## A Clockwork Tale

#### Daniele Teresi

daniele.teresi@ulb.ac.be

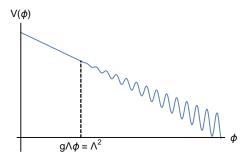
Service de Physique Théorique, Université Libre de Bruxelles, Belgium

MAX-PLANCK-INSTITUT FÜR KERNPHYSIK, HEIDELBERG, 06/11/17

# Prologue: Making numbers small dynamically

#### Naturalness from a relaxion [Graham, Kaplan, Rajendran, '15]

- slow-rolling pseudo NG boson: **relaxion**  $\phi$ , slope from  $g\Lambda^3\phi$
- relaxion-dependent H mass term:  $(\Lambda^2 g\Lambda\phi)H^\dagger H$
- backreaction when H vev:  $f_{\pi}^2 m_{\pi}^2(v) \cos(\phi/f)$
- $\implies v \ll \Lambda$  dynamically selected:



Daniele Teresi A Clockwork Tale 1/26

#### **Clockworking the relaxion**

- it requires tiny g, e.g.  $g = O(v/\Lambda)^4 \approx 10^{-50}$
- technically natural, NG shift symmetry for  $g \to 0$ , but still . . .
- it requires trans-planckian  $\Delta \phi$
- Solution: [Choi, Im , '15; Kaplan, Rattazzi, '15]
- g from much larger period  $F \gg f$ :

$$-\mathcal{L} \supset \left[\Lambda^2 - \Lambda^2 \cos\left(\frac{\phi}{F} + \alpha\right)\right] H^{\dagger} H - \Lambda^4 \cos\left(\frac{\phi}{F} + \alpha\right) - m_{BR}^4(v) \cos\frac{\phi}{f}$$

$$\Longrightarrow g = \Lambda/F$$

•  $F = 3^N f$  from **clockwork** chain:

$$-\mathcal{L} \supset \epsilon \left( \Phi_0^{\dagger} \Phi_1^3 + \Phi_1^{\dagger} \Phi_2^3 + \ldots + \Phi_{N-1}^{\dagger} \Phi_N^3 \right) + \frac{\phi_1}{\hat{f}} G \widetilde{G} + \frac{\phi_N}{\hat{f}} G \widetilde{G}$$

Daniele Teresi A Clockwork Tale 2 / 26

#### Clockworking the relaxion

- it requires tiny g, e.g.  $g = O(v/\Lambda)^4 \approx 10^{-50}$
- technically natural, NG shift symmetry for  $g \to 0$ , but still . . .
- it requires trans-planckian  $\Delta \phi$
- Solution: [Choi, Im , '15; Kaplan, Rattazzi, '15]
- g from much larger period  $F \gg f$ :

$$-\mathcal{L} \supset \left[\Lambda^2 - \Lambda^2 \cos\left(\frac{\phi}{F} + \alpha\right)\right] H^{\dagger} H - \Lambda^4 \cos\left(\frac{\phi}{F} + \alpha\right) - m_{BR}^4(v) \cos\frac{\phi}{f}$$

$$\implies g = \Lambda/F$$

•  $F = 3^N f$  from **clockwork** chain:

$$-\mathcal{L} \supset \epsilon \left(\Phi_0^\dagger \Phi_1^3 + \Phi_1^\dagger \Phi_2^3 + \ldots + \Phi_{N-1}^\dagger \Phi_N^3 
ight) \ + \ rac{\phi_1}{\hat{f}} \, G \, \widetilde{G} \ + \ rac{\phi_N}{\hat{f}} \, G \, \widetilde{G}$$

Daniele Teresi A Clockwork Tale 2 / 26

#### Get rid of the relaxion, keep the clockwork

- the last step itself is sufficient to generate hierarchies! [Giudice, McCullough, '16]
- clockwork mechanism  $\rightarrow$  an elegant and economical way to generate tiny numbers/large hierarchies X with only  $\mathcal{O}(1)$  couplings and  $N \sim \log X$  fields
- a framework for model building: [Giudice, McCullough, '16; Giudice, McCullough, DT, in prep.]
  - low-scale invisible axions [Giudice, McCullough, '16; Farina, Pappadopulo, Rompineve, Tesi, '16]
  - hierarchy problem [Giudice, McCullough, '16]
  - neutrino physics [Hambye, DT, Tytgat, '16; Carena, Li, Machado × 2, Wagner, '17]
  - inflation [Kehagias, Riotto, '16]
  - SUSY [Giudice, McCullough, DT, in prep.] and SUGRA [Kehagias, Riotto, '17; Antoniadis, Delgado, Markou, Pokorski, '17]
  - ... [not cited here for brevity]
  - dark matter [Hambye, DT, Tytgat, '16] (used in this talk to explain main features)
- dark matter cosmologically stable if decays by dim-5 ( $\Lambda \gg M_{PL}$ ), dim-6 ( $\Lambda \sim M_{GUT}$ ), tiny couplings  $\Longrightarrow$  all difficult to test
- clockwork mechanism → dark matter cosmologically stable although it decays into SM via O(1) interactions with TeV-scale particles!
- large interactions  $\Longrightarrow$  dark matter is a thermal relic, i.e. a WIMP

Daniele Teresi A Clockwork Tale 3/26

#### Get rid of the relaxion, keep the clockwork

- the last step itself is sufficient to generate hierarchies! [Giudice, McCullough, '16]
- clockwork mechanism  $\rightarrow$  an elegant and economical way to generate tiny numbers/large hierarchies X with only  $\mathcal{O}(1)$  couplings and  $N \sim \log X$  fields
- a framework for model building: [Giudice, McCullough, '16; Giudice, McCullough, DT, in prep.]
  - low-scale invisible axions [Giudice, McCullough, '16; Farina, Pappadopulo, Rompineve, Tesi, '16]
  - hierarchy problem [Giudice, McCullough, '16]
  - neutrino physics [Hambye, DT, Tytgat, '16; Carena, Li, Machado × 2, Wagner, '17]
  - inflation [Kehagias, Riotto, '16]
  - SUSY [Giudice, McCullough, DT, in prep.] and SUGRA [Kehagias, Riotto, '17; Antoniadis, Delgado, Markou, Pokorski, '17]
  - ... [not cited here for brevity]
  - dark matter [Hambye, DT, Tytgat, '16] (used in this talk to explain main features)
- dark matter cosmologically stable if decays by dim-5 ( $\Lambda \gg M_{PL}$ ), dim-6 ( $\Lambda \sim M_{GUT}$ ), tiny couplings  $\Longrightarrow$  all difficult to test
- clockwork mechanism → dark matter cosmologically stable although it decays into SM via O(1) interactions with TeV-scale particles!
- large interactions  $\Longrightarrow$  dark matter is a thermal relic, i.e. a WIMP

#### Get rid of the relaxion, keep the clockwork

- the last step itself is sufficient to generate hierarchies! [Giudice, McCullough, '16]
- clockwork mechanism  $\rightarrow$  an elegant and economical way to generate tiny numbers/large hierarchies X with only  $\mathcal{O}(1)$  couplings and  $N \sim \log X$  fields
- a framework for model building: [Giudice, McCullough, '16; Giudice, McCullough, DT, in prep.]
  - low-scale invisible axions [Giudice, McCullough, '16; Farina, Pappadopulo, Rompineve, Tesi, '16]
  - hierarchy problem [Giudice, McCullough, '16]
  - neutrino physics [Hambye, DT, Tytgat, '16; Carena, Li, Machado × 2, Wagner, '17]
  - inflation [Kehagias, Riotto, '16]
  - SUSY [Giudice, McCullough, DT, in prep.] and SUGRA [Kehagias, Riotto, '17; Antoniadis, Delgado, Markou, Pokorski, '17]
  - ... [not cited here for brevity]
  - dark matter [Hambye, DT, Tytgat, '16] (used in this talk to explain main features)
- dark matter cosmologically stable if decays by dim-5 ( $\Lambda \gg M_{PL}$ ), dim-6 ( $\Lambda \sim M_{GUT}$ ), tiny couplings  $\Longrightarrow$  all difficult to test
- clockwork mechanism → dark matter cosmologically stable although it decays into SM via O(1) interactions with TeV-scale particles!
- large interactions  $\Longrightarrow$  dark matter is a thermal relic, i.e. a WIMP

### Chapter 1:

How to do multiplications in QFT

#### The clockwork mechanism

#### Based on the simple observation that:

$$1/2 \times 1/2 \times 1/2 \times 1/2 \times 1/2 \times \dots \times 1/2$$
 can easily be tiny

Use a chain of N fields

$$\phi_0 \stackrel{1/q}{=} \phi_1 \stackrel{1/q}{=} \phi_2 \stackrel{1/q}{=} \phi_3 \stackrel{1/q}{=} \dots \stackrel{1/q}{=} \phi_N \longrightarrow SN$$

if clever symmetry 
$$\longrightarrow$$
  $\phi_{light} \approx \phi_0 \implies \phi_{light} - SM \sim 1/q^N \quad (q > 1)$ 

For **fermions** use chiral symmetries

$$R_0 \stackrel{m}{=} \underbrace{L_1 \quad R_1}_{qm} \stackrel{m}{=} \underbrace{L_2 \quad R_2}_{qm} \stackrel{m}{=} \underbrace{L_3 \quad R_3}_{qm} \stackrel{m}{=} \dots \stackrel{m}{=} \underbrace{L_N \quad R_N}_{qm} \stackrel{}{=} \underbrace{L_{SM}}_{qm}$$

light 
$$N \approx R_0 \implies N - L_{SM} \sim 1/q^N$$

Daniele Teresi A Clockwork Tale 4 / 26

#### The clockwork mechanism

Based on the simple observation that:

$$1/2 \times 1/2 \times 1/2 \times 1/2 \times 1/2 \times \dots \times 1/2$$
 can easily be tiny

Use a chain of N fields

$$\phi_0 = \frac{1/q}{q} \phi_1 = \frac{1/q}{q} \phi_2 = \frac{1/q}{q} \phi_3 = \frac{1/q}{q} \cdots = \frac{1/q}{q} \phi_N = SN$$

if clever symmetry 
$$\longrightarrow$$
  $\phi_{light} \approx \phi_0 \implies \phi_{light} - \text{SM} \sim 1/\text{q}^{\text{N}}$   $(q > 1)$ 

For **fermions** use chiral symmetries

$$R_0 \stackrel{m}{=} \underbrace{L_1 \quad R_1}_{qm} \stackrel{m}{=} \underbrace{L_2 \quad R_2}_{qm} \stackrel{m}{=} \underbrace{L_3 \quad R_3}_{qm} \stackrel{m}{=} \dots \stackrel{m}{=} \underbrace{L_N \quad R_N}_{qm} \stackrel{}{=} \underbrace{L_{SM}}_{qm}$$

light 
$$N \approx R_0 \implies N - L_{SM} \sim 1/q^N$$

Daniele Teresi A Clockwork Tale 4 / 26

#### The clockwork mechanism

Based on the simple observation that:

$$1/2 \times 1/2 \times 1/2 \times 1/2 \times 1/2 \times \dots \times 1/2$$
 can easily be tiny

Use a chain of N fields

$$\phi_0 = \frac{1/q}{q} \phi_1 = \frac{1/q}{q} \phi_2 = \frac{1/q}{q} \phi_3 = \frac{1/q}{q} \cdots = \frac{1/q}{q} \phi_N = SN$$

if clever symmetry  $\longrightarrow$   $\phi_{light} \approx \phi_0 \implies \phi_{light} - \text{SM} \sim 1/\text{q}^{\text{N}}$  (q > 1)

For **fermions** use chiral symmetries

$$R_0 \stackrel{m}{=} \underbrace{L_1 \quad R_1}_{qm} \stackrel{m}{=} \underbrace{L_2 \quad R_2}_{qm} \stackrel{m}{=} \underbrace{L_3 \quad R_3}_{qm} \stackrel{m}{=} \cdots \stackrel{m}{=} \underbrace{L_N \quad R_N}_{qm} \stackrel{L}{=} \underbrace{L_{SM}}_{qm}$$

light 
$$N \approx R_0 \implies N - L_{SM} \sim 1/q^N$$

Daniele Teresi A Clockwork Tale 4 / 26

#### Clockwork scalar

- For scalars, use a chain of N+1 symmetries:  $U(1)_0 \times U(1)_1 \times ... \times U(1)_N$
- broken by N spurions  $m_k^2 \equiv m^2$  with  $Q_k(m_k^2) = 1$ ,  $Q_{k+1}(m_k^2) = -q$  (q > 1)
- $\bullet \ \mathcal{L} = -\frac{f^2}{2} \sum_{k=0}^{N} |\partial U_k|^2 + \frac{m^2 f^2}{2} \sum_{k=0}^{N-1} \left( U_k^{\dagger} U_{k+1}^q + h.c. \right)$
- for the Goldstones  $\phi_k$ ,  $U_k \propto e^{i\phi_k/f}$ :  $-\mathcal{L} \supset \frac{m^2}{2} \sum_{k=0}^{N-1} \left(\phi_k q\phi_{k+1}\right)^2$
- ullet unbroken U(1) with  $\mathcal{Q} = \sum_k rac{\mathcal{Q}_k}{q^k} \quad \Longrightarrow \quad ext{massless } arphi_0 = \mathcal{N} \sum_k rac{\phi_k}{q^k}$
- For instance, if  $\mathcal{L} \supset \frac{\phi_N}{16\pi^2 f} \, G \, \widetilde{G} \quad \Longrightarrow \quad \frac{\varphi_0}{16\pi^2 F} \, G \, \widetilde{G} \quad \text{with} \quad F = f \, \frac{q^N}{\mathcal{N}} \gg f$

Daniele Teresi A Clockwork Tale 5 / 26

#### Clockwork scalar

- For scalars, use a chain of N+1 symmetries:  $U(1)_0 \times U(1)_1 \times ... \times U(1)_N$
- broken by N spurions  $m_k^2 \equiv m^2$  with  $Q_k(m_k^2) = 1$ ,  $Q_{k+1}(m_k^2) = -q$  (q > 1)
- $\bullet \ \mathcal{L} = -\frac{f^2}{2} \sum_{k=0}^{N} |\partial U_k|^2 + \frac{m^2 f^2}{2} \sum_{k=0}^{N-1} \left( U_k^{\dagger} U_{k+1}^q + h.c. \right)$
- for the Goldstones  $\phi_k$ ,  $U_k \propto e^{i\phi_k/f}$ :  $-\mathcal{L} \supset \frac{m^2}{2} \sum_{k=0}^{N-1} \left(\phi_k q\phi_{k+1}\right)^2$
- unbroken U(1) with  $\mathcal{Q} = \sum_k \frac{\mathcal{Q}_k}{q^k} \implies \text{massless } \varphi_0 = \mathcal{N} \sum_k \frac{\phi_k}{q^k}$
- For instance, if  $\mathcal{L} \supset \frac{\phi_N}{16\pi^2 f} \, G \, \widetilde{G} \quad \Longrightarrow \quad \frac{\varphi_0}{16\pi^2 F} \, G \, \widetilde{G} \quad \text{with} \quad F = f \, \frac{q^N}{\mathcal{N}} \gg f$

Daniele Teresi A Clockwork Tale 5 / 26

#### Clockwork dark matter [Hambye, DT, Tytgat, '16]

chiral symmetry group:

$$U(1)_{R_0}\times U(1)_{L_1}\times U(1)_{R_1}\times \ldots \times U(1)_{L_N}\times U(1)_{R_N} \quad \text{ with } \quad U(1)_{R_N}\equiv U(1)_{L_{SM}}$$

scalars:

$$S_i \sim (-1,1) \text{ under } U(1)_{R_i} \times U(1)_{L_{i+1}}$$
  $C_i \sim (1,-1) \text{ under } U(1)_{L_i} \times U(1)_{R_i}$ 

chain of fields:

$$\mathbf{R}_0 \stackrel{S_1}{=} L_1 \stackrel{C_1}{=} R_1 \stackrel{S_2}{=} L_2 \stackrel{C_2}{=} \dots \stackrel{C_N}{=} R_N$$

• clockwork mechanism when scalars acquire a vev:

$$m = v_S \langle S_i \rangle$$
  $qm = v_C \langle C_i \rangle$ 

• Majorana mass  $m_N$  for  $R_0$ , eigenstate  $N \approx R_0$  is the dark-matter candidate

Daniele Teresi A Clockwork Tale 6 / 26

#### Clockwork dark matter [Hambye, DT, Tytgat, '16]

chiral symmetry group:

$$U(1)_{R_0}\times U(1)_{L_1}\times U(1)_{R_1}\times \ldots \times U(1)_{L_N}\times U(1)_{R_N} \quad \text{ with } \quad U(1)_{R_N}\equiv U(1)_{L_{SM}}$$

scalars:

$$S_i \sim (-1,1) \text{ under } U(1)_{R_i} \times U(1)_{L_{i+1}}$$
  $C_i \sim (1,-1) \text{ under } U(1)_{L_i} \times U(1)_{R_i}$ 

chain of fields:

$$\mathbf{R_0} \stackrel{S_1}{=} L_1 \stackrel{C_1}{=} R_1 \stackrel{S_2}{=} L_2 \stackrel{C_2}{=} \dots \stackrel{C_N}{=} R_N \stackrel{\mathbf{L_{SM}}}{=} \mathbf{L_{SM}}$$

clockwork mechanism when scalars acquire a vev:

$$m = y_S \langle S_i \rangle$$
  $qm = y_C \langle C_i \rangle$ 

• Majorana mass  $m_N$  for  $R_0$ , eigenstate  $N \approx R_0$  is the dark-matter candidate

Daniele Teresi A Clockwork Tale 6 / 26

#### **Clockwork fermion**

the Lagrangian is

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{kinetic} - \sum_{i=1}^{N} (y_S S_i \bar{L}_i R_{i-1} - y_C C_i \bar{L}_i R_i + h.c.)$$
$$- (y \bar{L}_{SM} \tilde{H} R_N + h.c.) - \frac{1}{2} (m_N \overline{R_0^c} R_0 + h.c.)$$

• after the scalars acquire vevs  $m = y_S \langle S_i \rangle$ ,  $qm = y_C \langle C_i \rangle$ :

$$\mathcal{L}\supset -m\sum_{i=1}^N\left(ar{L}_iR_{i-1}-q\,ar{L}_iR_i
ight)-rac{m_N}{2}\,\overline{R_0^c}\,R_0+h.c.$$

- for  $m_N=0$ , the "right-handed" mass matrix satisfies  $M^\dagger M\equiv M_{scalar}^2$
- clockwork mechanism for  $m_N \lesssim qm$  (for  $q \gg 1$ )

Daniele Teresi A Clockwork Tale 7 / 26

Chapter 2:

How clockwork matter became dark

#### The spectrum

#### Take $q \gg 1$ for simplicity

• the dark-matter Majorana fermion N with mass  $\approx m_N$ :

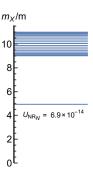
$$N \approx R_0 + \frac{1}{q^1}R_1 + \frac{1}{q^2}R_2 + \ldots + \frac{1}{q^N}R_N$$

• a band of N pseudo-Dirac  $\psi_i$  with mass  $\approx qm$ :

$$\psi_i pprox rac{1}{\sqrt{N}} \sum_k \mathcal{O}(1) L_k + \mathcal{O}(1) R_k$$

 N scalars S<sub>i</sub> and C<sub>i</sub> expected in the same mass range (not necessarily dynamic, but not discussed here)

N = 15, q = 10., 
$$m_N/m = 5.0$$



Relevant sizeable interactions:

Daniele Teresi A Clockwork Tale 8 / 26

#### The spectrum

Take  $q \gg 1$  for simplicity

• the dark-matter Majorana fermion N with mass  $\approx m_N$ :

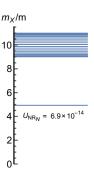
$$N \approx R_0 + \frac{1}{q^1}R_1 + \frac{1}{q^2}R_2 + \dots + \frac{1}{q^N}R_N$$

• a band of N pseudo-Dirac  $\psi_i$  with mass  $\approx qm$ :

$$\psi_i pprox rac{1}{\sqrt{N}} \sum_k \mathcal{O}(1) L_k + \mathcal{O}(1) R_k$$

 N scalars S<sub>i</sub> and C<sub>i</sub> expected in the same mass range (not necessarily dynamic, but not discussed here)

$$N = 15$$
,  $q = 10$ .,  $m_N/m = 5.0$ 



Relevant sizeable interactions:

$$N \searrow S_1, h$$

$$N \sum_{N} S_{1}, h$$

#### Cosmological (meta)stability of dark matter

$$\mathbf{R}_0 \xrightarrow{y_S \langle S_1 \rangle} L_1 \xrightarrow{y_C \langle C_1 \rangle} R_1 \xrightarrow{y_S \langle S_2 \rangle} L_2 \xrightarrow{y_C \langle C_2 \rangle} \dots \xrightarrow{y_C \langle C_N \rangle} R_N \xrightarrow{y_h} \mathbf{L}_{SM}$$

N can **decay**, e.g.  $N \rightarrow \nu h, \nu Z, lW$ , but

The coupling of dark matter to SM fermions is clockwork suppressed:

$$\mathcal{L}\supset -rac{y}{q^N}ar{L}_{SM}\widetilde{H}N_R$$

#### Dark matter cosmologically stable

The decay lifetime of N longer than the age of the Universe with  $\mathcal{O}(1)$  couplings and  $\leq$  **TeV-scale** states

- indirect detection  $\Longrightarrow q^{2N} > 1.5 \times 10^{50} \left(\frac{m_N}{\text{GeV}}\right) y^2$  for example:  $m_N \sim 100 \text{ GeV}$ ,  $y \sim 1$ ,  $q \sim 10$ ,  $N \sim 26$
- effect of clockwork gears  $\psi_i$  in loop diagrams also clockwork-suppressed

Daniele Teresi A Clockwork Tale 9 / 26

#### Cosmological (meta)stability of dark matter

$$\mathbf{R}_0 \xrightarrow{y_S\langle S_1\rangle} L_1 \xrightarrow{y_C\langle C_1\rangle} R_1 \xrightarrow{y_S\langle S_2\rangle} L_2 \xrightarrow{y_C\langle C_2\rangle} \dots \xrightarrow{y_C\langle C_N\rangle} R_N \xrightarrow{y_h} \mathbf{L}_{SM}$$

N can **decay**, e.g.  $N \rightarrow \nu h, \nu Z, lW$ , but

The coupling of dark matter to SM fermions is clockwork suppressed:

$$\mathcal{L}\supset -\,rac{y}{q^N}\,ar{L}_{\mathit{SM}}\widetilde{H}N_{\mathit{R}}$$

#### Dark matter cosmologically stable

The decay lifetime of N longer than the age of the Universe with  $\mathcal{O}(1)$  couplings and  $\leq$  **TeV-scale** states

- indirect detection  $\Longrightarrow q^{2N} > 1.5 \times 10^{50} \left(\frac{m_N}{\text{GeV}}\right) y^2$  for example:  $m_N \sim 100 \text{ GeV}$ ,  $y \sim 1$ ,  $q \sim 10$ ,  $N \sim 26$
- effect of clockwork gears  $\psi_i$  in loop diagrams also clockwork-suppressed

Daniele Teresi A Clockwork Tale 9 / 26

#### Cosmological (meta)stability of dark matter

$$\mathbf{R}_0 \xrightarrow{y_S\langle S_1\rangle} L_1 \xrightarrow{y_C\langle C_1\rangle} R_1 \xrightarrow{y_S\langle S_2\rangle} L_2 \xrightarrow{y_C\langle C_2\rangle} \dots \xrightarrow{y_C\langle C_N\rangle} R_N \xrightarrow{y_h} \mathbf{L}_{SM}$$

N can **decay**, e.g.  $N \rightarrow \nu h, \nu Z, lW$ , but

The coupling of dark matter to SM fermions is clockwork suppressed:

$$\mathcal{L}\supset -\,rac{y}{q^N}\,ar{L}_{\mathit{SM}}\widetilde{H}N_{\mathit{R}}$$

#### Dark matter cosmologically stable

The decay lifetime of N longer than the age of the Universe with  $\mathcal{O}(1)$  couplings and  $\leq$  **TeV-scale** states

- indirect detection  $\Longrightarrow q^{2N} > 1.5 \times 10^{50} \left(\frac{m_N}{\text{GeV}}\right) y^2$  for example:  $m_N \sim 100 \text{ GeV}$ ,  $y \sim 1$ ,  $q \sim 10$ ,  $N \sim 26$
- effect of clockwork gears  $\psi_i$  in loop diagrams also clockwork-suppressed

Daniele Teresi A Clockwork Tale 9 / 26

#### Scenario A: $m_S < m_N$

#### Dominant process:



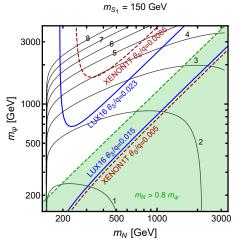
from  $N \sim R_0, \ \Psi_j \supset L_1$ and  $y_S S_1 \overline{L}_1 R_0$ 

#### not clockwork-suppressed!

⇒ N is a WIMP

perturbative  $y_S < \sqrt{4\pi} \simeq 3.5$ 

 $\Longrightarrow$  N and  $\psi_i$  light enough

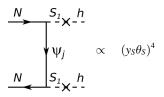


 $y_S$  needed for correct  $\Omega_{DM}$ 

Daniele Teresi A Clockwork Tale 10 / 26

#### Scenario B: $m_N < m_S$ and $2m_N < m_S + m_h$

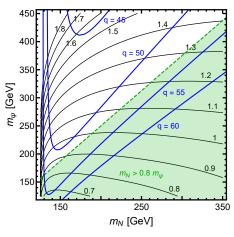
#### Dominant process:



#### $\theta_{\rm S} \lesssim 0.4$ from colliders

 $y_S$  non-perturbative for universal  $\theta_S$ :  $\theta_S \lesssim 0.4/\sqrt{N}$ 

it works also near the h and S resonances, for universal  $\theta_S$  too



 $y_S\theta_S$  needed for correct  $\Omega_{DM}$ 

Daniele Teresi A Clockwork Tale 11 / 26

#### Other limits and prospects

#### • Indirect detection: annihilation is p-wave, but decays $N \to h\nu$ monochromatic

- ullet  $\psi_j$  in the hundreds of GeV range, coupled via  $yL_{SM}HR_N$  and  $\psi_j\supset R_N$ 
  - ⇒ pseudo-Dirac RH neutrinos in the observable range, y sizeable
    - EWPT:  $|B_{l\psi}|^2 \equiv y^2 v^2/(2m_{gb}^2) \lesssim 10^{-3}$
    - LFV:  $BR(\mu \to e\gamma) \approx 8 \times 10^{-4} |B_{e\Psi}|^2 |B_{\mu\Psi}|^2 < 4.2 \times 10^{-13}$
    - direct L-conserving searches: up to  $m_{\psi} \approx 200$  GeV with 300 fb<sup>-1</sup> [Das, Dev, Okada, '14]
    - if  $m_N \ll m_\psi$  L-violating searches: up to  $m_\psi \approx 300$  GeV with 300 fb<sup>-1</sup> [Depoisch Dev. Pilattsis, '15]
- In scenario B  $S_1$  needs to have large mixing with h, in A it can
  - ⇒ limits and searches for scalar singlets [Falkowski, Gross, Lebedev, '15; Robens, Stefaniak, '15]
    - for  $m_S < 500 \,\mathrm{GeV}$ :  $\theta_S < 0.3 0.4 \,\mathrm{from}$  direct searches
    - for  $m_S > 500 \, \text{GeV}$ :  $\theta_S \lesssim 0.3 0.4 \, \text{from EWPT}$

Daniele Teresi A Clockwork Tale 12 / 26

#### Other limits and prospects

- Indirect detection: annihilation is p-wave, but decays  $N \to h \nu$  monochromatic
- $\psi_j$  in the hundreds of GeV range, coupled via  $y \, \overline{L}_{SM} HR_N$  and  $\psi_j \supset R_N$ 
  - ⇒ pseudo-Dirac RH neutrinos in the observable range, y sizeable
    - EWPT:  $|B_{l\psi}|^2 \equiv y^2 v^2/(2m_{\psi}^2) \lesssim 10^{-3}$
    - LFV:  $BR(\mu \to e\gamma) \approx 8 \times 10^{-4} |B_{e\Psi}|^2 |B_{\mu\Psi}|^2 < 4.2 \times 10^{-13}$
    - ullet direct L-conserving searches: up to  $m_\psi pprox 200$  GeV with 300 fb $^{-1}$  [Das, Dev, Okada, '14]
    - if  $m_N \not\ll m_\psi$  L-violating searches: up to  $m_\psi \approx 300$  GeV with 300 fb<sup>-1</sup> [Deppisch, Dev, Pilaftsis, '15]
- In scenario B  $S_1$  needs to have large mixing with h, in A it can
  - $\Longrightarrow$   $\mathsf{limits}$  and  $\mathsf{searches}$  for  $\mathsf{scalar}$   $\mathsf{singlets}$  [Falkowski, Gross, Lebedev, '15; Robens, Stefaniak, '15
    - for  $m_S < 500 \,\mathrm{GeV}$ :  $\theta_S < 0.3 0.4 \,\mathrm{from}$  direct searches
    - for  $m_S > 500 \, \text{GeV}$ :  $\theta_S \lesssim 0.3 0.4 \, \text{from EWPT}$

Daniele Teresi A Clockwork Tale 12 / 26

#### Other limits and prospects

- Indirect detection: annihilation is p-wave, but decays  $N \to h \nu$  monochromatic
- $\psi_j$  in the hundreds of GeV range, coupled via  $y \, \overline{L}_{SM} HR_N$  and  $\psi_j \supset R_N$ 
  - ⇒ pseudo-Dirac RH neutrinos in the observable range, y sizeable
    - EWPT:  $|B_{l\psi}|^2 \equiv y^2 v^2/(2m_{\psi}^2) \lessapprox 10^{-3}$
    - LFV:  $BR(\mu \to e\gamma) \approx 8 \times 10^{-4} |B_{e\Psi}|^2 |B_{\mu\Psi}|^2 < 4.2 \times 10^{-13}$
    - direct L-conserving searches: up to  $m_{\psi} \approx 200~{
      m GeV}$  with 300  ${
      m fb}^{-1}$  [Das, Dev, Okada, '14]
    - if  $m_N \not\ll m_\psi$  L-violating searches: up to  $m_\psi \approx 300$  GeV with 300 fb<sup>-1</sup> [Deppisch, Dev, Pilaftsis, '15]
- In scenario B  $S_1$  needs to have large mixing with h, in A it can
  - ⇒ limits and searches for scalar singlets [Falkowski, Gross, Lebedev, '15; Robens, Stefaniak, '15]
    - for  $m_S < 500 \,\mathrm{GeV}$ :  $\theta_S < 0.3 0.4 \,\mathrm{from}$  direct searches
    - for  $m_S > 500 \,\mathrm{GeV}$ :  $\theta_S \lesssim 0.3 0.4 \,\mathrm{from}$  EWPT

Daniele Teresi A Clockwork Tale 12 / 26

#### Majorana neutrino masses [Hambye, DT, Tytgat, '16]

- SM leptons interact with TeV-scale  $\psi_i$  with large Yukawas  $\Longrightarrow$  huge  $m_{\nu}$ ???
- Clockwork at work: if there were no  $R_0 \Longrightarrow$  no chiral partner for  $\nu$ s but effect of  $R_0$  has to go through the **whole clockwork chain**:

$$m_
u \simeq rac{m_D^2}{q^{2N} m_N}$$

- suppression here is smaller than for DM:  $q = 10, m_N = 1 \text{ TeV} \implies N \approx 7$
- > 2 nonzero  $m_{\nu} \Longrightarrow$  at least 2 clockwork chains
- a suggestive possibility: 1 chain for dark matter, 2 chains for neutrino masses
- + resonant leptogenesis? Starting from 2 degenerate, a small mass splitting can be generated by interactions with the clockwork gears...
- many model-building variants (not discussed here)

Daniele Teresi A Clockwork Tale 13 / 26

#### Majorana neutrino masses [Hambye, DT, Tytgat, '16]

- SM leptons interact with TeV-scale  $\psi_i$  with large Yukawas  $\Longrightarrow$  huge  $m_{\nu}$ ???
- Clockwork at work: if there were no  $R_0 \Longrightarrow$  no chiral partner for  $\nu$ s but effect of  $R_0$  has to go through the **whole clockwork chain**:

$$m_
u \simeq rac{m_D^2}{q^{2N} m_N}$$

- suppression here is smaller than for DM:  $q = 10, m_N = 1 \text{ TeV} \implies N \approx 7$
- $\geq$  2 nonzero  $m_{\nu} \Longrightarrow$  at least 2 clockwork chains
- a suggestive possibility: 1 chain for dark matter, 2 chains for neutrino masses
- + resonant leptogenesis? Starting from 2 degenerate, a small mass splitting can be generated by interactions with the clockwork gears...
- many model-building variants (not discussed here)

Daniele Teresi A Clockwork Tale 13 / 26

# One more dimension

Chapter 3:

#### Clockwork from a flat extra dimension [Hambye, DT, Tytgat, '16]

- the clockwork Lagrangian can come from a discretized 5th dimension
- flat-spacetime construction for fermion:
- 1 Dirac fermion with mass M in the 5D bulk  $\rightarrow L_i, R$

$$\mathcal{L}_5 \supset \bar{\psi}(\overrightarrow{i} \overrightarrow{\partial}_D - M) \psi = i \overline{\psi} \gamma^\mu \partial_\mu \psi \, + \, \left[ \frac{1}{2} \Big( \overline{L} \, \partial_Z R - (\partial_Z \overline{L}) R \Big) \, - M \overline{L} R \, + h.c. \right]$$

- + Wilson term  $-\frac{a}{2} \partial_z \overline{\psi} \partial_z \psi = -\frac{a}{2} \partial_z \overline{L} \partial_z R$  removes 1 hopping direction
- to get light mode: orbifold with Dirichlet b.c. L(0) = 0 (1 spare **chiral fermion** on one brane  $\rightarrow R_0$ )
- ullet SM chiral leptons on the other brane  $o L_{SM}$
- discretized Lagrangian  $\mathcal{L} \supset \sum_{i=0}^{N-1} \frac{1}{a} \, \overline{L}_{i+1} R_i \sum_{i=1}^{N} \left( \frac{1}{a} + M \right) \overline{L}_i R_i$
- clockwork with  $m=\frac{1}{a}, qm=\frac{1}{a}+M, q^N=\left(1+\frac{\pi RM}{N}\right)^N \rightarrow e^{\pi RM}$

Daniele Teresi A Clockwork Tale 14 / 26

#### Clockwork from a flat extra dimension [Hambye, DT, Tytgat, '16]

- the clockwork Lagrangian can come from a discretized 5th dimension
- flat-spacetime construction for fermion:
- 1 Dirac fermion with mass M in the 5D bulk  $\rightarrow L_i, R_i$

$$\mathcal{L}_5 \supset \bar{\psi}(i\overleftrightarrow{\partial_D} - M)\psi = i\overline{\psi}\gamma^{\mu}\partial_{\mu}\psi + \left[\frac{1}{2}\Big(\overline{L}\,\partial_Z R - (\partial_Z\overline{L})R\Big) - M\overline{L}R + h.c.\right]$$

- ullet + Wilson term  $-rac{a}{2}\,\partial_Z\overline{\psi}\,\partial_Z\psi=-rac{a}{2}\,\partial_Z\overline{L}\,\partial_Z R$  removes 1 hopping direction
- to get light mode: orbifold with Dirichlet b.c. L(0) = 0 (1 spare **chiral fermion** on one brane  $\rightarrow R_0$ )
- SM chiral leptons on the other brane  $\rightarrow L_{SM}$
- discretized Lagrangian  $\mathcal{L} \supset \sum_{i=0}^{N-1} \frac{1}{a} \, \overline{L}_{i+1} R_i \sum_{i=1}^N \left( \frac{1}{a} + M \right) \overline{L}_i R_i$
- clockwork with  $m=\frac{1}{a}, \quad qm=\frac{1}{a}+M, \quad q^N=\left(1+\frac{\pi RM}{N}\right)^N \ \to \ e^{\pi RM}$

Daniele Teresi A Clockwork Tale 14 / 26

#### Clockwork from a flat extra dimension [Hambye, DT, Tytgat, '16]

- the clockwork Lagrangian can come from a discretized 5th dimension
- flat-spacetime construction for fermion:
- 1 Dirac fermion with mass M in the 5D bulk  $\rightarrow L_i, R_i$

$$\mathcal{L}_5 \supset \bar{\psi}(\overrightarrow{i}\overrightarrow{\partial}_D - M)\psi = i\overline{\psi}\gamma^{\mu}\partial_{\mu}\psi + \left[\frac{1}{2}\left(\overline{L}\partial_Z R - (\partial_Z\overline{L})R\right) - M\overline{L}R + h.c.\right]$$

- + Wilson term  $-\frac{a}{2} \partial_Z \overline{\psi} \partial_Z \psi = -\frac{a}{2} \partial_Z \overline{L} \partial_Z R$  removes 1 hopping direction
- to get light mode: orbifold with Dirichlet b.c. L(0) = 0(1 spare chiral fermion on one brane  $\rightarrow R_0$ )
- SM chiral leptons on the other brane  $\rightarrow L_{SM}$
- discretized Lagrangian  $\mathcal{L} \supset \sum_{i=1}^{N-1} \frac{1}{a} \overline{L}_{i+1} R_i \sum_{i=1}^{N} \left(\frac{1}{a} + M\right) \overline{L}_i R_i$
- clockwork with  $m=rac{1}{a}, \quad qm=rac{1}{a}+M, \quad q^N=\left(1+rac{\pi RM}{N}
  ight)^N \ 
  ightarrow \ e^{\pi RM}$

A Clockwork Tale 14 / 26 **Daniele Teresi** 

#### Clockwork from the metric [Giudice, McCullough, '16]

- curved-spacetime construction for scalar:
- curved metric  $ds^2 = X(|Z|) dx^2 + Y(|Z|) dZ^2$
- massless scalar in the 5D bulk:

$$\mathcal{S} = -2\int_0^R\!\!dZ\int\!\!d^4x\,\sqrt{-g}\,\frac{1}{2}\,g^{MN}\partial_M\phi\,\partial_N\phi = -\int_0^R\!\!dZ\int\!\!d^4x\,X^2Y^{1/2}\left[\frac{(\partial_\mu\phi)^2}{X} + \frac{(\partial_Z\phi)^2}{Y}\right]$$

discretized Lagrangian:

$$\mathcal{L} \supset \sum_{j=0}^{N-1} m_j^2 (\phi_j - q_j \phi_{j+1})^2$$
 with  $m_j^2 = \frac{X_j}{a^2 Y_j}$ ,  $q_j = \frac{X_j^{1/2} Y_j^{1/4}}{X_{j+1}^{1/2} Y_{j+1}^{1/4}}$ 

• clockwork if  $X_j \propto Y_j$ finite for  $N \to \infty$  if  $X_j \propto Y_j \propto e^{-\frac{4}{3}k a_j}$ 

• 
$$m = \frac{1}{a}$$
,  $q = e^{ka}$ ,  $q^N = e^{\pi kR}$ 

Daniele Teresi A Clockwork Tale 15 / 26

#### Clockwork from the metric [Giudice, McCullough, '16]

- curved-spacetime construction for scalar:
- curved metric  $ds^2 = X(|Z|) dx^2 + Y(|Z|) dZ^2$
- massless scalar in the 5D bulk:

$$\mathcal{S} = -2\int_0^R\!\!dZ\int\!\!d^4x\,\sqrt{-g}\,\frac{1}{2}\,g^{MN}\partial_M\phi\,\partial_N\phi = -\int_0^R\!\!dZ\int\!\!d^4x\,X^2Y^{1/2}\left[\frac{(\partial_\mu\phi)^2}{X} + \frac{(\partial_Z\phi)^2}{Y}\right]$$

discretized Lagrangian:

$$\mathcal{L} \supset \sum_{j=0}^{N-1} m_j^2 (\phi_j - q_j \phi_{j+1})^2$$
 with  $m_j^2 = \frac{X_j}{a^2 Y_j}$ ,  $q_j = \frac{X_j^{1/2} Y_j^{1/4}}{X_{j+1}^{1/2} Y_{j+1}^{1/4}}$ 

- clockwork if  $X_j \propto Y_j$ finite for  $N \to \infty$  if  $X_j \propto Y_j \propto e^{-\frac{4}{3}k aj}$
- $\bullet \ m = \frac{1}{a}, \quad q = e^{ka}, \quad q^N = e^{\pi kR}$

Daniele Teresi A Clockwork Tale 15 / 26

### The clockwork metric

- in the continuum:  $ds^2 = e^{\frac{4}{3}k|Z|} (dx^2 + dZ^2)$
- Kaluza-Klein modes for massless scalar

$$\psi_0(Z) \simeq \sqrt{k\pi R} \, e^{-k\pi R}$$
  $\Longrightarrow \frac{dP}{dZ} \propto e^{2kZ}$   $\psi_n(Z) = e^{-kZ} \times \text{oscillatory}$   $\Longrightarrow \frac{dP}{dZ} = \text{oscillatory}$ 

what about Large Extra Dimension or Randall-Sundrum?

LED	$\frac{1}{a}$	1
RS	$\frac{1}{a} e^{-\hat{k} a j}$	$e^{\hat{k}a}$
clockwork	$\frac{1}{a}$	$e^{\hat{k}a}$

Daniele Teresi A Clockwork Tale 16 / 26

### The clockwork metric

- in the continuum:  $ds^2 = e^{\frac{4}{3}k|Z|} (dx^2 + dZ^2)$
- Kaluza-Klein modes for massless scalar

$$\psi_0(Z) \simeq \sqrt{k\pi R} \, e^{-k\pi R}$$
  $\Longrightarrow \frac{dP}{dZ} \propto e^{2\,k\,Z}$   $\psi_n(Z) = e^{-k\,Z} imes ext{oscillatory}$   $\Longrightarrow \frac{dP}{dZ} = ext{oscillatory}$ 

what about Large Extra Dimension or Randall-Sundrum?

	$m_j$	$q_{j}$
LED	$\frac{1}{a}$	1
RS	$\frac{1}{a}e^{-\hat{k}aj}$	$e^{\hat{k}a}$
clockwork	$\frac{1}{a}$	$e^{\hat{k}a}$

Daniele Teresi A Clockwork Tale 16 / 26

# Clockwork naturalness

Chapter 4:

### **Clockwork graviton**

- discrete clockwork: N+1 copies of 4D gravity  $g_i^{\mu\nu}$
- linear approximation:  $g_j^{\mu\nu}=\eta_j^{\mu\nu}+2h_j^{\mu\nu}/M_j$
- clockwork Pauli-Fierz mass terms

$$\mathcal{L} = -rac{m^2}{2} \sum_{i=0}^{N-1} \left( \left[ h_j^{\mu 
u} - q \, h_{j+1}^{\mu 
u} \right]^2 - \left[ \eta_{\mu 
u} (h_j^{\mu 
u} - q \, h_{j+1}^{\mu 
u}) \right]^2 \right)$$

- $\bullet \ \ \text{invariant under} \ h_j^{\mu\nu} \to h_j^{\mu\nu} + \frac{1}{q^j} (\partial^\mu A^\nu + \partial^\nu A^\mu)$
- $\Longrightarrow$  massless graviton  $\mathfrak{h}_0^{\mu\nu}$  localized at j=0:

$$\frac{1}{M_N} h_N^{\mu\nu} T_{\mu\nu} \longrightarrow \frac{1}{M_P} \mathfrak{h}_0^{\mu\nu} T_{\mu\nu} \quad \text{with} \quad M_P = \frac{q^N M_N}{\mathcal{N}}$$

but... multi-gravity theories are dodgy → continuum limit

Daniele Teresi A Clockwork Tale 17 / 26

### **Clockwork graviton**

- ullet discrete clockwork: N+1 copies of 4D gravity  $g_{j}^{\mu 
  u}$
- linear approximation:  $g_j^{\mu\nu}=\eta_j^{\mu\nu}+2h_j^{\mu\nu}/M_j$
- clockwork Pauli-Fierz mass terms

$$\mathcal{L} = -rac{m^2}{2} \sum_{i=0}^{N-1} \left( \left[ h_j^{\mu 
u} - q \, h_{j+1}^{\mu 
u} \right]^2 - \left[ \eta_{\mu 
u} (h_j^{\mu 
u} - q \, h_{j+1}^{\mu 
u}) \right]^2 \right)$$

- invariant under  $h_j^{\mu 
  u} o h_j^{\mu 
  u} + rac{1}{q^j} (\partial^\mu A^
  u + \partial^
  u A^\mu)$
- $\Longrightarrow$  massless graviton  $\mathfrak{h}_0^{\mu\nu}$  localized at j=0:

$$rac{1}{M_N}\,h_N^{\mu
u}T_{\mu
u}\;\longrightarrow\;rac{1}{M_P}\,\mathfrak{h}_0^{\mu
u}T_{\mu
u}\quad ext{with}\quad M_P=rac{q^N\,M_N}{\mathcal{N}}$$

but... multi-gravity theories are dodgy → continuum limit

Daniele Teresi A Clockwork Tale 17 / 26

### **Clockwork graviton**

- ullet discrete clockwork: N+1 copies of 4D gravity  $g_{j}^{\mu 
  u}$
- linear approximation:  $g_j^{\mu\nu}=\eta_j^{\mu\nu}+2h_j^{\mu\nu}/M_j$
- clockwork Pauli-Fierz mass terms

$$\mathcal{L} = -rac{m^2}{2} \sum_{i=0}^{N-1} \left( \left[ h_j^{\mu 
u} - q \, h_{j+1}^{\mu 
u} \right]^2 - \left[ \eta_{\mu 
u} (h_j^{\mu 
u} - q \, h_{j+1}^{\mu 
u}) \right]^2 \right)$$

- invariant under  $h_j^{\mu 
  u} o h_j^{\mu 
  u} + rac{1}{q^j} (\partial^\mu A^
  u + \partial^
  u A^\mu)$
- $\Longrightarrow$  massless graviton  $\mathfrak{h}_0^{\mu\nu}$  localized at j=0:

$$rac{1}{M_N} \, h_N^{\mu 
u} T_{\mu 
u} \, \longrightarrow \, rac{1}{M_P} \, \mathfrak{h}_0^{\mu 
u} T_{\mu 
u} \quad ext{with} \quad M_P = rac{q^N \, M_N}{\mathcal{N}}$$

but... multi-gravity theories are dodgy → continuum limit

Daniele Teresi A Clockwork Tale 17 / 26

### The metric from the linear dilaton

- we want massless 5D gravity with a clockwork metric  $ds^2 = e^{\frac{4}{3}k|Z|} (dx^2 + dZ^2)$
- clockwork gravity → metric should not be treated as a background
- can we obtain the metric?
- linear dilaton model (Jordan frame):

$$S = \int d^4x \, dZ \sqrt{-g} \, \frac{M_5^3}{2} \, e^S (\mathcal{R} + g^{MN} \partial_M S \, \partial_N S + 4k^2) + \text{brane } \Lambda s$$

- k breaks global Weyl  $g_{MN} \rightarrow e^{-2\alpha} g_{MN}, S \rightarrow S + 3\alpha$
- go to Einstein frame, solve EoMs: S = 2k|Z|,  $ds^2 = e^{\frac{4}{3}k|Z|}(dx^2 + dZ^2)$
- in Jordan frame  $g_{MN} = \eta_{MN}$ , but Planck mass exponentially varying

Daniele Teresi A Clockwork Tale 18 / 26

### The metric from the linear dilaton

- ullet we want massless 5D gravity with a clockwork metric  $ds^2=e^{rac{4}{3}k|Z|}\left(dx^2+dZ^2
  ight)$
- clockwork gravity → metric should not be treated as a background
- can we obtain the metric?
- linear dilaton model (Jordan frame):

$$S = \int d^4x \, dZ \sqrt{-g} \, \frac{M_5^3}{2} \, e^S (\mathcal{R} \, + \, g^{MN} \partial_M S \, \partial_N S \, + \, 4k^2) \, + \, \text{brane} \, \Lambda s$$

- k breaks global Weyl  $g_{MN} \rightarrow e^{-2\alpha} g_{MN}, S \rightarrow S + 3\alpha$
- go to Einstein frame, solve EoMs: S = 2k|Z|,  $ds^2 = e^{\frac{4}{3}k|Z|}(dx^2 + dZ^2)$
- in Jordan frame  $g_{MN} = \eta_{MN}$ , but Planck mass exponentially varying

Daniele Teresi A Clockwork Tale 18 / 26

### The metric from the linear dilaton

- ullet we want massless 5D gravity with a clockwork metric  $ds^2=e^{rac{4}{3}k|Z|}\left(dx^2+dZ^2
  ight)$
- clockwork gravity → metric should not be treated as a background
- can we obtain the metric?
- linear dilaton model (Jordan frame):

$$S = \int d^4x \, dZ \sqrt{-g} \, \frac{M_5^3}{2} \, e^S (\mathcal{R} + g^{MN} \partial_M S \, \partial_N S + 4k^2) + \text{brane } \Lambda s$$

- k breaks global Weyl  $g_{MN} \rightarrow e^{-2\alpha} g_{MN}, S \rightarrow S + 3\alpha$
- go to Einstein frame, solve EoMs: S = 2k|Z|,  $ds^2 = e^{\frac{4}{3}k|Z|}(dx^2 + dZ^2)$
- in Jordan frame  $g_{MN} = \eta_{MN}$ , but Planck mass exponentially varying

Daniele Teresi A Clockwork Tale 18 / 26

### A solution to the hierarchy problem

- effective 4D Planck mass:  $M_P^2 = 2M_5^2 \int_0^{\pi R} dZ \, e^{2kZ} = \frac{M_5^3}{k} (e^{2\pi kR} 1)$
- 4D graviton fluctuations:  $ds^2 = e^{\frac{4}{3}k|Z|} \left[ \left( \eta_{\mu\nu} + \frac{2}{M_5^{3/2}} h_{\mu\nu} \right) dx^\mu dx^\nu + dZ^2 \right]$
- $\bullet \text{ action: } \mathcal{S} = -\frac{1}{2} \int \! d^4x \, dZ \, e^{2 \, k |Z|} \Big[ (\partial_\lambda h_{\mu\nu}) (\partial^\lambda h^{\mu\nu}) \, + \, (\partial_Z h_{\mu\nu}) (\partial_Z h^{\mu\nu}) \Big]$
- $\mathcal{L}_{SM}$  at Z=0:  $\frac{h^{\mu\nu}(x,Z=0)\,T^{SM}_{\mu\nu}(x)}{M_5^{3/2}} \longrightarrow \sum_n \frac{\mathfrak{h}_n^{\mu\nu}(x)\,T^{SM}_{\mu\nu}(x)}{\Lambda_n}$  with

$$\Lambda_0 = M_P$$
 ,  $\Lambda_n^2 = M_5^3 \, \pi R \, (1 + k^2 R^2 / n^2)$ 

- ullet the cutoff is  $M_5 \implies m_h = O(M_5) \ll M_P \rightarrow {f solution}$  to hierarchy problem
- for k=1 TeV,  $M_5=10$  TeV  $\rightarrow kR \simeq 10$

Daniele Teresi A Clockwork Tale 19 / 26

### A solution to the hierarchy problem

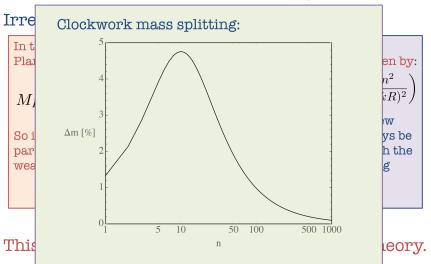
- effective 4D Planck mass:  $M_P^2 = 2M_5^2 \int_0^{\pi R} dZ \, e^{2kZ} = \frac{M_5^3}{k} (e^{2\pi kR} 1)$
- 4D graviton fluctuations:  $ds^2 = e^{\frac{4}{3}k|Z|} \left[ \left( \eta_{\mu\nu} + \frac{2}{M_5^{3/2}} h_{\mu\nu} \right) dx^{\mu} dx^{\nu} + dZ^2 \right]$
- $\bullet \text{ action: } \mathcal{S} = -\frac{1}{2} \int \! d^4x \, dZ \, e^{2\,k|Z|} \Big[ (\partial_\lambda h_{\mu\nu}) (\partial^\lambda h^{\mu\nu}) \, + \, (\partial_Z h_{\mu\nu}) (\partial_Z h^{\mu\nu}) \Big]$
- $\bullet \ \mathcal{L}_{\mathit{SM}} \ \text{at} \ Z = 0 \colon \quad \frac{h^{\mu\nu}(x,Z=0) \ T^{\mathit{SM}}_{\mu\nu}(x)}{M_{5}^{3/2}} \ \longrightarrow \ \sum_{n} \frac{\mathfrak{h}_{n}^{\mu\nu}(x) \ T^{\mathit{SM}}_{\mu\nu}(x)}{\Lambda_{n}} \quad \text{with}$

$$\Lambda_0 = M_P$$
,  $\Lambda_n^2 = M_5^3 \pi R (1 + k^2 R^2 / n^2)$ 

- the cutoff is  $M_5 \implies m_h = O(M_5) \ll M_P \rightarrow$  solution to hierarchy problem
- for k=1 TeV,  $M_5=10$  TeV  $\rightarrow kR \simeq 10$

Daniele Teresi A Clockwork Tale 19 / 26

## Phenomenology

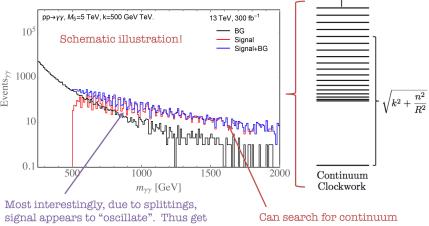


[ slide by M. McCullough ]

Daniele Teresi A Clockwork Tale 20 / 26

### Phenomenology

### At colliders would look something like:



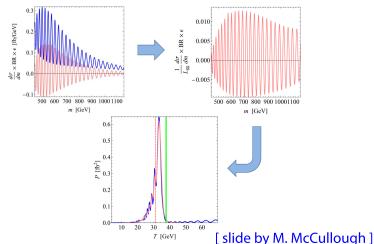
signal appears to "oscillate". Thus get extra sensitivity by doing spectral analysis... The "power spectrum" of LHC data!

Can search for continuum spectrum at high energies.

[ slide by M. McCullough ]

### Phenomenology

The fourier transform would then exhibit a peak near the inverse radius:



Daniele Teresi A Clockwork Tale 22 / 26

# Epilogue:

Scrambling the clockwork

### Disassembling the clockwork? [Craig, Garcia Garcia, Sutherland, '17]

- Disclaimer: I'm simplifying the argument
- Definition of clockwork: a theory with no exponential hierarchies in fundamental parameters that gives rise to exponentially suppressed site-dependent couplings to symmetry-protected zero mode
- Claim 1: no clockwork from geometry
- for a scalar in curved clockwork metric:  $\psi_0 = \text{const.} \equiv \mathcal{C}_0$
- coupling on a brane at  $Z=Z_0$ :  $\frac{\phi}{16\pi^2 f_{5D}} G \, \widetilde{G} \implies \frac{\varphi_0}{16\pi^2 F} G \, \widetilde{G}$  with  $F=f_{5D}^{3/2}/\mathcal{C}_0=M_{PL}\left(\frac{f_{5D}}{Z_0}\right)^{3/2}$  independent on  $Z_0 \implies$  no clockwold independent independent on  $Z_0 \implies$  no clockwold independen
- Claim 1b: clockwork only for flat-spacetime construction with bulk mass and boundary terms (they do it for scalar, but essentially a re-discovery of construction in (Nambus DT Tutest 156).)

Daniele Teresi A Clockwork Tale 23 / 26

### Disassembling the clockwork? [Craig, Garcia, Garcia, Sutherland, '17]

- Disclaimer: I'm simplifying the argument
- Definition of clockwork: a theory with no exponential hierarchies in fundamental parameters that gives rise to exponentially suppressed site-dependent couplings to symmetry-protected zero mode
- Claim 1: no clockwork from geometry
- for a scalar in curved clockwork metric:  $\psi_0 = \text{const.} \equiv \mathcal{C}_0$
- ullet coupling on a brane at  $Z=Z_0$ :  $\dfrac{\phi}{16\pi^2f_{5D}}G\widetilde{G} \implies \dfrac{arphi_0}{16\pi^2F}G\widetilde{G}$

with 
$$F = f_{5D}^{3/2}/\mathcal{C}_0 = M_{PL} \left(\frac{f_{5D}}{M_5}\right)^{3/2}$$
 independent on  $Z_0 \Longrightarrow$  no clockwork

 Claim 1b: clockwork only for flat-spacetime construction with bulk mass and boundary terms (they do it for scalar, but essentially a re-discovery of construction in [Hambye, DT, Tytgat, '16] )

Daniele Teresi A Clockwork Tale 23 / 26

### Disassembling the clockwork? [Craig, Garcia, Garcia, Sutherland, '17]

- Disclaimer: I'm simplifying the argument
- Definition of clockwork: a theory with no exponential hierarchies in fundamental parameters that gives rise to exponentially suppressed site-dependent couplings to symmetry-protected zero mode
- Claim 1: no clockwork from geometry
- for a scalar in curved clockwork metric:  $\psi_0 = \text{const.} \equiv \mathcal{C}_0$
- ullet coupling on a brane at  $Z=Z_0$ :  $\dfrac{\phi}{16\pi^2f_{5D}}G\widetilde{G} \implies \dfrac{arphi_0}{16\pi^2F}G\widetilde{G}$

with 
$$F = f_{5D}^{3/2}/\mathcal{C}_0 = M_{PL} \left(\frac{f_{5D}}{M_5}\right)^{3/2}$$
 independent on  $Z_0 \Longrightarrow$  no clockwork

 Claim 1b: clockwork only for flat-spacetime construction with bulk mass and boundary terms (they do it for scalar, but essentially a re-discovery of construction in [Hambye, DT, Tytgat, '16] )

Daniele Teresi A Clockwork Tale 23 / 26

- Claim 2: no non-Abelian clockwork (including gravity)
- non-Abelian Yang-Mills clockwork chain
- kinetic terms:

$$-\mathcal{L}_{kin} = \sum_{j=0}^{N} \frac{1}{4g_{j}^{2}} F_{j} F_{j} = \sum_{j=0}^{N} \frac{1}{4g_{j}^{2}} \Big( F_{j}^{abelian} F_{j}^{abelian} + 4f A_{j} A_{j} \partial A_{j} + f f A_{j} A_{j} A_{j} A_{j} \Big)$$

• if, in terms of zero mode  $A_0$ ,  $A_i = c_i A_0 + \dots$ 

$$-\mathcal{L}_{\mathit{kin}} \supset \sum_{j=0}^{N} \frac{c_{j}^{2}}{4g_{j}^{2}} \mathcal{F}_{0}^{\mathit{abelian}} \mathcal{F}_{0}^{\mathit{abelian}} + \sum_{j=0}^{N} \frac{c_{j}^{3}}{4g_{j}^{2}} 4f \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \partial \mathcal{A}_{0} \, + \sum_{j=0}^{N} \frac{c_{j}^{4}}{4g_{j}^{2}} ff \, \mathcal{A}_{0} \, \mathcal{A}_{0}$$

- gauge invariance for  $\mathcal{A}_0$ :  $\sum_{j=0}^N \frac{c_j^2}{g_j^2} = \sum_{j=0}^N \frac{c_j^3}{g_j^2} = \sum_{j=0}^N \frac{c_j^4}{g_j^2} \equiv \frac{1}{g_{eff}^2}$
- $\bullet \implies c_i \in \{0,1\} \rightarrow \mathsf{no} \mathsf{clockwork}$
- $g_i = \text{const.} \implies g_{eff} \sim g_i \text{ no exponential suppression}$

Daniele Teresi A Clockwork Tale 24 / 26

- Claim 2: no non-Abelian clockwork (including gravity)
- non-Abelian Yang-Mills clockwork chain
- kinetic terms:

$$-\mathcal{L}_{kin} = \sum_{j=0}^{N} \frac{1}{4g_{j}^{2}} F_{j} F_{j} = \sum_{j=0}^{N} \frac{1}{4g_{j}^{2}} \left( F_{j}^{abelian} F_{j}^{abelian} + 4f A_{j} A_{j} \partial A_{j} + f f A_{j} A_{j} A_{j} A_{j} \right)$$

• if, in terms of zero mode  $A_0$ ,  $A_j = c_j A_0 + \dots$ 

$$-\mathcal{L}_{\textit{kin}} \supset \sum_{j=0}^{N} \frac{c_{j}^{2}}{4g_{j}^{2}} \mathcal{F}_{0}^{\textit{abelian}} \mathcal{F}_{0}^{\textit{abelian}} + \sum_{j=0}^{N} \frac{c_{j}^{3}}{4g_{j}^{2}} \, 4f \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \partial \mathcal{A}_{0} \, + \sum_{j=0}^{N} \frac{c_{j}^{4}}{4g_{j}^{2}} ff \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \partial \mathcal{A}_{0} + \sum_{j=0}^{N} \frac{c_{j}^{4}}{4g_{j}^{2}} \, ff \, \mathcal{A}_{0} \, \mathcal{A}$$

- gauge invariance for  $A_0$ :  $\sum_{j=0}^{N} \frac{c_j^2}{g_j^2} = \sum_{j=0}^{N} \frac{c_j^3}{g_j^2} = \sum_{j=0}^{N} \frac{c_j^4}{g_j^2} \equiv \frac{1}{g_{eff}^2}$
- $\bullet \implies c_i \in \{0,1\} \rightarrow \mathsf{no} \mathsf{clockwork}$
- $g_i = \text{const.} \implies g_{eff} \sim g_i$  no exponential suppression

Daniele Teresi A Clockwork Tale 24 / 26

- Claim 2: no non-Abelian clockwork (including gravity)
- non-Abelian Yang-Mills clockwork chain
- kinetic terms:

$$-\mathcal{L}_{kin} = \sum_{j=0}^{N} \frac{1}{4g_{j}^{2}} F_{j} F_{j} = \sum_{j=0}^{N} \frac{1}{4g_{j}^{2}} \left( F_{j}^{abelian} F_{j}^{abelian} + 4f A_{j} A_{j} \partial A_{j} + f f A_{j} A_{j} A_{j} A_{j} \right)$$

• if, in terms of zero mode  $A_0$ ,  $A_j = c_j A_0 + \dots$ 

$$-\mathcal{L}_{\textit{kin}} \supset \sum_{j=0}^{N} \frac{c_{j}^{2}}{4g_{j}^{2}} \mathcal{F}_{0}^{\textit{abelian}} \mathcal{F}_{0}^{\textit{abelian}} + \sum_{j=0}^{N} \frac{c_{j}^{3}}{4g_{j}^{2}} \, 4f \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \partial \mathcal{A}_{0} \, + \sum_{j=0}^{N} \frac{c_{j}^{4}}{4g_{j}^{2}} ff \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \mathcal{A}_{0} \, \partial \mathcal{A}_{0} + \sum_{j=0}^{N} \frac{c_{j}^{4}}{4g_{j}^{2}} \, ff \, \mathcal{A}_{0} \, \mathcal{A}$$

- gauge invariance for  $A_0$ :  $\sum_{j=0}^{N} \frac{c_j^2}{g_j^2} = \sum_{j=0}^{N} \frac{c_j^3}{g_j^2} = \sum_{j=0}^{N} \frac{c_j^4}{g_j^2} \equiv \frac{1}{g_{eff}^2}$
- $\bullet \implies c_i \in \{0,1\} \rightarrow \mathsf{no} \mathsf{clockwork}$
- $g_i = \text{const.} \implies g_{eff} \sim g_i$  no exponential suppression

Daniele Teresi A Clockwork Tale 24 / 26

### Reassembling the clockwork [Giudice, McCullough, '17]

- Disclaimer: I'm simplifying the argument
- answer to Claim 2 (here for scalar, given also for gravity):

- in this basis the unbroken symmetry is  $\pi_k \to \pi_k + \alpha$ , rather than  $\phi_k \to \phi_k + q^{-k}\alpha$  $\implies c_k = 1 \to \text{ argument for non-abelian disappears}$
- in this basis  $g_k = g_0 q^{-k} \neq \text{const.} \implies g_{\text{eff}} \approx g_N = g_0 q^{-N}$
- in this basis the theory does not look like clockwork, but it's the same theory (physics is the same)

Daniele Teresi A Clockwork Tale 25 / 26

### Reassembling the clockwork [Giudice, McCullough, '17]

- Disclaimer: I'm simplifying the argument
- answer to Claim 2 (here for scalar, given also for gravity):

- in this basis the unbroken symmetry is  $\pi_k \to \pi_k + \alpha$ , rather than  $\phi_k \to \phi_k + q^{-k}\alpha$   $\implies c_k = 1 \to \text{argument for non-abelian disappears}$
- in this basis  $g_k = g_0 q^{-k} \neq \text{const.} \implies g_{eff} \approx g_N = g_0 q^{-N}$
- in this basis the theory does not look like clockwork, but it's the same theory (physics is the same)

Daniele Teresi A Clockwork Tale 25 / 26

- answer to Claim 1b:
- the curved and flat-spacetime scalar constructions are the same! (related by field redefinition  $\phi = e^{-kZ}\pi$ )
- summary of the answer so far: 2 theories with same Lagrangian are the same
- answer to Claim 1:
- definition of [Craig, Garcia Garcia, Sutherland, '17] is (UV) model dependent, e.g.

$$S \supset \int d^5x \, \delta(Z - Z_0) \, e^{nS/2} \frac{\phi}{16\pi^2 f} G \, \widetilde{G}$$

 $Z_0$ -profile of coupling depends on n, spurion charge of f under global Weyl

- for model building and hierarchy problem, relevant definition:

   a theory with no exponential hierarchies in fundamental parameters that gives rise to exponentially suppressed ratios between the zero-mode and the clockwork-gears couplings
- $F = f_{5D}^{3/2}/C_0 = M_{PL} \left(\frac{f_{5D}}{M_5}\right)^{3/2} \simeq \frac{f}{\sqrt{2\pi R}} e^{k\pi R}$

Daniele Teresi A Clockwork Tale 26 / 26

- answer to Claim 1b:
- the curved and flat-spacetime scalar constructions are the same! (related by field redefinition  $\phi = e^{-kZ}\pi$ )
- summary of the answer so far: 2 theories with same Lagrangian are the same
- answer to Claim 1:
- definition of [Craig, Garcia Garcia, Sutherland, '17] is (UV) model dependent, e.g.:

$$\mathcal{S} \supset \int d^5 x \, \delta(Z-Z_0) \, e^{nS/2} rac{\phi}{16\pi^2 f} G \, \widetilde{G}$$

 $Z_0$ -profile of coupling depends on n, spurion charge of f under global Weyl

- for model building and hierarchy problem, relevant definition:

   a theory with no exponential hierarchies in fundamental parameters that gives rise to exponentially suppressed ratios between the zero-mode and the clockwork-gears couplings
- $F = f_{5D}^{3/2}/C_0 = M_{PL} \left(\frac{f_{5D}}{M_5}\right)^{3/2} \simeq \frac{f}{\sqrt{2\pi R}} e^{k\pi R}$

- answer to Claim 1b:
- the curved and flat-spacetime scalar constructions are the same! (related by field redefinition  $\phi = e^{-kZ}\pi$ )
- summary of the answer so far: 2 theories with same Lagrangian are the same
- answer to Claim 1:
- definition of [Craig, Garcia Garcia, Sutherland, '17] is (UV) model dependent, e.g.:

$$\mathcal{S} \supset \int d^5 x \, \delta(Z-Z_0) \, e^{nS/2} rac{\phi}{16\pi^2 f} G \, \widetilde{G}$$

 $Z_0$ -profile of coupling depends on n, spurion charge of f under global Weyl

- for model building and hierarchy problem, relevant definition:

   a theory with no exponential hierarchies in fundamental parameters that gives rise to exponentially suppressed ratios between the zero-mode and the clockwork-gears couplings
- $F = f_{5D}^{3/2}/\mathcal{C}_0 = M_{PL} \left(\frac{f_{5D}}{M_5}\right)^{3/2} \simeq \frac{f}{\sqrt{2\pi R}} e^{k\pi R}$

- answer to Claim 1b:
- the curved and flat-spacetime scalar constructions are the same! (related by field redefinition  $\phi = e^{-kZ}\pi$ )
- summary of the answer so far: 2 theories with same Lagrangian are the same
- answer to Claim 1:
- definition of [Craig, Garcia Garcia, Sutherland, '17] is (UV) model dependent, e.g.:

$$\mathcal{S} \supset \int d^5 x \, \delta(Z-Z_0) \, e^{nS/2} rac{\phi}{16\pi^2 f} G \, \widetilde{G}$$

 $Z_0$ -profile of coupling depends on n, spurion charge of f under global Weyl

- for model building and hierarchy problem, relevant definition:

   a theory with no exponential hierarchies in fundamental parameters that gives rise to exponentially suppressed ratios between the zero-mode and the clockwork-gears couplings
- $F = f_{5D}^{3/2}/C_0 = M_{PL} \left(\frac{f_{5D}}{M_5}\right)^{3/2} \simeq \frac{f}{\sqrt{2\pi R}} e^{k\pi R}$

# The End?



The relation between curved- and flat-spacetime constructions opens a Pandora box...

[Giudice, McCullough, DT, in preparation]