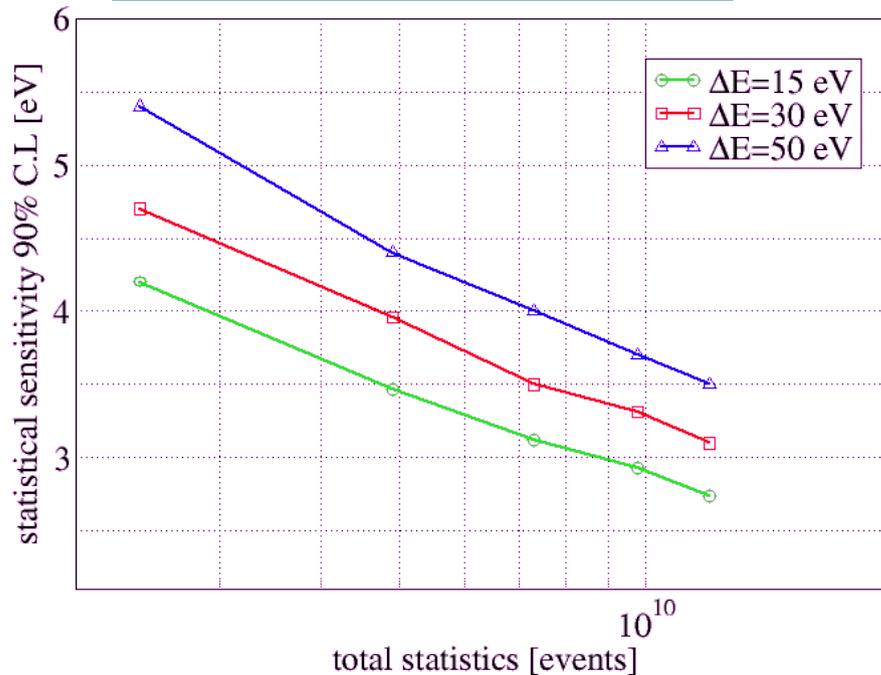
A run aimed to test the performance of this setup is ongoing, after which the absorbers will be glued also on the second array. With two arrays, a sensitivity of 4.5 eV at 90% C.L. is expected in three years running time.



- Working temperature $T \approx 88 \text{ mK}$
- $\Delta E \approx 24 \text{ eV @ } 1.5 \text{ keV}$
- $\tau_R \sim 1 \text{ ms}$

MARE 1 in Milano: sensitivity

MonteCarlo approach



- setup designed to host up to 8 arrays
- 288 AgReO_4 crystals
- start with 2 arrays (72 ch.)
- gradual deployment



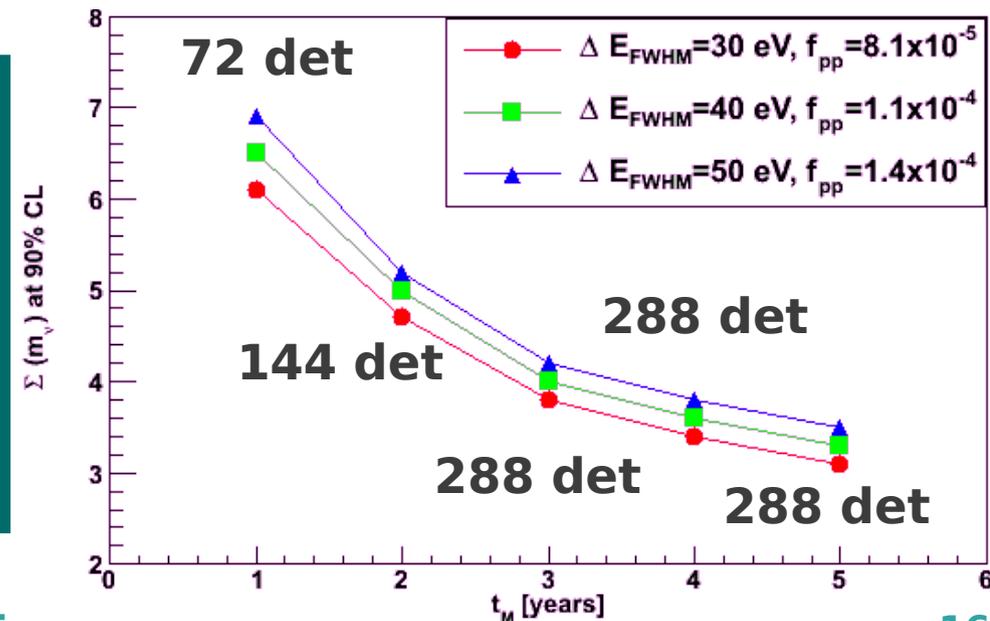
Since only two arrays are available up to now, it is useful to estimate the sensitivity on neutrino mass over the years by increasing the detectors number from year to year.

Analytic approach (1st order)

Detectors

$\Delta E_{\text{FWHM}} \sim 50$ eV and $\tau_R \sim 500$ μs
 1 year and 72 channels $\rightarrow \Sigma(m_\nu) \sim 7\text{eV}$
 3 years and 288 channels $\rightarrow \Sigma(m_\nu) \sim 4.2\text{eV}$

$\Delta E_{\text{FWHM}} \sim 30$ eV and $\tau_R \sim 300$ μs
 1 year and 72 channels $\rightarrow \Sigma(m_\nu) \sim 6\text{eV}$
 3 years and 288 channels $\rightarrow \Sigma(m_\nu) \sim 3.8\text{eV}$

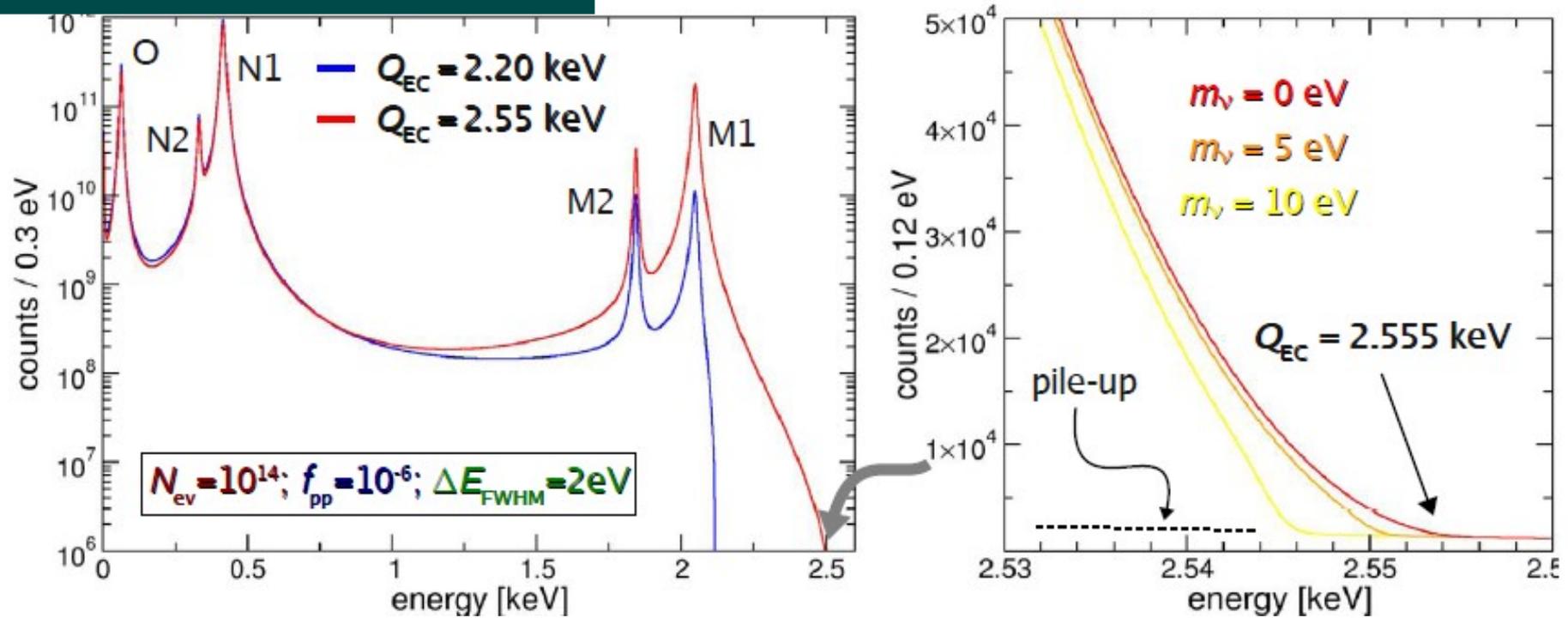


MARE extensions: ^{163}Ho EC measurement

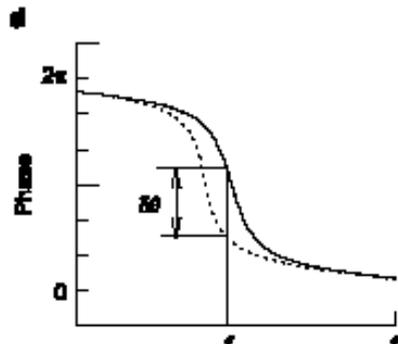
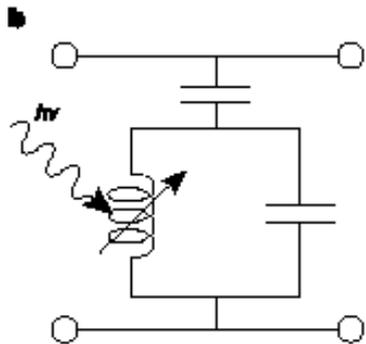
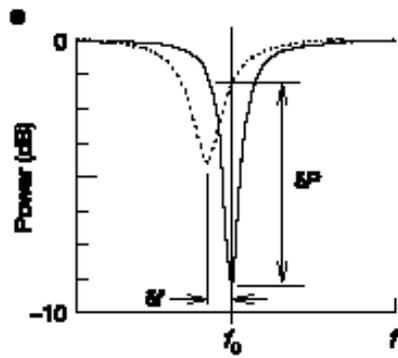
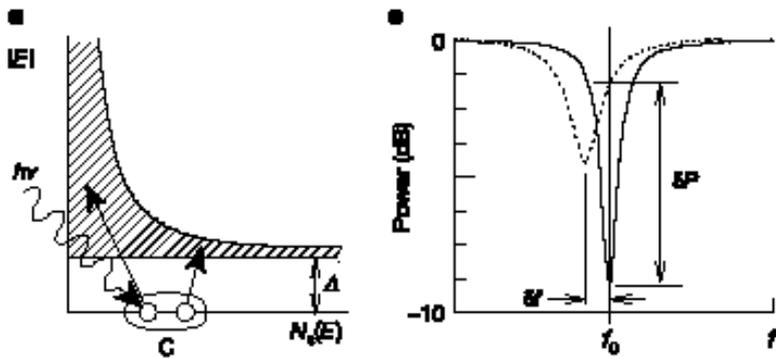


electron capture from shell \geq M1

A. De Rujula and M. Lusignoli, Phys. Lett. B 118 (1982) 429



- Calorimetric measurement of non-radiative Dy atomic de-excitations
- Breit Wigner M,N,O lines have an end-point at the Q value
- rate at end-point may be as high as for ^{187}Re but depends on Q_{EC}
 - Q_{EC} ? Measured: $Q_{\text{EC}} = 2.3 \div 2.8$ keV. Recommended: $Q_{\text{EC}} = 2.555$ keV
- $\tau_{1/2} \approx 4570$ years: few active nuclei are needed
 - can be implanted in any suitable microcalorimeter absorber
- ^{163}Ho production by neutron irradiation of ^{162}Er enriched Er



- resonator exploiting the T dependence of inductance in a superconducting film
 - **detectors** suitable for large absorbers
 - **Good time resolution** (low pile-up f_{pp})
 - **high energy resolution**
 - **multiplexing** for very large number of pixel

Sensitivity

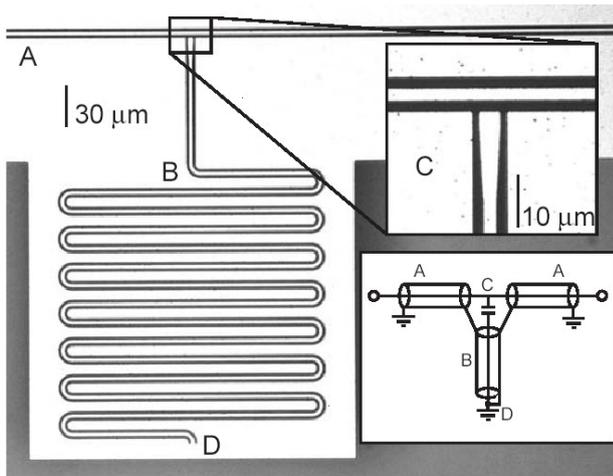
$$\Delta E = 5 \text{ eV}$$

$$t_M = 36000 \text{ detectors} \times 3 \text{ years}$$

$$A_\beta = 20 \text{ c/s/det}$$

$$\tau_{\text{rise}} = 1 \mu\text{s} \Rightarrow m_\nu < 0.2 \text{ eV}$$

$$\tau_{\text{rise}} = 100 \mu\text{s} \Rightarrow m_\nu < 0.4 \text{ eV}$$



application to bulky absorber still requires further efforts

MKIDs for ^{163}Ho EC decay end point measurement



4-12 GHz
cryo amp

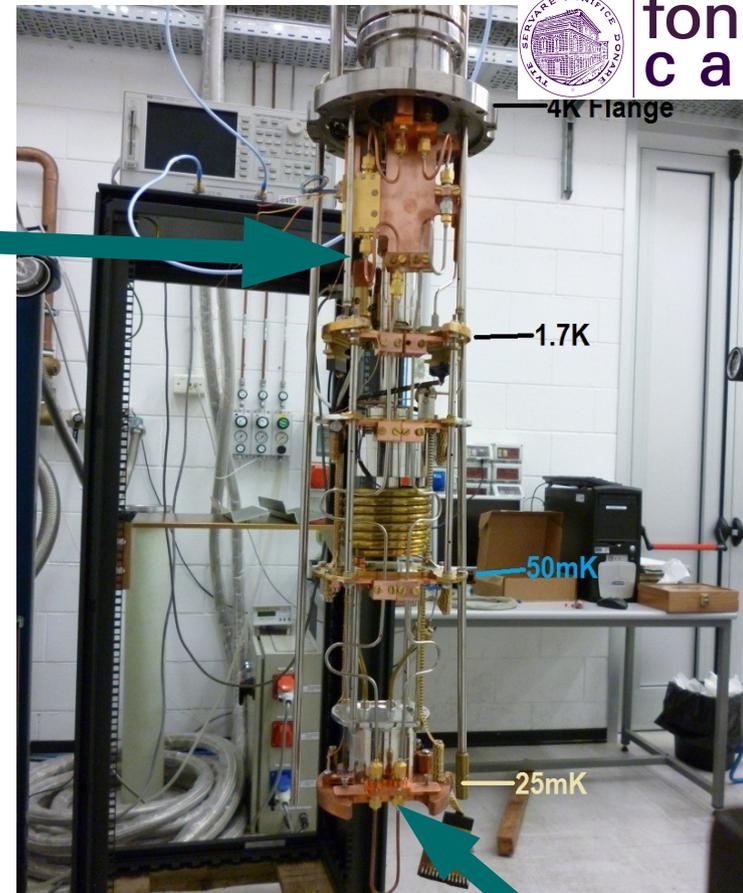
So far tested stoichiometric TiN ($T_c = 4,6\text{K}$) films and Ti/TiN multilayer (produced by FBK), which behaves like a sub-stoichiometric TiN film ($T_c = 1,6\text{K}$)

Gap parameter:

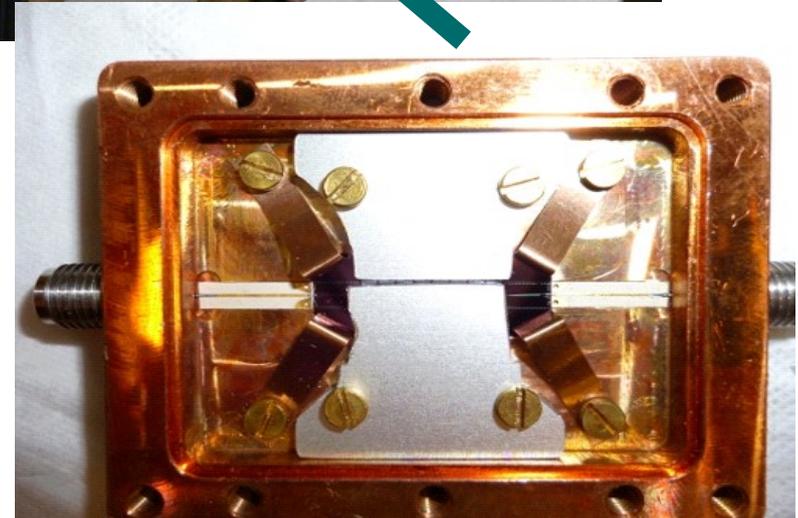
- TiN 0.8 meV
- Ti/TiN 0.26 meV

The devices were tested with ^{55}Fe (6keV) and Al X-ray (1,5keV) and the first pulses were acquired

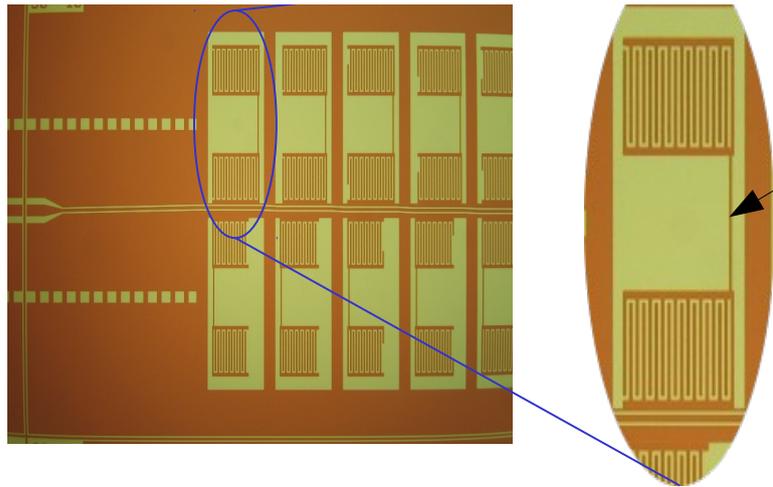
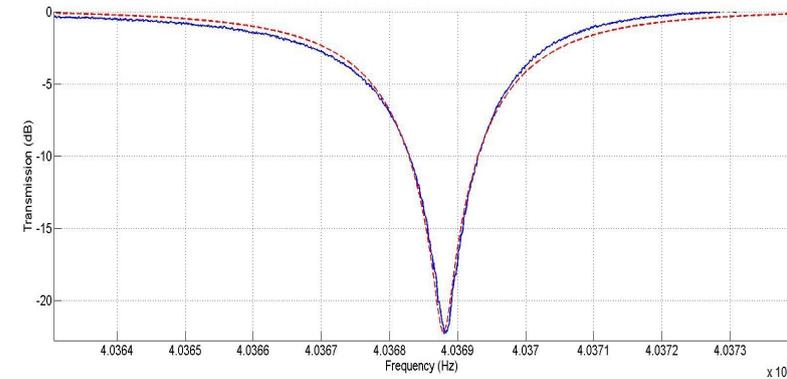
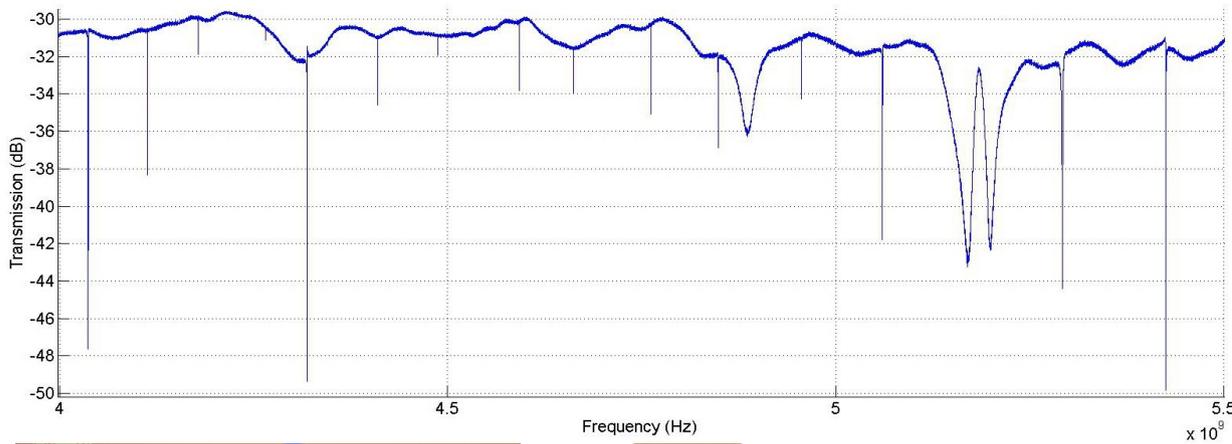
Not resolving yet because of events interacting in the Si substrate under the superconductor



fondazione
cariplo



MKIDs for ^{163}Ho EC decay end point



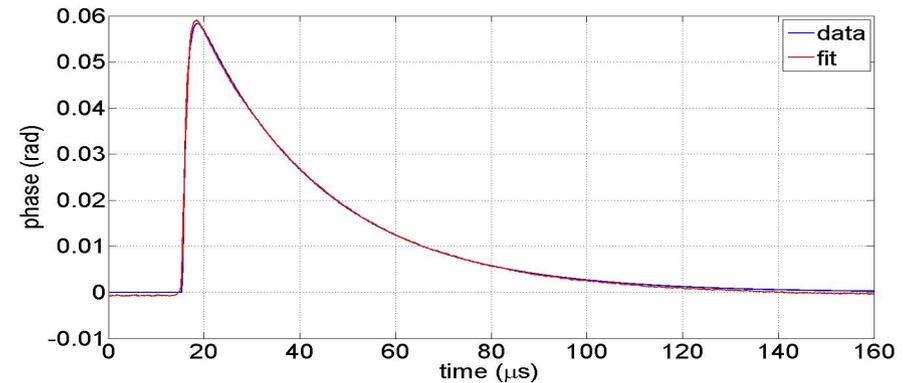
The ^{163}Ho will be embedded in the center of the inductive part of the resonator, deep enough to ensure low escape probability. A thickness of $<500\text{nm}$ will be enough

10^{12} Ho nuclei are needed for a count rate of 10 Hz

theoretical resolution

$$\Delta E_{\text{th}} = 2\text{keV}/N_{\text{qp}}^{1/2} = 1.5 \text{ eV}$$

This work is supported by Fondazione Cariplo through the project "Development of Microresonator Detectors for Neutrino Physics" (grant 2010-2351).



Conclusion

First array of MARE-1 has been assembled

→ 31 thermistors are equipped with AgReO₄ absorbers

The goal performances of the detectors have been achieved

→ first spectra were acquired obtaining a resolution of $\sim 28\text{eV}$ @ $1,5\text{keV}$

The next step will be to assemble the detectors on the second array

→ start of data taking in the next months

In the meanwhile new detector technology under investigation

→ Ho EC measurement with MKDs