BEGe detectors in GERDA Phase I: performance, physics analysis and surface events

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Chair for Experimental Physics and Astroparticle Physics ПШ

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Frankfurt HK-DPG, 17th March 2014 BEGe in GERDA Phase I

Background from external β -emitter

 PSD for $\mathsf{n}+\mathsf{surface}$ events

BEGe PSD in GERDA Phase I

Toward Phase II

BEGe in Phase I

Background study

BEGe data set: 10% the exposure (2.4 of \sim 20 kg \cdot yr)



Background index w/o PSD: 4.2 10^{-2} cts/(keV \cdot kg \cdot y) ⁴²K on the surface is the dominant component in ROI.



BEGe in Phase I

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⁴²Ar cosmogenic, long half-life

 42 K beta emitter Q = 3.5 MeV

Locate on detector surfaces

Phase I solution



Surface β PSD





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Lower A/E than SSE

Small gap at high energies

Same region of MSE

Performances depends on A/E resolution



β rejection in vacuum cryostat



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Gerda A/E resolution

BEGe in vacuum cryostat

BEGe in GERDA Phase I

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The goal for Phase II is FWHM < 1.0%

PSD cut definition



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PSD cut definition



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Toward Phase II

New frond end apparatus:

Optimization for BEGe and PSD

Close pre-amp, less noise

Details in the next talk (HK 15.5)



Improving PSD:

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A/E based on a locale information increase sensitivity on pulse tail distinguish MSE and surface events



New Classification [WORK IN PROGRESS]



Second derivative charge signal:

Resolve the number of interaction:

- Δt max–min
- trigger on MSE-oscillations

Ratio minimum/maximum:

- max \propto energy (for SSE)
- min relate to the tail shape

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The first set of BEGe detectors have been operated in ${\rm GERDA}.$ No unforeseen contaminations have been measured.

With a conservative PSD, 92% acceptance of $0\nu\beta\beta$, we reach:

$$5 \cdot 10^{-3} \text{ cts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$$

⁴²K is the challenge for Phase II.
A/E method will reduce it below Phase II goal.
Other background sources will be further veto by liquid Ar scintillation.

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