



Status of Daya Bay

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For

The Daya Bay Collaboration

LAUNCH Workshop at MPI, Heidelberg, Germany, 23 March, 2007

Significance of θ_{13}

U_{MNSP} Matrix

Maki, Nakagawa, Sakata, Pontecorvo

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} = \begin{pmatrix} 0.8 & 0.5 & U_{e3} \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix} \quad ?$$

$$= \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}}_{\text{atmospheric, accelerator}} \times \underbrace{\begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix}}_{\text{reactor, accelerator}} \times \underbrace{\begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{SNO, solar SK, KamLAND}} \times \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha/2} & 0 \\ 0 & 0 & e^{i\alpha/2+i\beta} \end{pmatrix}}_{\text{O}\nu\beta\beta}$$

atmospheric,
accelerator

$$\theta_{23} \approx 45^\circ$$

$$\Delta m_{32}^2 \approx 2.4 \times 10^{-3} \text{ eV}^2$$

reactor,
accelerator

$$\theta_{13} = ?$$

SNO, solar SK,
KamLAND

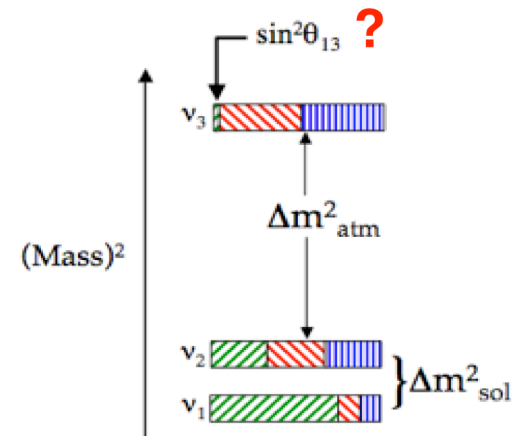
$$\theta_{12} \approx 32^\circ$$

$$\Delta m_{21}^2 \approx 7.9 \times 10^{-5} \text{ eV}^2$$

$$\Delta m_{31}^2 \approx 2.5 \times 10^{-3} \text{ eV}^2$$

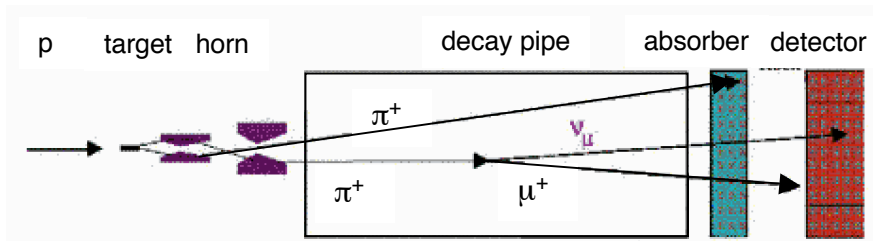
- What is ν_e fraction of ν_3 ?
- U_{e3} is the gateway to CP violation in neutrino sector:

$$P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \propto \sin(2\theta_{12})\sin(2\theta_{23})\cos^2(\theta_{13})\sin(2\theta_{13})\sin\delta$$



Some Methods For Determining θ_{13}

Method 1: Accelerator Experiments



$$P_{\mu e} \approx \sin^2 2\theta_{13} \sin^2 2\theta_{23} \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E_\nu} \right) + \dots$$

- $\nu_\mu \rightarrow \nu_e$ appearance experiment
- need other mixing parameters to extract θ_{13}
- baseline $O(100-1000 \text{ km})$, matter effects present
- expensive

Method 2: Reactor Experiments



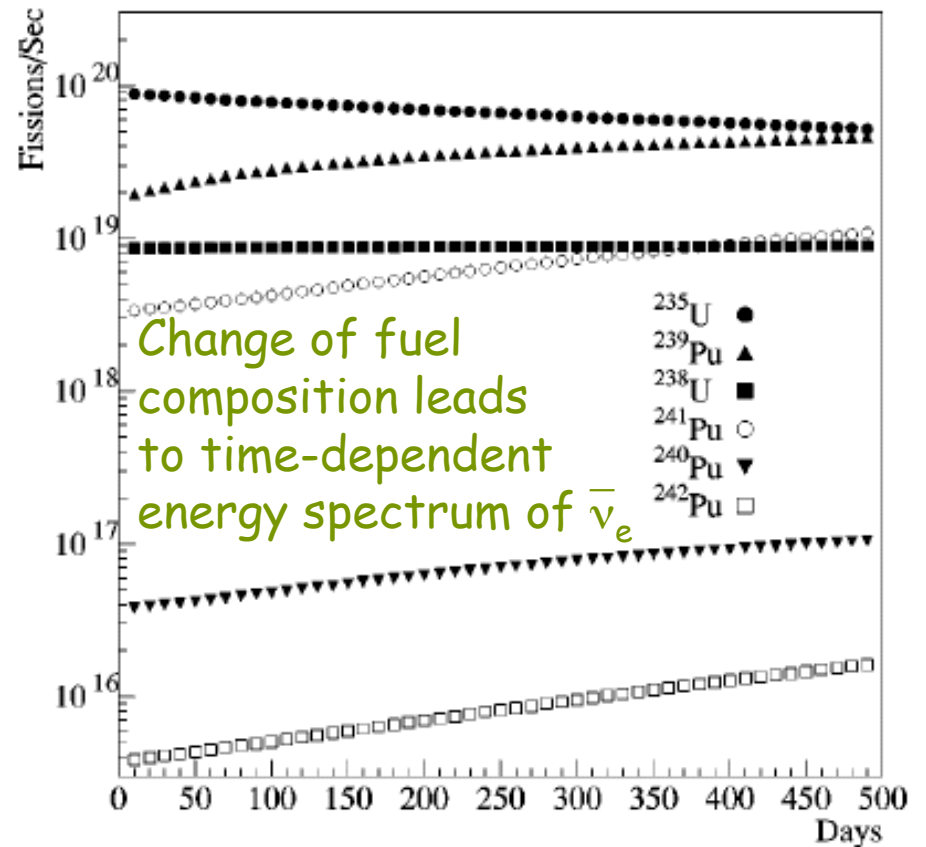
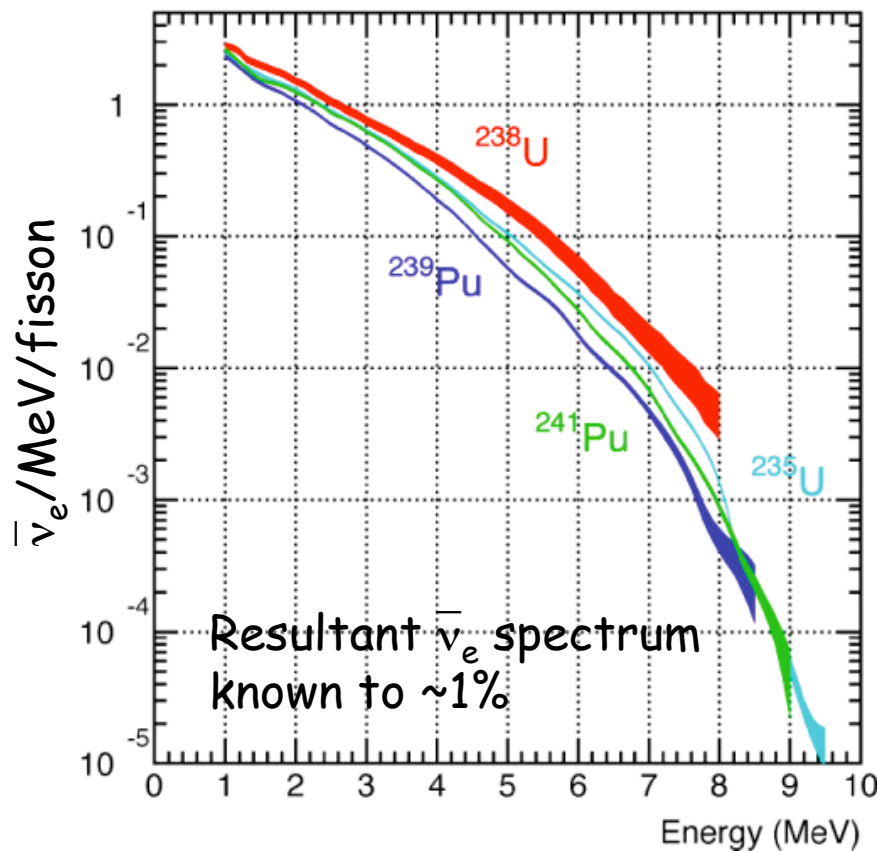
$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E_\nu} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E_\nu} \right)$$

- $\bar{\nu}_e \rightarrow X$ disappearance experiment
- baseline $O(1 \text{ km})$, no matter effect, no ambiguity
- relatively cheap

Reactor $\bar{\nu}_e$

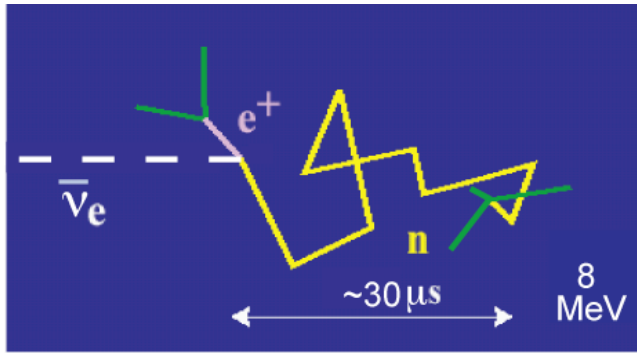
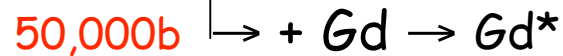
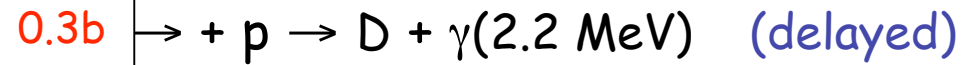
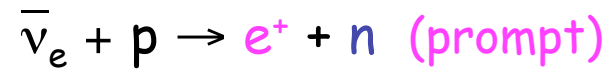
- Fission processes in nuclear reactors produce huge number of low-energy $\bar{\nu}_e$:

3 GW_{th} generates 6×10^{20} $\bar{\nu}_e$ per sec



Detecting $\bar{\nu}$ With Liquid Scintillator

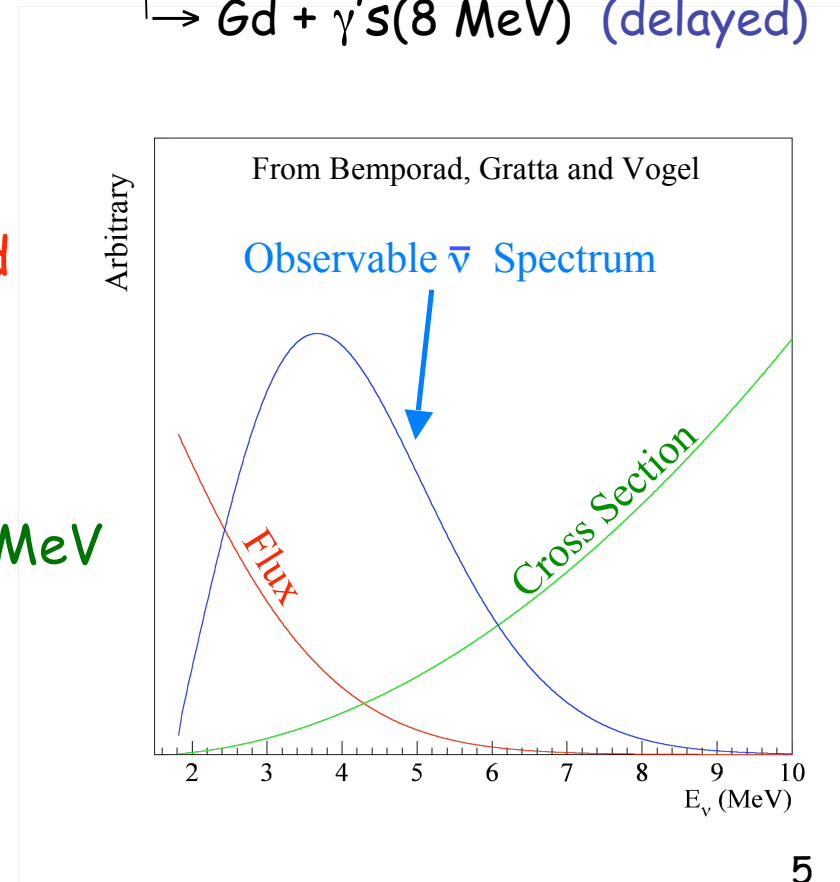
- Use the inverse β -decay reaction in 0.1% Gd-doped liquid scintillator:



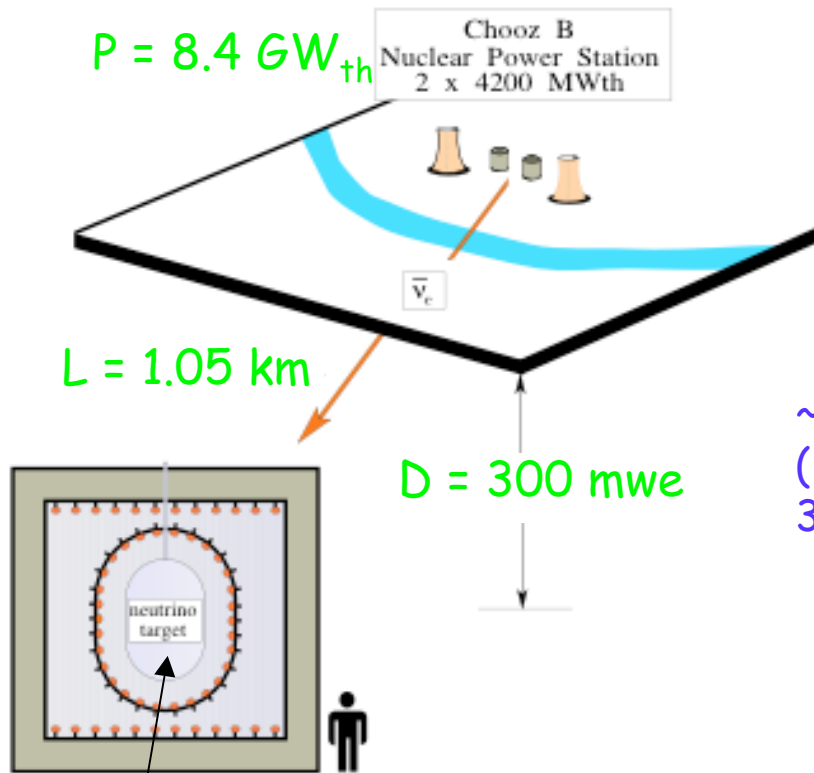
- Time- and energy-tagged signal is a good tool to suppress background events.
- Energy of $\bar{\nu}_e$ is given by:

$$E_{\bar{\nu}} \approx T_{e^+} + T_n + (m_n - m_p) + m_{e^+} \approx T_{e^+} + 1.8 \text{ MeV}$$

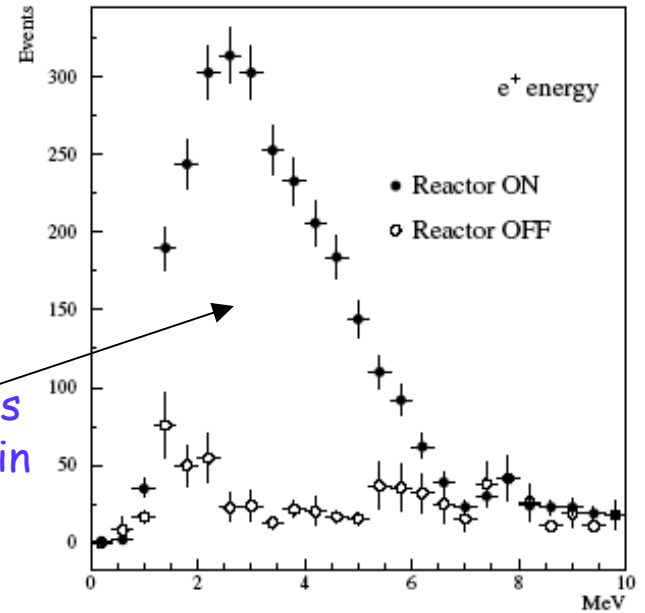
10-40 keV



Chooz: Finding θ_{13}



$\sim 3000 \bar{\nu}_e$ candidates (included 10% bkg) in 335 days



5-ton 0.1% Gd-loaded liquid scintillator to detect $\bar{\nu}_e + p \rightarrow e^+ + n$

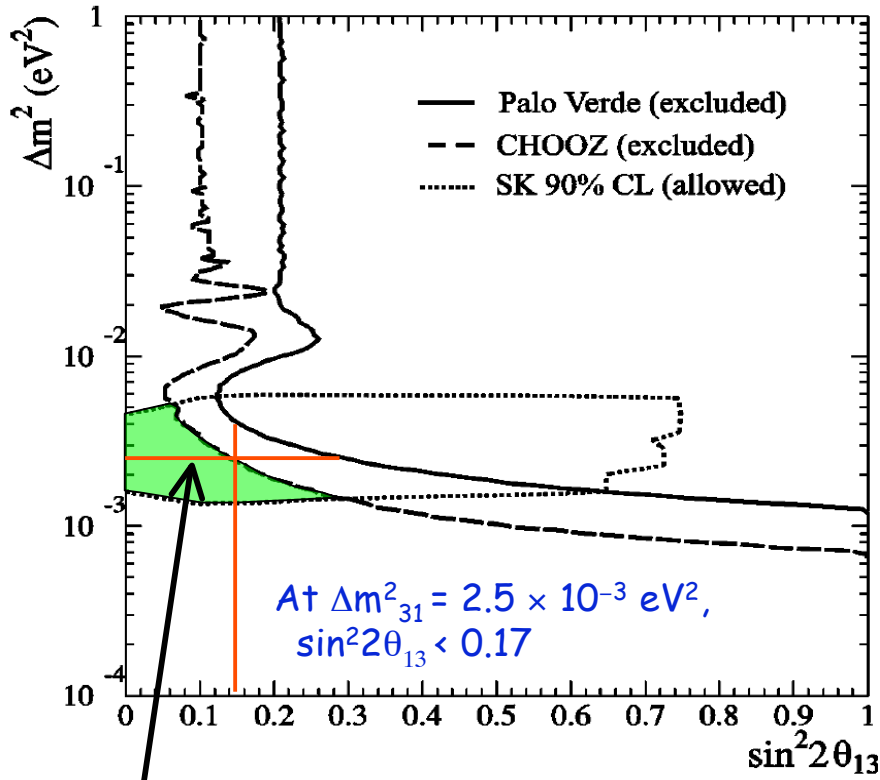
Rate:
 ~ 5 evts/day/ton (full power)
 including 0.2-0.4 bkg/day/ton

Systematic uncertainties

parameter	relative uncertainty (%)
reaction cross section	1.9
number of protons	0.8
detection efficiency	1.5
reactor power	0.7
energy released per fission	0.6
combined	2.7

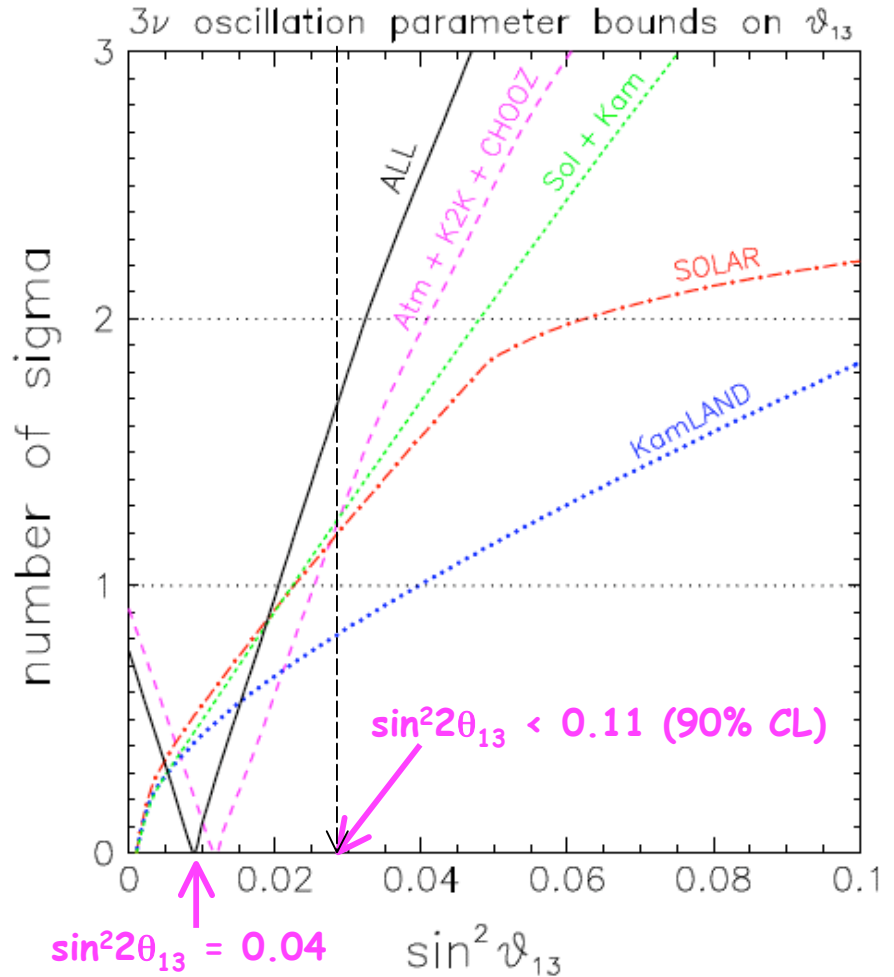
Current Knowledge of θ_{13}

Direct search



allowed region

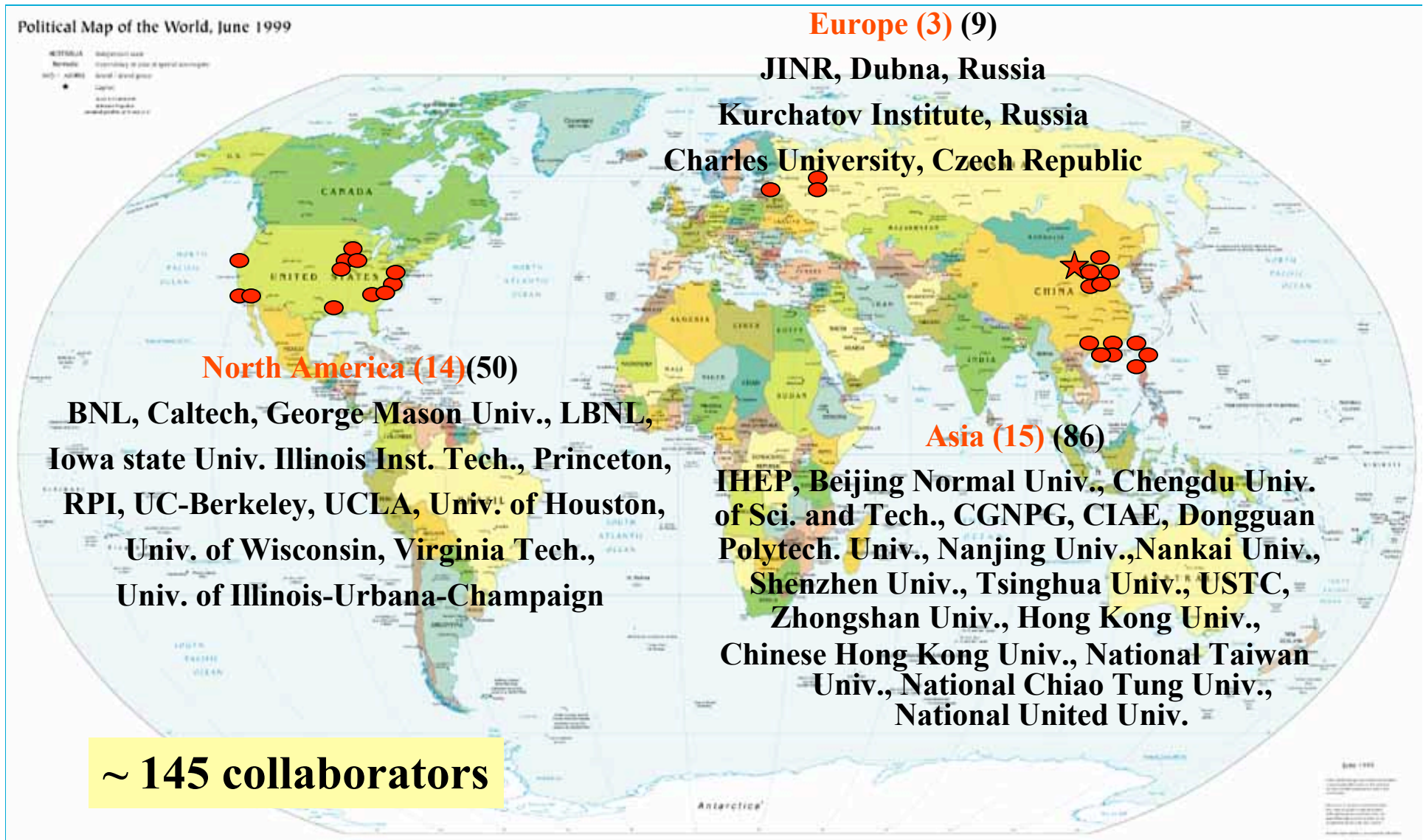
Global fit



Best fit value of $\Delta m^2_{32} = 2.4 \times 10^{-3} eV^2$

Fogli et al., hep-ph/0506083

Daya Bay Collaboration

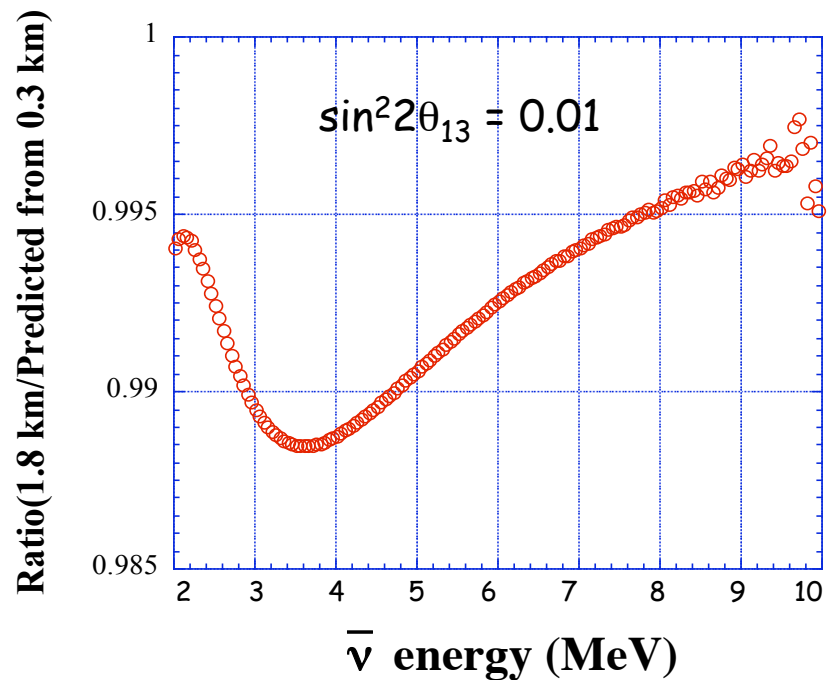




12-14 Jan, 2007 at HKU

Daya Bay: Goal And Approach

- Determine $\sin^2 2\theta_{13}$ with a sensitivity of ≤ 0.01 by measuring deficit in $\bar{\nu}_e$ rate and spectral distortion.



- Recommendation of the APS Neutrino Study Group:
 - *An expeditiously deployed multidetector reactor experiment with sensitivity to $\bar{\nu}_e$ disappearance down to $\sin^2 2\theta_{13} = 0.01$, an order of magnitude below present limits.*

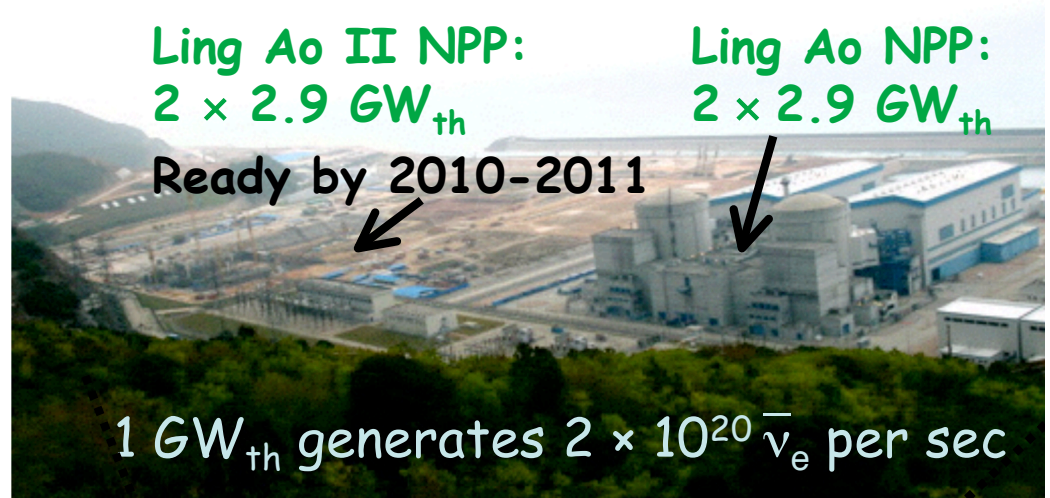
How To Reach A Precision of 0.01 in Daya Bay?

- **Increase statistics:**
 - Use more powerful nuclear reactors
 - Utilize larger target mass, hence larger detectors
- **Suppress background:**
 - Go deeper underground to gain overburden for reducing cosmogenic background
 - Use active shield around the target
- **Reduce systematic uncertainties:**
 - **Reactor-related:**
 - Optimize baseline for best sensitivity and smaller residual reactor-related errors
 - Near and far detectors to minimize reactor-related errors
 - **Detector-related:**
 - Use "Identical" pairs of detectors to do *relative* measurement
 - Comprehensive program in calibration/monitoring of detectors
 - Interchange near and far detectors (optional)



The Daya Bay Nuclear Power Complex

- 12th most powerful in the world ($11.6 \text{ GW}_{\text{th}}$)
- Fifth most powerful by 2011 ($17.4 \text{ GW}_{\text{th}}$)
- Adjacent to mountain, easy to construct tunnels to reach underground labs with sufficient overburden to suppress cosmic rays



Where To Place The Detectors ?

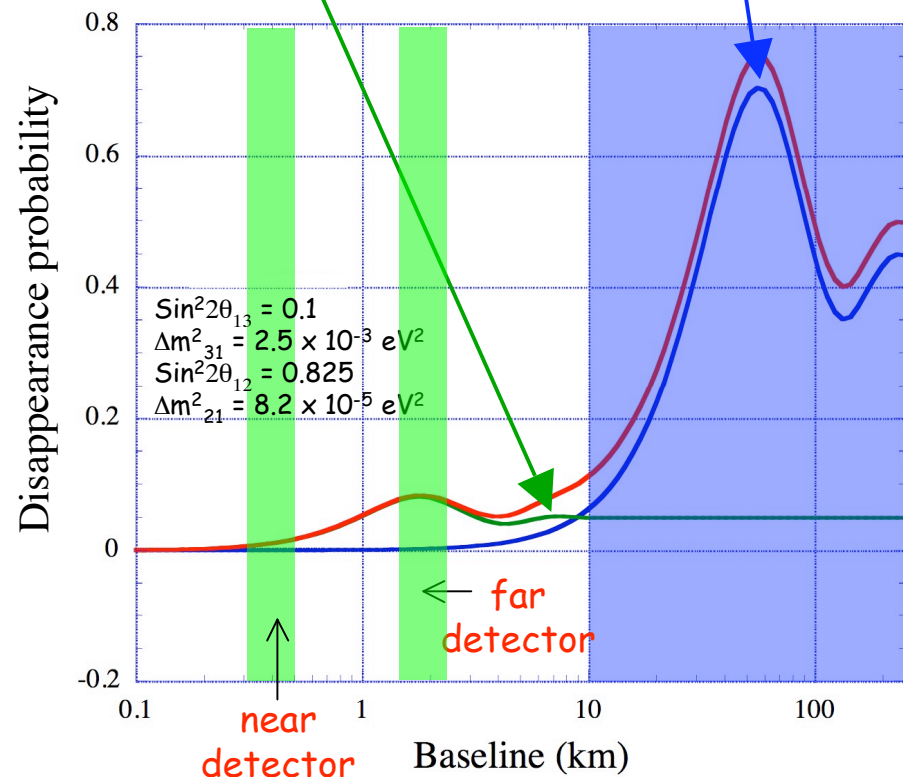
- Since reactor $\bar{\nu}_e$ are low-energy, it is a disappearance experiment:

$$P(\bar{\nu}_e \rightarrow x) \approx \sin^2 2\theta_{13} \sin^2\left(\frac{\Delta m_{31}^2 L}{4E}\right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2\left(\frac{\Delta m_{21}^2 L}{4E}\right)$$

- Place **near detector**(s) close to reactor(s) to measure flux and spectrum of $\bar{\nu}_e$ for normalization, hence reducing reactor-related systematic
- Position a **far detector** near the first oscillation maximum to get the highest sensitivity, and also be less affected by θ_{12}

Small-amplitude oscillation due to θ_{13} integrated over E

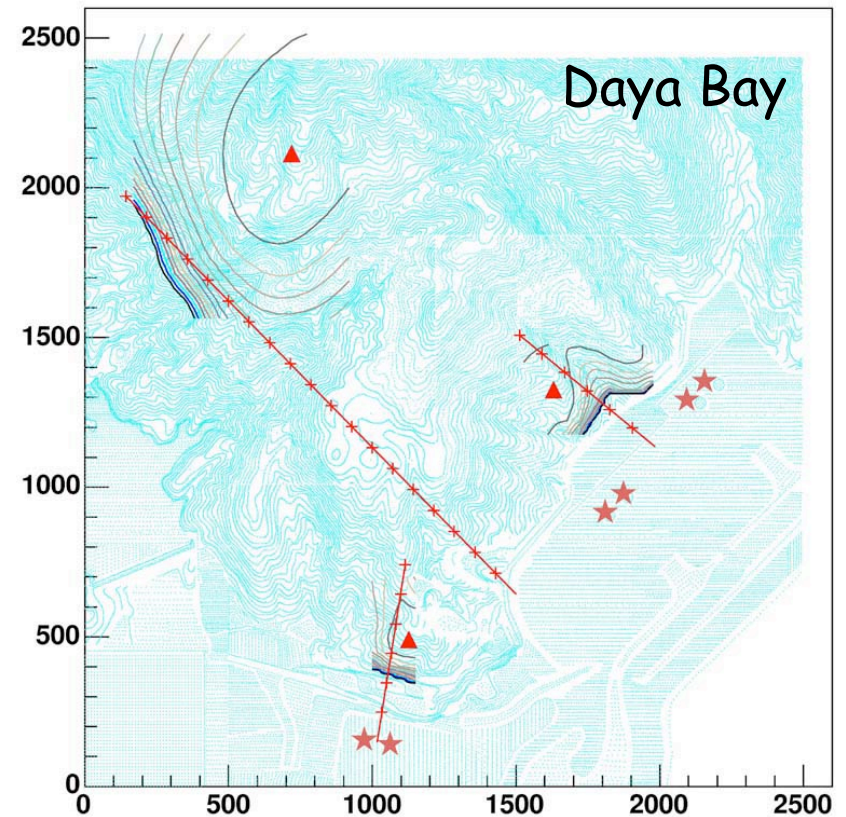
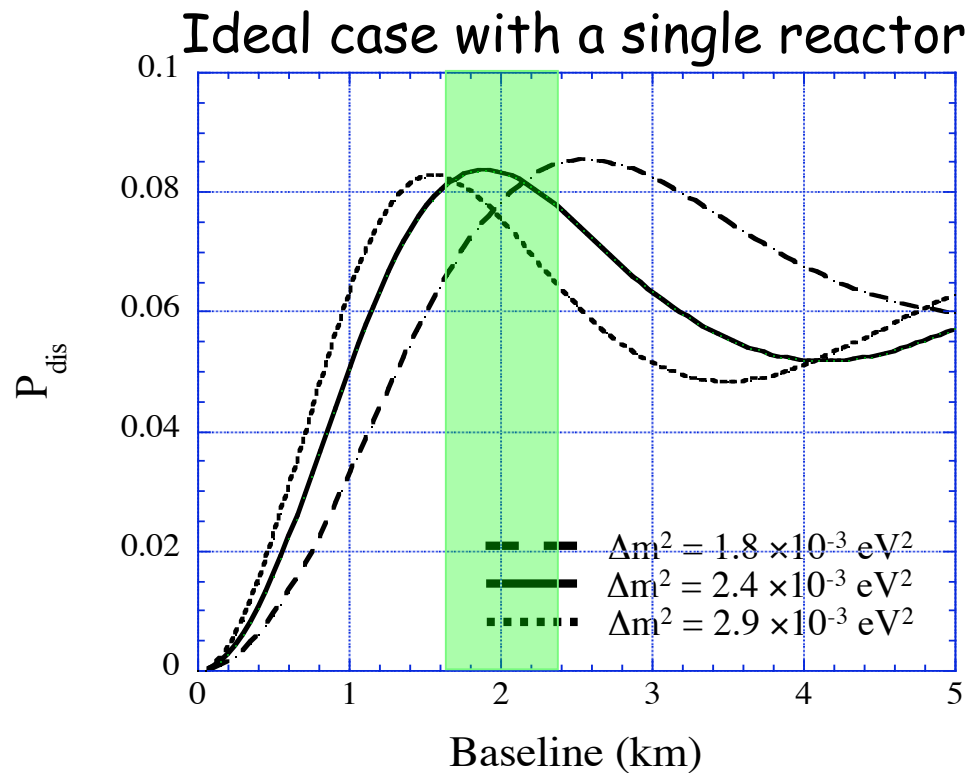
Large-amplitude oscillation due to θ_{12}

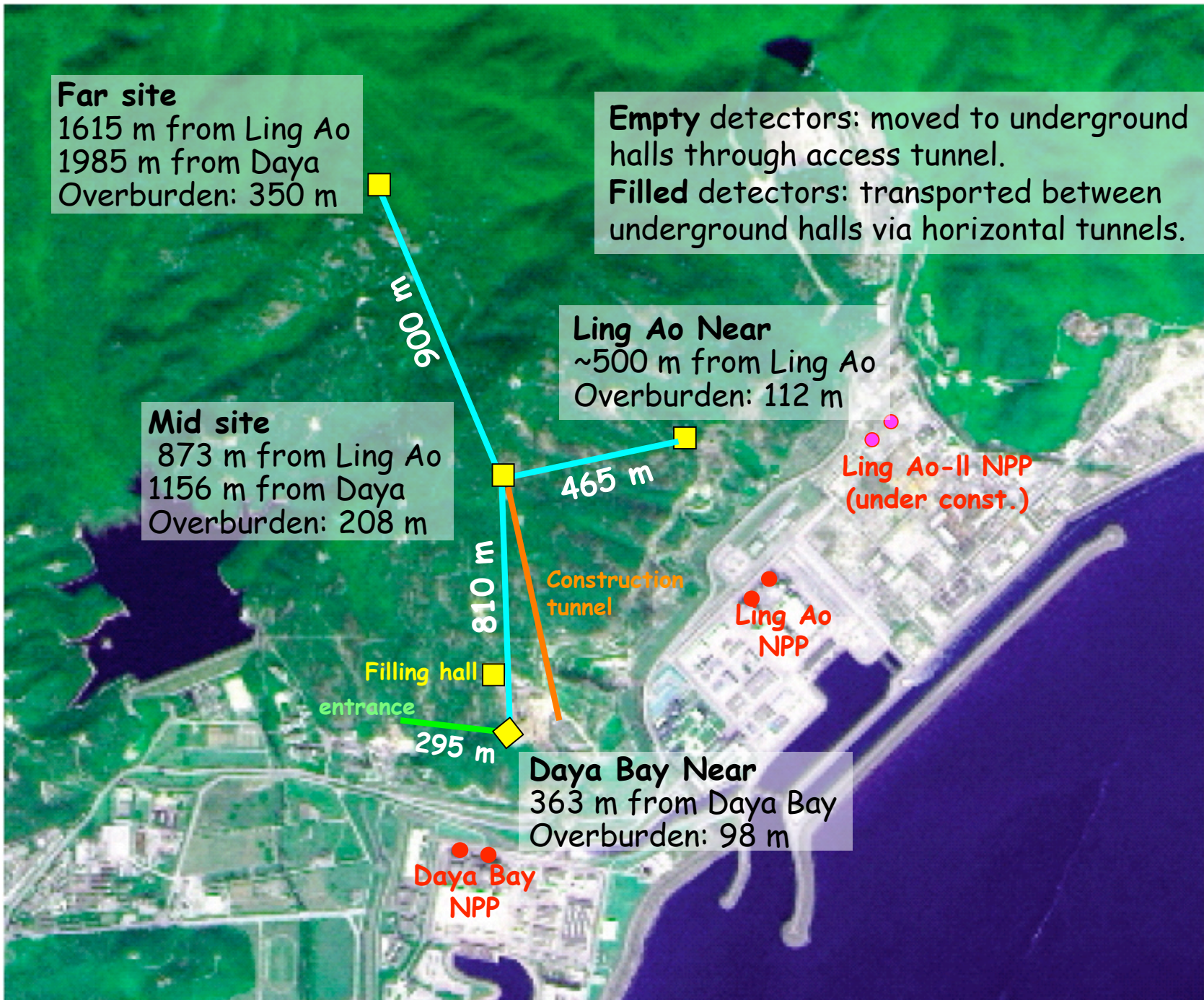


Baseline optimization and site selection

Inputs to the process:

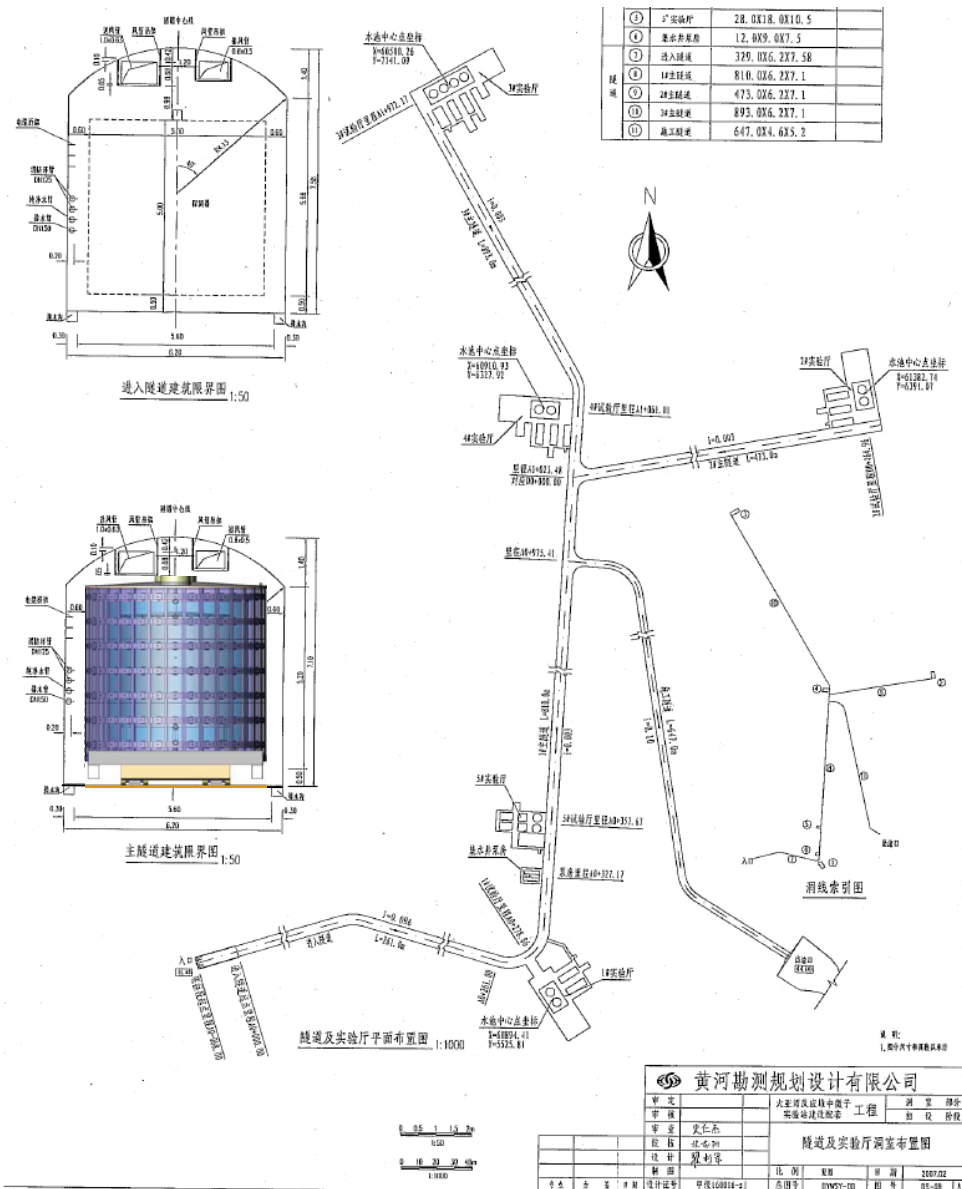
- Flux and energy spectrum of reactor antineutrino
- Systematic uncertainties of reactors and detectors
- Ambient background and uncertainties
- Position-dependent rates and spectra of cosmogenic neutrons and ${}^9\text{Li}$





Empty detectors: moved to underground halls through access tunnel.
Filled detectors: transported between underground halls via horizontal tunnels.

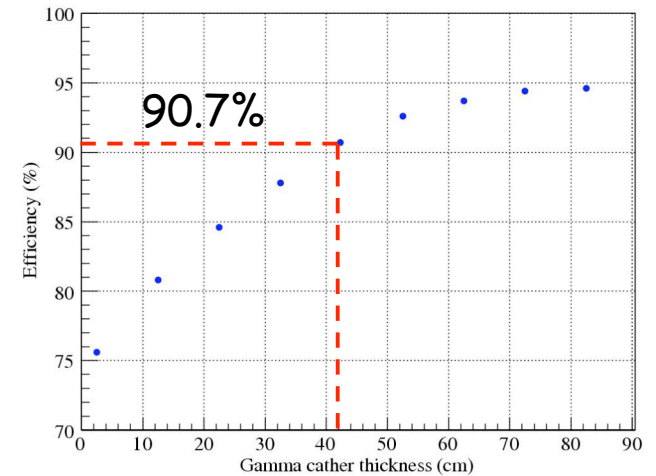
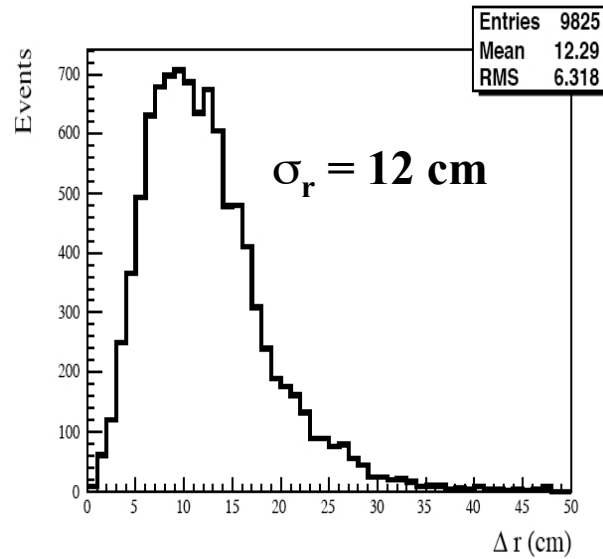
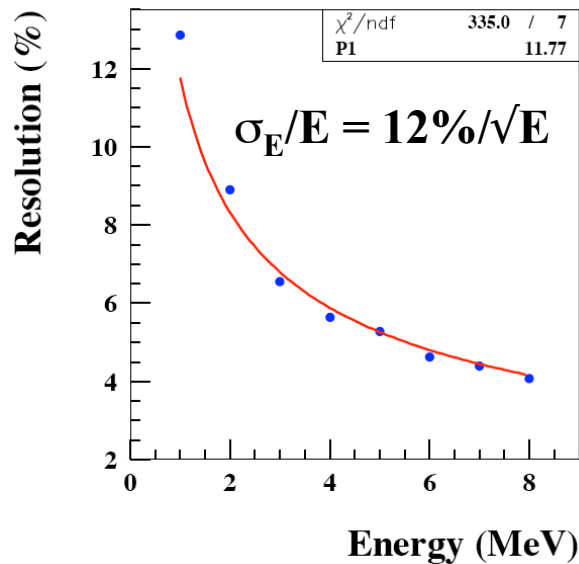
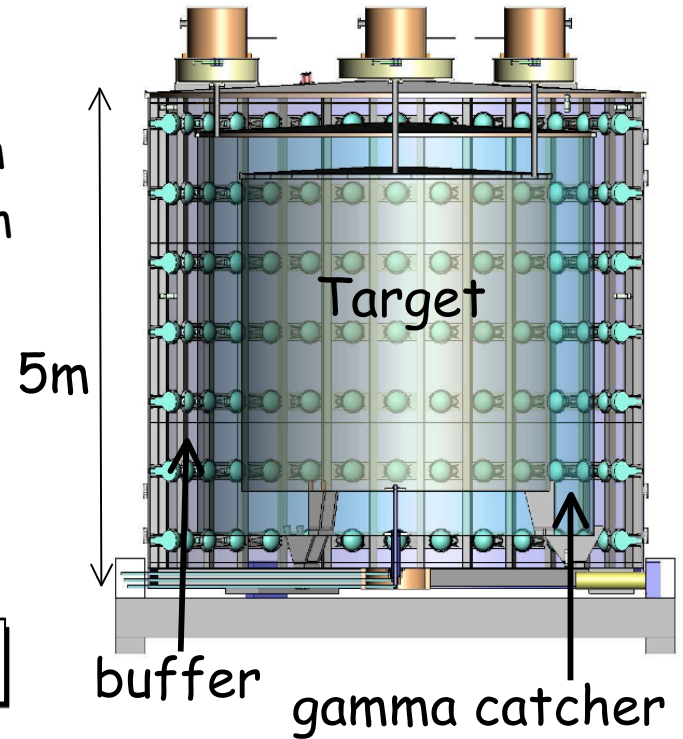
Civil Design Almost Complete



- Total tunnel length is about 3100 m
- Estimated total construction cost is ~\$11 M
- Construction will start this summer
- First experimental hall is available after 13 months
- Construction time is ~22 months

Antineutrino Detectors

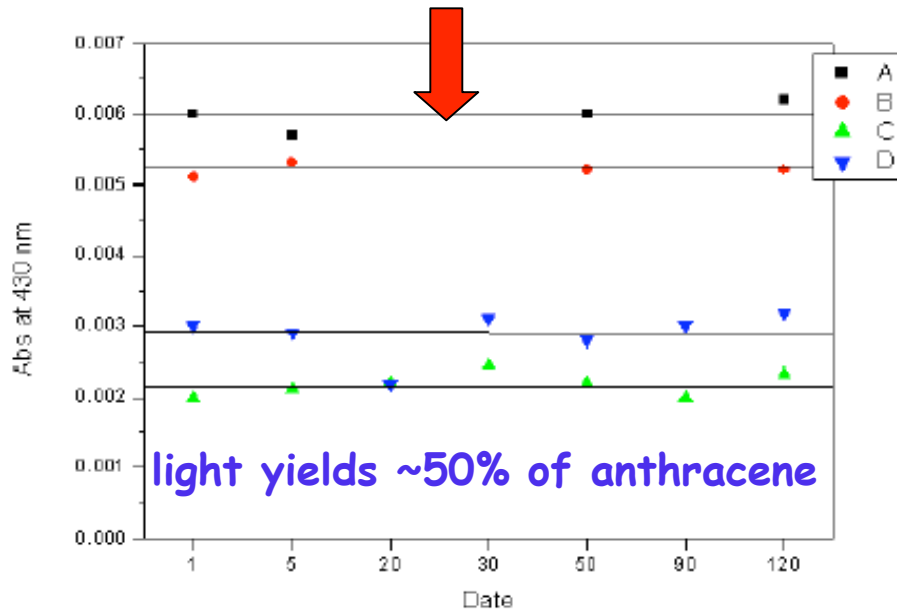
- Three-zone cylindrical detector design
 - Target: 20 T (0.1% Gd-LS), radius = 1.55 m
 - Gamma catcher: 20 T (LS), thickness = 0.42 m
 - Buffer : 40 T (mineral oil) , thickness = 0.48 m
- Low-background 8" PMT: 192
- Reflectors at top and bottom
- Photocathode coverage: 5.6 % \rightarrow 12% (with reflectors)



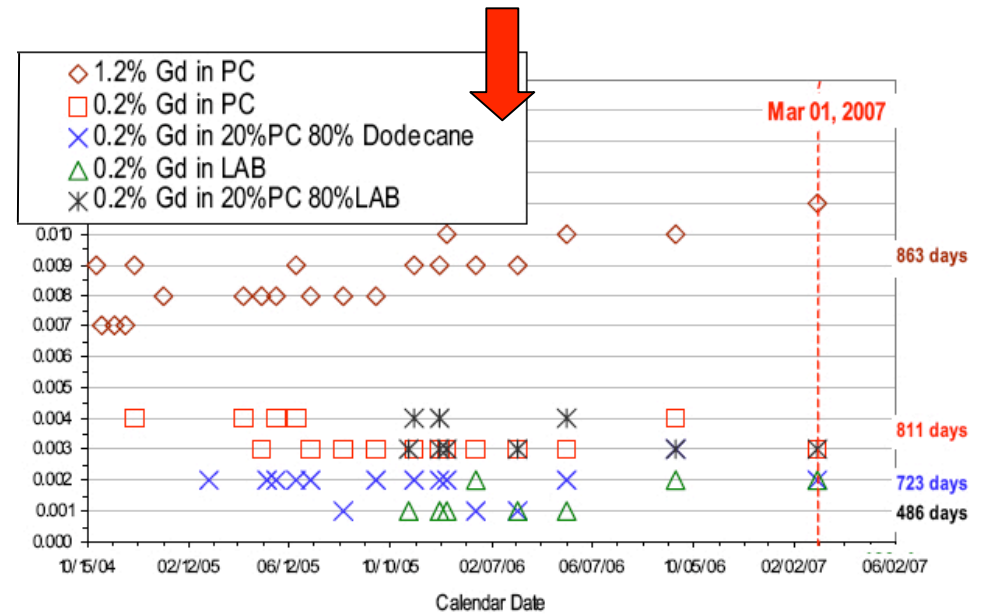
- Eight 'identical' detector modules

Gd-Liquid Scintillator

IHEP Gd-LS



BNL Gd-LS



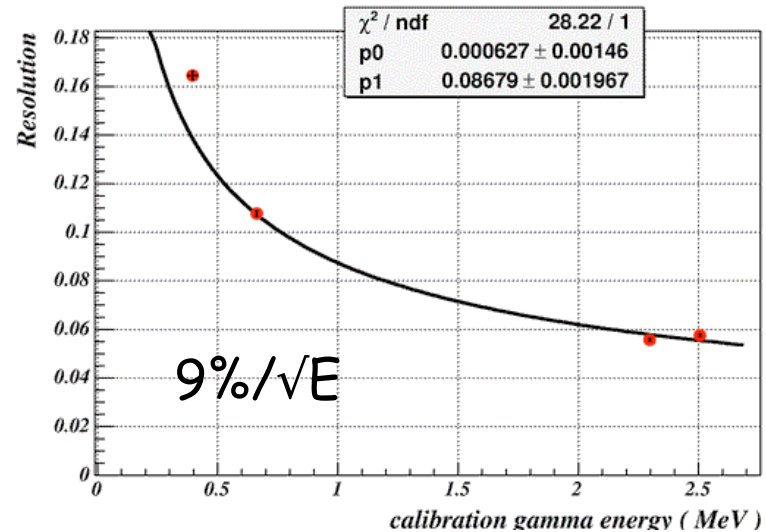
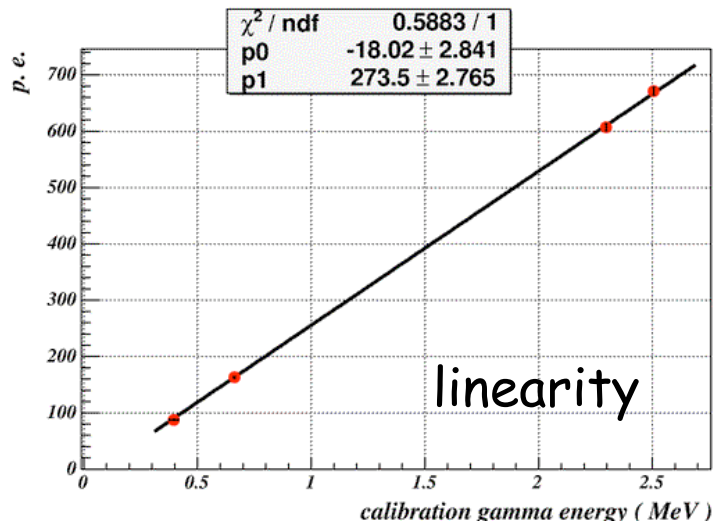
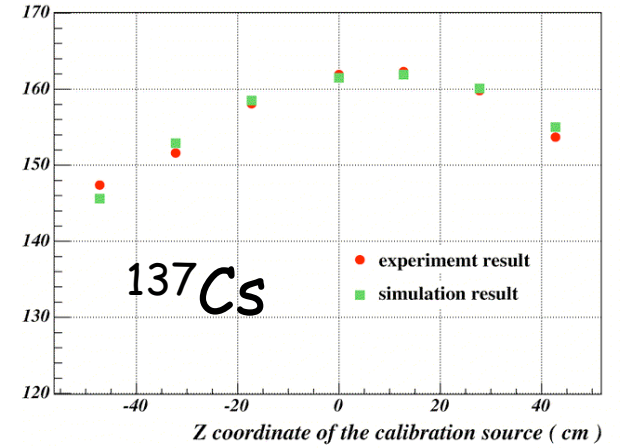
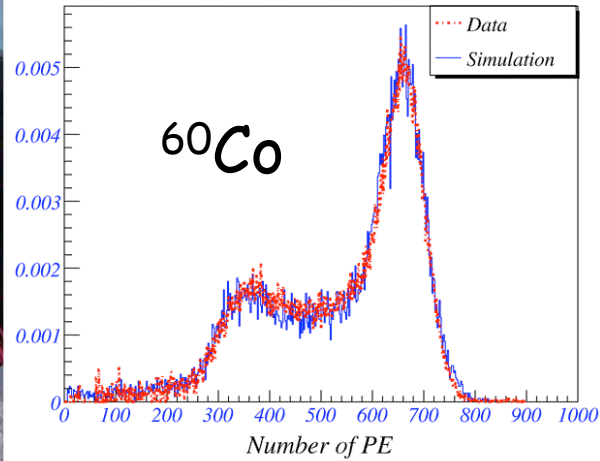
- A. Gd-isonanoate in 4: 6 **mesitylene**/dodecane
- B. Gd-ethylhexanoate in 2: 8 **mesitylene**/dodecane
- C. Gd-isonanoate in **LAB**
- D. Gd-ethylhexanoate in 2: 8 **mesitylene**/LAB

- Use LAB (Linear Alkyl Benzene):
 - good stability
 - high flash point
 - used in bio. sc.
 - high light yield
 - no EH&S issues
 - cheap (raw material for making detergent)



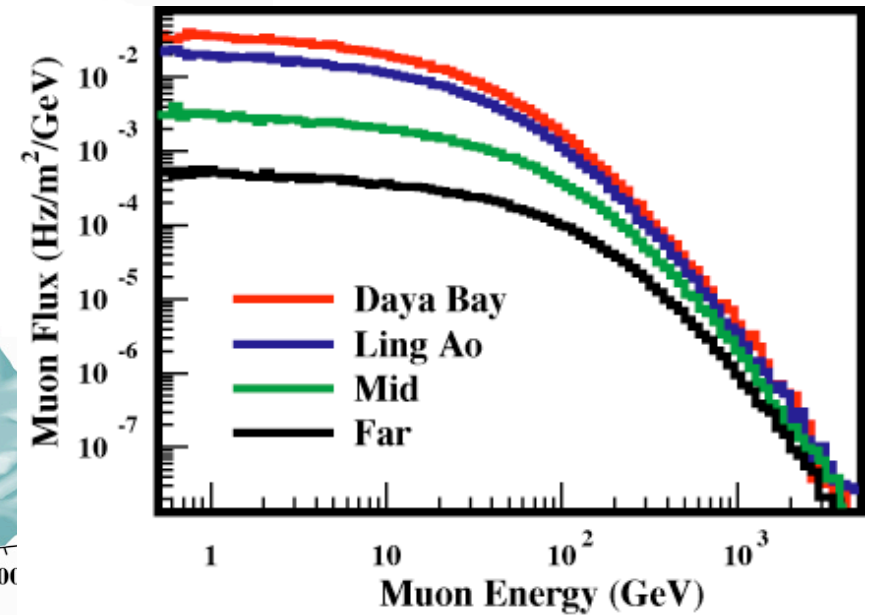
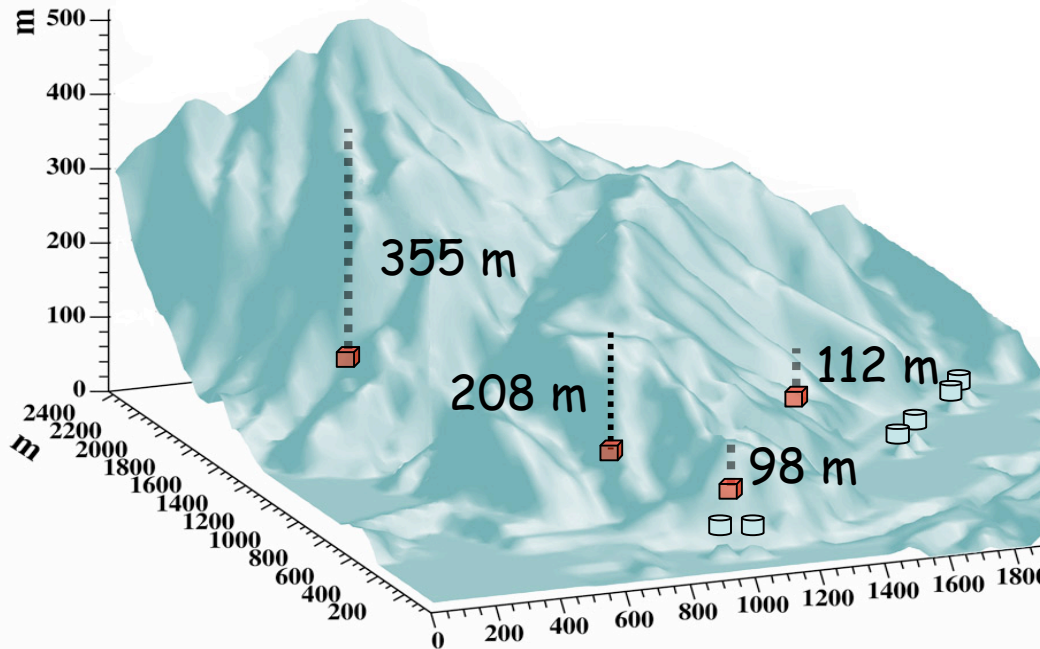
2-zone Prototype at IHEP

- 0.5 ton unloaded LS (now replaced with Gd-LS)
- 45 8" PMTs with reflecting top and bottom



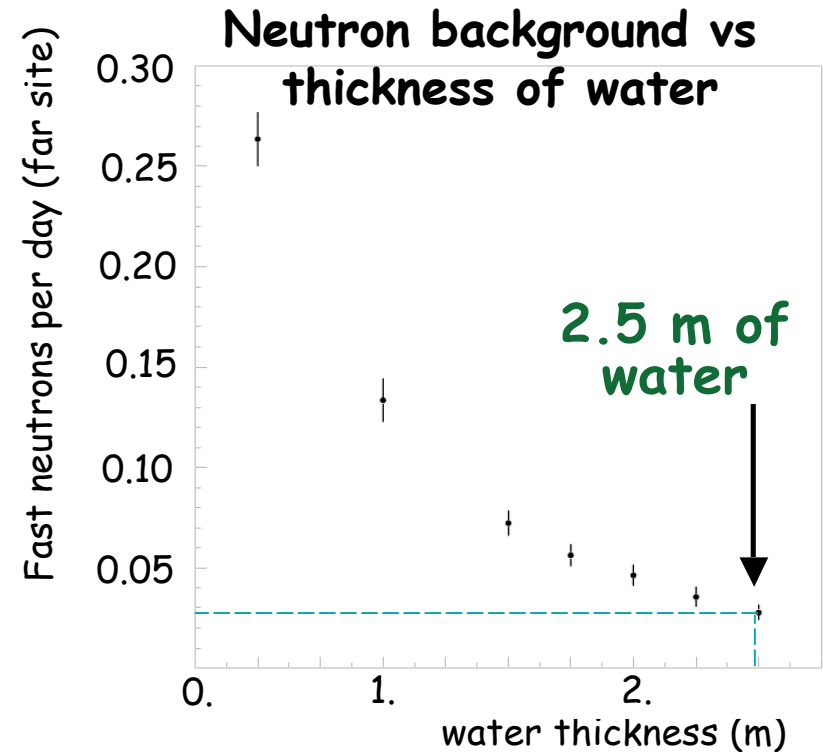
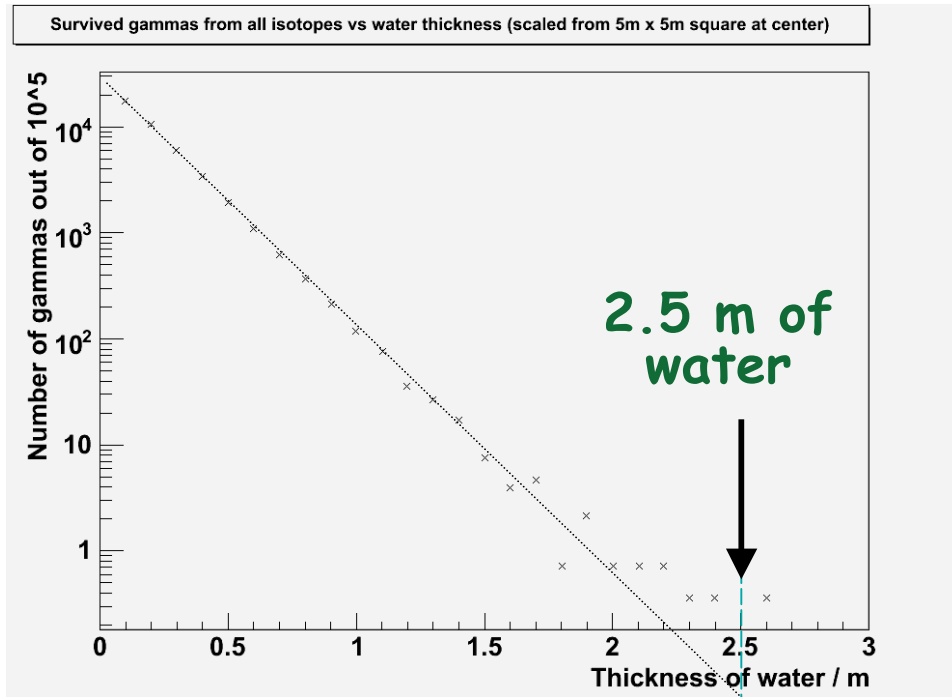
Cosmic-ray Muon

- Use a modified Gaisser parametrization for cosmic-ray flux at surface
- Apply MUSIC and mountain profile to estimate muon intensity & energy



	DYB	LingAo	Mid	Far
Overburden (m)	98	112	208	355
Muon intensity (Hz/m ²)	1.16	0.73	0.17	0.041
Mean Energy (GeV)	55	60	97	138

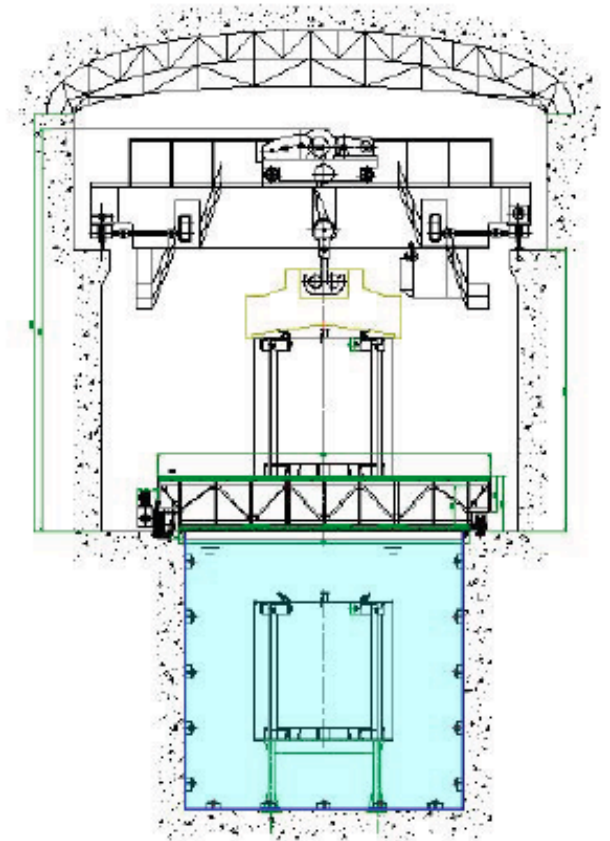
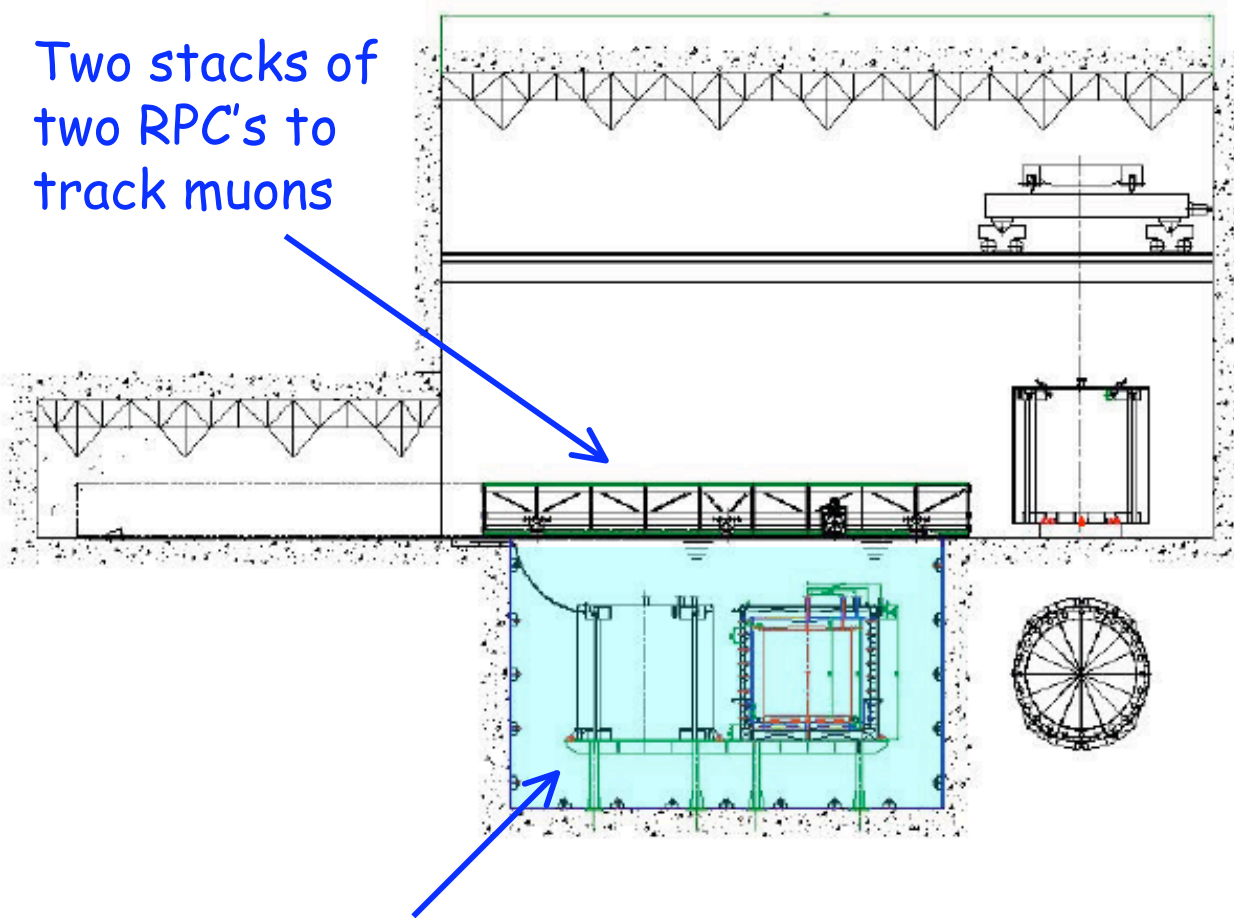
Shielding Antineutrino Detectors



- Detector modules enclosed by 2.5 m of water to shield energetic neutrons produced by cosmic-ray muons and gamma-rays from the surrounding rock

An Example of the Muon System

Two stacks of two RPC's to track muons

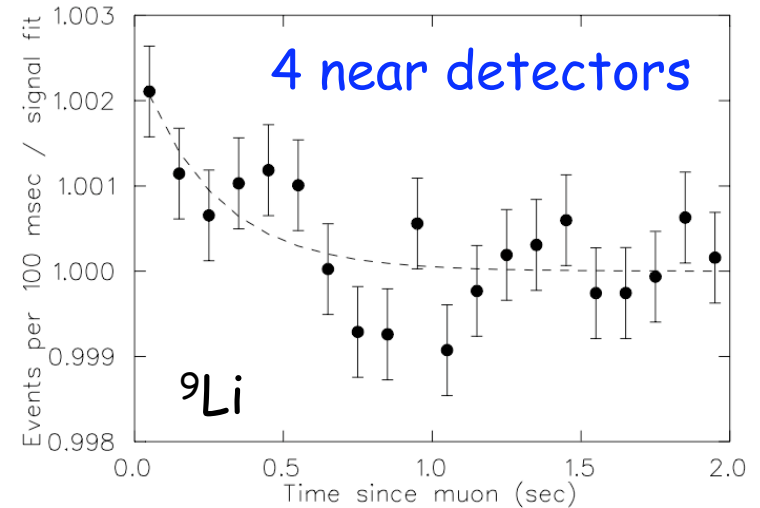


Water shield also serves as Cherenkov counter for tagging muons

- Combined efficiency of Cherenkov and tracker $> 99.5\%$ with error measured to better than 0.25%

Background

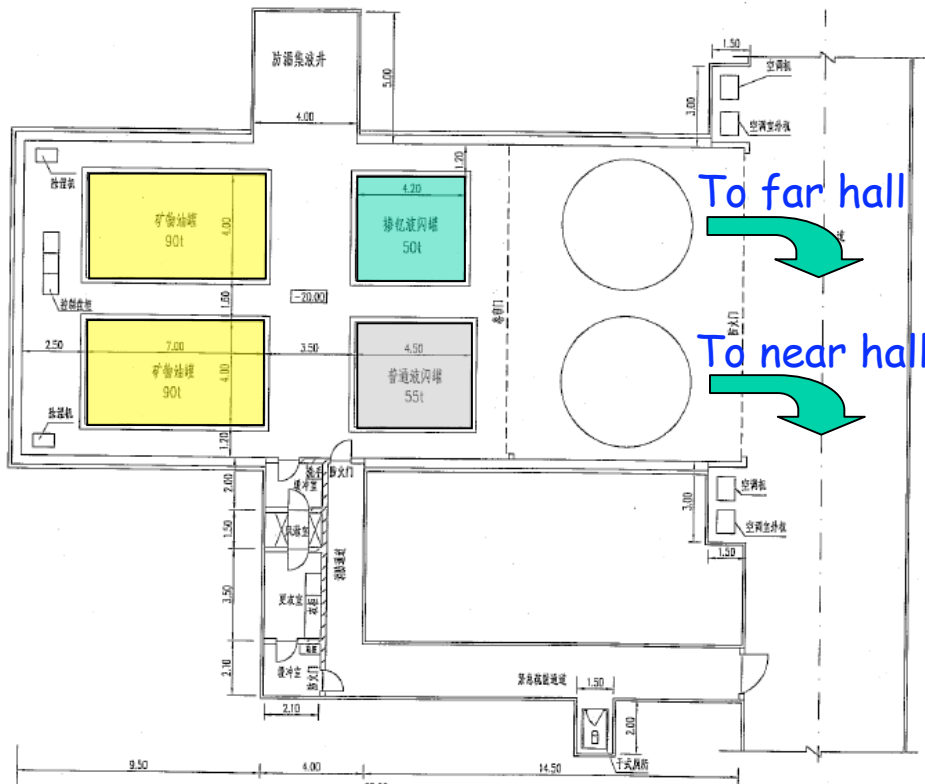
- Uncorrelated background:
Sources: U/Th/K/Rn/neutron
Single gamma rate @ 0.9MeV < 50Hz/module
Single neutron rate < 1000/day/module
- Correlated backgrounds: $n \propto E_{\mu}^{0.75}$
Fast Neutrons: double coincidence
 ${}^8\text{He}/{}^9\text{Li}$: beta-neutron emitting decays
- All these background events can be measured and corrected



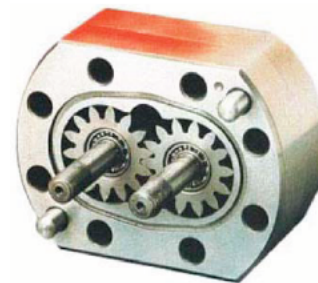
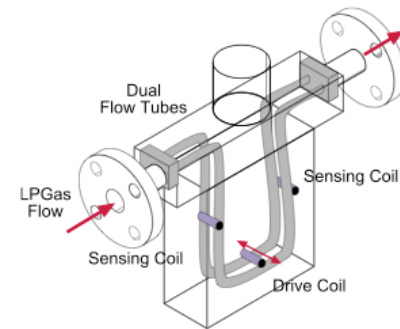
Per module	Daya Bay Near	Ling Ao Near	Far Hall
Baseline (m)	363	481 from Ling Ao 526 from Ling Ao II	1985 from Daya Bay 1615 from Ling Ao's
Overburden (m)	98	112	350
Radioactivity (Hz)	<50	<50	<50
Muon rate (Hz)	36	22	1.2
Antineutrino Signal (events/day)	930	760	90
Accidental Background/Signal (%)	<0.2	<0.2	<0.1
Fast neutron Background/Signal (%)	0.1	0.1	0.1
${}^8\text{He}+{}^9\text{Li}$ Background/Signal (%)	0.3	0.2	0.2

Controlling Target Mass & Composition

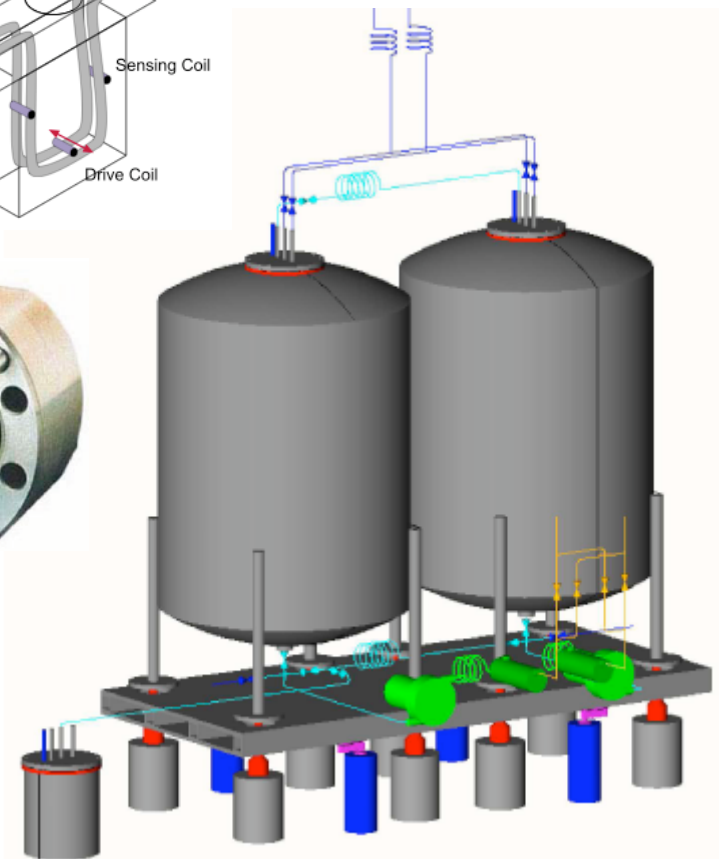
- Store one batch of pure LAB in a 200-T tank
- Mix and fill a pair of detector at a time underground with controls of **Coriolis mass flow meters** (reproducible at 0.1%), **load cells** (0.008% accuracy), and **volume flow meters** (good to 0.02%), as well as temperature monitoring :



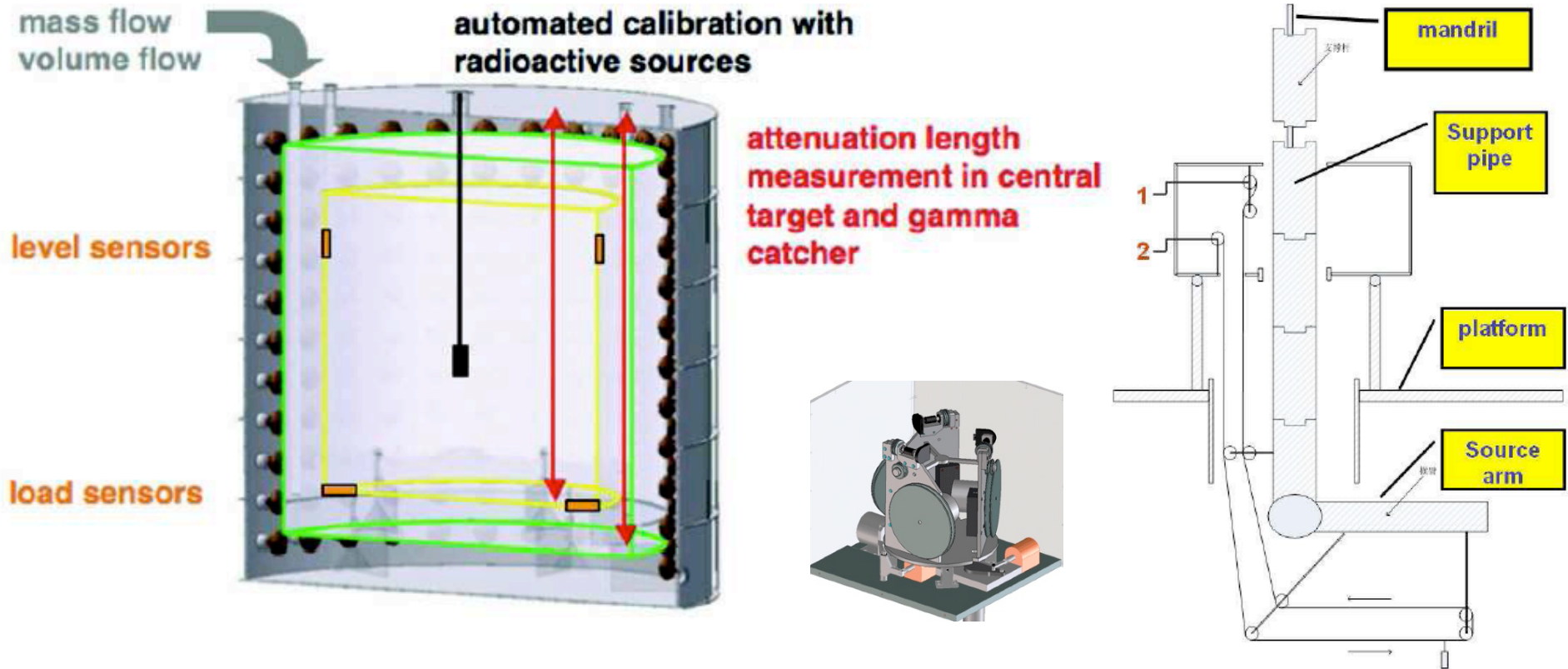
Kam-Biu Luk



Daya Bay

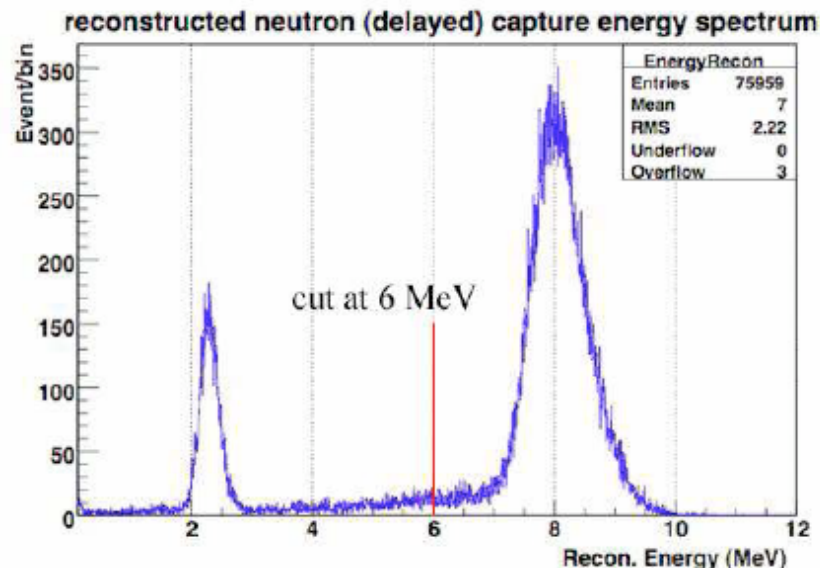
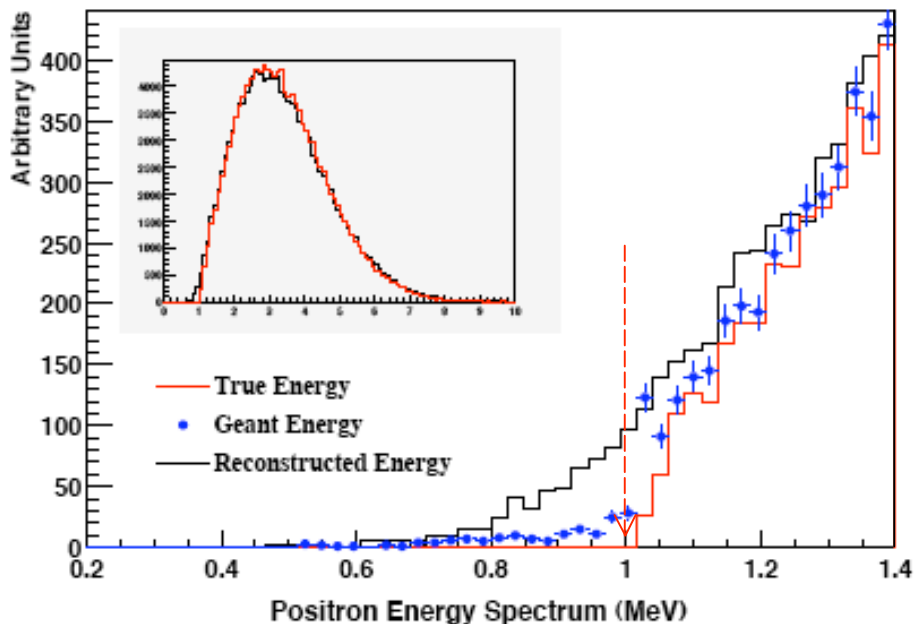


Calibration And Monitoring

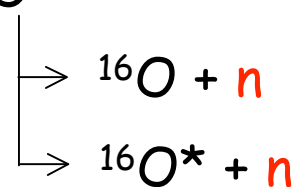
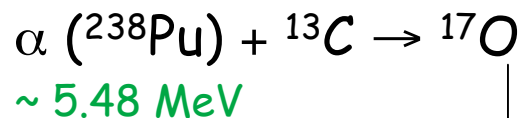


- Use a manual system to do full-volume calibrations
- Monitor level, load and temperature sensors regularly
- Weekly calibration with radioactive sources and LED along vertical axes with an automated calibration system

Calibrating Energy Cuts



- Routine calibration with
 - ^{68}Ge : e^+ annihilation yields two 0.511 MeV γ 's
 - ^{60}Co : 2.6 MeV in γ 's
 - ^{238}Pu - ^{13}C (α -n source):



Systematic Uncertainties

- Reactor-related:

Near detectors
really help!

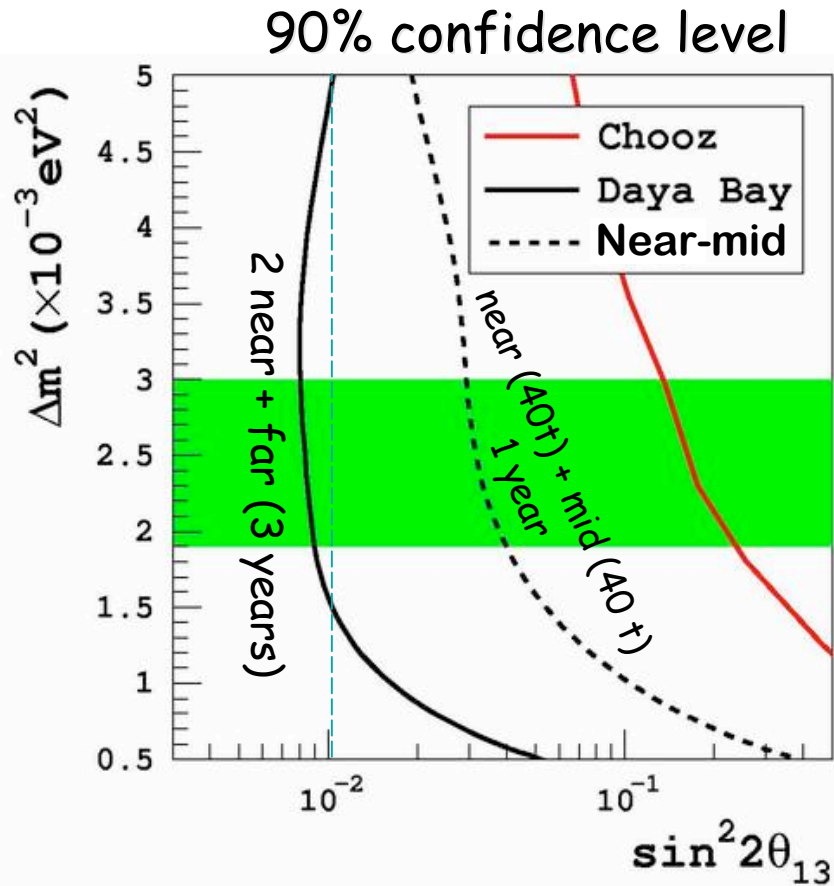
Number of cores	α	$\sigma_\rho(\text{power})$	$\sigma_\rho(\text{location})$	$\sigma_\rho(\text{total})$
4	0.338	0.035%	0.08%	0.087%
6	0.392	0.097%	0.08%	0.126%

- Detector-related:

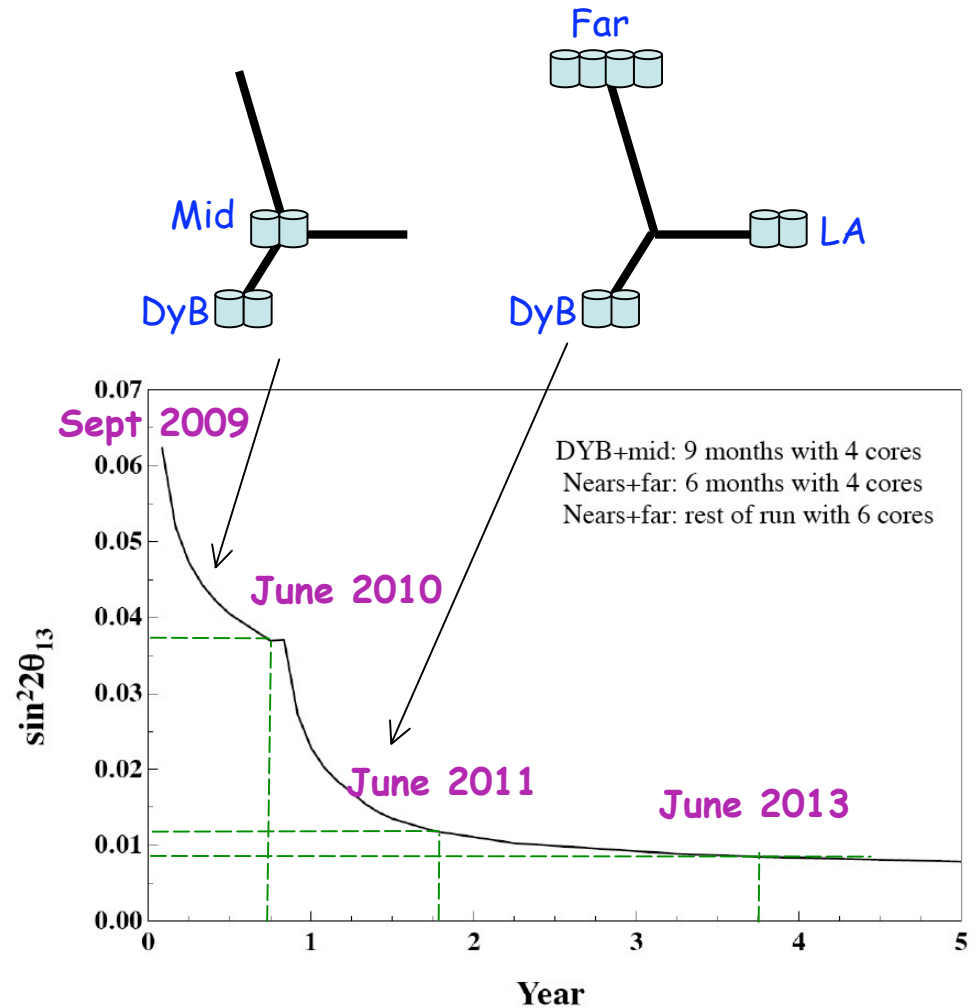
Source of uncertainty		Chooz (<i>absolute</i>)	Daya Bay (<i>relative</i>)		
			Baseline	Goal	Goal w/Swapping
# protons	H/C ratio	0.8	0.2	0.1	0
	Mass	-	0.2	0.02	0.006
Detector Efficiency	Energy cuts	0.8	0.2	0.1	0.1
	Position cuts	0.32	0.0	0.0	0.0
	Time cuts	0.4	0.1	0.03	0.03
	H/Gd ratio	1.0	0.1	0.1	0.0
	n multiplicity	0.5	0.05	0.05	0.05
	Trigger	0	0.01	0.01	0.01
	Live time	0	< 0.01	< 0.01	< 0.01
Total detector-related uncertainty		1.7%	0.38%	0.18%	0.12%

↑
anticipated
with R&D

Sensitivity in $\sin^2 2\theta_{13}$



- Use rate and spectral shape
- input relative detector systematic error of 0.38%



Sensitivity of 9 months running with near-mid, 3 years with near-far

Funding And Support

- Funding Committed from
 - Chinese Academy of Sciences,
 - Ministry of Science and Technology
 - Natural Science Foundation of China
 - China Guangdong Nuclear Power Group
 - Shenzhen municipal government
 - Guangdong provincial government
- Strong support from:
 - China Guangdong Nuclear Power Group
 - China atomic energy agency
 - China nuclear safety agency
- Supported by BNL/LBNL LDRD funds
- Supported by DOE \$1M R&D to date
- Support by funding agencies from other countries & regions
- China plans to provide civil construction and ~half of the detector systems; U.S.plans to bear ~half of the detector cost



Summary

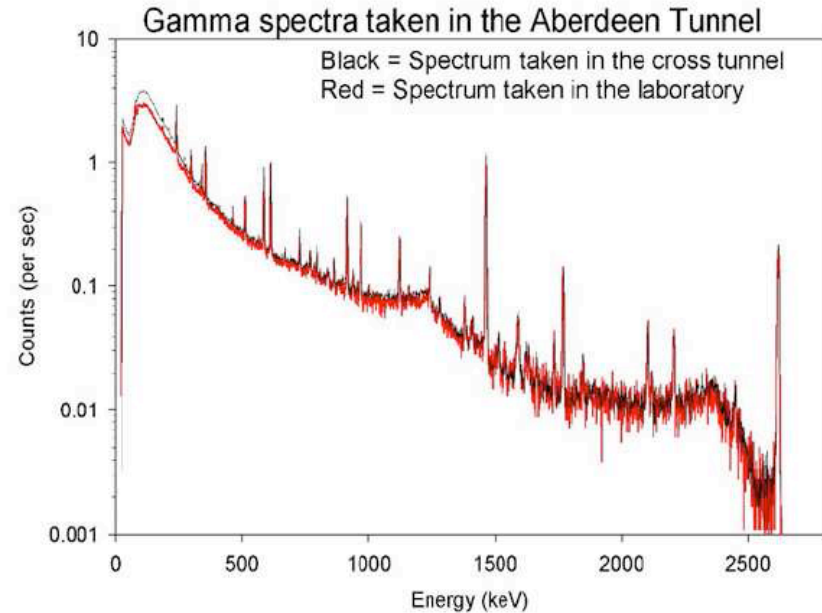
- The Daya Bay experiment will reach a sensitivity of ≤ 0.01 for $\sin^2 2\theta_{13}$
- Design of detectors is in progress and R&D is ongoing
- Completed DOE scientific review, and is preparing for DOE CD-1 review in April, 2007
- Detailed design of tunnels and infrastructures is in progress
- The Daya Bay Collaboration continues to grow
- Received commitments from Chinese agencies and power plant
- Plan to start civil construction in 2007, deploying detectors in 2009, begin full operation in 2010, and reach the designed sensitivity by 2013

Schedule

- Will begin civil construction Jun 2007
- Bring up the first pair of detectors Jun 2009
- Begin data taking with the Near-Mid configuration Sept 2009
- Begin data taking with the Near-Far configuration Jun 2010
- Reach $\sin^2 2\theta_{13} < 0.01$ Jun 2013

Aberdeen Tunnel Experiment (Hong Kong)

For studying cosmic muons, spallation neutrons, and gamma background



similar overburden/geology
 between Aberdeen
 and Daya Bay

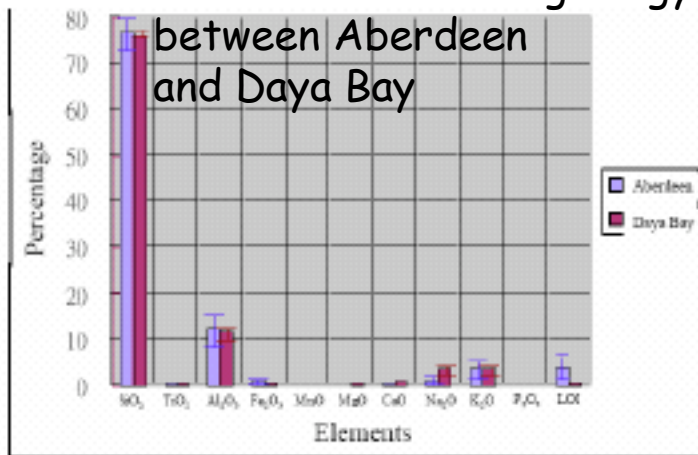


Fig. 1 A comparison of rock compositions at Daya Bay and Aberdeen

