

COSMOGENIC RADIONUCLIDES ***in stainless steel and copper***

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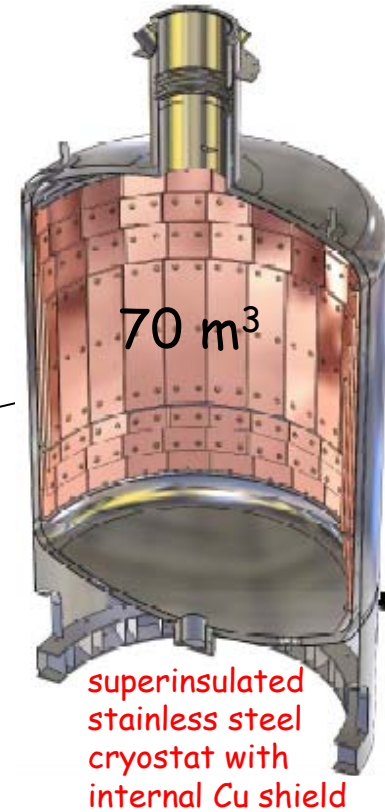
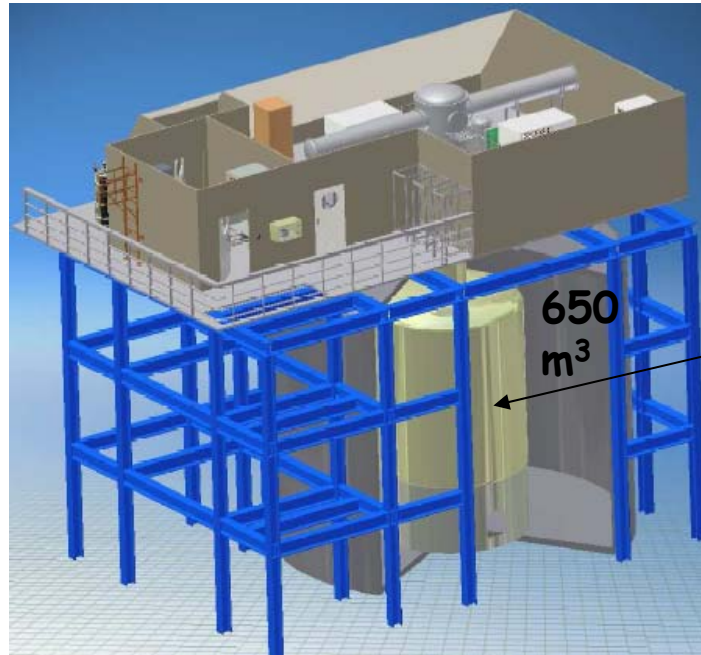
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- *cosmic ray activation analysis at LNGS surface and irradiation ages*
- *interpretation of deviation from secular equilibrium in U/Th chain*
- *comparison of cosmogenic production rates in Cu with Monte Carlo simulation*

see also Maneschg et al., NIM A 593 (2008), 448-453

GERDA

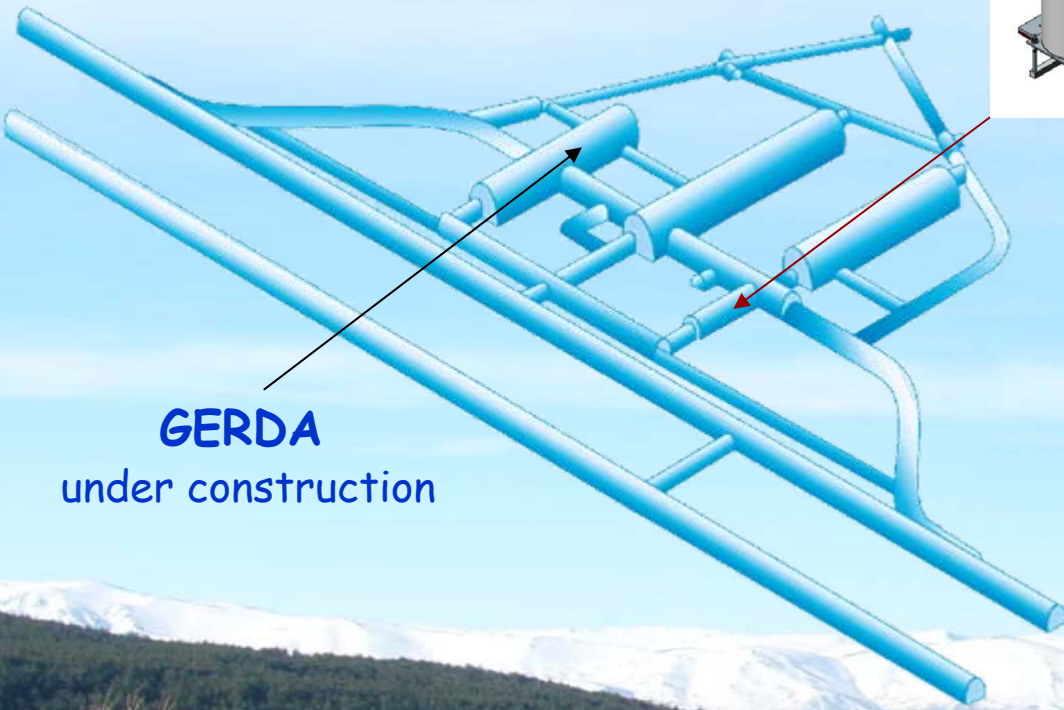
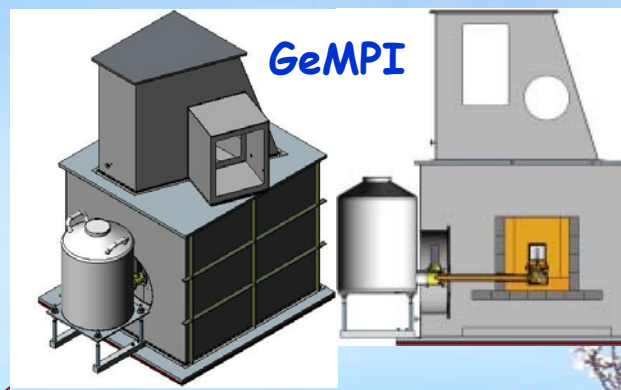
GERmanium Detector Array for the search of neutrinoless double beta decay of ^{76}Ge



low radioactivity stainless steel
($^{228}\text{Th} \leq 5 \text{ mBq/kg}$)

Laboratori Nazionali del GranSasso

(3800 mw.e. - 10^6 muon flux suppression)



GERDA
under construction



Cosmic ray induced isotopes in stainless steel (FeCrNiMo) measured for GERDA

| Steel 457.1 | activity [mBq/kg] | | | | | |
|---|-------------------------|--|---|--|---|---|
| sample | cosmogenic radionuclide | | | | | |
| $T_{1/2} \rightarrow$ | ^7Be 53.3 d | ^{54}Mn 312.2 d | ^{58}Co 70.9 d | ^{56}Co 77.3 d | ^{46}Sc 83.8 d | ^{48}V 16.0 d |
| Production \rightarrow channels | spallation | $^{56}\text{Fe}(n,p2n)$ ($\mu^-, \nu 2n$) | $^{60}\text{Ni}(n,p2n)$ ($\mu^-, \nu 2n$) $^{58}\text{Ni}(n,p)$ | $^{58}\text{Ni}(n,p2n)$ ($\mu^-, \nu 2n$) | $^{48}\text{Ti}(n,p2n)$ ($\mu^-, \nu 2n$) spallation on Fe | $^{50}\text{Cr}(n,p2n)$ ($\mu^-, \nu 2n$) spallation on Fe |
| G1 | ≤ 3.9 | 1.3 ± 0.4 | 0.67 ± 0.34 | ≤ 0.32 | ≤ 0.35 | 0.30 ± 0.11 |
| G2 | ≤ 3.0 | 1.5 ± 0.1 | 0.99 ± 0.12 | 0.17 ± 0.06 | 0.24 ± 0.06 | 0.36 ± 0.07 |
| G3 | ≤ 5.7 | 0.92 ± 0.24 | 0.56 ± 0.23 | ≤ 0.62 | ≤ 0.54 | 0.27 ± 0.11 |
| G4 | 9.6 ± 2.9 | 2.0 ± 0.3 | 0.71 ± 0.26 | ≤ 0.71 | ≤ 0.67 | 0.31 ± 0.13 |
| G5 | 4.8 ± 1.7 | 1.7 ± 0.2 | 0.69 ± 0.16 | 0.28 ± 0.10 | 0.47 ± 0.14 | 0.22 ± 0.09 |
| G6 | 13.6 ± 2.5 | 1.4 ± 0.2 | 0.59 ± 0.20 | ≤ 0.42 | ≤ 0.31 | 0.40 ± 0.12 |
| G7 | ≤ 5.9 | 1.6 ± 0.3 | 0.54 ± 0.27 | ≤ 0.6 | 0.61 ± 0.26 | 0.39 ± 0.13 |
| P. Rate sea level [[$10^3 \text{ sec}^{-1} \text{ kg}^{-1}$]] | 4.5 ± 0.7 | 2.7 ± 0.3 | 0.6 ± 0.09 | 0.24 ± 0.04 | 0.22 ± 0.04 | 0.4 ± 0.04 |

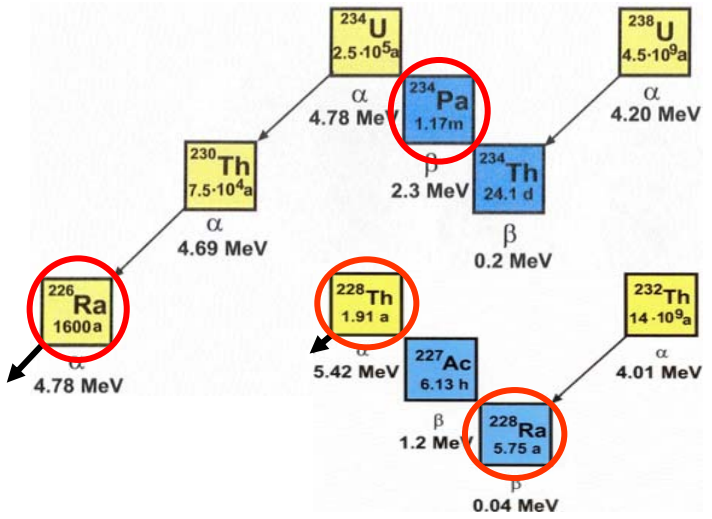
^{51}Cr : 2.0 ± 0.7 ^{52}Mn : 3.2 mBq/kg (Heusser, 1994) ^{56}Ni : 0.17 ± 0.05 [mBq/kg]

^{60}Co before: 11.1 ± 0.5 after: 11.5 ± 0.6 [Bq/kg]

irradiation time
 ← at sea level underground than exposed for 314 d at LNGS surface ($\approx 907 \text{ g/cm}^2$)
 $\approx 200 \text{ d}$
 $\approx 600 \text{ d}$
 all others are compatible with
 \rightarrow ^{56}Ni produced in situ, but contamination from filtration dust (up to 40 kBq/kg , Wershofen PTB)
 ← Normalized to sea level ($\times 0.12$)

Attempt to date steel production by disequilibrium in the Th-chain

| sample | primordial radionuclides [mBq/kg] and ratios | | | | | |
|----------------|--|-------------------|------------------------------------|-------------------|-------------------|-----------------------------------|
| | ^{234m}Pa | ^{226}Ra | $^{234m}\text{Pa}/^{226}\text{Ra}$ | ^{228}Th | ^{228}Ra | $^{228}\text{Th}/^{228}\text{Ra}$ |
| old ship steel | 5.7 ± 1.4 | 0.15 ± 0.02 | 38 ± 11 | 0.46 ± 0.07 | 0.47 ± 0.05 | 0.98 ± 0.18 |
| G5 | 54 ± 16 | 1.0 ± 0.6 | 54 ± 36 | 1.5 ± 0.2 | 1.0 ± 0.5 | 1.5 ± 0.77 |
| G7 | ≤ 56 | 3.9 ± 1.6 | $\leq 10-24$ | 5.2 ± 0.5 | 1.9 ± 1.0 | 2.7 ± 1.5 |

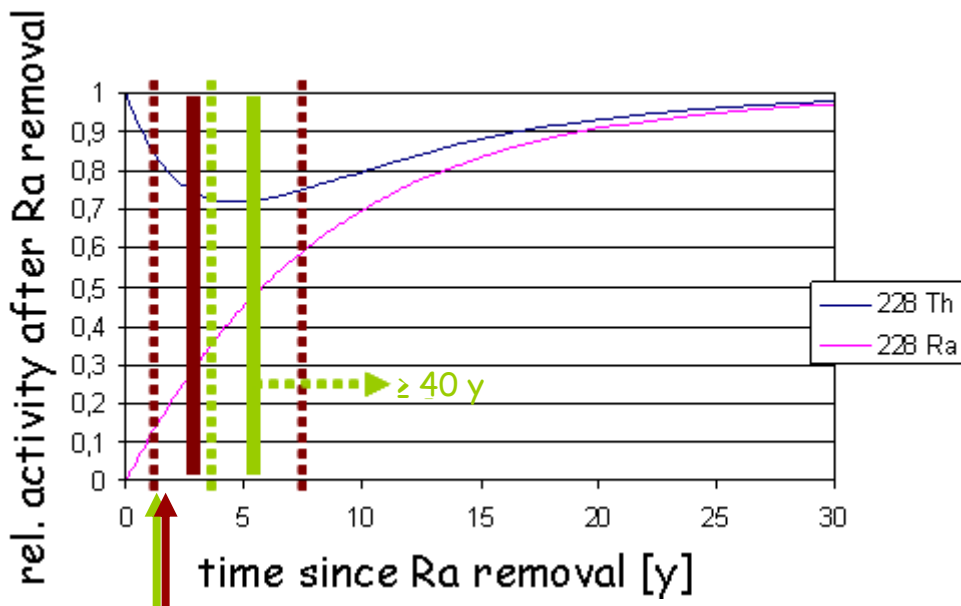


according to experts most likely the disequilibrium is introduced through additives in the iron-steel conversion.

Since electric furnace melting uses a large fraction of scrap, the mean age will be higher

Ra removal seems not be connected with steel production

sea level exposure age via ^{54}Mn



Cosmogenic activation of Cu during 270 days at LNGS surface ($\approx 930 \text{ g/cm}^2$)

| radionuclide | half-life | (saturation) activity [$\mu\text{Bq/kg}$] | | |
|------------------------|-----------------------|---|----------------------------------|-------------------|
| | | exposed | estimated from Cu used in HDM | estimated |
| ^{56}Co | 77.31 d | 230 ± 30 | | 557×2 |
| ^{57}Co | 271.83 d | 1800 ± 400 | 2100 | $2147 \checkmark$ |
| ^{58}Co | 70.86 d | 1650 ± 90 | 3600 | 3878×2 |
| ^{60}Co | 5.27 y | 2100 ± 190 | 2100 | $2367 \checkmark$ |
| ^{54}Mn | 312.15 d | 215 ± 21 | 700 | 791×3 |
| ^{59}Fe | 44.5 d | 455 ± 120 | | $157 : 4$ |
| ^{46}Sc | 83.79 d | 53 ± 18 | | $93 \checkmark$ |
| ^{48}V | 15.97 d | 110 ± 40 | | |
| primordial | | | | |
| ^{226}Ra (U) | 1600 y | < 35 | < 16 | |
| ^{228}Th (Th) | 1.91 y | < 20 | < 12 | |
| ^{40}K | 1.277×10^9 y | < 120 | < 110 | |

Monte Carlo simulation of hadronic interaction
Cebrian S., 2006. (IDEA-Projekt),
http://idea.dipscfm.uninsubria.it/frontend/docs/reports/report_upgrade_codes.pdf

applying:
 $\text{activity} = \text{PR}/2.1 \times (1 - e^{-\lambda t})$
 $\Rightarrow \leq 37$ days exposed at sea level

summary

- cosmic activation of metals is measurable with high sensitive Ge-spectrometry
- the exposure history of the investigated samples is accessible
- deviation from secular equilibrium in the primordial decay chains cannot be used for dating stainless steel production, but seems to reflect a mean age of the scrap material
- cosmogenic production rates in Cu can be rather well estimated by the Monte Carlo method

Further possible metal targets:

Al, Au, Cr, Co, Hg, Mn, Ni, Ti, V, W and Zn

enlarge the database on cosmic induced production rates
also for testing the Monte Carlo methods

