Search for Neutrinoless Double Beta Decay in the GERDA Experiment

Andrea Kirsch — on behalf of the Gerda Collaboration —

Max-Planck Institut für Kernphysik, Heidelberg

Xth Recontres du Vietnam Flavour Physics Conference

Quy Nhon — 31st July 2014



1 / 17

ββ

GERDA

Outline

- 1 Neutrinos and Double Beta Decay $(2\nu\beta\beta$ vs. $0\nu\beta\beta)$
- 2 GERDA experimental design
- Operation of the second state of the second
- 4 Background modeling
- 5 Pulse shape discrimination
- 6 Unblinding and results of $0\nu\beta\beta$ analysis
- Outlook on Phase II

Flavour eigenstates ν_{α} (with $\alpha = e, \mu, \tau$) as linear superposition of mass eigenstates ν_i

 $|\nu_{\alpha}\rangle = \sum_{i=1}^{3} U_{\alpha i} |\nu_{i}\rangle$

Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix



where
$$s_{ij} = \sin \theta_{ij}$$
 and $c_{ij} = \cos \theta_{ij}$



What we know

- squared mass differences Δm^2_{12} and $|\Delta m^2_{23}|$
- mixing angles θ_{12} , θ_{23} and θ_{12}

from e.g. solar neutrino + long baseline reactor, amospheric neutrino + long baseline accelerator or short baseline reactor / accelerator experiments

What we do not know

- new physics behind SM
- absolute mass scale
- mass hierarchy
- nature of the neutrino (Dirac or Majorana?)

Flavour eigenstates ν_{α} (with $\alpha = e, \mu, \tau$) as linear superposition of mass eigenstates ν_i

 $|\nu_{\alpha}\rangle = \sum_{i=1}^{3} U_{\alpha i} |\nu_{i}\rangle$

Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix





What we know

- squared mass differences $\left(\Delta m_{12}^2\right)$ and $\left|\Delta m_{23}^2\right|$
- mixing angles (θ_{12}) , θ_{23} and θ_{12}

from e.g. (solar neutrino + long baseline reactor), amospheric neutrino + long baseline accelerator or short baseline reactor / accelerator experiments

What we do not know

- new physics behind SM
- absolute mass scale
- mass hierarchy
- nature of the neutrino (Dirac or Majorana?)

Flavour eigenstates ν_{α} (with $\alpha = e, \mu, \tau$) as linear superposition of mass eigenstates ν_i

 $|\nu_{\alpha}\rangle = \sum_{i=1}^{3} U_{\alpha i} |\nu_{i}\rangle$

Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix





What we know

- squared mass differences Δm^2_{12} and $\left|\Delta m^2_{23}\right|$
- mixing angles θ_{12} , (θ_{23}) and θ_{12}

from e.g. solar neutrino + long baseline reactor,

amospheric neutrino + long baseline accelerator

or short baseline reactor / accelerator experiments

What we do not know

- new physics behind SM
- absolute mass scale
- mass hierarchy
- nature of the neutrino (Dirac or Majorana?)

3 / 17

Flavour eigenstates ν_{α} (with $\alpha = e, \mu, \tau$) as linear superposition of mass eigenstates ν_i

 $|\nu_{\alpha}\rangle = \sum_{i=1}^{3} U_{\alpha i} |\nu_{i}\rangle$

Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix



where
$$s_{ij} = \sin \theta_{ij}$$
 and $c_{ij} = \cos \theta_{ij}$



What we know

- squared mass differences Δm^2_{12} and $|\Delta m^2_{23}|$
- mixing angles θ_{12} , θ_{23} and $\left(\theta_{13}\right)$

from e.g. solar neutrino + long baseline reactor, amospheric neutrino + long baseline accelerator or short baseline reactor/accelerator experiments

What we do not know

- new physics behind SM
- absolute mass scale
- mass hierarchy
- nature of the neutrino (Dirac or Majorana?)

Unveiling the nature of the neutrino



4 / 17

Unveiling the nature of the neutrino



Double Beta ($\beta\beta$) Decay:

- rare second order nuclear transision
- occurs between 2 even-even isobars
- only if single β decay energetically forbidden or ΔJ large



VS.

Majorana: $\nu = \overline{\nu}$



allows for an unambigous test

- 35 canditates in nature
- predominantly used isotopes for experimental $\beta\beta$ search:
 - ⁴⁸Ca, ⁷⁶Ge, ⁸²Se, ⁹⁶Zr, ¹⁰⁰Mo, ¹¹⁶Cd, ¹²⁸Te. ¹³⁰Te, ¹³⁶Xe, ¹⁵⁰Nd, ...

Andrea Kirsch (MPIK)

Quy Nhon, 31st July 2014

4 / 17

Double Beta Decay

 $2\nu\beta\beta$: $(A,Z) \rightarrow (A,Z+2) + 2e^- + 2\overline{\nu}_e$

- so far observed in 12 nuclei
- half lifes in the range of $10^{19} 10^{24}$ yr with $T_{1/2}^{2\nu}({}^{76}\text{Ge}) = 1.84_{-0.10}^{+0.14} \cdot 10^{21} \text{yr}$ J. Phys. G: Nucl. Part. Phys. 40 (2013) 035110
- $\Delta L = 0$: Lepton-number conserved
- allowed by Standard Model



- component
- still hunted process

 $0\nu\beta\beta$:



- note: one claim by subgroup of HdM with $T_{1/2}^{0\nu}({}^{76}\text{Ge}) = 1.19_{-0.23}^{+0.37} \cdot 10^{25} \text{yr}$ Phys. Rev. Lett. B 586, 198-212 (2004)
- $\Delta L = 2$: Lepton-number violation
- not allowed by Standard Model



Andrea Kirsch (MPIK)

Double Beta Decay

 $2\nu\beta\beta$: $(A,Z) \rightarrow (A,Z+2) + 2e^- + 2\overline{\nu}_e$

- so far observed in 12 nuclei
- half lifes in the range of $10^{19} 10^{24} \mathrm{yr}$ with $T_{1/2}^{2\nu}({}^{76}\text{Ge}) = 1.84_{-0.10}^{+0.14} \cdot 10^{21} \text{yr}$ J. Phys. G: Nucl. Part. Phys. 40 (2013) 035110
- $\Delta L = 0$: Lepton-number conserved
- allowed by Standard Model



- Majorana mass component
- still hunted process

 $0\nu\beta\beta$:



note: one claim by subgroup of HdM with $T_{1/2}^{0\nu}({}^{76}\text{Ge}) = 1.19_{-0.23}^{+0.37} \cdot 10^{25} \text{yr}$ Phys. Rev. Lett. B 586, 198-212 (2004)

can be mediated by e.g. light Majorana ν , R-handed weak currents, SUSY particles, ...



Andrea Kirsch (MPIK)

Neutrinoless Double Beta Decay





Values taken from: Phys. Rev. D86 (2012) 013012; JCAP 07 (2012) 053

Neutrinoless Double Beta Decay



Neutrinoless Double Beta Decay



Search in ⁷⁶Ge (with well established semiconductor technology)

Advantages

- source = detector \rightarrow high ϵ
- High Purity Ge → low intrinsic BI
- FWHM @ $Q_{\beta\beta} \sim 0.2\% \rightarrow \text{excellent } \Delta E$
- test of claim for $0\nu\beta\beta$ observation by HdM without depending on NME

Disadvantages

- low $Q_{\beta\beta}$ -value \rightarrow possible external BIfrom e.g. ²⁰⁸Tl + small $G^{0\nu}(Q_{\beta\beta}, Z)$
- a=7.8% for ${}^{76}\text{Ge} \rightarrow \text{enrichment needed}$
- rather long and costly process to get large active detector mass

GER manium Detector Array Eur. J. Phys. C73 (2013) 2330





Background reduction techniques from construction of setup:

- located in Hall A of LNGS underground laboratory, Italy \rightarrow cosmic μ flux reduced by 10^6
- novel idea: Ge detectors are operated bare in LAr as coolant
- only minimal amount of (screened) material close to detectors
- rock shielded by ultra-pure water → copper lined steel cryostat → LAr
- active μ -veto by using water tank instrumented with PMTs and plastic scintillator above the cryostat neck

Andrea Kirsch (MPIK)

Search for $0\nu\beta\beta$ in GERDA

7 / 17

GERDA Phase I data

Experiment	Phase	Mass	Aspired BI	Livetime	$T^{0 u}_{1/2}$ Sensitivity	Phase II \rightarrow
proceeds in	T Hase	[kg]	[cts/(keV·kg·yr)]	[yr]	[yr]	liquid argon
two phases:	I	15	10^{-2}	1	$2.4 \cdot 10^{25}$	scintillation
		35	10^{-3}	3	$1.4\cdot 10^{26}$	veto upgrade

$8 \times$ semi-coaxial detectors

- inherited from HdM (ANG1-5) & IGEX (RG1-3) experiments; all reprocessed at Canberra
- enrichement fraction of $^{76}\mathrm{Ge:}$ ${\sim}86\%$
- data taking: November 2011 May 2014
- 2 detectors not considered (high leakage current)
- ${ullet}$ total mass used for analysis: $14.6~{\rm kg}$

- 100 mm



5× Broad Energy Germanium (BEGe) detectors

- from ${\sim}30$ newly processed Phase II detectors
- enrichement fraction of $^{76}\text{Ge:}$ $\sim 88\%$
- data taking (first test): July 2012 May 2014
- 1 detector not considered (unstable behaviour)
- total mass used for analysis: 3.0 kg



Additional background reduction (by off-line analysis)

Signal

Ge ββ_φ



- $\beta\beta$ events: range for 1MeV electrons in Ge about $\sim 1 \text{ mm}$
- interaction via ionization or exitation of absorber atoms
- one drift of electrons and holes originated close-by in located charge cloud
- \rightarrow single-site event (SSE)

Background

- γ events: range of 1MeV γ's in Ge ≤10× larger (compared to electrons)
- interaction via compton scattering, e⁺e⁻ pair creation or photelectric absorption
- sum of several separated electron-hole drifts
- \rightarrow multi-site event (MSE)





Ramo-Shockley theorem: $Q(t) = -q \times [\phi(\mathbf{r}_{b}(t)) - \phi(\mathbf{r}_{e}(t))]$

- distinguish between SSE's and MSE's
- different PSD for semi-coaxial and BEGe detectors

Signal processing: diode \rightarrow amplifier \rightarrow FADC \rightarrow digital filter \rightarrow E, pulse shape, ...

- anti-coincidence between detectors: ${\sim}20\%$ rejected @ $Q_{\beta\beta}$
- quality cuts (e.g. T, E instabilities): ~9% rejected @ $Q_{\beta\beta}$
- pulse shape discrimination (PSD): $\sim 50\%$ rejected @ $Q_{\beta\beta}$

Andrea Kirsch (MPIK)

Calibration, time stability and energy resolution



Blinded physics spectrum



Main background components:

- β -spectrum of ³⁹Ar (with Q = 565 keV)
- $2\nu\beta\beta$ -spectrum of 76 Ge
- γ -lines from 40 K, 42 K, 60 Co, 214 Bi, 212 Bi, 208 Tl & 228 Ac
- α-spectrum of ²³⁸U chain (in semi-coaxial detectors)

Division in 3 data sets:

- semi-coaxial data splitted in two sets ("golden", "silver") according to BI
- "BEGe" set kept seperated due to different resolution and background

data set	Exposure [kg·yr]	FWHM @ $Q_{\beta\beta}$ [keV]
golden	17.9	4.8 ± 0.2
silver	1.3	4.8 ± 0.2
BEGe	2.4	3.2 ± 0.2

Andrea Kirsch (MPIK)

11 / 17

Background model Eur. Phys. J. C74 (2014) 2764

General procedure:

- simulation of known (from material screening) and observed (from detector operation) background sources
- spectral fit with combination of all components in the energy region between 570 keV and 7500 keV on the 3 data sets
- 2 extremes: "minimum" (all known + visible contributions) & "maximum" (additional contributions from other possible locations)

Results:

- no $\gamma\text{-line expected in the blinded}$ window around $Q_{\beta\beta}$
- flat background between 1930 keV and 2190 keV (= ROI) excluding known peaks @ 2103 keV (²⁰⁸TI SEP) as well as 2119 keV (²¹⁴Bi)

•
$$BI = (1.76 - 2.38) \cdot 10^{-2} \frac{\text{cts}}{\text{kg·keV·yr}}$$

(depending on assumption of source location)



Andrea Kirsch (MPIK)

Search for $0\nu\beta\beta$ in GERDA

Background model Eur. Phys. J. C74 (2014) 2764

General procedure:

- simulation of known (from material screening) and observed (from detector operation) background sources
- spectral fit with combination of all components in the energy region between 570 keV and 7500 keV on the 3 data sets
- 2 extremes: "minimum" (all known + visible contributions) & "maximum" (additional contributions from other possible locations)

Results:

- no $\gamma\text{-line expected in the blinded}$ window around $Q_{\beta\beta}$
- flat background between 1930 keV and 2190 keV (= ROI) excluding known peaks @ 2103 keV (²⁰⁸TI SEP) as well as 2119 keV (²¹⁴Bi)

•
$$BI = (1.76 - 2.38) \cdot 10^{-2} \frac{\text{cts}}{\text{kg} \cdot \text{keV} \cdot \text{yr}}$$

(depending on assumption of source location)

Partial unblinding @ $Q_{\beta\beta} \pm 20 \text{ keV} \rightarrow \pm 5 \text{ keV}$ with 8.6-10.3 expected and 13 observed events



Pulse shape discrimination (PSD) Eur. Phys. J. C73 (2013) 2583

semi-coaxial: artificial neural network (ANN)

- TMVA / TMIpANN with 2 hidden layers of $n_{\rm var}$ and $n_{\rm var}{+1}$ nodes
- Input: time when charge pulse reaches 1%, 3%,...,90% of maximum (n_{var}=50)
- training using ²²⁸Th calibration data
 → SSE: ²⁰⁸TI DEP @ 1620.7 keV
 → MSE: ²¹²Bi FEP @ 1592.5 keV
- cut defined such that acceptance of 208 TI DEP @ 2614.5 keV is fixed to 90%
- $0\nu\beta\beta$ acceptance = 90^{+5}_{-9} %; background acceptance @ $Q_{\beta\beta} \sim 55\%$



BEGe: mono-parametric A/E

- A = amplitude of current pulse
- E = energy
- high capability of distinguising SSE from MSE and surface p⁺/n⁺ events
- tuned using 208 TI DEP (per definition A/E=1) \rightarrow keep 0.965 < A < 1.07
- well tested and understood method
- $0\nu\beta\beta$ acceptance = $(92 \pm 2)\%$ (determined from ²⁰⁸TI DEP and MC); background acceptance @ $Q_{\beta\beta} \leq 20\%$



Andrea Kirsch (MPIK)

Search for $0\nu\beta\beta$ in GERDA

Quy Nhon, 31st July 2014

Unblinding @ $Q_{\beta\beta} \pm 5 \text{ keV}$ Phys. Rev. Let. 111 (2013) 122503



vobs
5
2
L
L
L
)

Unblinding @ $Q_{\beta\beta} \pm 5 \text{ keV}$ Phys. Rev. Let. 111 (2013) 122503



Andrea Kirsch (MPIK)

Unblinding @ $Q_{\beta\beta} \pm 5~{\rm keV}$ Phys. Rev. Let. 111 (2013) 122503



Andrea Kirsch (MPIK)

Phase I finished \rightarrow currently preparing Phase II

Goal:

- Phase I: 20 kg·yr with a BI of $\sim 10^{-2}$ cts/(kg·keV·yr)
- Phase II: 100 kg·yr with a BI of $\sim 10^{-3}$ cts/(kg·keV·yr)

Upgrade:

- 2× detector mass
 (~20 kg BEGe + ~15 kg semi-coax)
- liquid argon veto instrumented with PMTs
- optimized readout electronics
- low mass detector suspension + wedge bond contact (made of clean material)

$T_{1/2}^{0 u} \propto f_{76} \cdot \epsilon \cdot \sqrt{rac{M \cdot t}{BI \cdot \Delta E}}$ (if background is present)

- increased statistic (higher $\beta\beta$ emitter mass M and longer data taking t)
- lower background index BI (by about factor 10)
- ${\bullet}~$ better energy resolution ΔE
- \rightarrow expected sensitivity $\simeq 1.4 \cdot 10^{26}$ yr (90% C.L.)



Conclusion: Phase I (2011 - 2013)



blind analysis performed (first time in this field)

- unprecedented BI of ~10⁻² cts/(kg·keV·yr) after PSD (order of magnitude lower than previous experiments)
- observed 3 events @ $Q_{\beta\beta} \pm 5$ keV compared to expected background of $2.5 \pm 0.3 \rightarrow$ no $0\nu\beta\beta$ signal

• upper half-life limit from profile likelihod fit: $T_{1/2}^{0\nu}(90\%\text{C.L.}) > 2.1 \cdot 10^{25} \text{ yr}$ with GERDA alone \rightarrow HdM claim (2004) rejected @ 99% level $T_{1/2}^{0\nu}(90\%\text{C.L.}) > 3.0 \cdot 10^{25} \text{ yr}$ with GERDA + HdM[1] + IGEX[2] [1] Euro Phys J A12 (2001) 147 [2] Phys Rev D65 (2002) 092007

Outlook: Phase II (start scheduled end 2014)• new BEGe detectors of additional $\sim 20 \text{ kg} \rightarrow \text{available}$ • upgrade of infrastructure (lock system, glove box, ...) \rightarrow finished• liquid argon scintillation veto \rightarrow currently installed• last integration tests (new contacting, electronics, ...,) \rightarrow ongoing

The Collaboration



Andrea Kirsch (MPIK)

Search for $0\nu\beta\beta$ in GERDA

Quy Nhon, 31st July 2014 17 / 17

The Collaboration

... and the people behind the GERDA experiment.



Picture taken during last GERDA Meeting in June 2014 hosted by the Max-Planck-Institut für Kernphysik @ Heidelberg, Germany

Andrea Kirsch (MPIK)

Search for $0\nu\beta\beta$ in GERDA