

KATRIN and Mare: direct neutrino mass experiments

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- Introduction
- The ^{187}Re decay experiment MARE
- The Karlsruhe TRItium Neutrino experiment KATRIN
- KATRIN's background suppression, statistics & systematics
- Summary



bmb+f - Förderschwerpunkt

Astroteilchenphysik

Großgeräte der physikalischen
Grundlagenforschung

3 complementary ways to the absolute ν mass scale

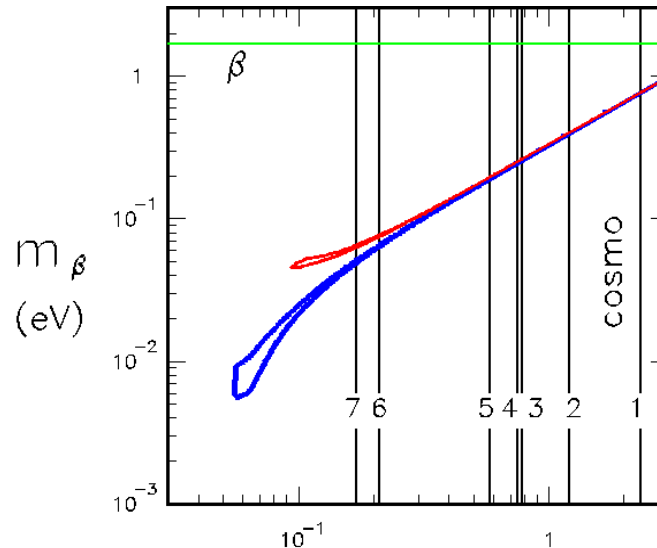
1) cosmology:
very sensitive $\Sigma m(\nu_i)$
but model dependent

2) $0\nu\beta\beta$:
evidence?
Majorana neutrinos

$$m_{ee}(\nu) = \left| \sum |U_{ei}|^2 e^{i\alpha(i)} m(\nu_i) \right|$$

3) direct neutrino mass
determination:
no further assumptions
no cancellations:

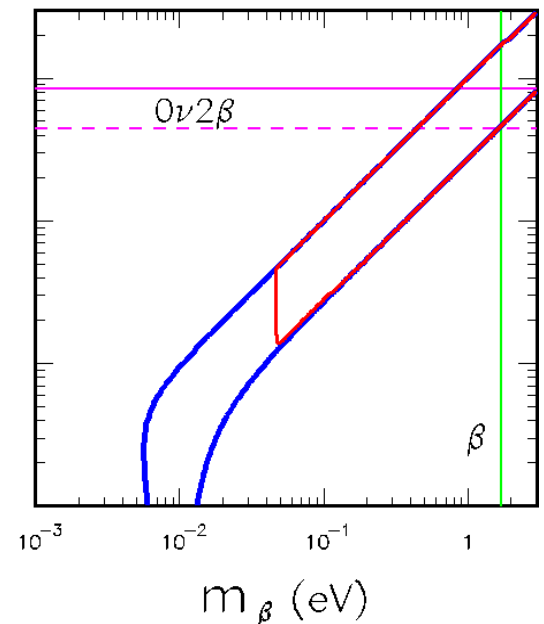
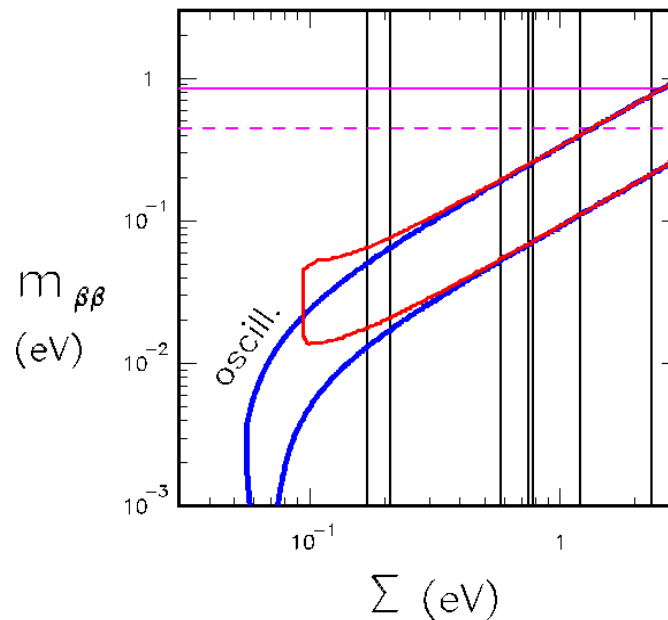
$$m^2(\nu_e) = \sum |U_{ei}|^2 m^2(\nu_i)$$



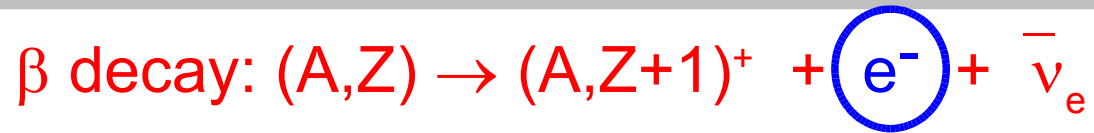
2σ bounds from :

- ν oscillation data
- β decay
- $0\nu 2\beta$ decay
- cosmology

— normal hierarchy
— inverted hierarchy



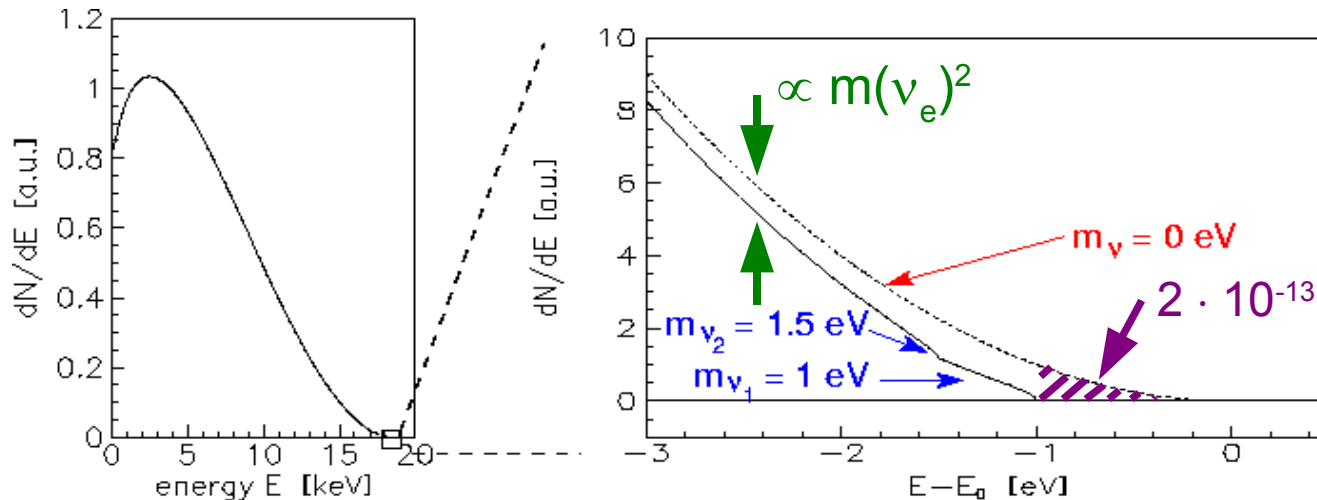
Direct determination of $m(\nu_e)$ from β decay



β electron energy spectrum:

$$dN/dE = K F(E,Z) p E_{\text{tot}} (E_0 - E_e) \sum |U_{ei}|^2 [(E_0 - E_e)^2 - m(\nu_i)^2]^{1/2}$$

(modified by electronic final states, recoil corrections, radiative corrections)



oscillation exp: small Δm_{ij}^2
 \Rightarrow see only average
 neutrino mass squared:
 $m(\nu_e)^2 := \sum |U_{ei}|^2 m(\nu_i)^2$

Need: low endpoint energy
 very high energy resolution &
 very high luminosity &
 very low background

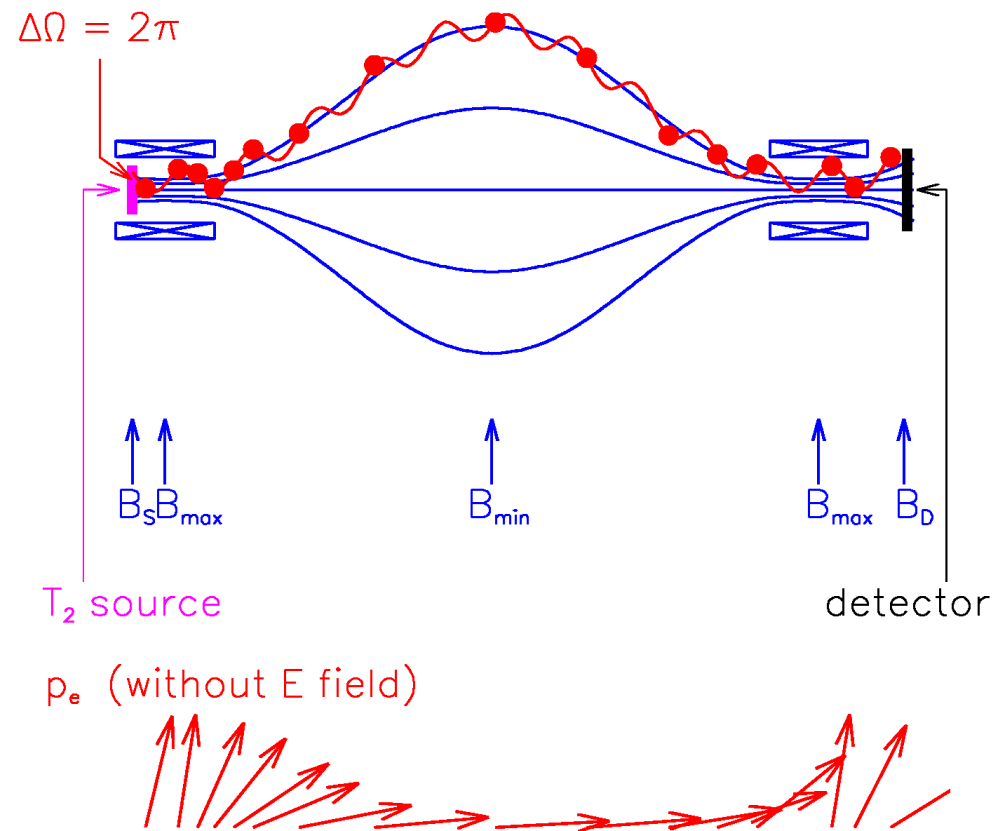
\Rightarrow Tritium ^3H , (^{187}Re)

\Rightarrow MAC-E-Filter
 (or bolometer for ^{187}Re)

Principle of the MAC-E-Filter

Magnetic Adiabatic Collimation + Electrostatic Filter
(A. Picard et al., Nucl. Instr. Meth. 63 (1992) 345)

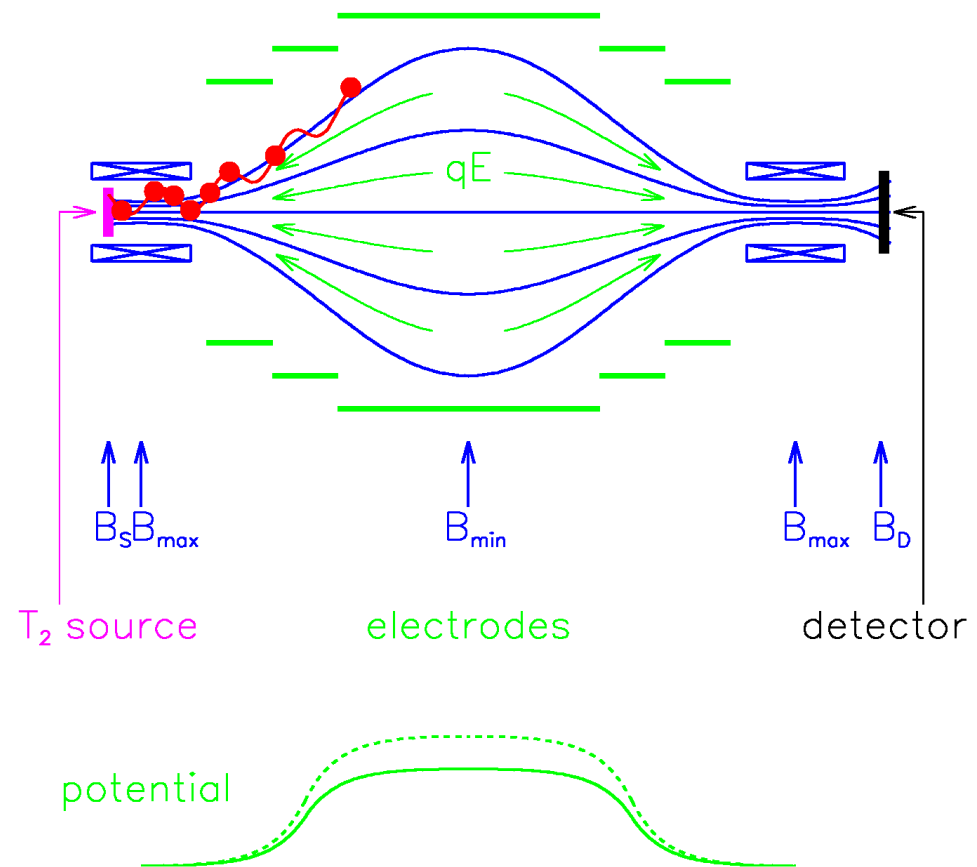
- Two supercond. solenoids compose magnetic guiding field
- Electron source (T_2) in left solenoid
- e^- in forward direction: magnetically guided
- adiabatic transformation:
 $\mu = E_{\perp}/B = \text{const.}$
 \Rightarrow parallel e^- beam



Principle of the MAC-E-Filter

Magnetic Adiabatic Collimation + Electrostatic Filter
(A. Picard et al., Nucl. Instr. Meth. 63 (1992) 345)

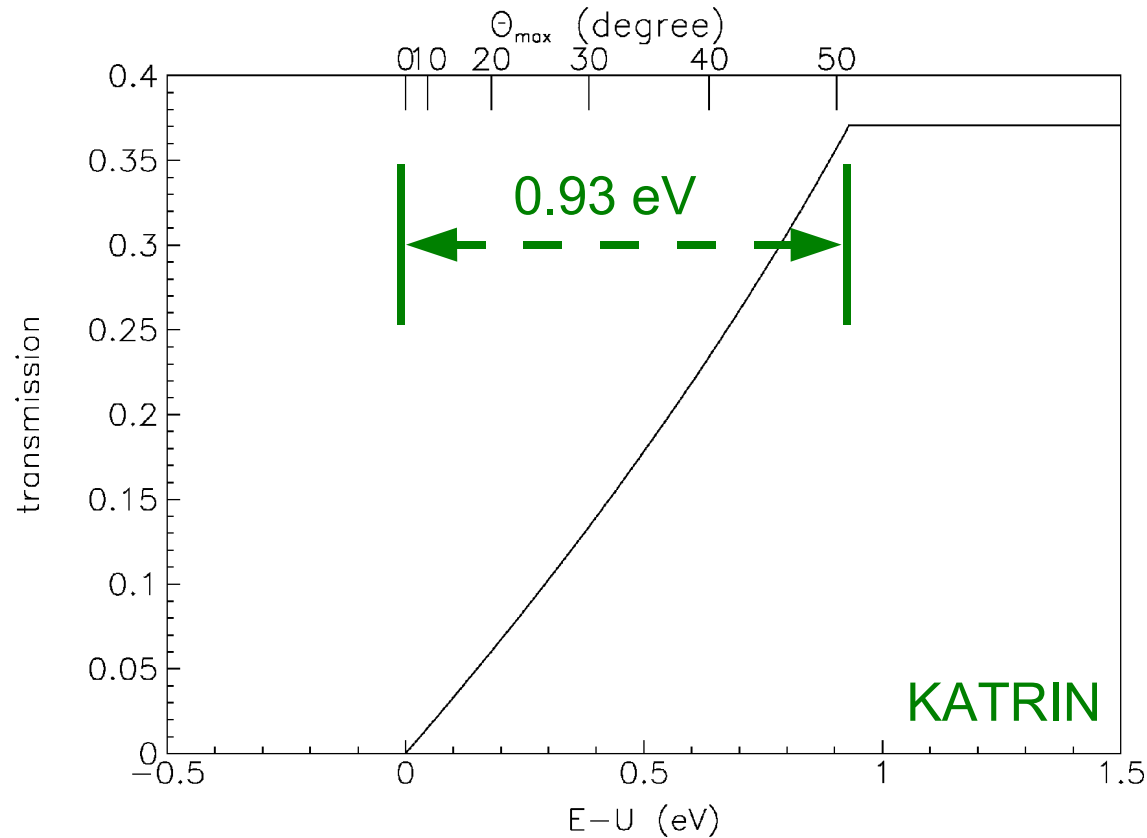
- Two supercond. solenoids compose magnetic guiding field
- Electron source (T_2) in left solenoid
- e^- in forward direction: magnetically guided
- adiabatic transformation: $\mu = E_{\perp}/B = \text{const.}$
 \Rightarrow parallel e^- beam
- Energy analysis by electrostat. retarding field
 $\Delta E = E \cdot B_{\min} / B_{\max} = E \cdot A_{s,\text{eff}} / A_{\text{analyse}} \approx 4.8 \text{ eV (Mainz)} = 0.93 \text{ eV (KATRIN)}$



$\approx 4.8 \text{ eV (Mainz)} = 0.93 \text{ eV (KATRIN)}$

Principle of the MAC-E-Filter

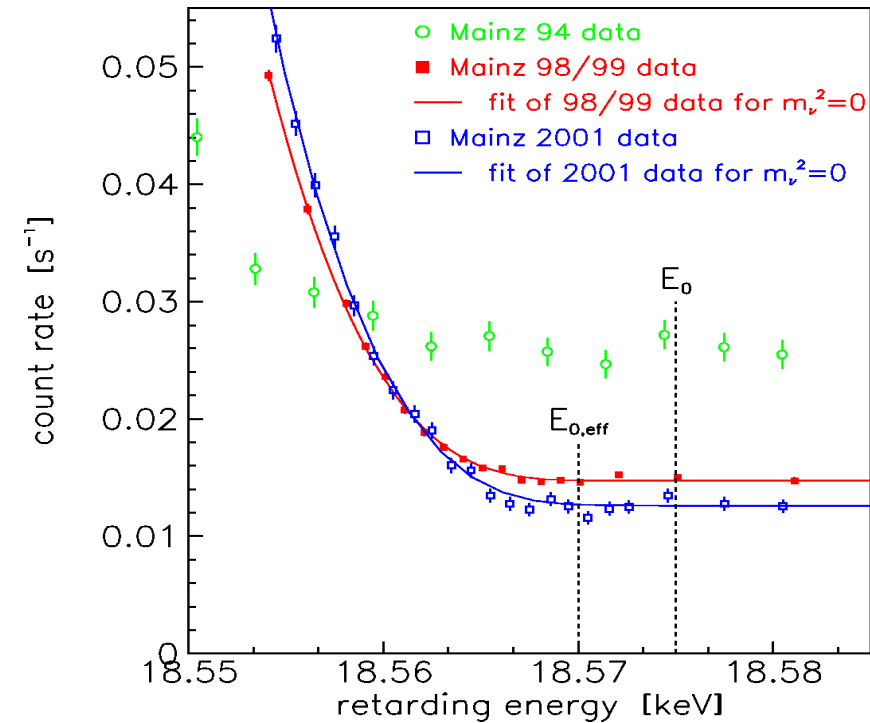
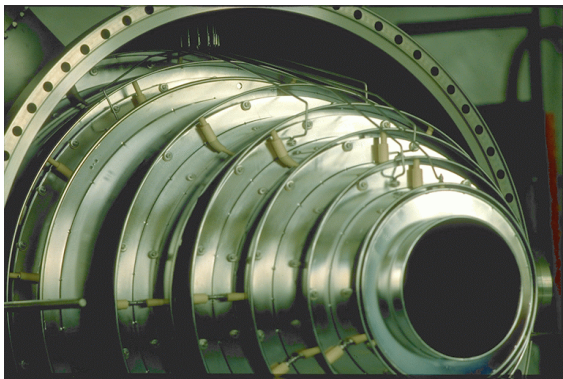
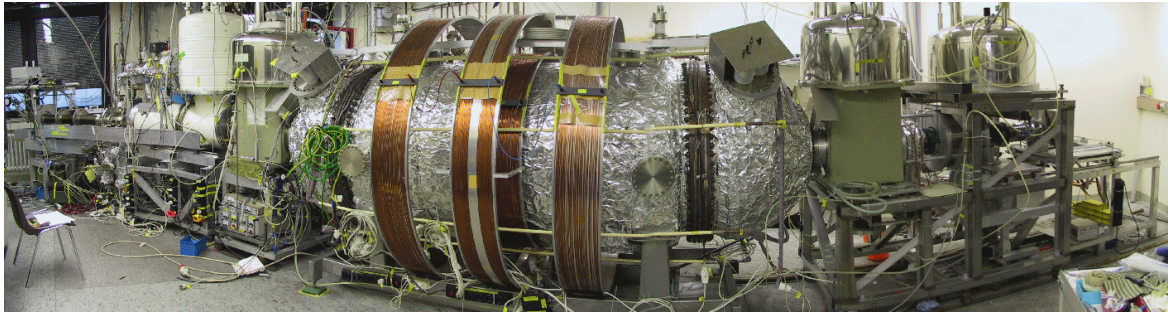
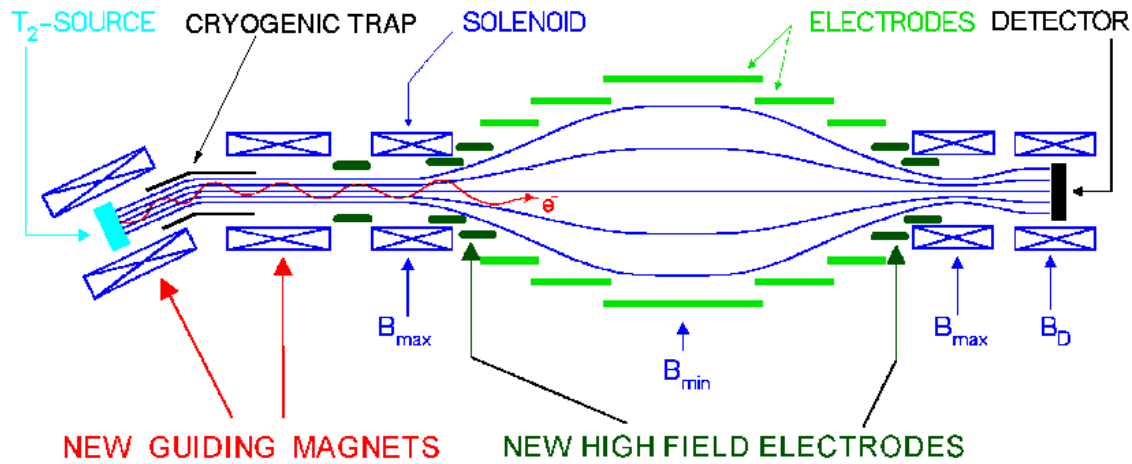
Magnetic Adiabatic Collimation + Electrostatic Filter
(A. Picard et al., Nucl. Instr. Meth. 63 (1992) 345)



⇒ sharp integrating transmission function without tails:

$$\Delta E = E \cdot \frac{B_{\min}}{B_{\max}} = E \cdot \frac{A_{s,\text{eff}}}{A_{\text{analyse}}} = 0.93 \text{ eV, KATRIN} \quad (4.8 \text{ eV, Mainz})$$

The Mainz Neutrino Mass Experiment Phase 2: 1997-2001



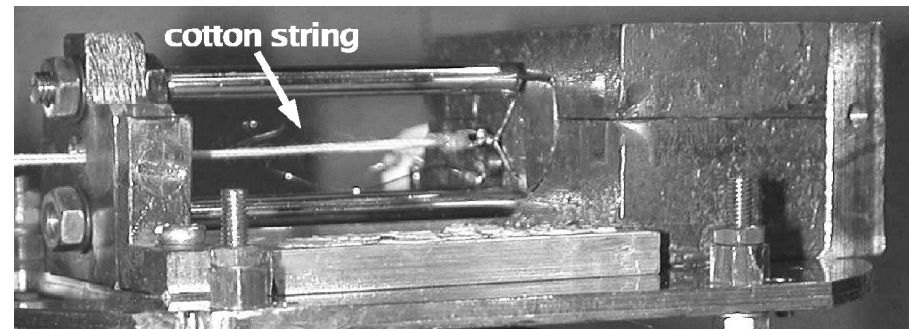
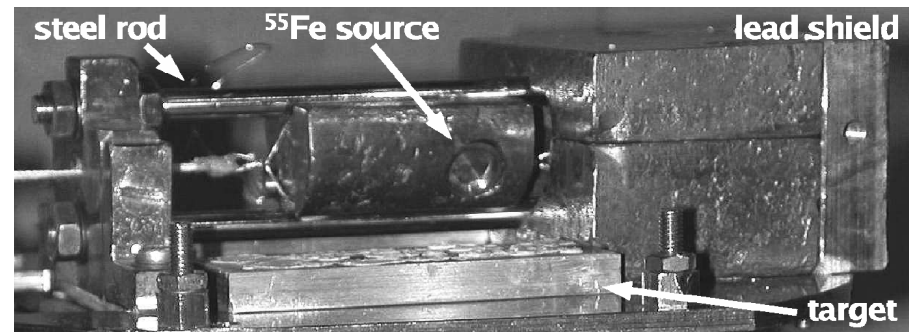
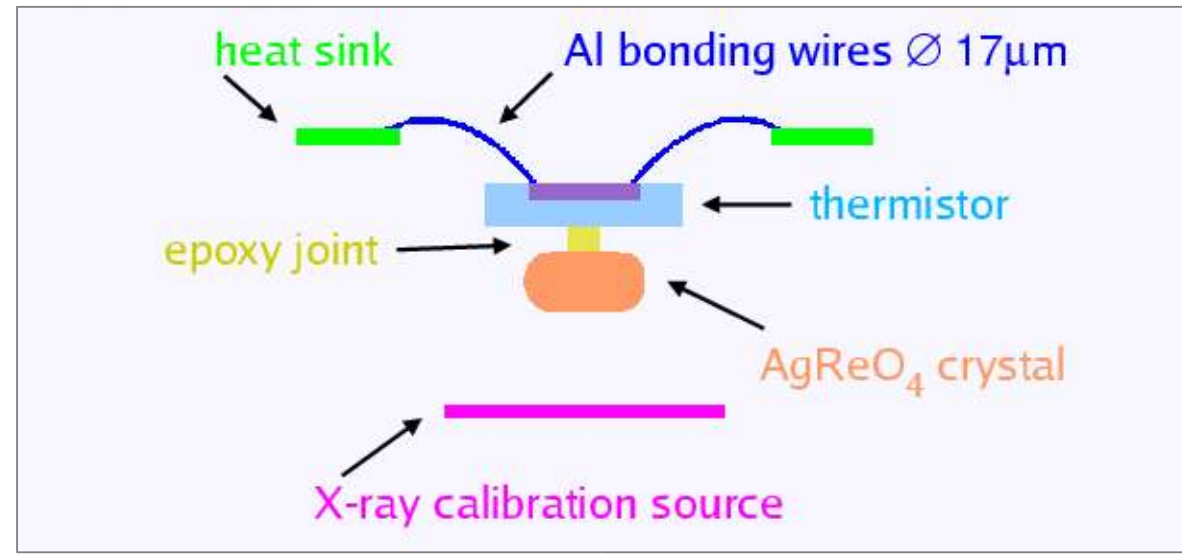
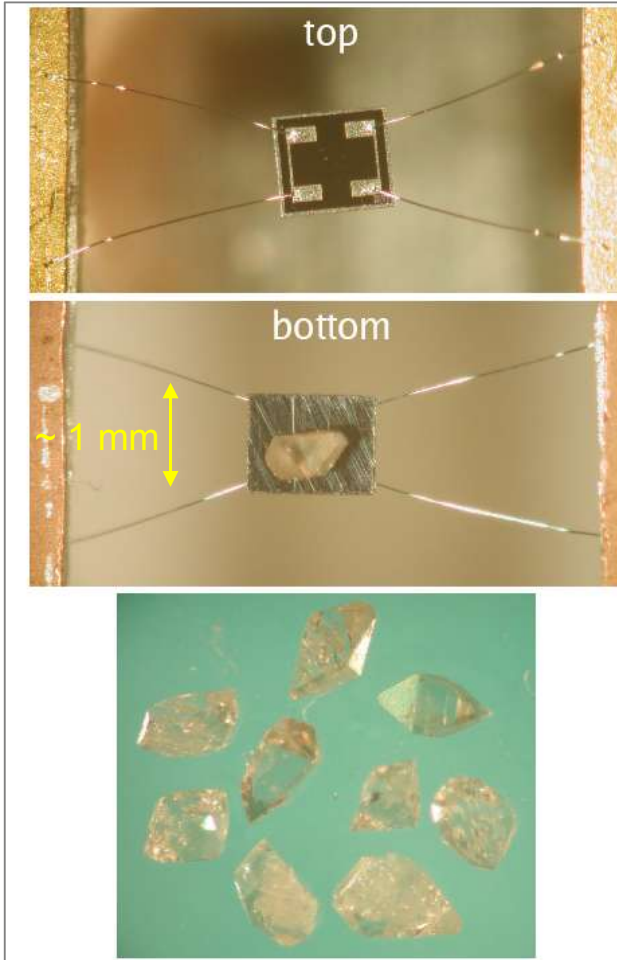
After all critical systematics measured by own experiment
(inelastic scattering, self-charging, neighbor excitation):

$$m^2(\nu) = -0.6 \pm 2.2 \pm 2.1 \text{ eV}^2 \Rightarrow m(\nu) < 2.3 \text{ eV (95\% C.L.)}$$

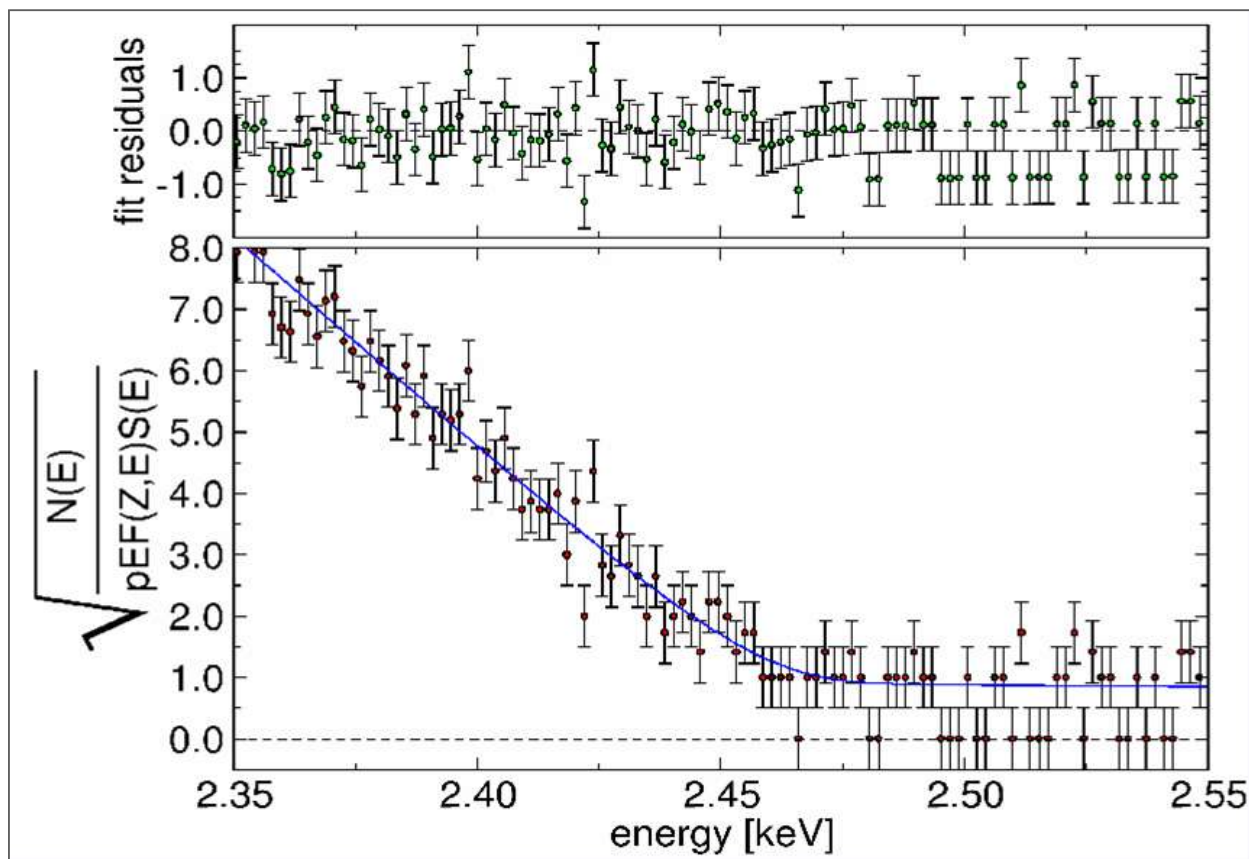
C. Kraus et al., Eur. Phys. J. C 40 (2005) 447

MIBETA (Milano/Como)

AgReO_4 (10 * 250 -350 mg)



MIBETA: final result



Parameters

detectors: 10

rate each: 0.13 1/s

energy res.: $\Delta E = 28$ eV

pile-up frac.: $1.7 \cdot 10^{-4}$

$$M_\nu^2 = -141 \pm 211_{\text{stat}} \pm 90_{\text{sys}} \text{ eV}^2$$

$$M_\nu < 15.6 \text{ eV (90\% c.l.)}$$

(M. Sisti et al., NIMA520 (2004) 125)

MANU (Genova)

- Re metallic crystal (1.5 mg)
- BEFS observed (F.Gatti et al., Nature 397 (1999) 137)
- sensitivity: $m(\nu) < 26$ eV (F.Gatti, Nucl. Phys. B91 (2001) 293)

Proposal: Microcalorimeter Arrays for a Rhenium Experiment (MARE)

Collaboration: Genova, Goddard Space Flight Center/NASA, Heidelberg, Como, Milano, Trento, U Wisconsin

Idea: 2nd and 3rd generation rhenium β decay experiment

MARE I: 300 detectors (MIBETA: 10)
 $\Delta E = 10$ eV (MIBETA: 28 eV)
 $\tau = 10^{-4}$ s (MIBETA: 10^{-3} s)
with semiconductor sensors (like MIBETA/MANU)

a

expected sensitivity on $m(\nu_e)$: 2-3 eV
(before full start of KATRIN in 2010)

MARE I looks realistic as a first step
really complementary to tritium β experiments

Proposal: Microcalorimeter Arrays for a Rhenium Experiment (MARE)

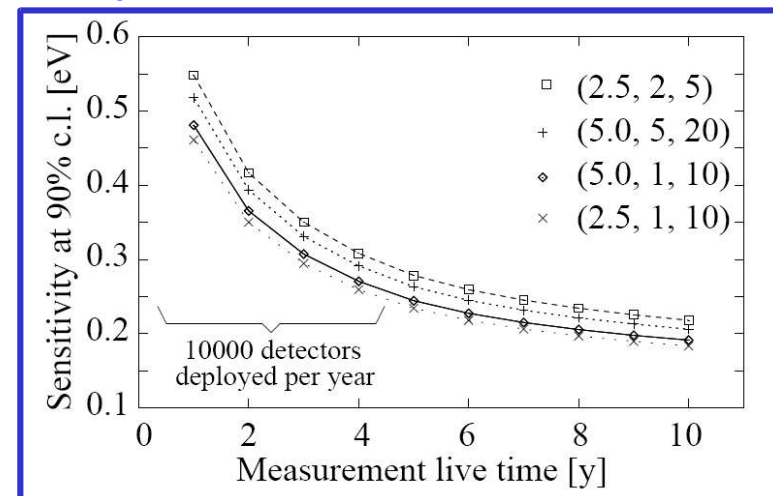
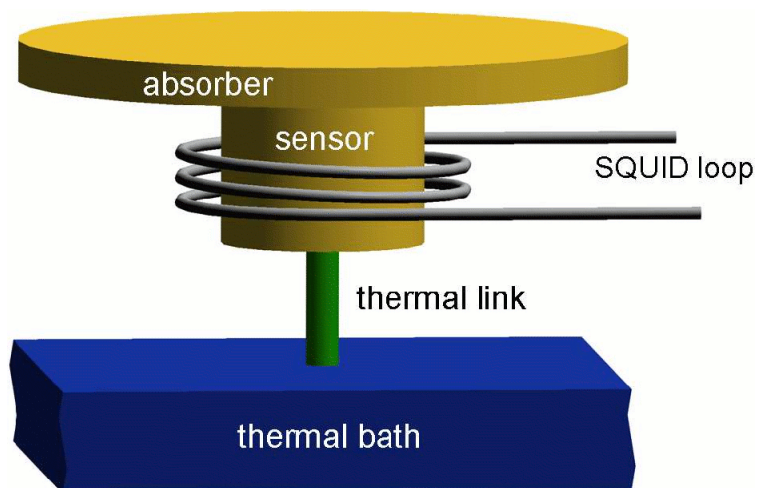
Collaboration: Genova, Goddard Space Flight Center/NASA, Heidelberg, Como, Milano, Trento, U Wisconsin

Idea: 2nd and 3rd generation rhenium β decay experiment

MARE II: 5000 – 50000 detectors (MIBETA: 10)
 $\Delta E = 2.5 - 5$ eV (MIBETA: 28 eV)
 $\tau = \text{a few } 10^{-6}$ s (MIBETA: 10^{-3} s)
 with superconducting transition edge sensors (TES) or
 with metallic magnetic temperature sensors (MMC)

a

expected sensitivity on $m(\nu_e)$: 0.2 eV



Proposal: Microcalorimeter Arrays for a Rhenium Experiment (MARE)

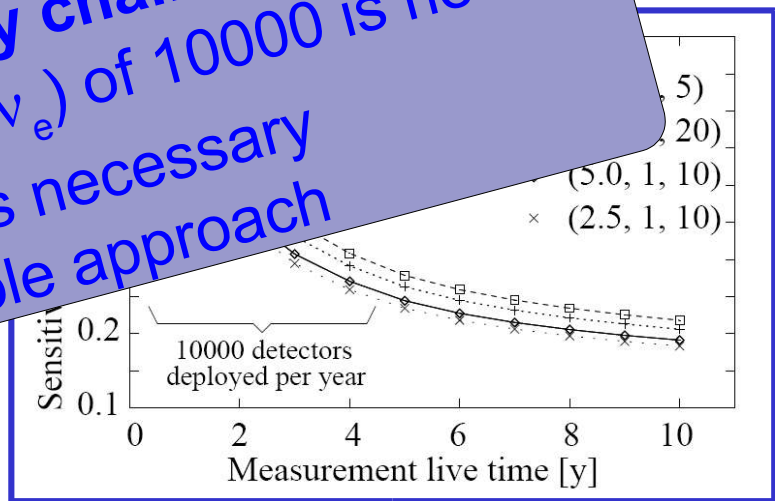
Collaboration: Genova, Goddard Space Flight Center/NASA, Heidelberg, Como, Milano, Trento, U Wisconsin

Idea: 2nd and 3rd generation rhenium β decay experiment

MARE II: 5000 – 50000 detectors (MIBETA: 10)
 $\Delta E = 5$ eV (MIBETA: 28 eV)
 $\tau =$ a few 10^{-6} s (MIBETA: 10^{-3} s)
 with superconducting transition edge sensors (TES) or
 with metallic magnetic temperature sensors (MTC)

expected sensitivity on $m^2(\nu_e)$

MARE II looks really challenging,
 since an improvement in $m^2(\nu_e)$ of 10000 is needed
 \Rightarrow a lot of R&D is necessary
 but it is a scalable approach





The Karlsruhe Tritium Neutrino experiment KATRIN



is being set up at the Forschungszentrum Karlsruhe

Physics Aim:

Improvement of sensitivity by 1 order of magnitude: $2.2 \text{ eV} \rightarrow 0.2 \text{ eV}$

- higher energy resolution: $\Delta E \approx 1 \text{ eV}$

since $E/\Delta E \sim A_{\text{spectrometer}}$

\Rightarrow larger spectrometer

- relevant region below endpoint becomes smaller

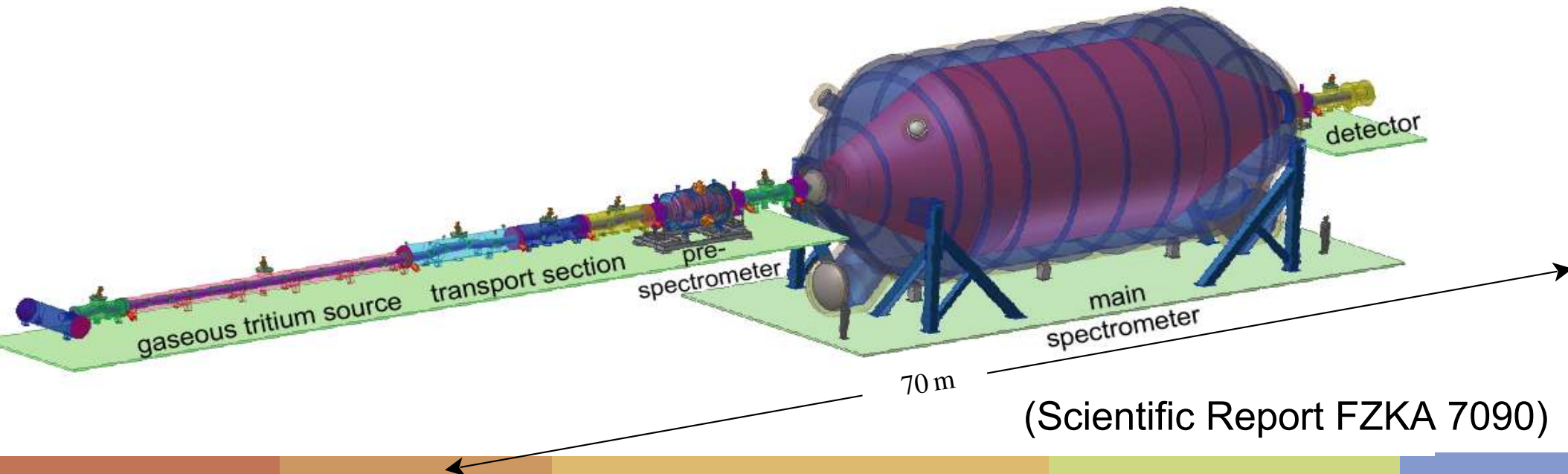
even less count rate $dN/dt \sim A_{\text{spectrometer}}$

\Rightarrow larger spectrometer

- much longer measurement time:

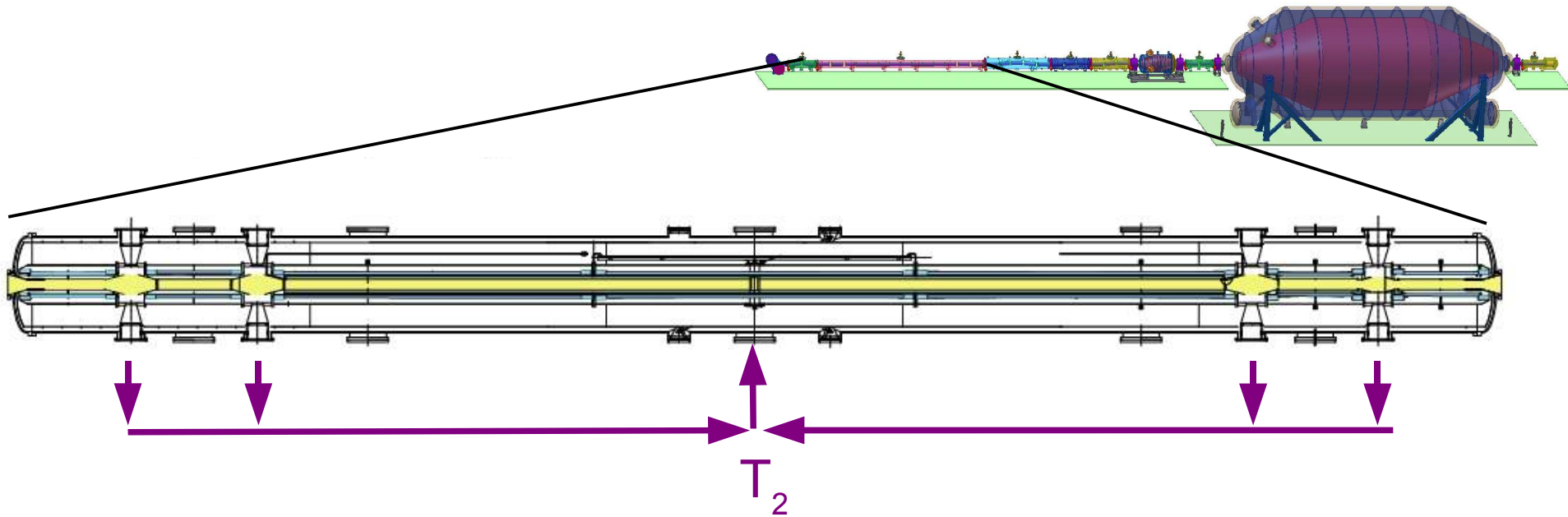
$100 \text{ d} \rightarrow 1000 \text{ d}$

} $\varnothing 10\text{m}$



(Scientific Report FZKA 7090)

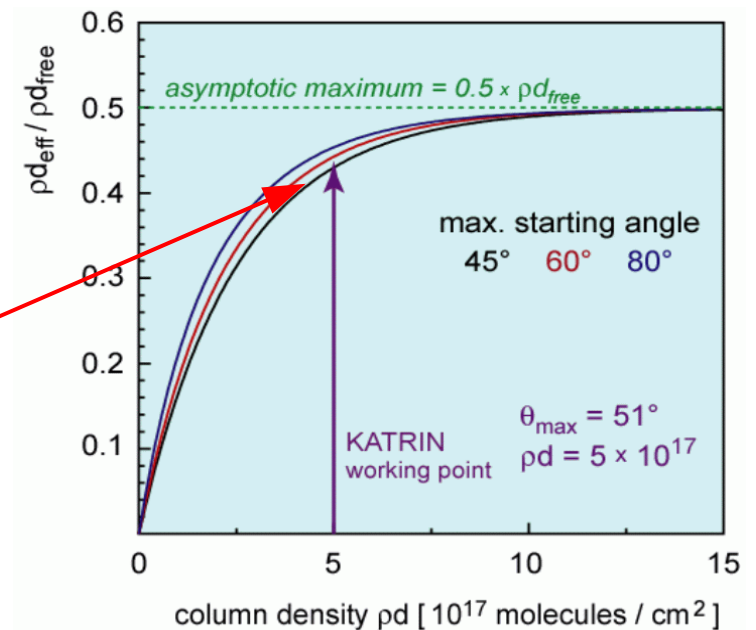
Molecular Windowless Gaseous Tritium Source WGTS



WGTS: tub in long superconducting solenoids
 \varnothing 9cm, length: 10m, T = 30 K

Tritium recirculation (and purification)
 $p_{inj} = 0.003$ mbar, $q_{inj} = 4.7$ Ci/s

allows to measure with near to
 maximum count rate using
 $\rho d = 5 \cdot 10^{17}/\text{cm}^2$
 with small systematics

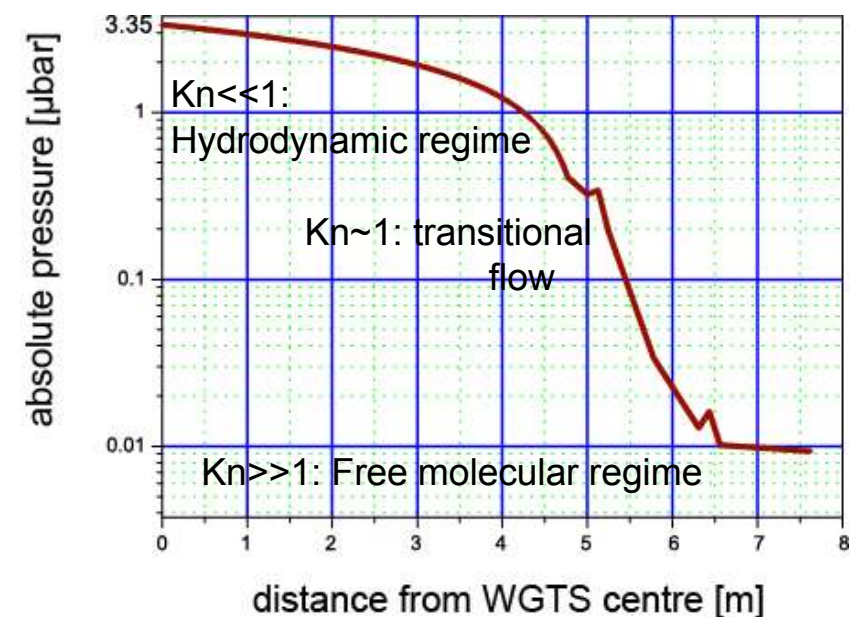
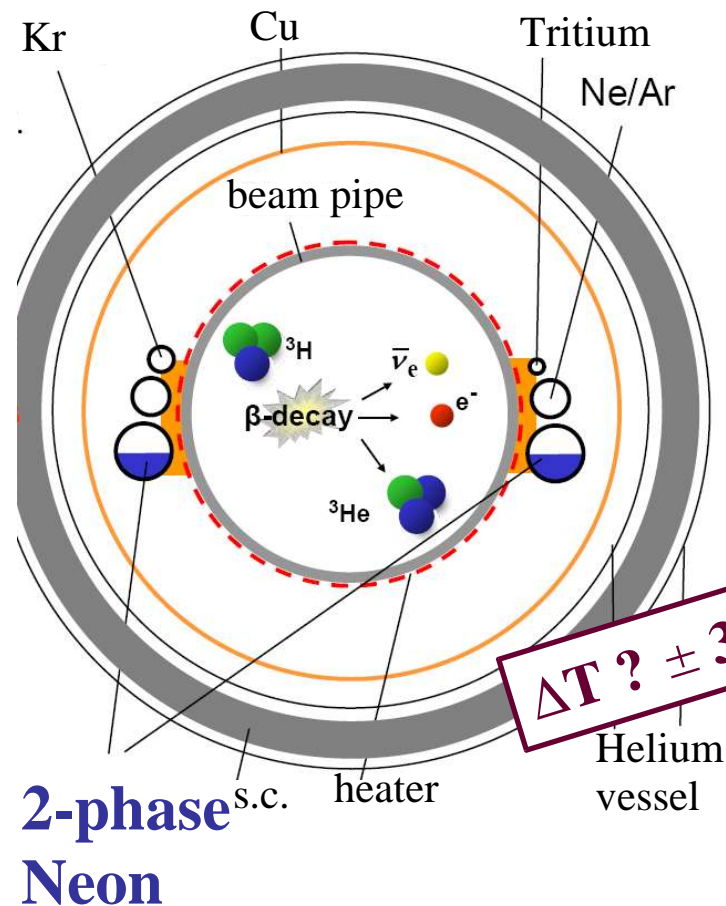
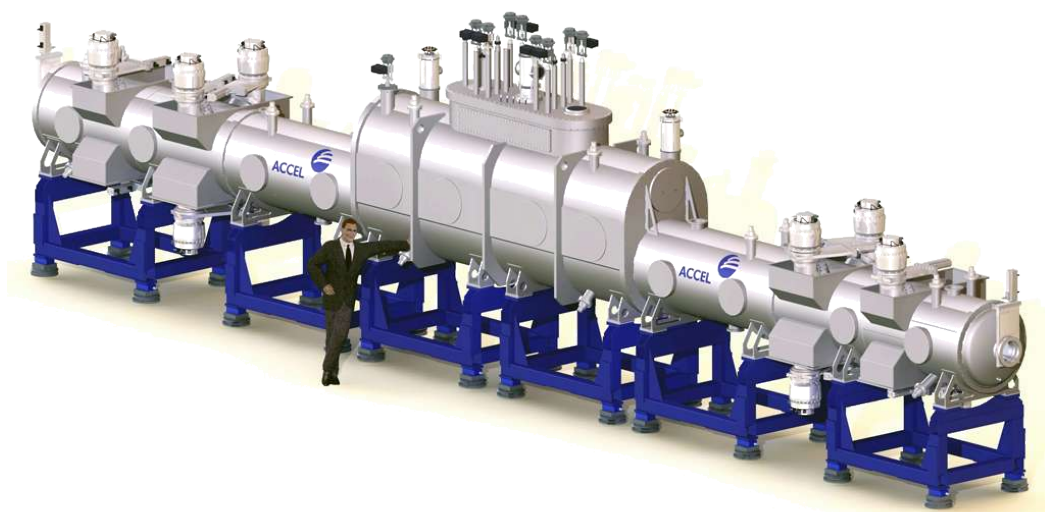


Molecular Windowless Gaseous Tritium Source WGTS

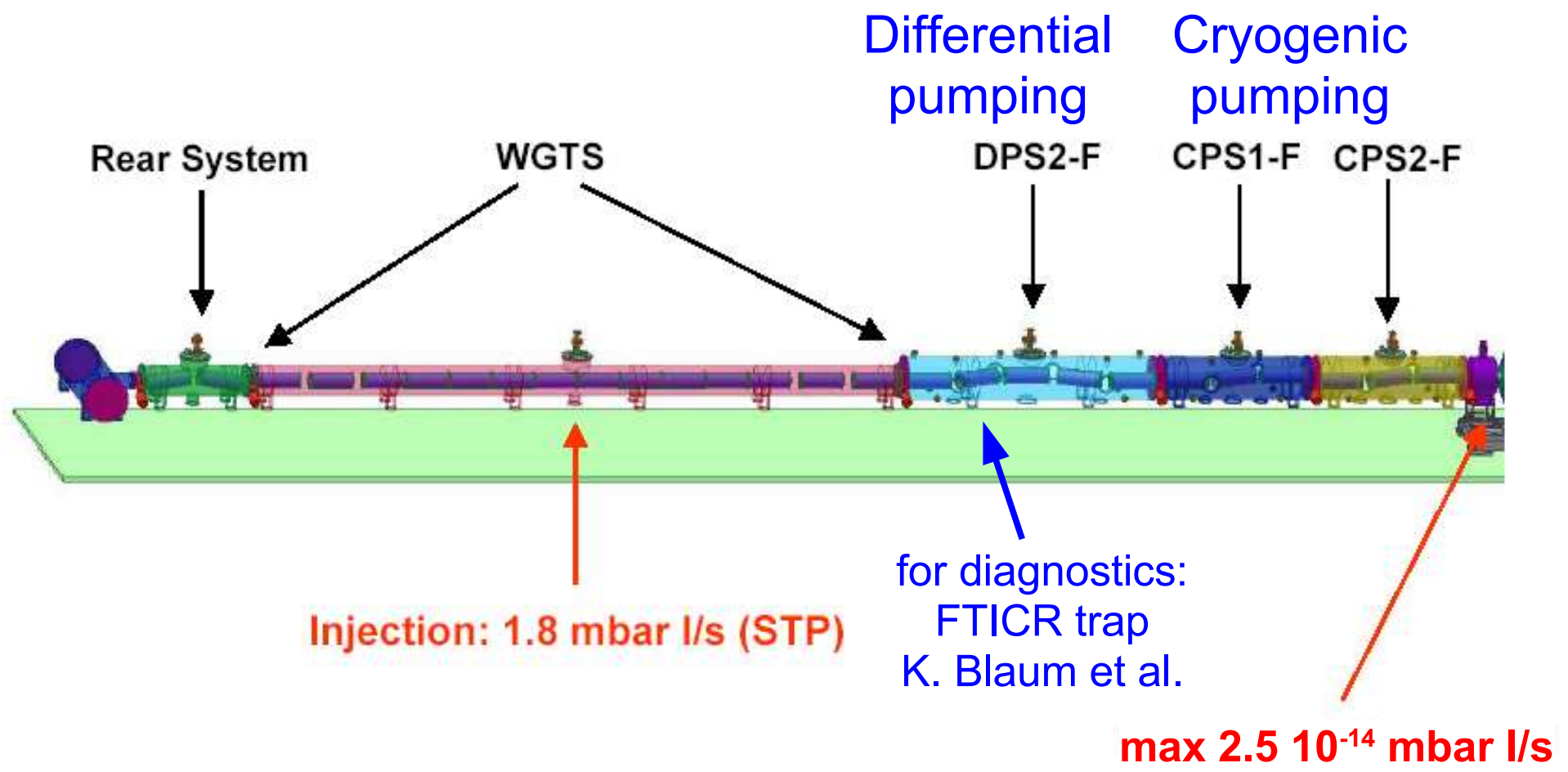
Conceptional design

2 phase Neon cooling with
operating temperature: 27–28 K

- **spatial** (homogeneity): $\pm 0.1\%$
- **time** (stability/hour): $\pm 0.1\%$



Transport and differential & cryo pumping sections



requirements:

- adiabatic electron guiding
- T_2 reduction factor of $\sim 10^{14}$

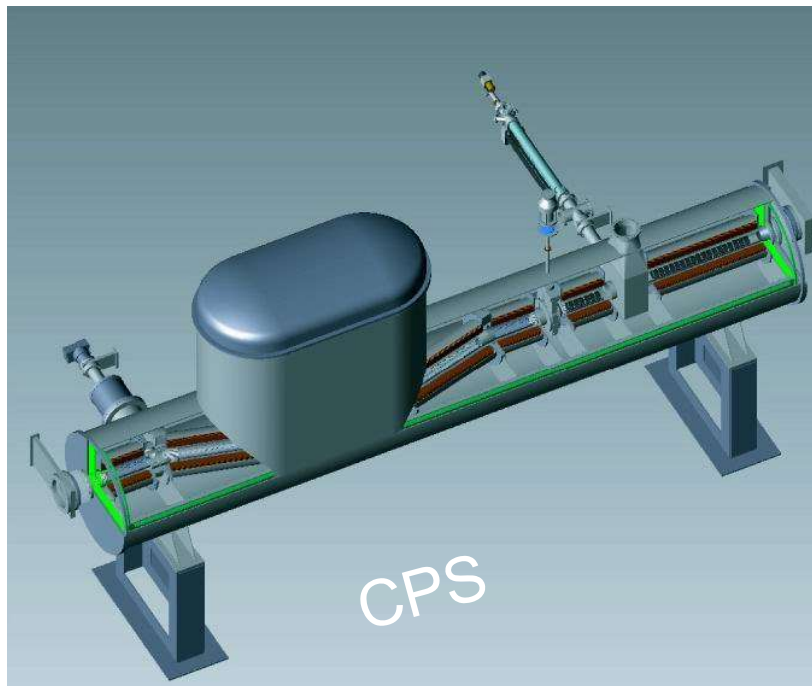
Cryogenic pumping section

Objective: retention of remaining tritium flux, reduction factor 10^7

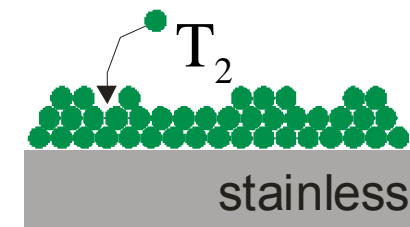
(tritium partial pressure in main spectrometer $p < 10^{-20}$ mbar)

method: **cryo-sorption** on condensed Ar-frost

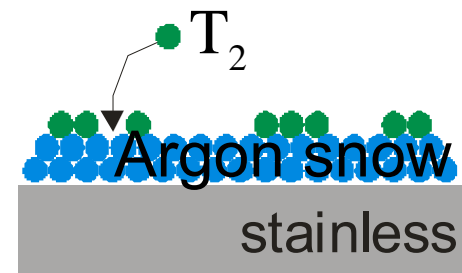
rate: <1 Ci T_2 in 60 days (regeneration with warm He-gas)



Cryocondensation



Cryosorption



Inlet flow rate: $\approx 10^{-6}$ mbar l/s

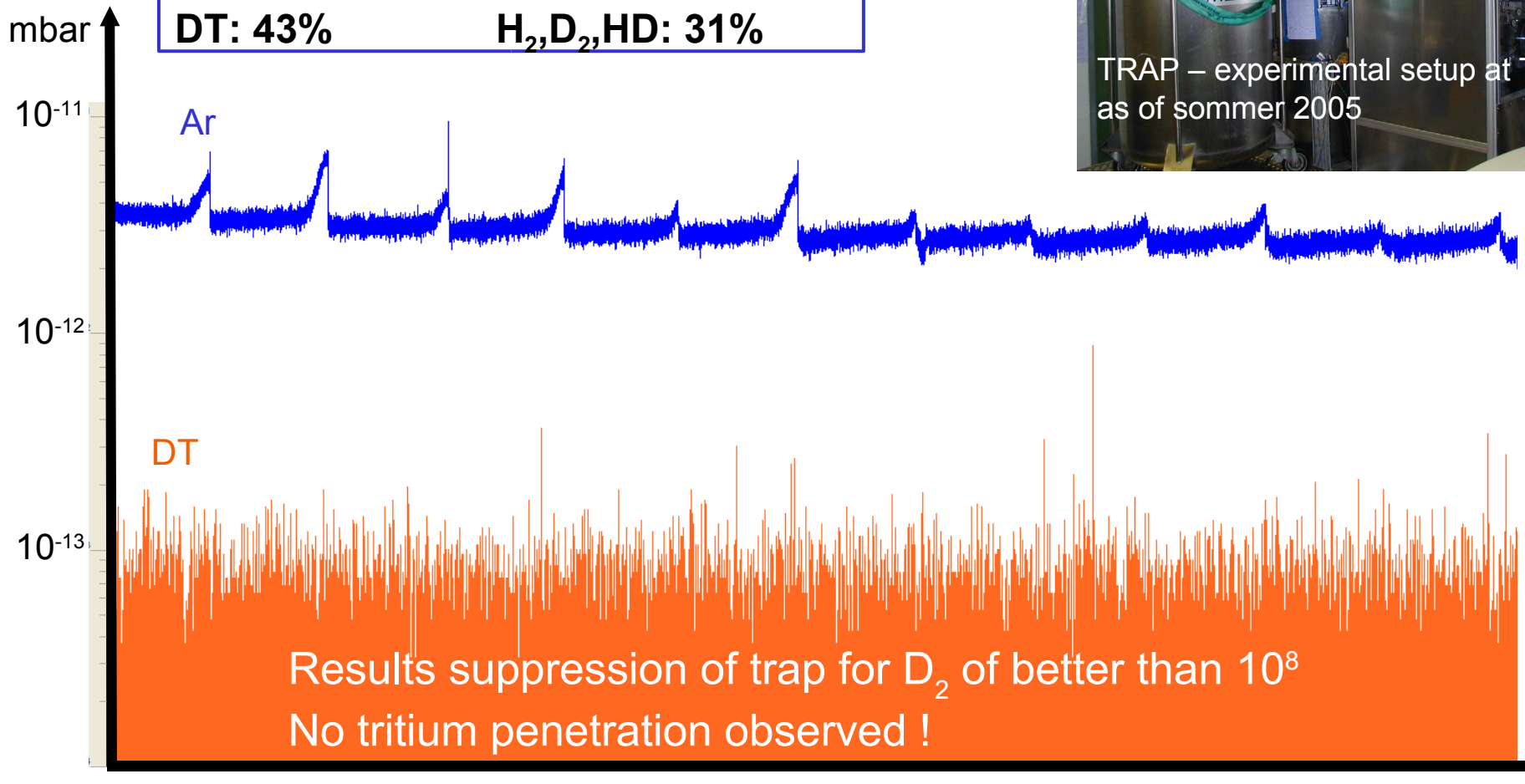
Isotopic composition (CAPER-GC):

HT: 7%

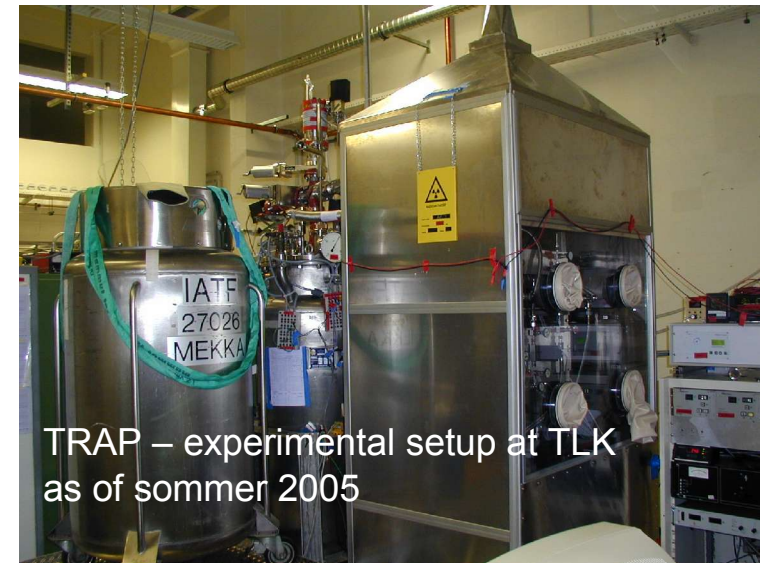
T₂: 19%

DT: 43%

H₂,D₂,HD: 31%



Results suppression of trap for D₂ of better than 10⁸
No tritium penetration observed !

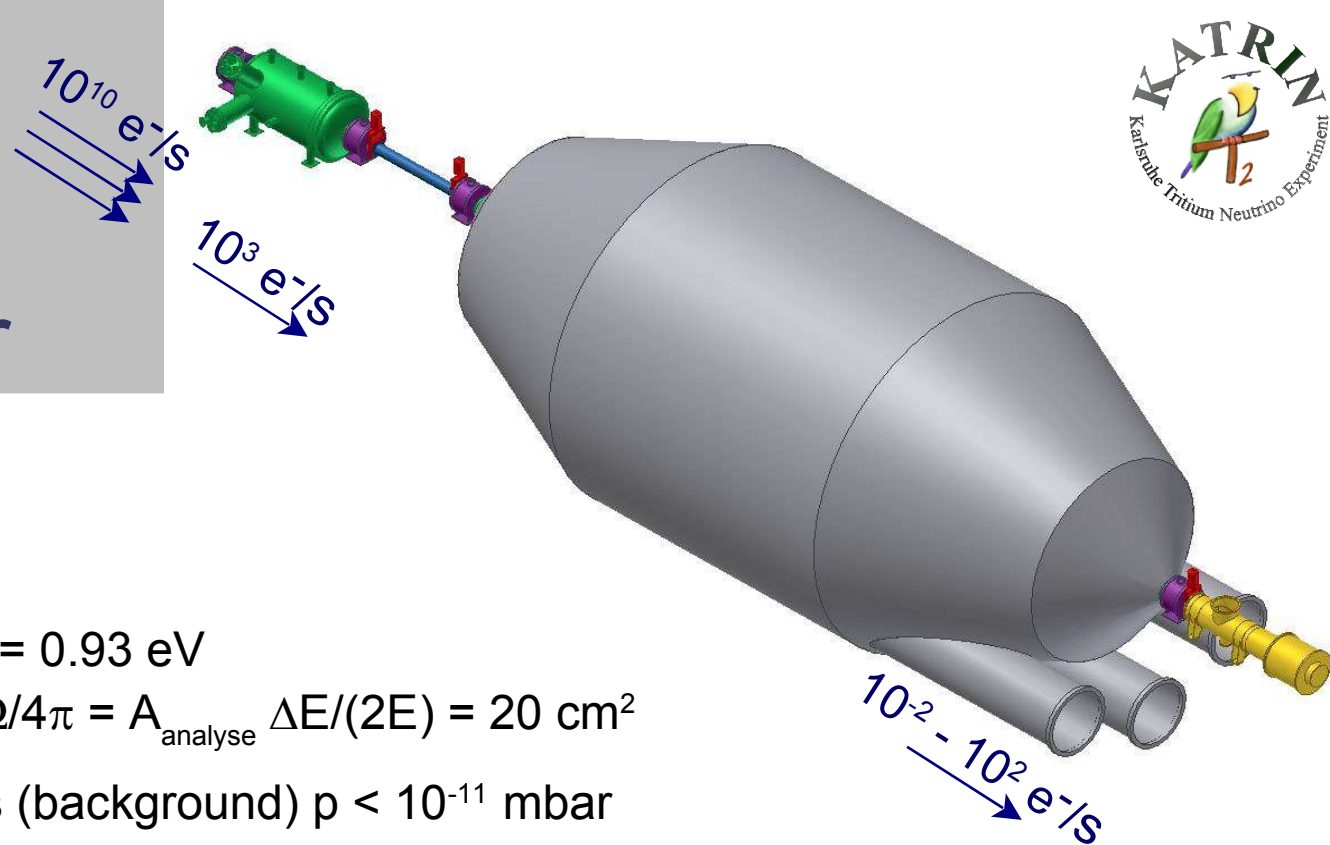


TRAP – experimental setup at TLK as of sommer 2005



6 days

Pre and main spectrometer



Main spectrometer:

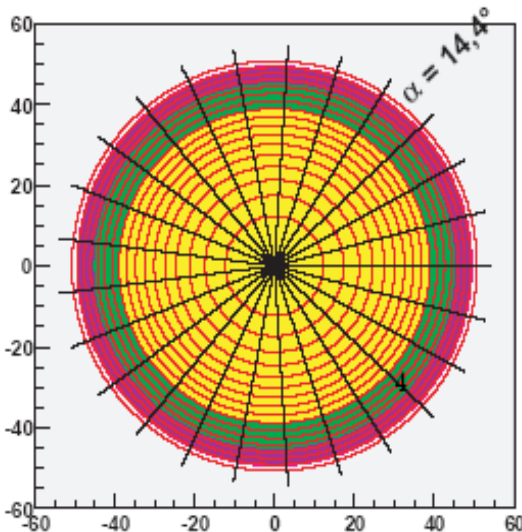
- $\varnothing 10\text{m}$, length 24m
 - \Rightarrow large energy resolution: $\Delta E = 0.93 \text{ eV}$
 - \Rightarrow high luminosity: $L = A_{\text{Seff}} \Delta\Omega/4\pi = A_{\text{analyse}} \Delta E/(2E) = 20 \text{ cm}^2$
- ultrahigh vacuum requirements (background) $p < 10^{-11} \text{ mbar}$
- „simple“ construction: vacuum vessel at HV + „massless“ screening electrode

Pre spectrometer

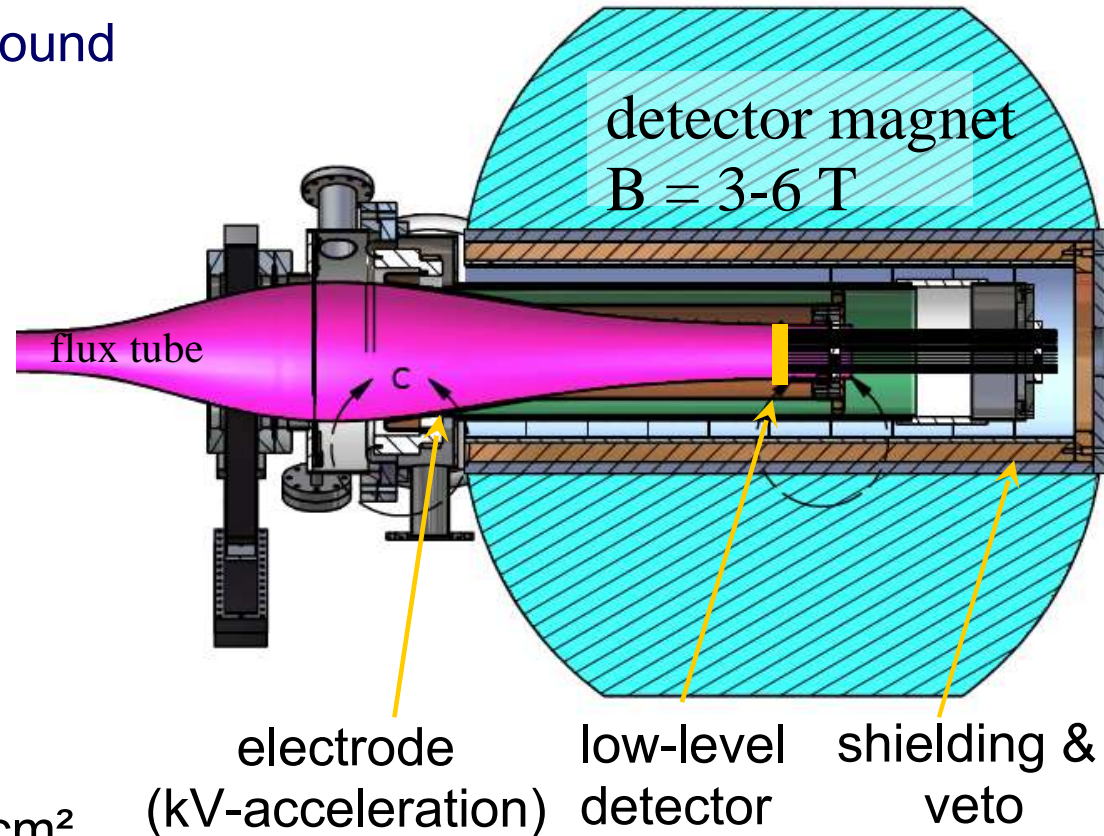
- Transmission of electron with highest energy only
 - (10^{-7} part in last 100 eV)
 - \Rightarrow Reduction of scattering probability in main spectrometer
 - \Rightarrow Reduction of background
- only moderate energy resolution required: $\Delta E = 80 \text{ eV}$
- test of new ideas (XHV, shape of electrodes, avoid and remove of trapped particles, ...)

Detector

task: **detection** of transmitted β -decay electrons
 with high energy resolution ($\Delta E = 1$ keV)
 record **radial profile** of flux tube
 aim: background minimisation, systematic effects
 \Rightarrow post-acceleration to place signal line
 at lower intrinsic background



design: radially segmented
 Si-PIN diode array
 ~150 pixels with $A=100$ cm²



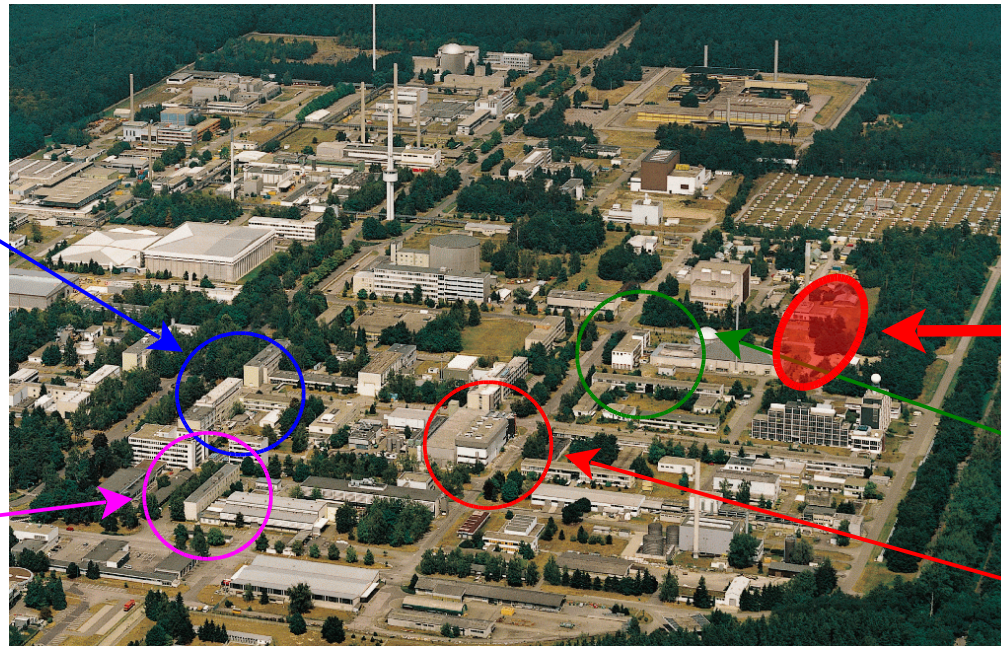
Technical challenges

- Recirculation and purification of tritium to a large extent (kCi)
- ≈ 30 superconducting solenoids
- UHV ($< 10^{-11}$ mbar) in huge volume (1000m²)
- HV calibration and stability on ppm level
- High resolution detectors
-

⇒ ideal place: Forschungszentrum Karlsruhe/Germany

Inst. f. Kernphysik
(IK)

Inst. f. Prozessdaten-
verarbeitung
und Elektronik (IPE)

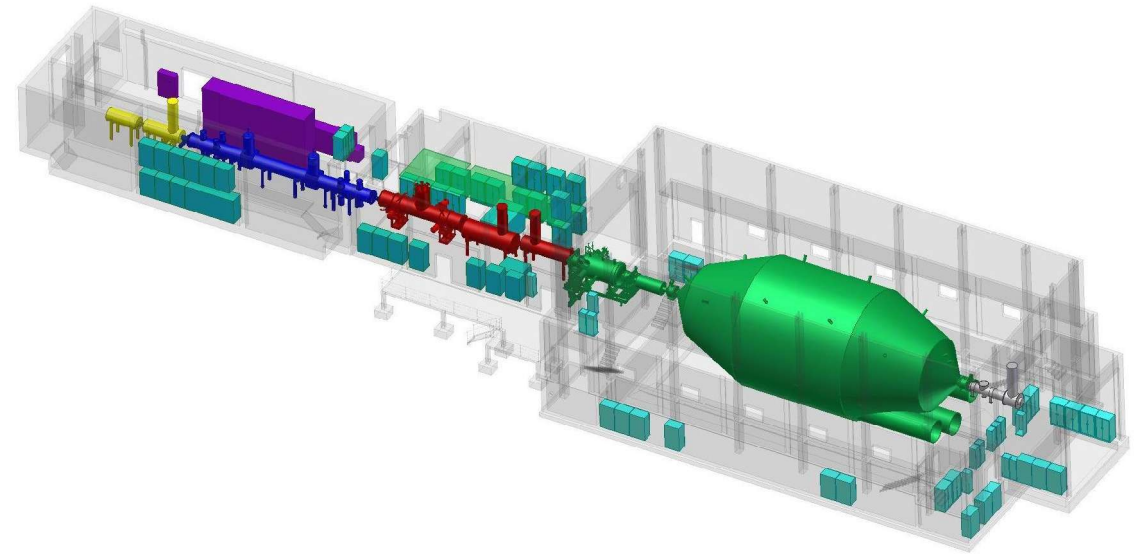


KATRIN

Tritiumlabor
Karlsruhe (TLK)

Institut für Technische
Physik (ITP)

KATRIN's location at Forschungszentrum Karlsruhe

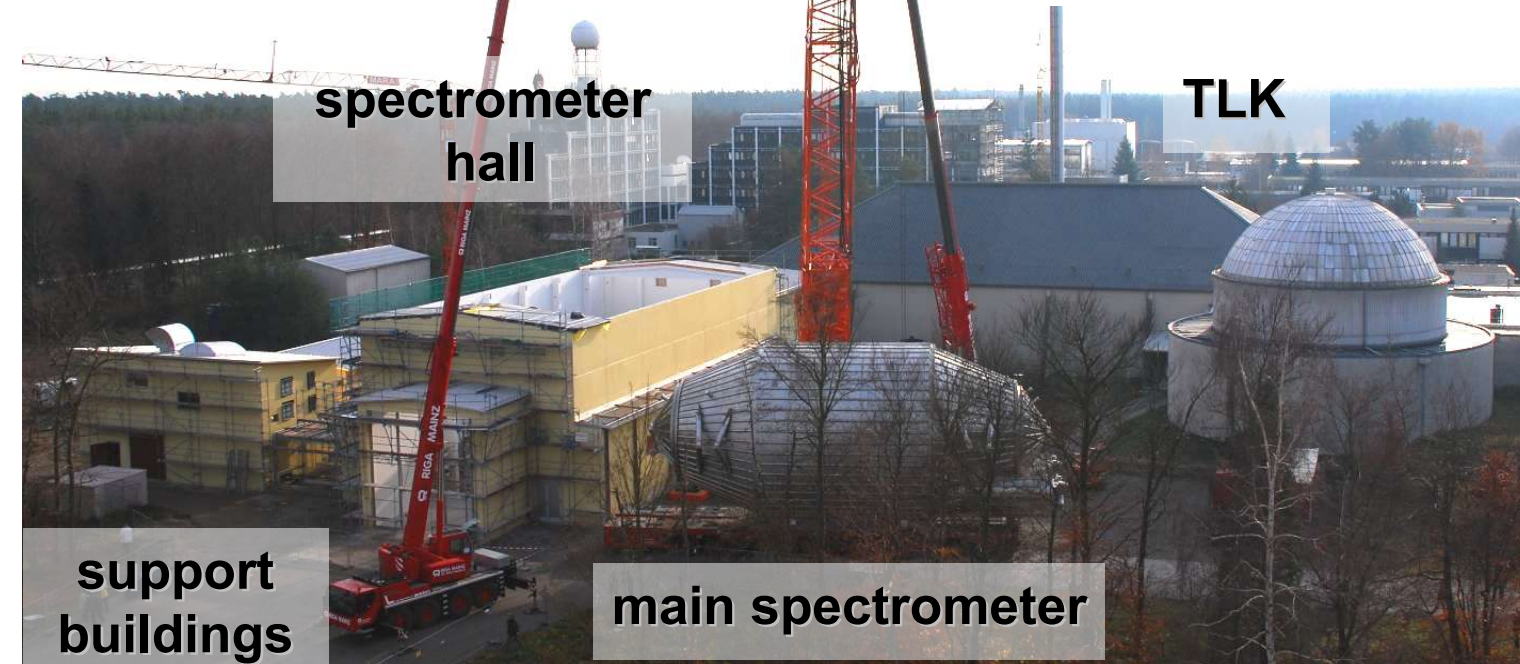


spectrometer
hall

TLK

support
buildings

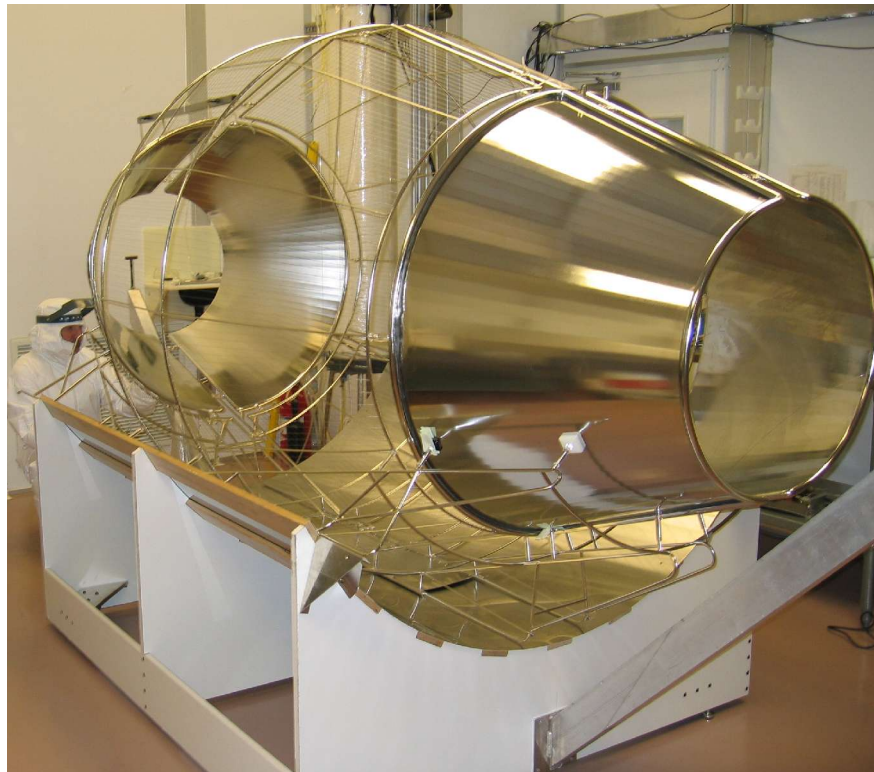
main spectrometer



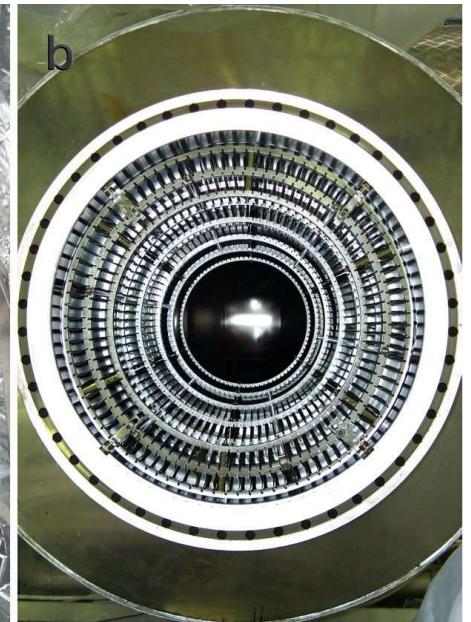
Vacuum tests and inner electrode system of pre spectrometer



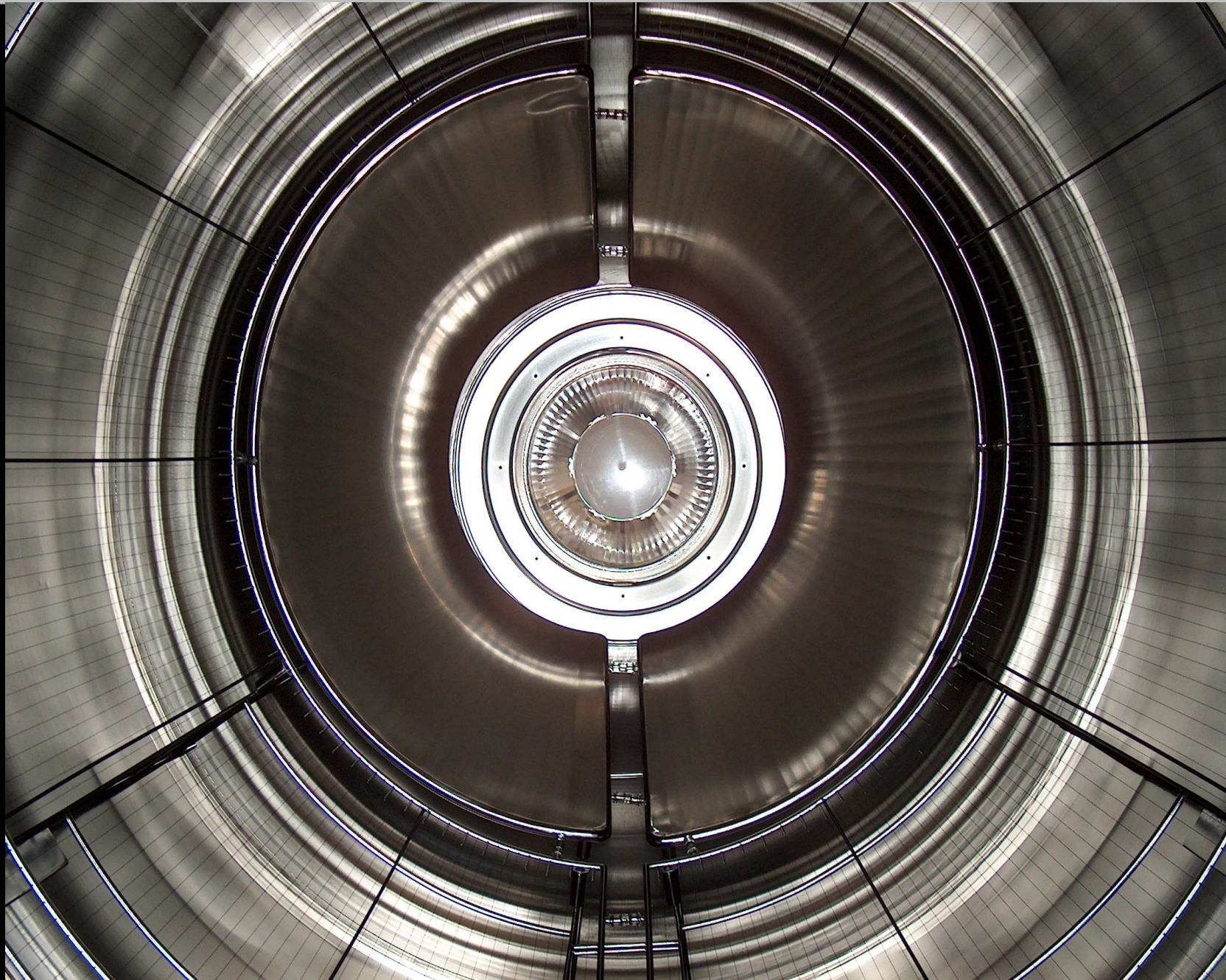
ground electrode
wire electrode + solid cones



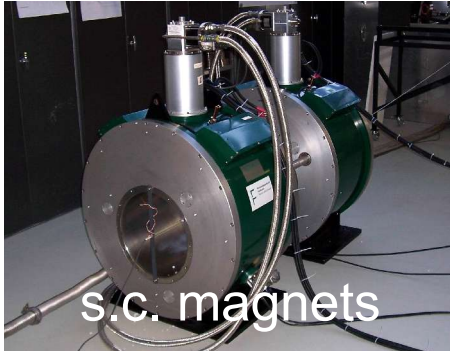
dry air compartment to allow cooling
at -20°C : outgasing rate $< 10^{-13}$ mbar l/s cm^2
with getter pumps (NEG, 25000 l/s): $p < 10^{-11}$ mbar
 \Rightarrow better than KATRIN requirements



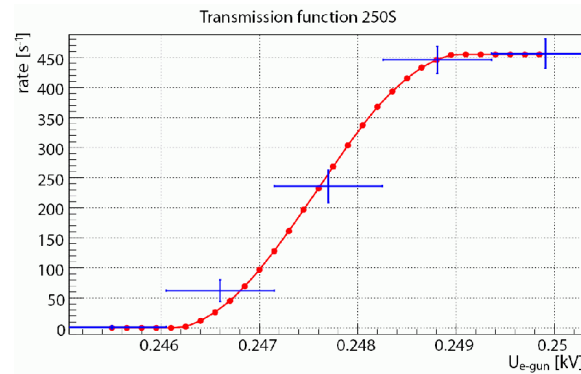
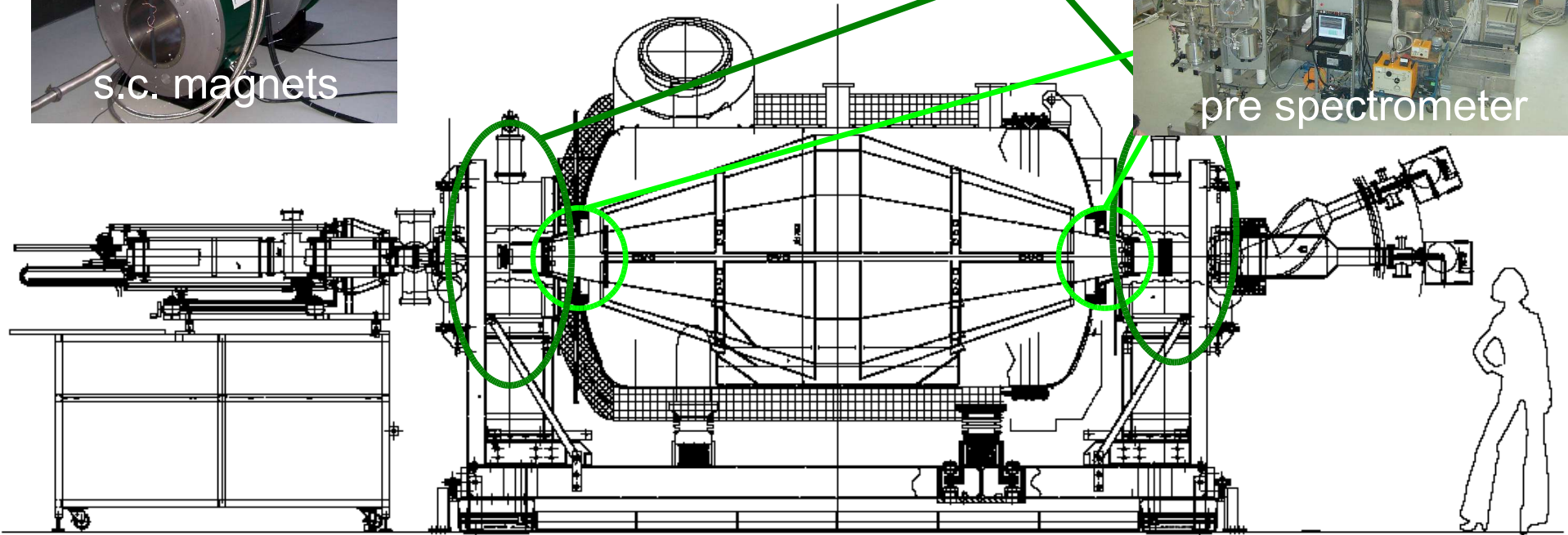
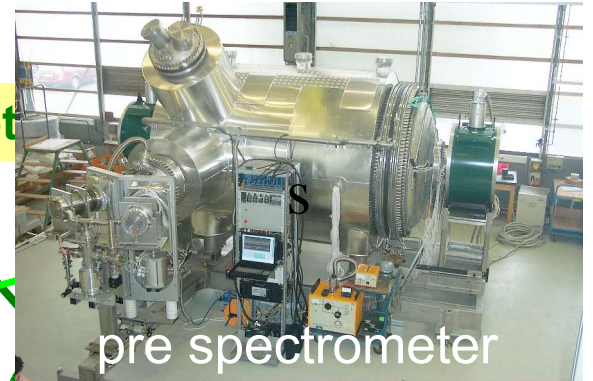
Wire electrode in pre spectrometer



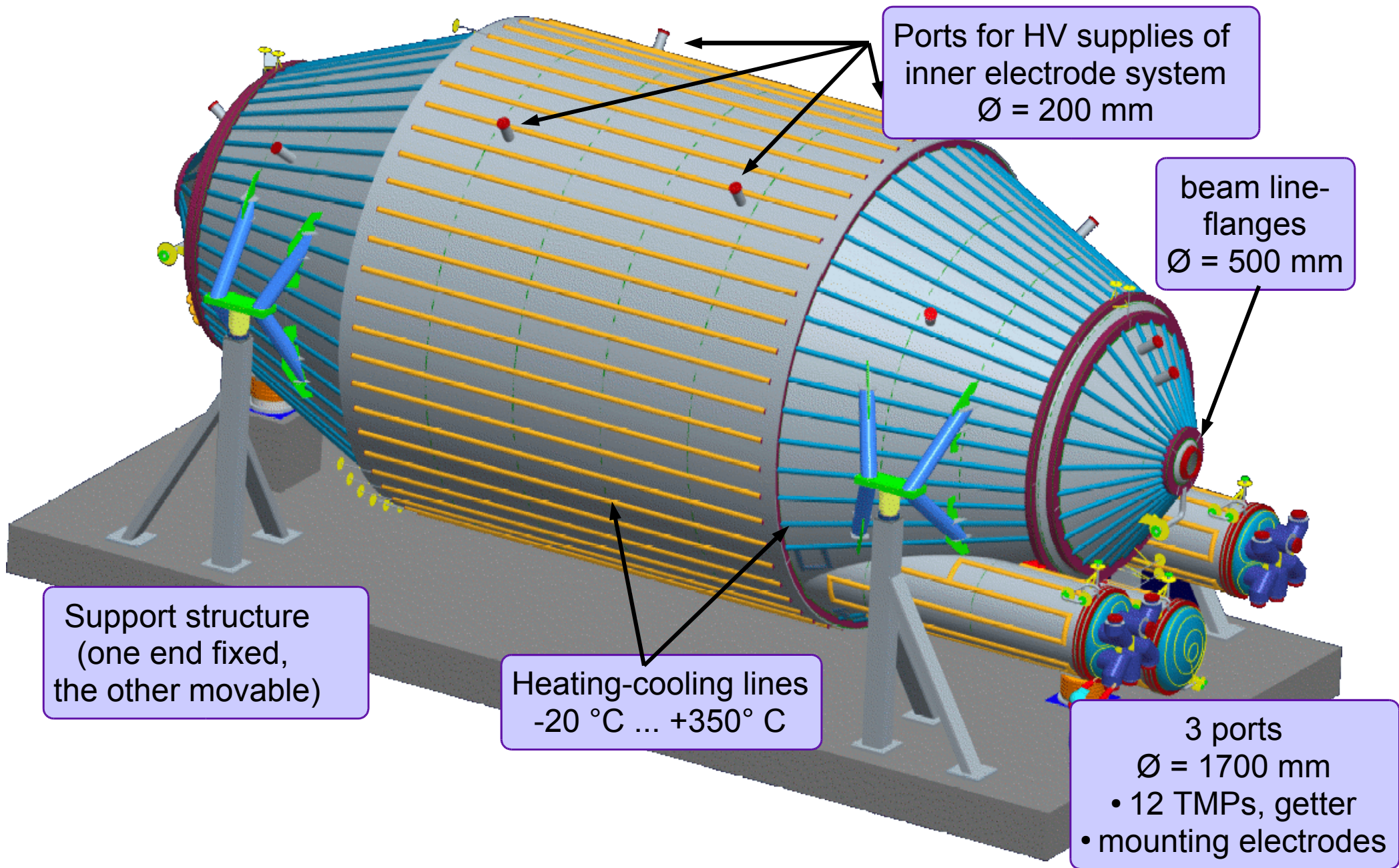
Electromagnetic design tests at the pre spectrometer have just started



s.c. magnet



Main spectrometer 3dim model with heating-cooling system



Main spectrometer vessel construction at MAN DWE



Main spectrometer vessel construction at MAN DWE



21/07/2006

Main spectrometer vessel construction at MAN DWE



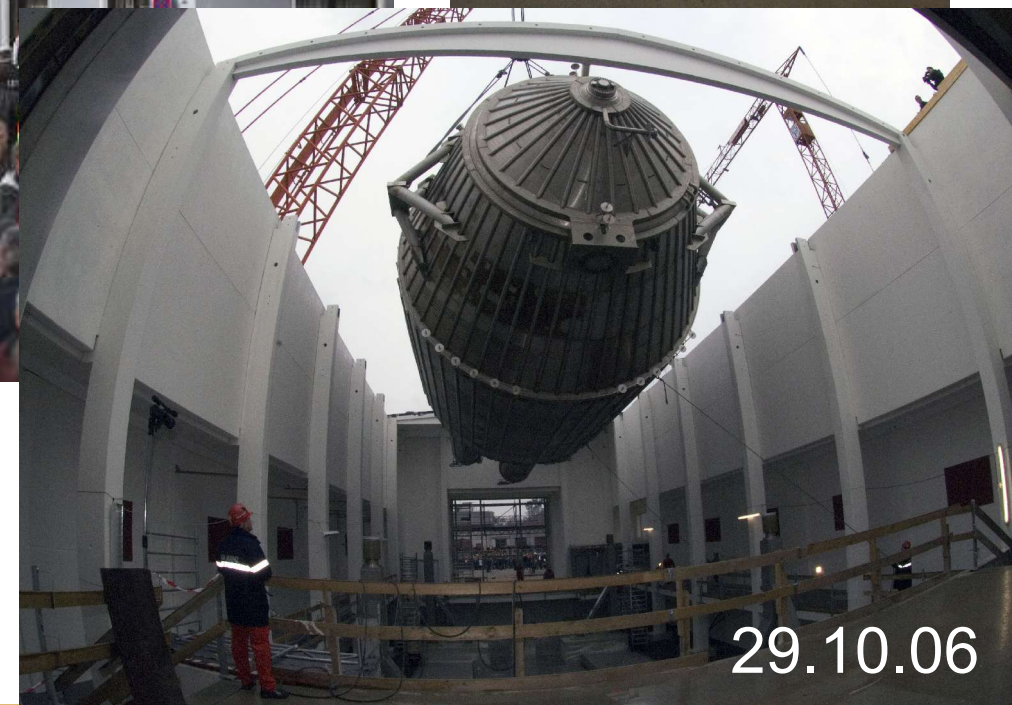
August 2006:

- construction of main spectrometer vessel has been finished
 - vessel has passed leak test successfully !
- ⇒ world`s biggest XHV vessel ever been build !

It is very big and heavy ...
⇒ a 8500 km long detour



Arrival of the Main Spectrometer Vessel: October 2006



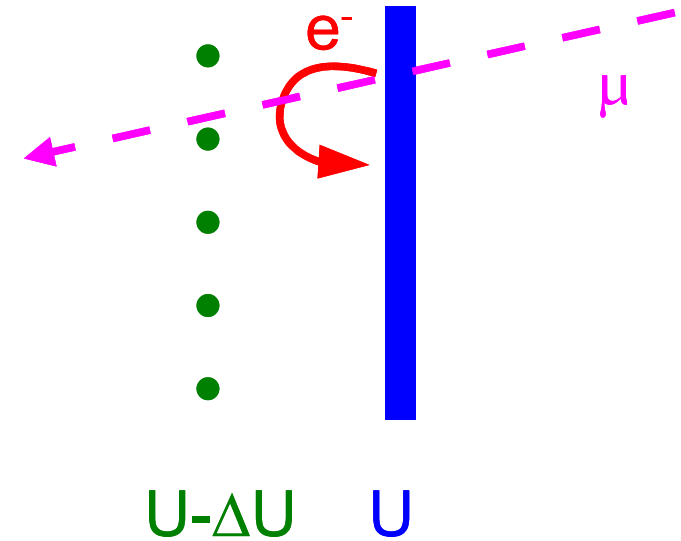
- 1) Low background: Mainz experiment:
most background from spectrometer
but KATRIN spectrometer is much bigger!
⇒ need something new !
- 2) Huge statistics: optimized source &
large spectrometer
- 3) Systematic uncertainties:
need to be very small !

Electric screening by a „massless“ wire electrode

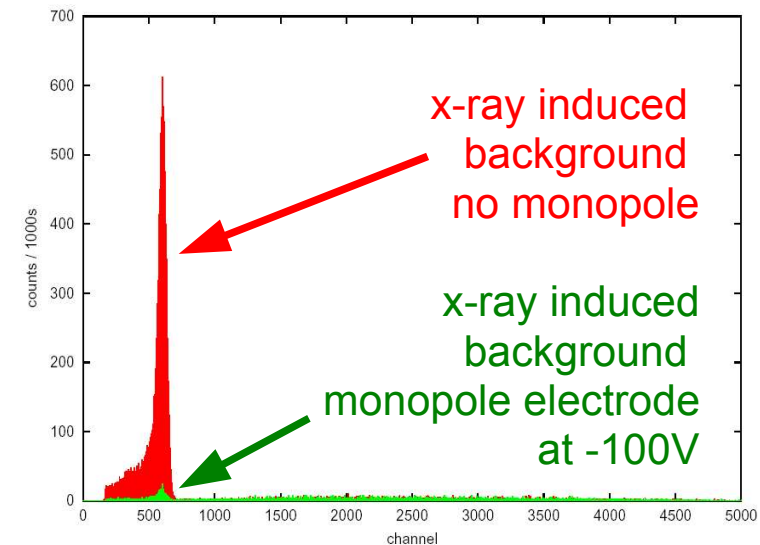
Secondary electrons from wall/electrode

by cosmic rays, environmental radioactivity, ...

wire electrode on slightly more negative potential

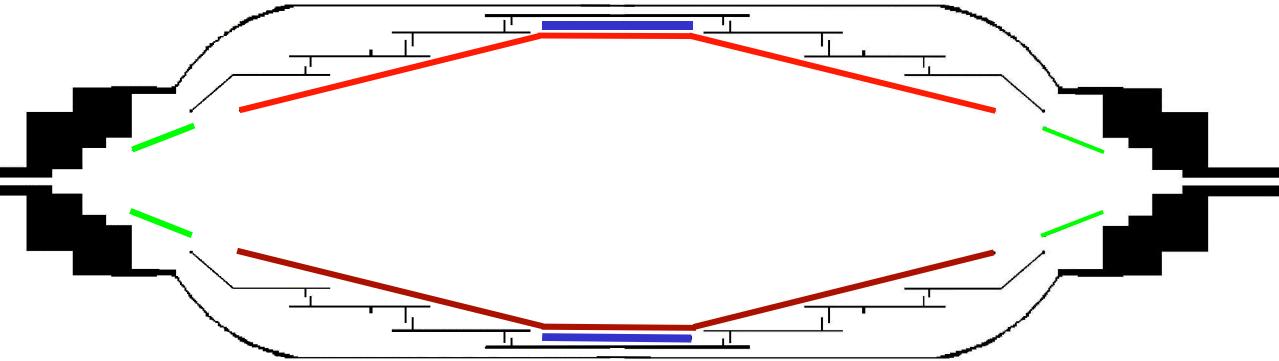


First realisation:
Mainz III
(diploma thesis
Beatrix Müller/Ostrick)



⇒ background suppression by factor 3

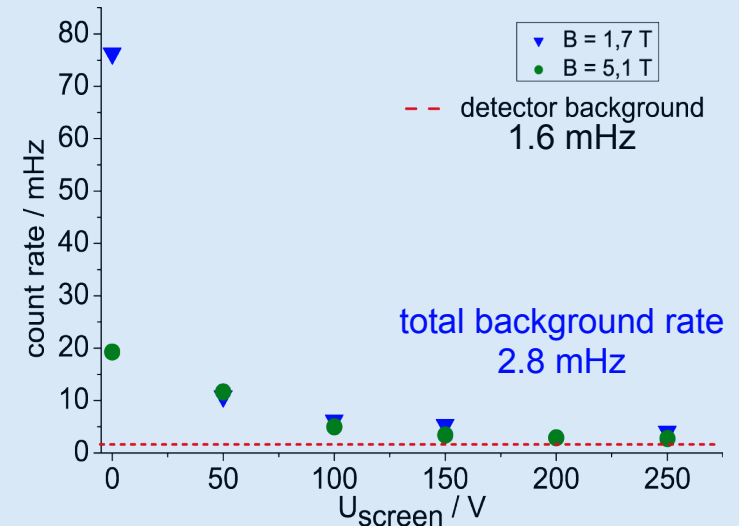
Even more progress with „massless“ wire electrode



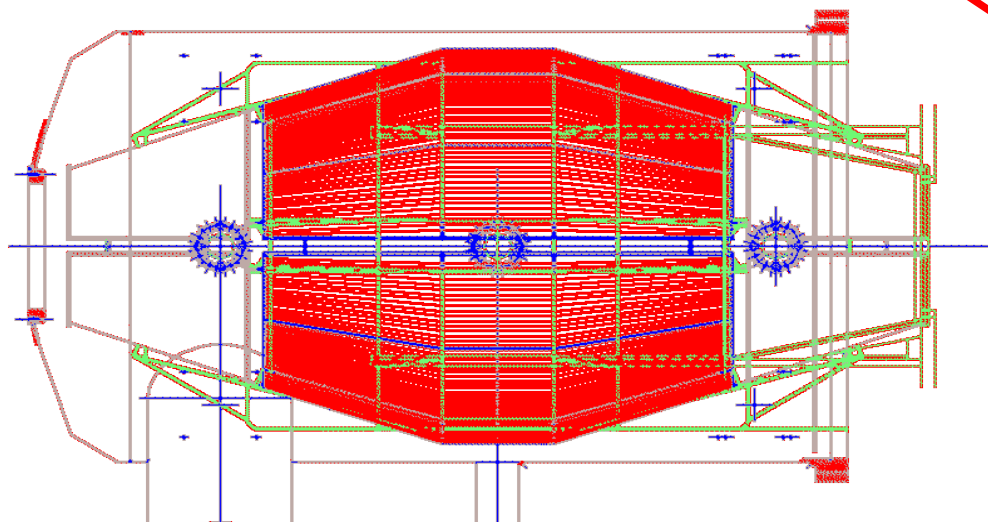
Mainz V (2004)
(PhD thesis Björn Flatt/Mainz)

**New record !
April 04**

Background suppression **successfully tested**
at the Mainz MAC-E filter:

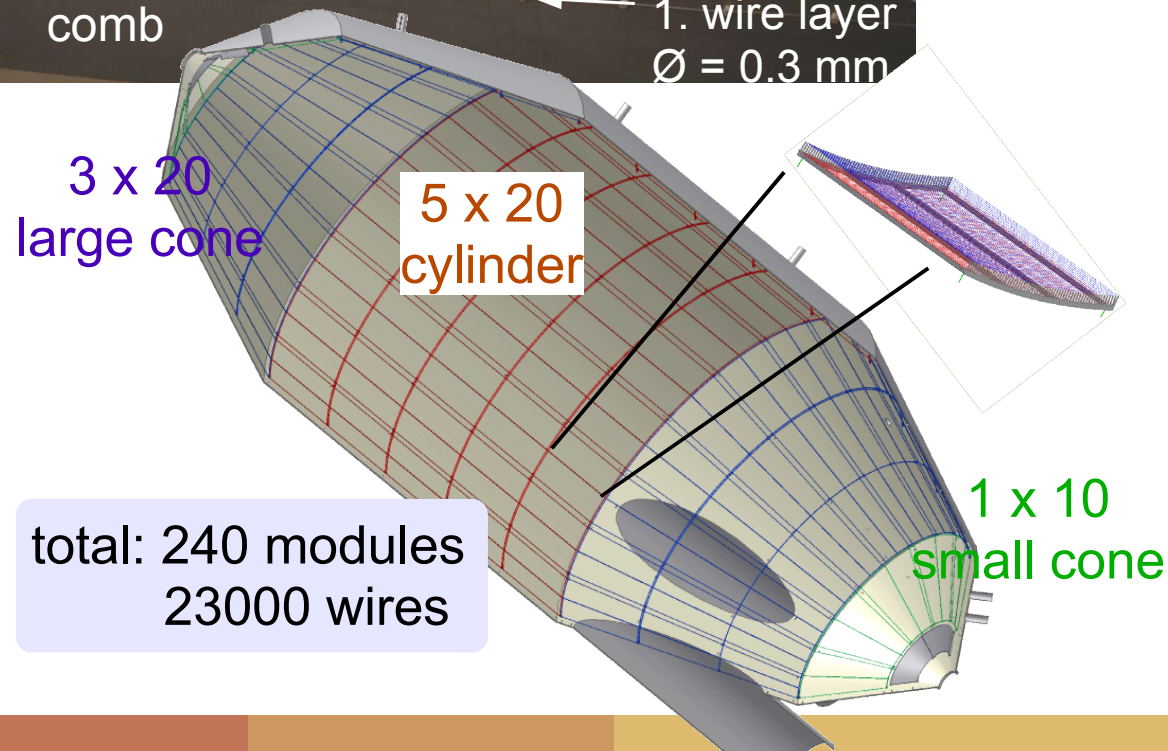
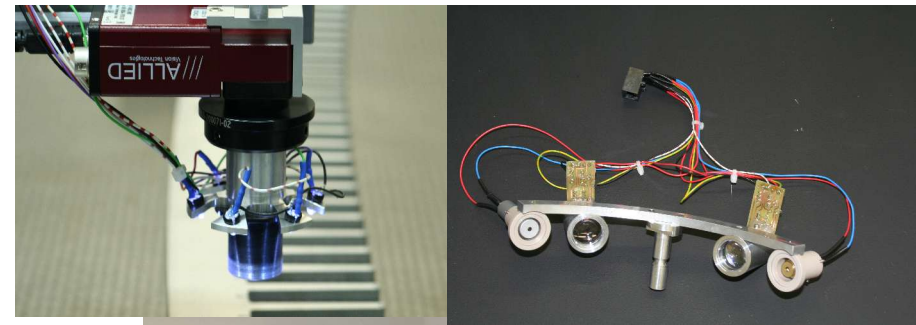
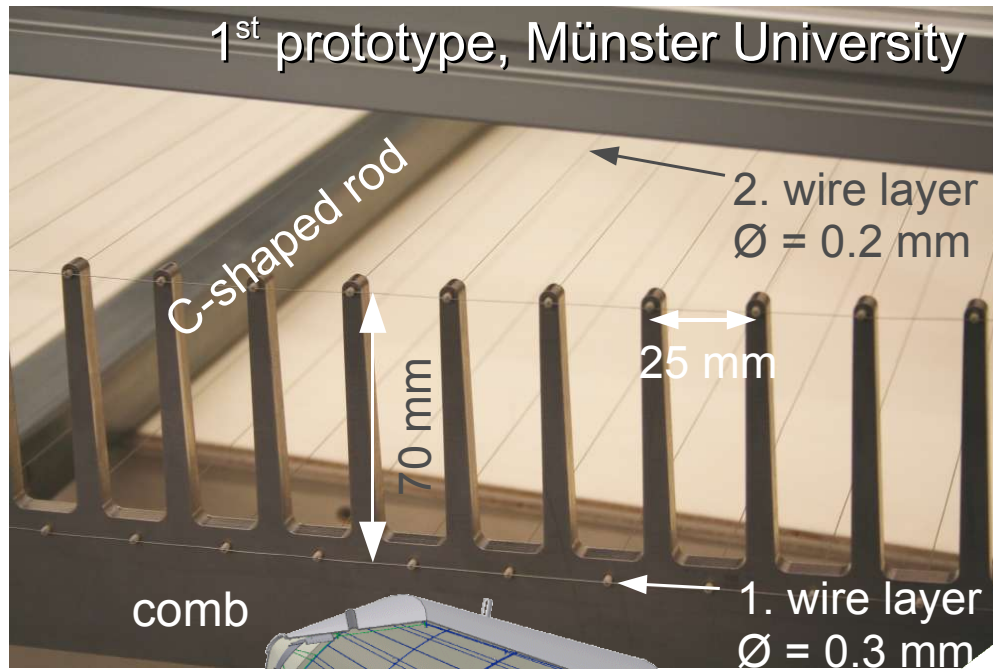


Dipl. thesis B. Ostrick (U Mainz, 2002),
PhD thesis B. Flatt (U Mainz, 2004)

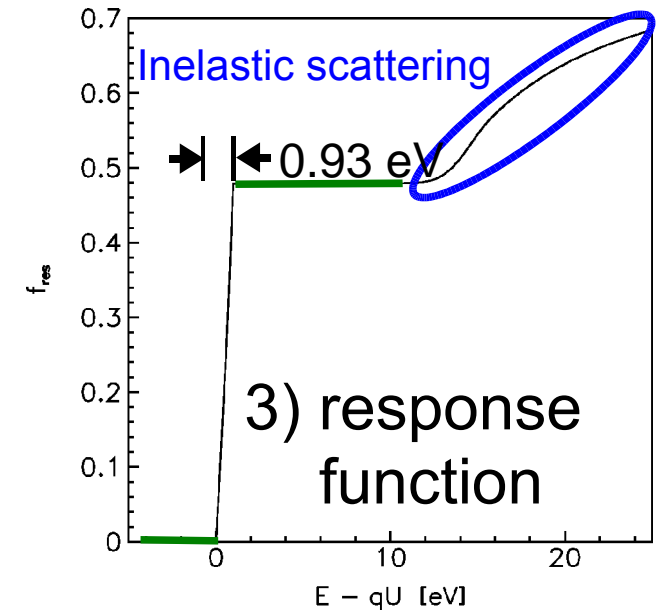
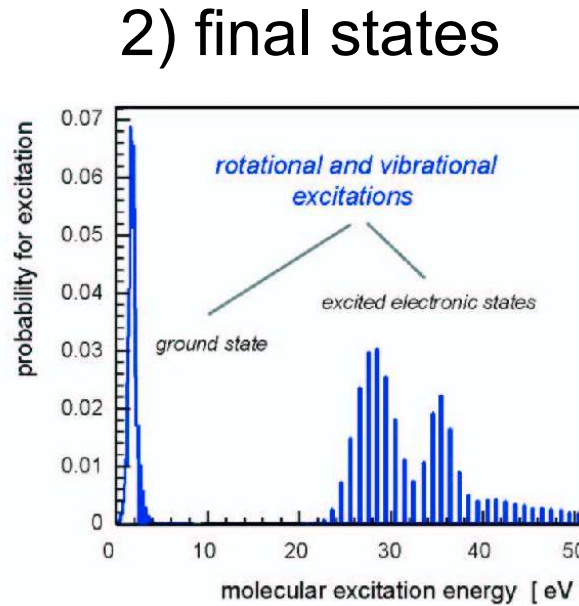
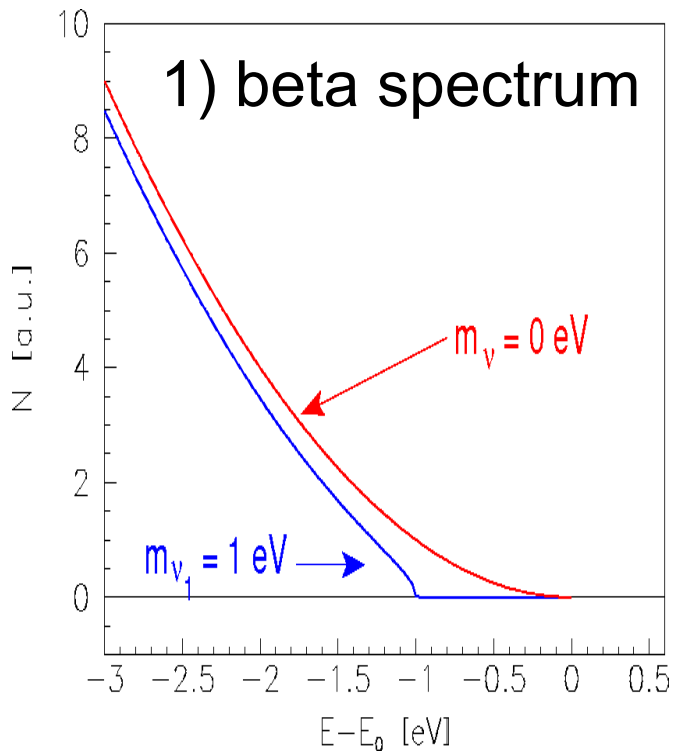


KATRIN pre spectrometer

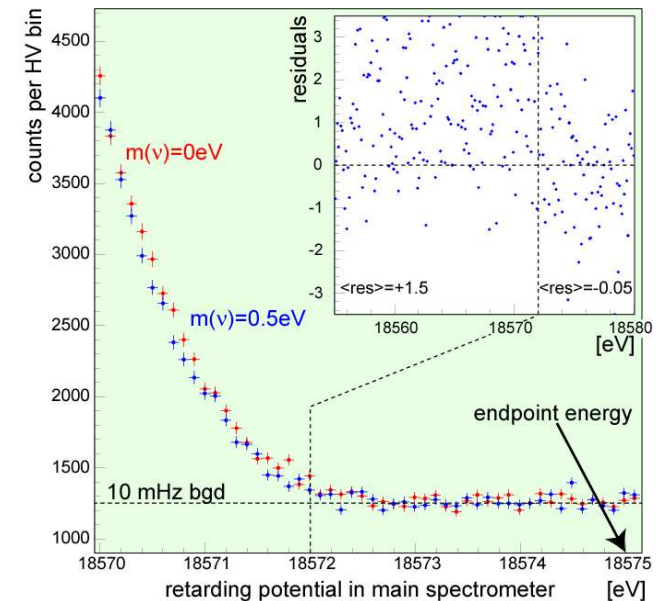
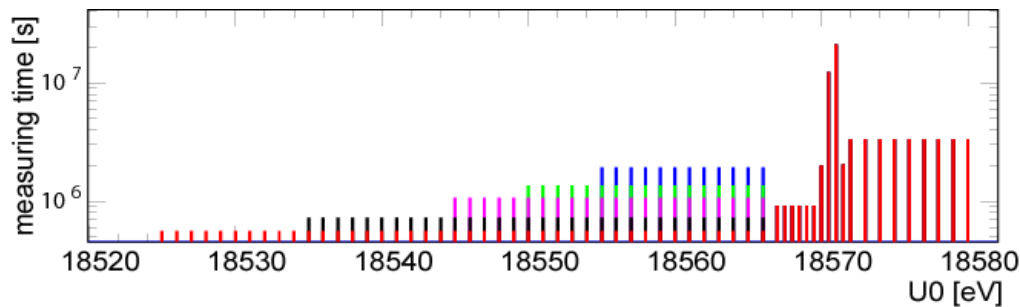
KATRIN: ≈ 240 double layer wire electrode modules

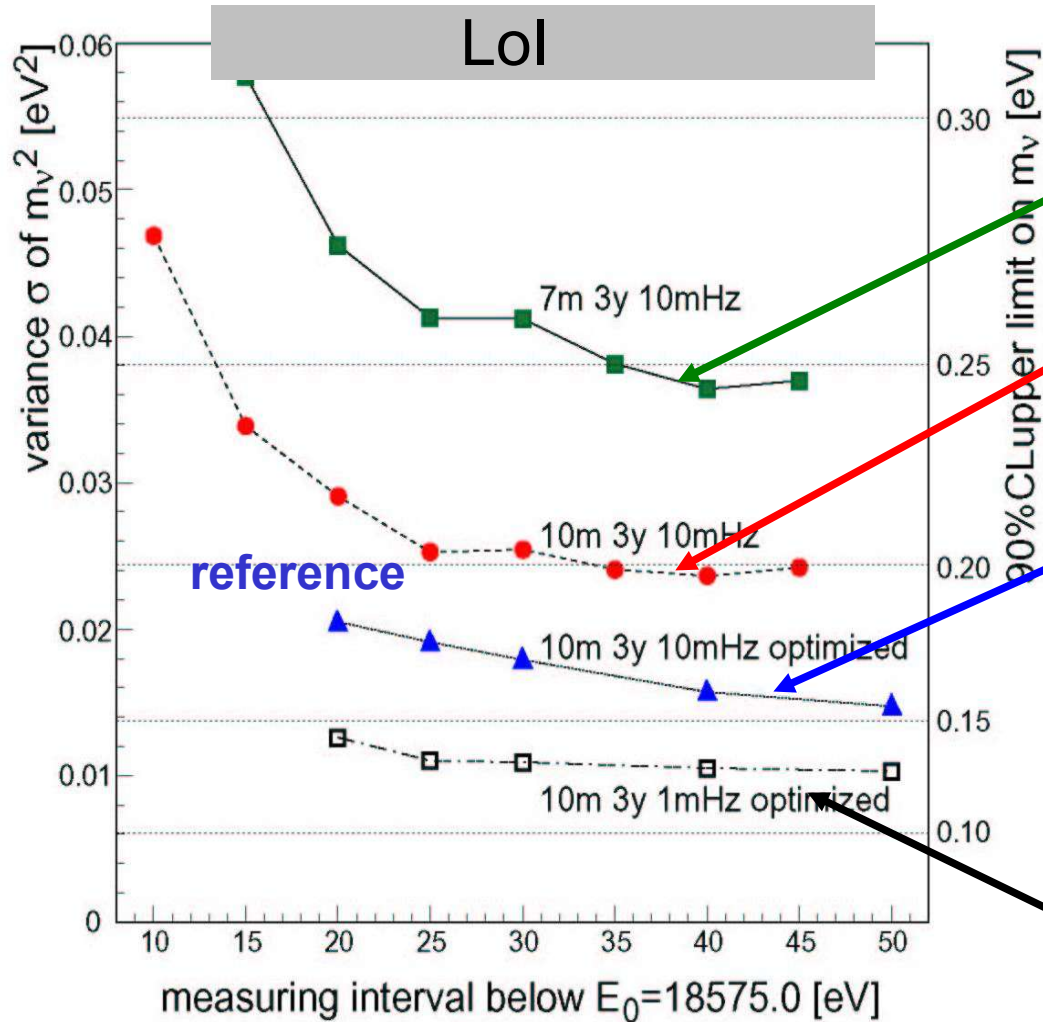


2) Statistics



4) measurement point distrib. \Rightarrow MC data





design optimisation '01 - '03

- tritium purity by tritium laboratory (>95%)

- 2× stronger gaseous source
($\varnothing=75\text{mm} \rightarrow \varnothing=90\text{mm}$)
requires $\varnothing=10\text{m}$ spectrometer)

- optimised measuring point
distribution (~ 5 eV below E_0)

- active background reduction by
inner electrode system, low
background detector
(needs further detailed tests)

3) Systematic uncertainties

As smaller $m(\nu)$

as smaller the region of interest below endpoint E_0

⇒ Excited electronic final states does not play a role ($\Delta E_{\text{exc}} > 27 \text{ eV}$)

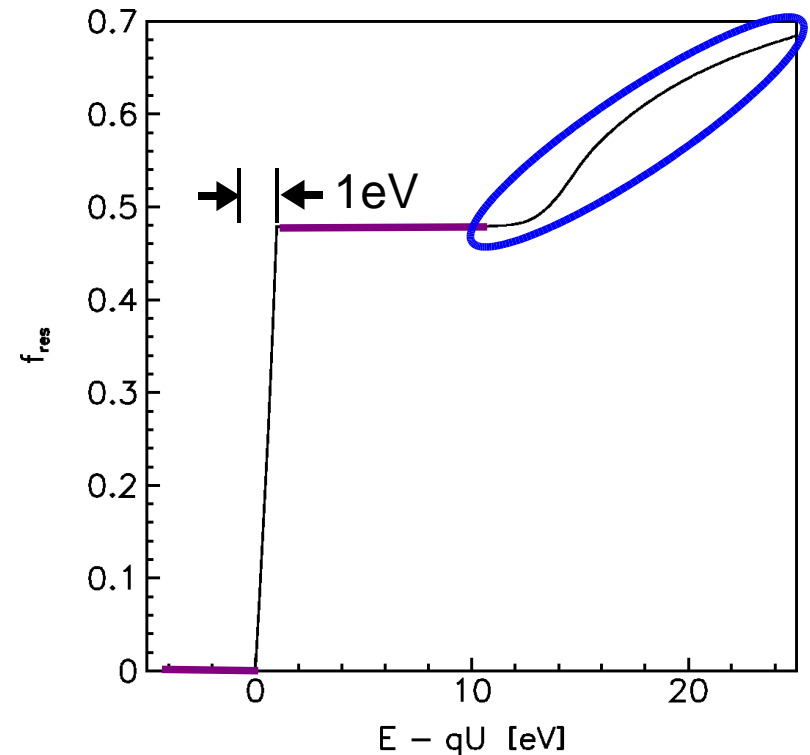
⇒ Inelastic scattering in T_2 is small

($\Delta E_{\text{inel.}} > 12 \text{ eV}$)

⇒ largest interval 25eV: 2%

⇒ One well-defined final state (similar to cryo detectors)

Is only true, since MAC-E-Filter response function has no tails



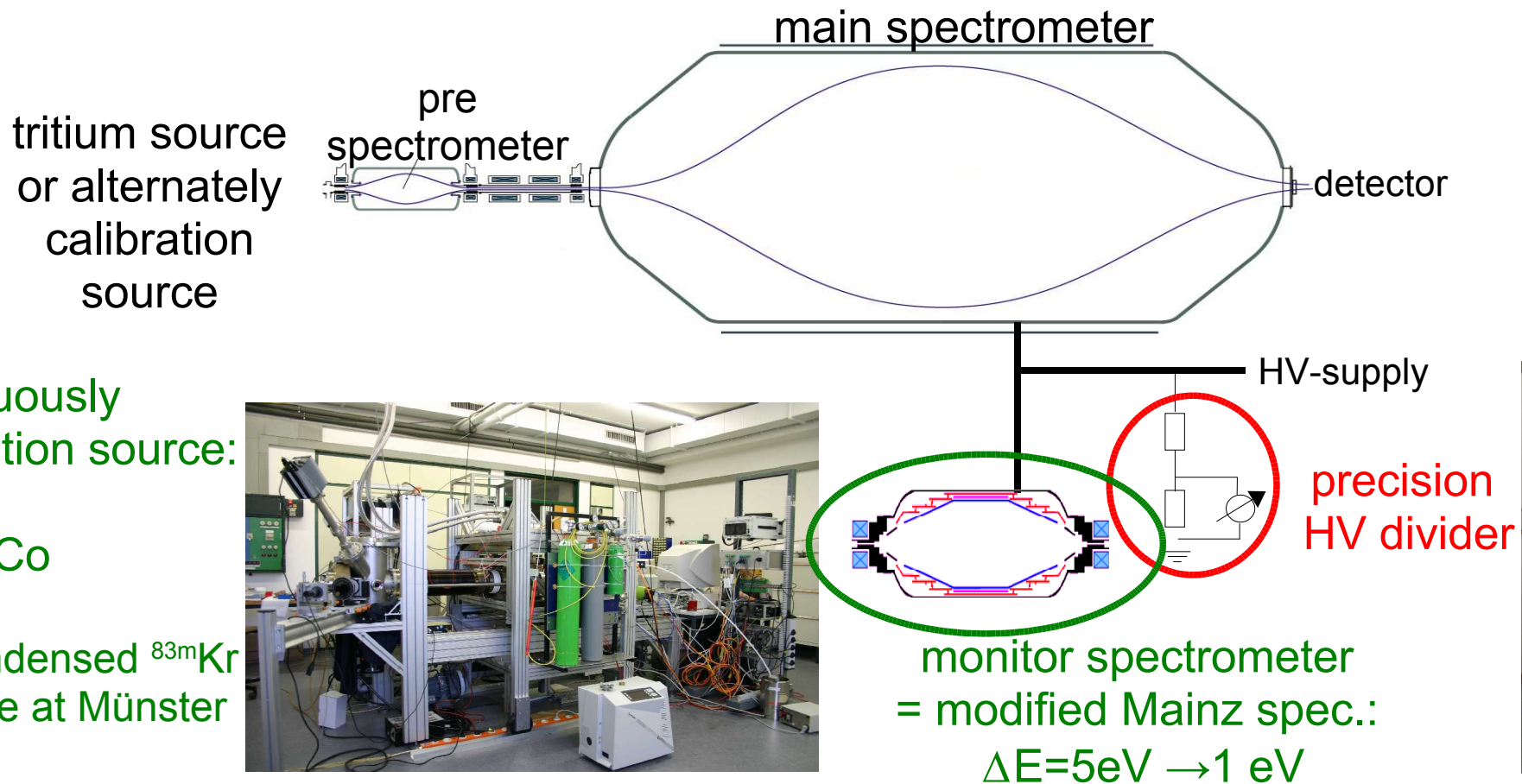
3) Systematic uncertainties

any not accounted variance σ^2 leads to negative shift of m_ν^2 : $\Delta m_\nu^2 = -2 \sigma^2$

1. inelastic scatterings of β 's inside WGTS
 - requires dedicated e-gun measurements, unfolding techniques for response fct.
2. fluctuations of WGTS column density (required $< 0.1\%$)
 - rear detector, Laser-Raman spectroscopy, T=30K stabilisation, e-gun measurements
3. transmission function
 - spatial resolved e-gun measurements
4. WGTS charging due to remaining ions (MC: $\phi < 20\text{mV}$)
 - inject low energy meV electrons from rear side, diagnostic tools available
5. final state distribution
 - reliable quantum chem. calculations
6. HV stability of retarding potential on $\sim 3\text{ppm}$ level required
 - precision HV divider (PTB), monitor spectrometer beamline

a few
contributions
with each:
 $\Delta m_\nu^2 \leq 0.007 \text{ eV}^2$

- Measure HV by precision HV divider
- Lock retarding HV by measuring energetically well-defined (atomic/nuclear standard) and sharp electron line with monitor spectrometer



continuously
calibration source:
 $^{83\text{m}}\text{Kr}$,
 $^{241}\text{Am/Co}$

condensed $^{83\text{m}}\text{Kr}$
source at Münster

KATRIN's sensitivity:

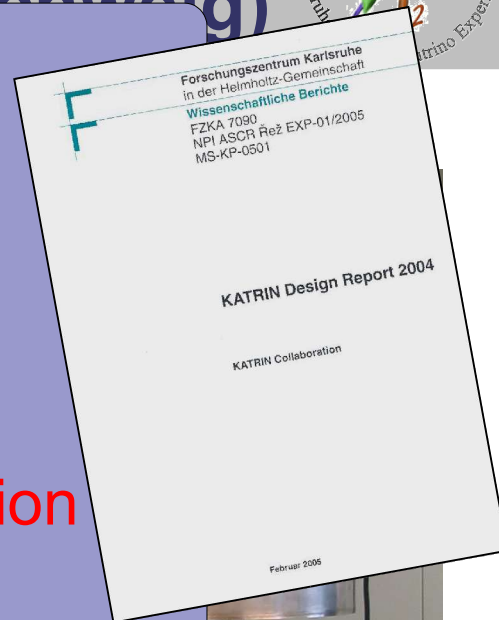
- higher T2 purity
- larger statistics
- optimized measurement point distribution
- smaller systematic uncertainties

⇒ sensitivity on $m(\nu_e)$
 $\approx 0.20 \text{ eV}/c^2$

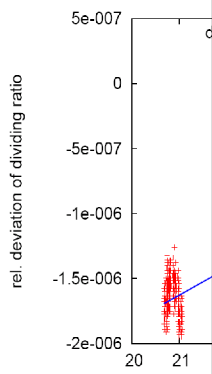
(about equal contribution from stat. and syst. uncertainties)
 (90% C.L. upper limit for $m(\nu_e) = 0$)

$m(\nu_e) = 0.30 \text{ eV}$ observable with 3σ

$m(\nu_e) = 0.35 \text{ eV}$ observable with 5σ



tempera



⇒ TC

Measurements at PTB Braunschweig

Absolute neutrino mass scale is needed

for particle physics and astrophysics/cosmology

by direct neutrino mass measurement (less model dependent & complementary)

Direct ν mass measurement from tritium β decay:

- Mainz experiment finished (all problems solved): $m(\nu_e) < 2.3 \text{ eV}$ (95% C.L.)

- KATRIN: A large tritium β neutrino mass experiment with sub-eV sensitivity

$m(\nu_e) < 0.2 \text{ eV}$ or $m(\nu_e) > 0 \text{ eV}$: for $m(\nu_e) \geq 0.30 \text{ eV}$ @ 3σ

start of data taking in 2010 ...



Direct ν mass measurement from rhenium β decay:

- Mibeta, Manu2 (1st generation of exp.): $m(\nu_e) < 15.6 \text{ eV}$ (90% C.L.)

- MARE: Proposal for a large array of improved rhenium micro calorimeters,
how far can they go?