

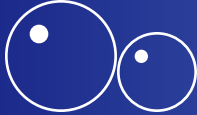


# Plans and prospects for the Hyper-Kamiokande detector

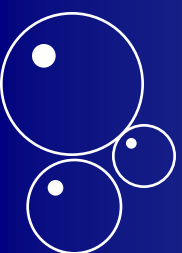
M.Miura

Kamioka Observatory, ICRR

BLV 2013, Heidelberg

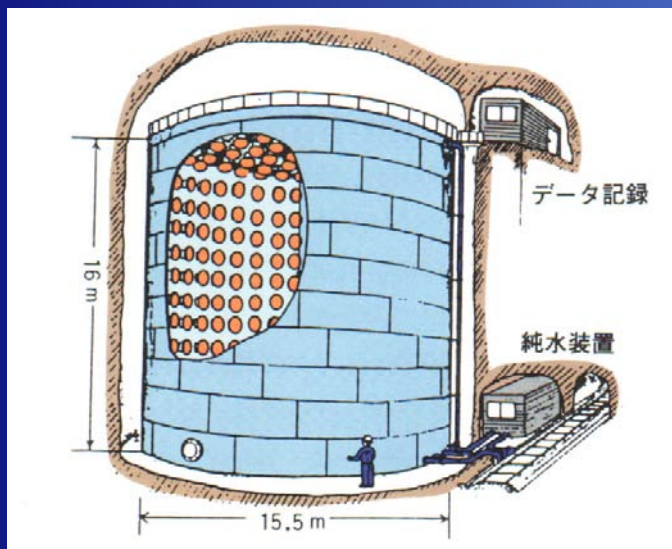


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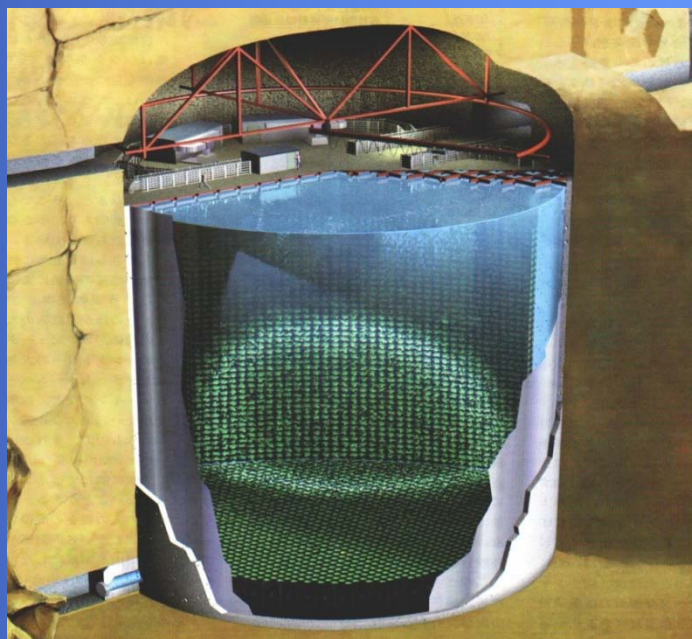
# 1. Introduction

Kamiokande  
(1983-1996)

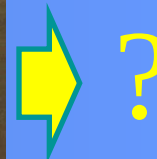


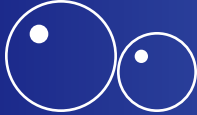
3000 ton, ~1000 PMTs  
- Observed Super Nova  
neutrino.

Super-Kamiokande  
(1996-Now)



50,000 ton, ~11,000 PMTs  
- Found neutrino  
oscillations





# Need larger detector

- Neutrino Oscillation

- $\theta_{13}$  is measured in 2012, non-zero and not so small.
- Long base line experiments have chance to measure  $\delta_{CP}$ .
  - use  $\bar{\nu}_{\mu}$ , cross section is lower than  $\nu_{\mu}$
- Mass hierarchy and Octant of  $\theta_{13}$  ← Atmospheric  $\nu$  with MSW effect.

→ Statistics is crucial !

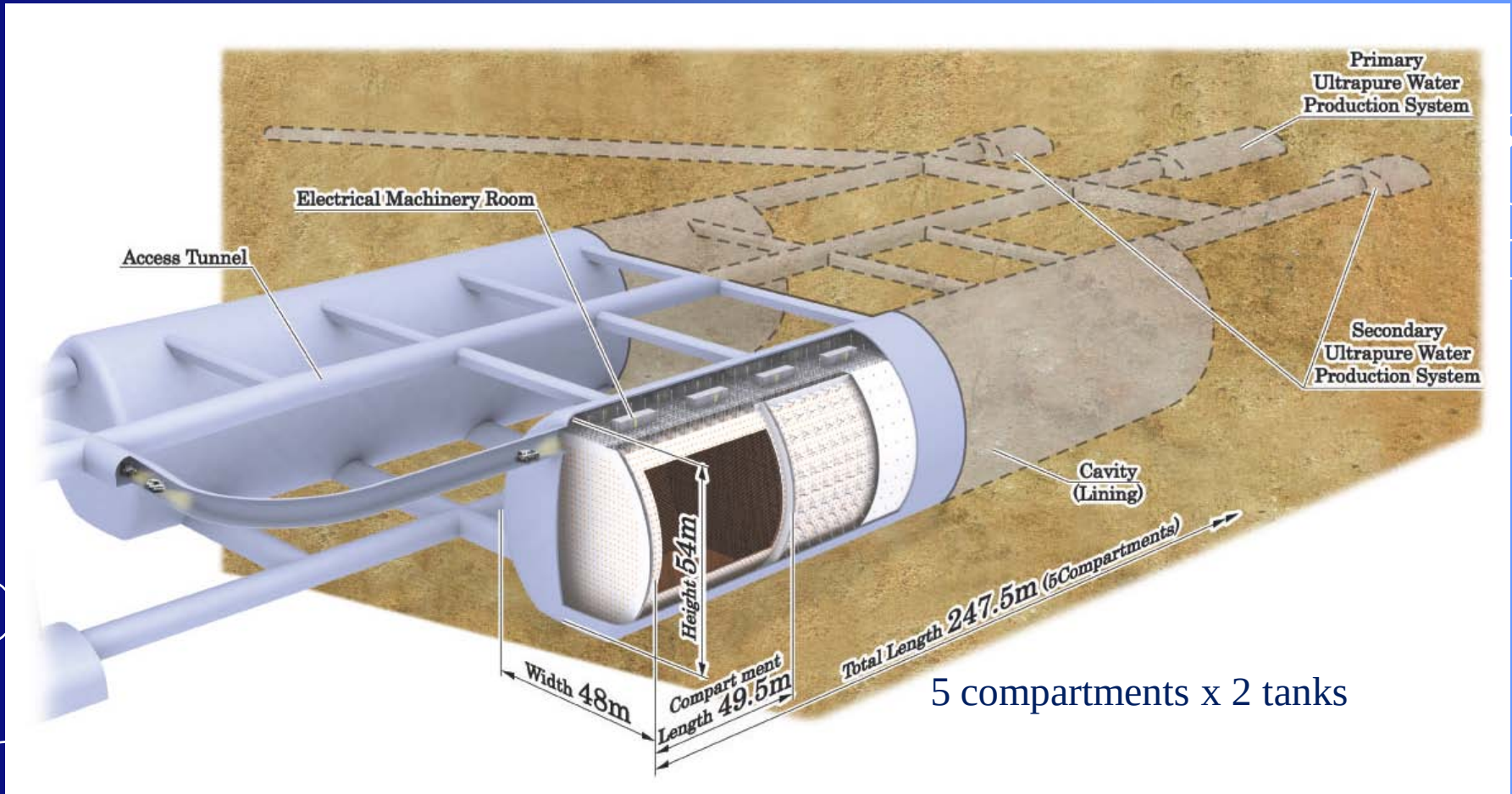


- Nucleon decay search

- Explored until 0.2 Mton·year exposure in Super-K.
- No evidence found yet.

→ Need more exposure !

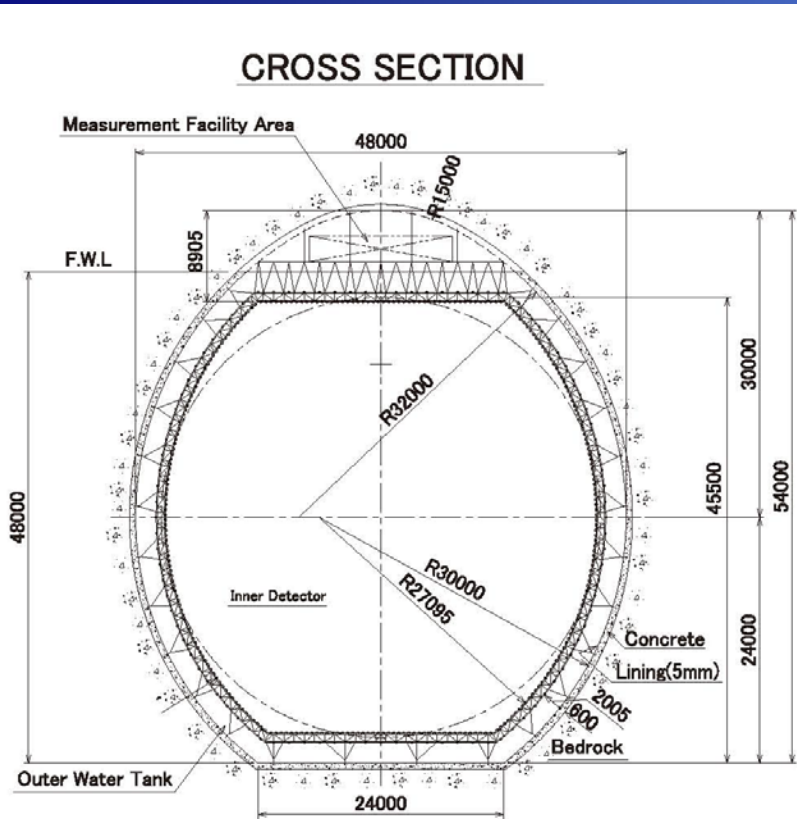
## 2. Hyper-Kamiokande



1 Mega ton Water Cherenkov Detector



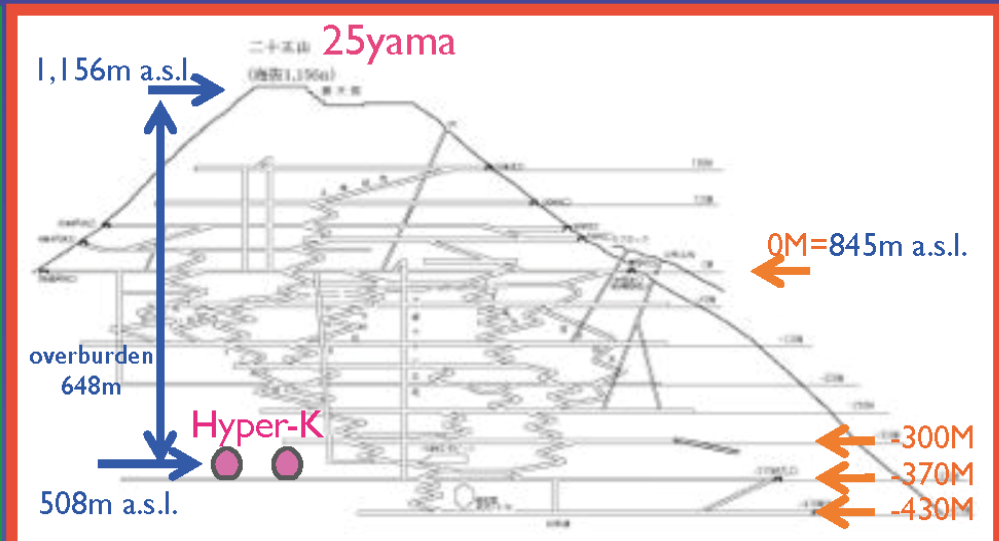
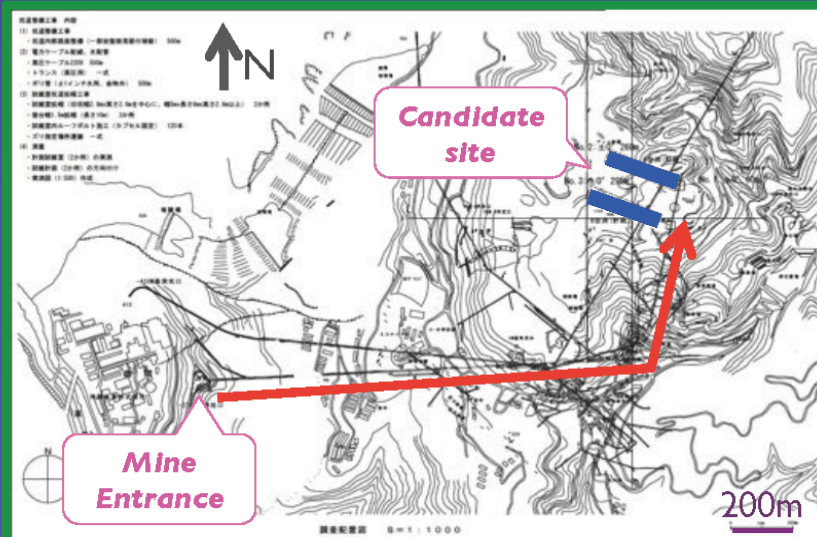
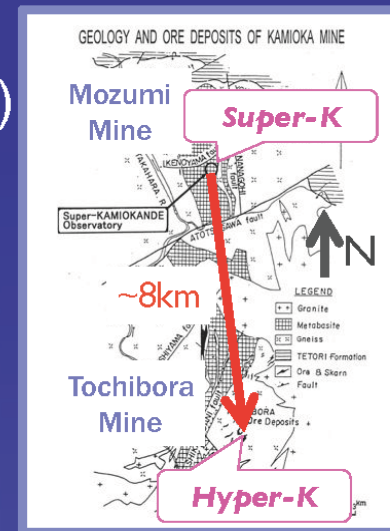
# Base design



- Total Volume: 0.99 Mton
- Inner Volume: 0.74 Mton
- Fiducial volume: 0.56 Mton  
(0.056 Mton x 10 compartments)
- Outer volume: 0.2 Mton
- Photo-sensors:
  - 99,000 20 inch PMT (ID)  
(20 % photo coverage)
  - 25,000 8 inch PMT (OD)

# Hyper-K candidate site

- ◆ 8km south from Super-K
- ◆ same T2K beam off-axis angle (2.5 degree)
- ◆ same baseline length (295km)
- ◆ 2.6km horizontal drive from entrance
- ◆ under the peak of Nijuugo-yama
- ◆ 648m of rock or 1,750 m.w.e. overburden
- ◆ 13,000 m<sup>3</sup>/day or 1 megaton/80days natural water



# Physics Topics in Hyper-K

- Accelerator  $\nu$  beam
- Atmospheric  $\nu$
- Solar  $\nu$
- Astrophysical  $\nu$  (Super Nova, Dark matter, e.t.c.)
- $\nu$  geophysics
- Nucleon decay

$\nu$  oscillation with CP violation

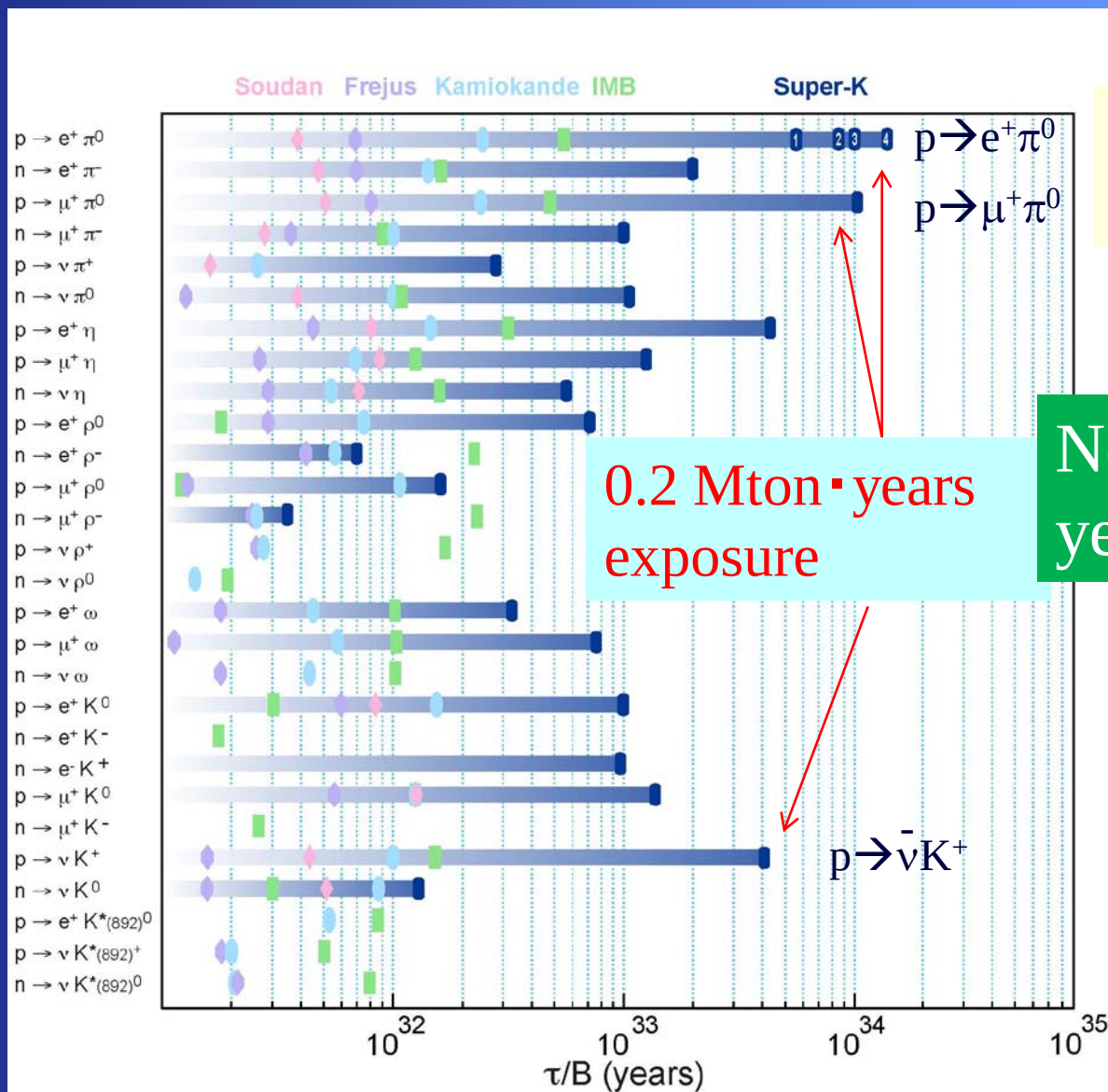
GUTs,  
B&L Violation



### 3. Prospect for nucleon decay by Hyper-K

- Nucleon decay experiment is the direct probe for GUTs.
- $p \rightarrow e^+ \pi^0$  (non-SUSY GUTs),  $p \rightarrow \bar{\nu} K^+$  (SUSY-GUTs) are regarded as the dominant mode.
- Why water cherenkov detector ?
  - Nucleon decay is extremely rare event. Need large number of nucleon.
    - ➔ (relatively) easy to enlarge detector size.
  - Two free protons, without interaction in nucleus.
    - ➔ High efficiency for  $p \rightarrow e^+ \pi^0$ , e.t.c. .

# How far we have explored by Super-K

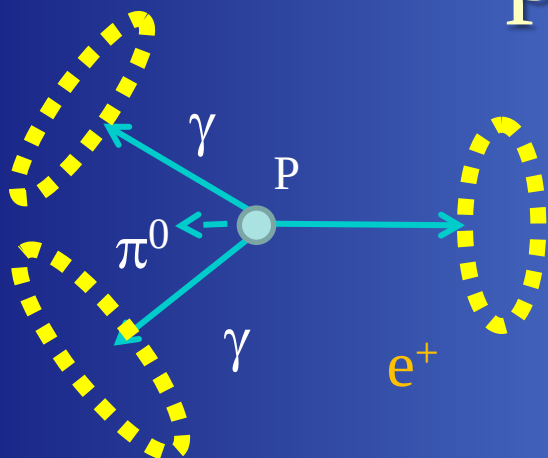


$> 1.3 \times 10^{34}$  years

No evidence yet

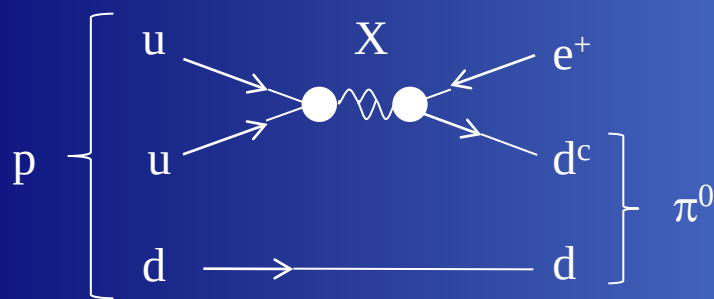
$> 4 \times 10^{33}$  years

$$p \rightarrow e^+ \pi^0$$



## Event features;

- $e^+$  and  $\pi^0$  are back-to-back (459 MeV/c)
- $\pi^0 \rightarrow 2 \gamma$  : all particles can be detectable.
- ➔ Reconstruct **proton mass and momentum.**



Minimal SU(5) model

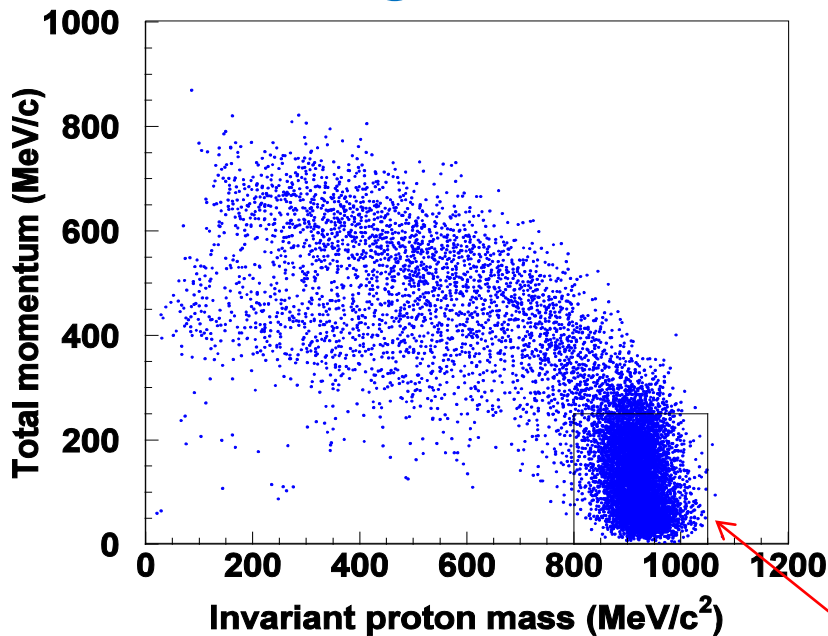
## Selection;

- Fully contained, VTX in fiducial volume.
- 2 or 3 ring .
- all e-like, w/o decay-e.
- $85 < M\pi^0 < 185$  MeV (for 3-ring event) .
- $800 < M_p < 1050$  MeV &  $P_{\text{tot}} < 250$  MeV/c

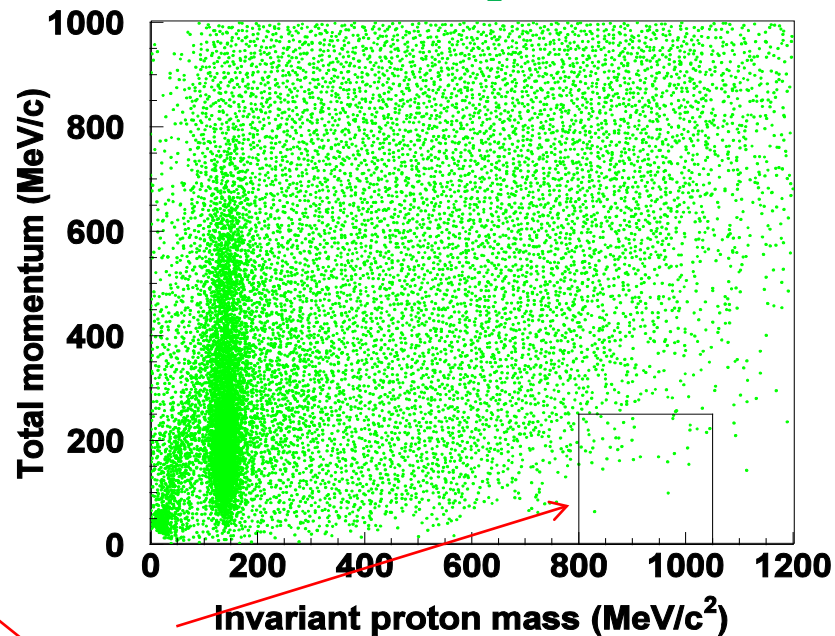
**All particles can be seen by water cherenkov detector**

# $M_{\text{tot}}$ vs $P_{\text{tot}}$ (after all other cuts)

Signal MC



BKG: atmospheric  $\nu$  MC



Signal Box

\* Inefficiency mostly due to  $\pi^0$  interaction in nucleus (absorption, scattering, charge exchange).

\* Dominant BKG: CC1 $\pi$   
 $\bar{\nu}_e p \rightarrow e^+ n \pi^0$   
 $\nu_e n \rightarrow e^- n \pi^+ + \text{charge exchange}$

# Performance of Hyper-K for $p \rightarrow e^+ \pi^0$

Efficiency: 45 % (87 % for free proton)

- \* photo coverage 20 % , but almost same performance is expected with current SK case (40 %). Phys.Rev.Lett. 102,141801(2009)

Background rate: 1.6 events/Mton $\cdot$ year

- \* Consistent with the estimation by using Acc.  $\nu$  data (K2K),  $1.63^{+0.62}_{-0.61}$  events/Mton $\cdot$ year.

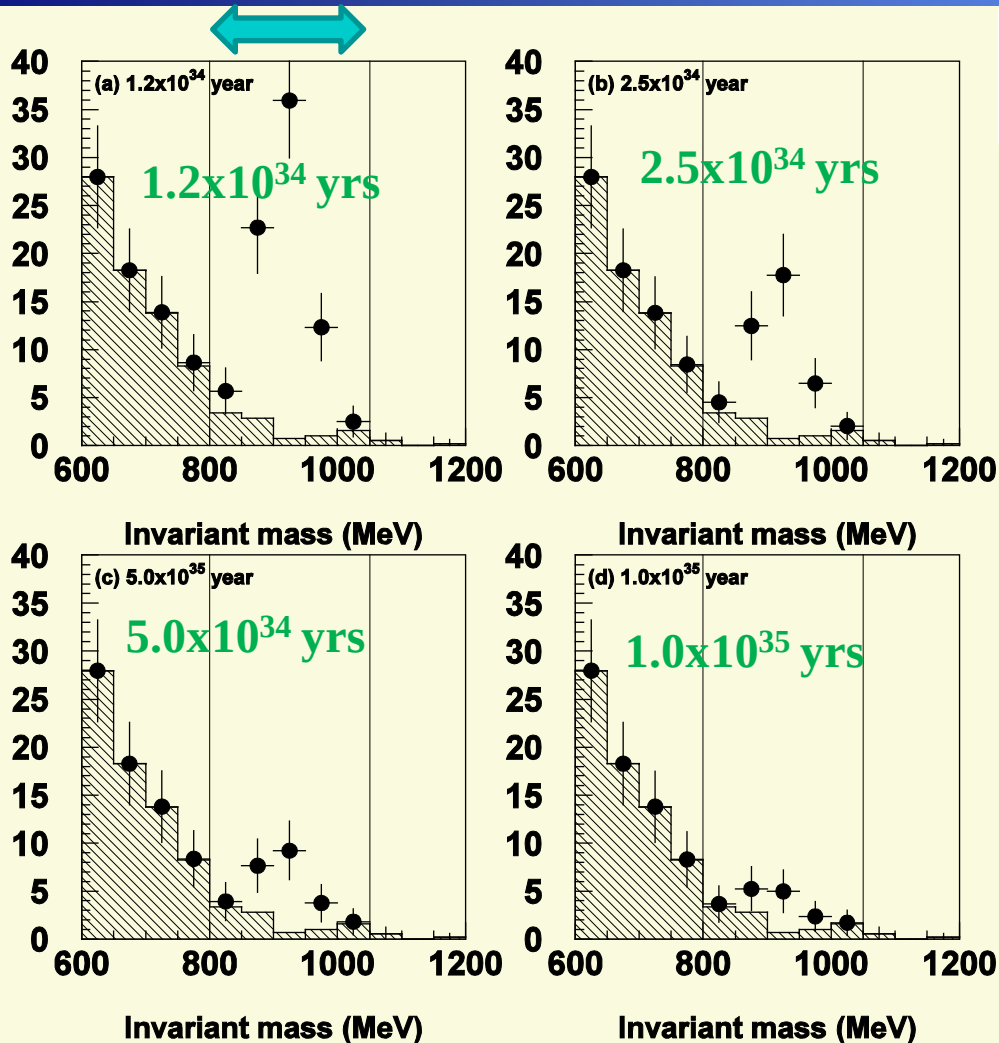
Phys.Rev.D77,032003 (2008)

- \* Room for improvement if **neutron can be tagged** (neutron is captured by water after a few 100  $\mu$ sec and 2.2 MeV  $\gamma$  is emitted).



# Discovery potential for $p \rightarrow e^+ \pi^0$

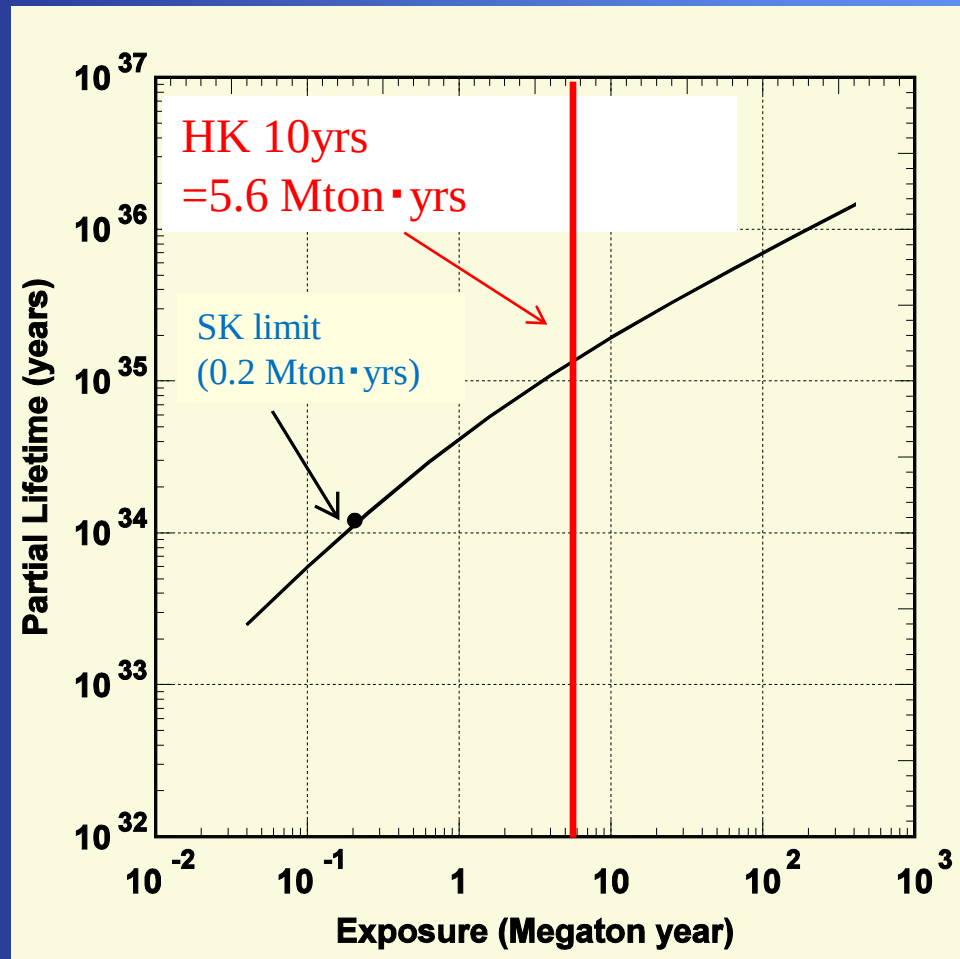
Total mass with 10yrs run assuming life time



Dot: Signal+BKG  
Hist: BKG

Discovery potential ( $3\sigma$ )  
 $< 5.6 \times 10^{34}$  years  
\* Counting event only.

# Sensitivity for $p \rightarrow e^+ \pi^0$ (90 %CL)



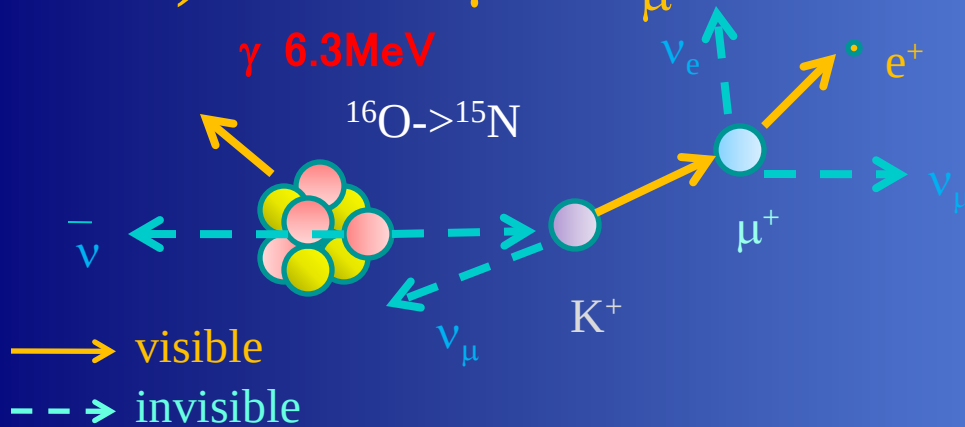
10 yrs run (5.6 Mton·yrs)

→ reach to  $1.3 \times 10^{35}$  yrs (90 %CL)

$$p \rightarrow \bar{\nu} + K^+$$

\* Cannot see  $K^+$ , stop in water  $\rightarrow$  use decay products

$$A) K^+ \rightarrow \mu^+ + \nu_\mu$$

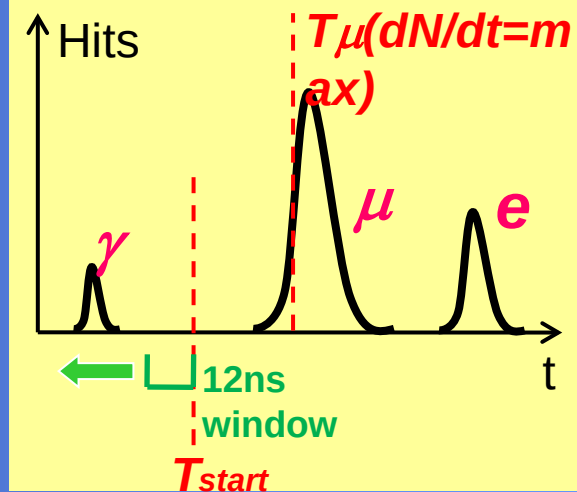


### Event features;

- 2 body decay ( $P_\mu = 236 \text{ MeV/c}$ ).
- $\rightarrow$  Excess in  $P_\mu$ .
- Proton in  $^{16}\text{O}$  decays and excited nucleus emits 6 MeV  $\gamma$  (Prob. 41%, not clear ring).
- $\rightarrow$  Tag  $\gamma$  to eliminate BKG.

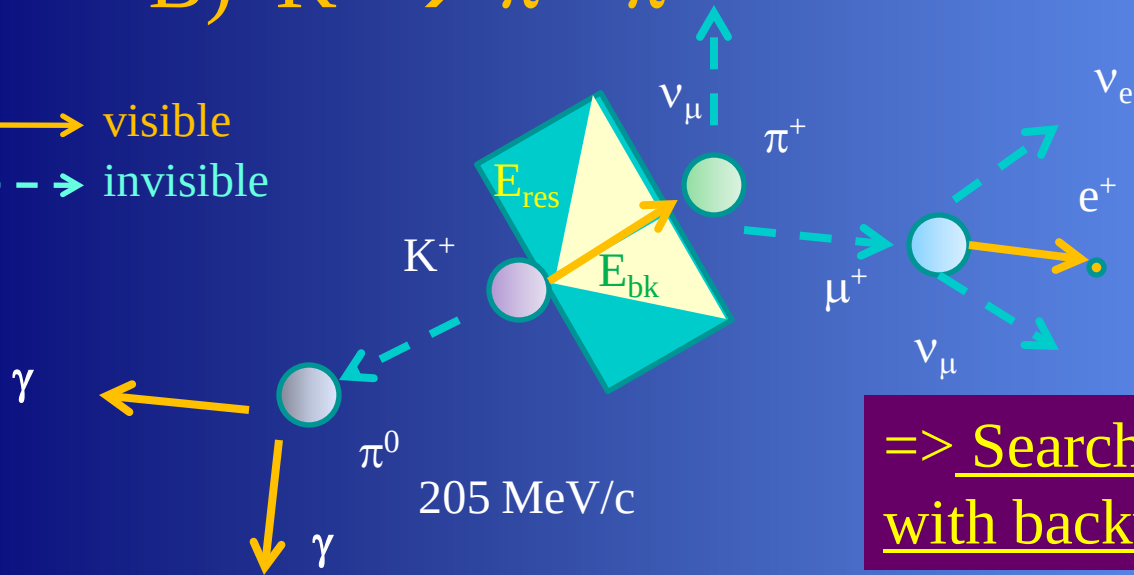
### Selection:

- 1  $\mu$ -like ring with decay-e.
- $215 < P_\mu < 260 \text{ MeV/c}$
- Search Max hit cluster by sliding time window (12ns width);  
 $4 < N_\gamma < 30 \text{ hits}$   
 $T_\mu - T_\gamma < 75 \text{ nsec}$





—→ visible  
 - - → invisible



### Event features;

- Br. 21 %.
- $\pi^0$  and  $\pi^+$  are back-to-back and have 205 MeV/c.
- $P\pi^+$  is just above  $\check{C}$  thres. (not clear ring).

=> Search for monochromatic  $\pi^0$  with backward activities.

### Selection:

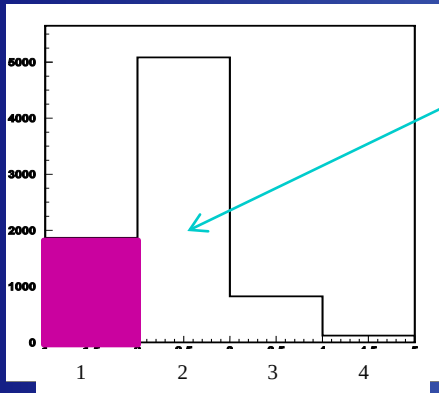
- 2 e-like rings with decay-e.
- $85 < M\pi^0 < 185$  MeV.
- $175 < P\pi^0 < 250$  MeV/c.
- $E_{bk}$ : visible energy sum in 140-180 deg. of  $\pi^0$  dir,

$E_{res}$ : in 90-140 deg.

$7 < E_{bk} < 17$  MeV &  $E_{res} < 12$  MeV

# Recent improvement(1): $p \rightarrow \bar{\nu} K^+, K^+ \rightarrow \pi^+ \pi^0$

# of Ring:  $K^+ \rightarrow \pi^+ \pi^0$



Judge as 1 ring if opening angle of 2  $\gamma$ s is small or momentum of one  $\gamma$  is small,

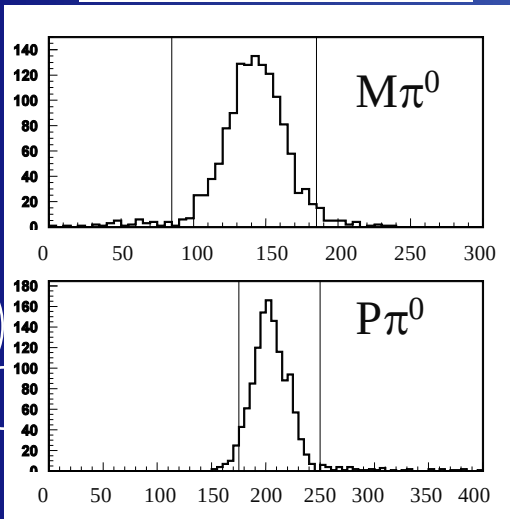


Use “ $\pi^0$  fitter”

- Make likelihood assuming  $\pi^0$  and search for missing ring.
- It is used for  $\nu_e$  appearance analysis of T2K to reduce BKG.

**It makes 1 ring sample available for this analysis!**

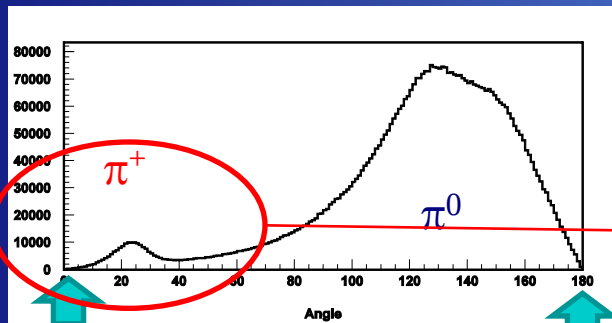
**→ efficiency increased.**





## Recent improvement(2): $p \rightarrow \bar{\nu} K^+, K^+ \rightarrow \pi^+ \pi^0$

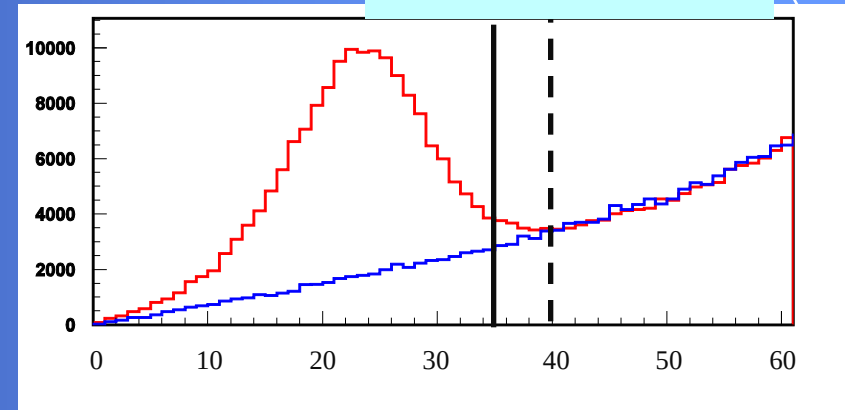
Charge distribution in angle



zoom

Red: PDK MC

Blue: BKG



Opposite of  $\pi^0$  dir  
=  $\pi^+$  dir

$\pi^0$  dir

Conventional method: Use charge sum in  $< 40^\circ$

New method: Use charge sum in  $< 35^\circ$

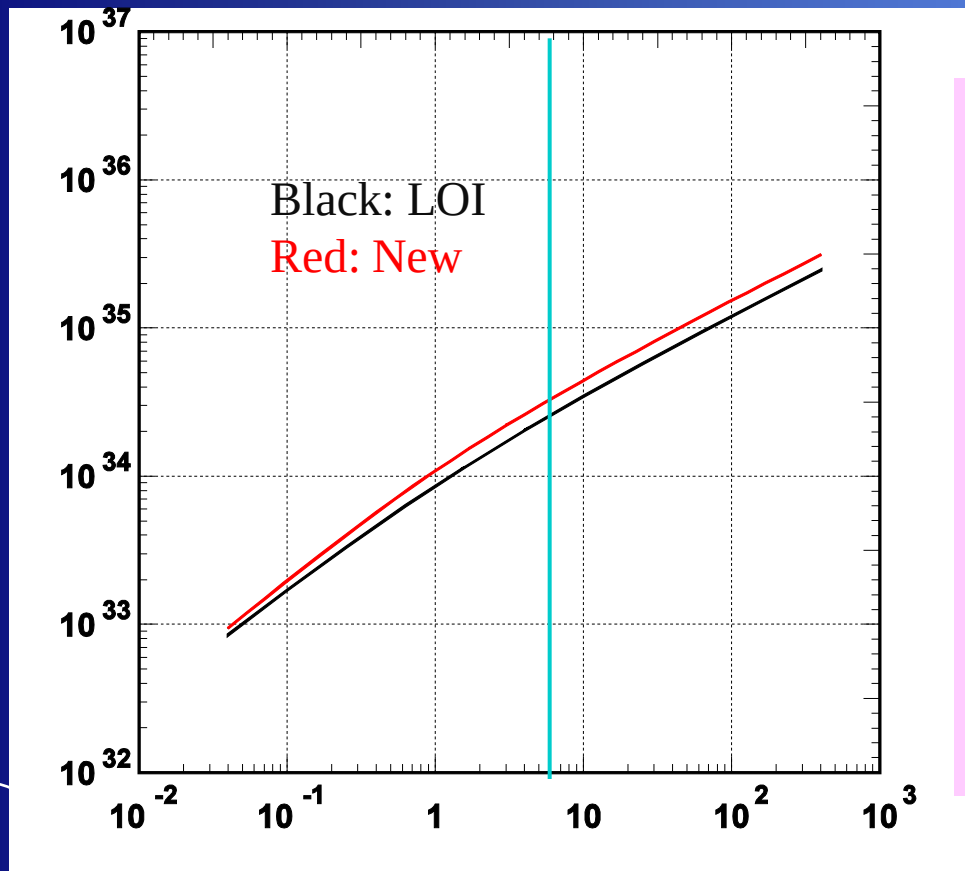
and compare shape by likelihood assuming signal and BKG.

## How much improved ?

|                                                   | Effi (%)                       | BKG/Mton·yr                     |
|---------------------------------------------------|--------------------------------|---------------------------------|
| $K^+ \rightarrow \pi^+ \pi^0$                     | 6.6 (LOI)<br>$\rightarrow 7.6$ | 6.6 (LOI )<br>$\rightarrow 1.8$ |
| $K^+ \rightarrow \mu\nu + \text{nuclear } \gamma$ | 7.1                            | 1.6                             |
| $K^+ \rightarrow \mu\nu, P\mu$                    | 43                             | 1940                            |

- Efficiency increased by 15 %, due to 1R sample.
- BKG reduced by 72 %, due to angle cut and shape likelihood.
- $\pi^+ \pi^0$  method become comparable to prompt  $\gamma$  method !

# Sensitivity curve for $p \rightarrow \bar{\nu} K^+$ (90 % C.L.)



Sensitivity for 5.6 Mton  $\cdot$  year exposure (10 years run)

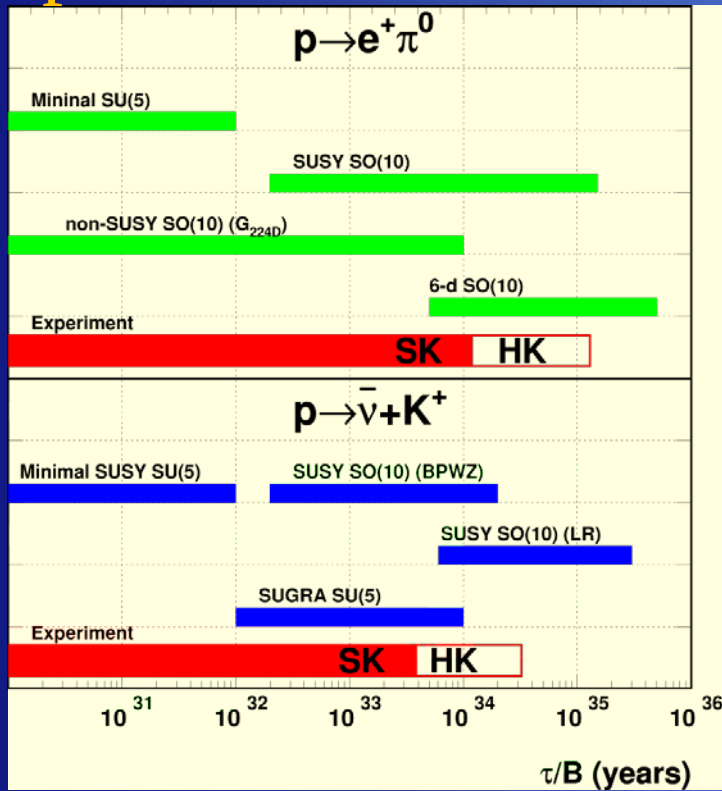
$2.5 \times 10^{34} \rightarrow 3.2 \times 10^{34} \text{ year}$

$3\sigma$  discovery potential

$0.95 \times 10^{34} \rightarrow 1.23 \times 10^{34} \text{ year}$

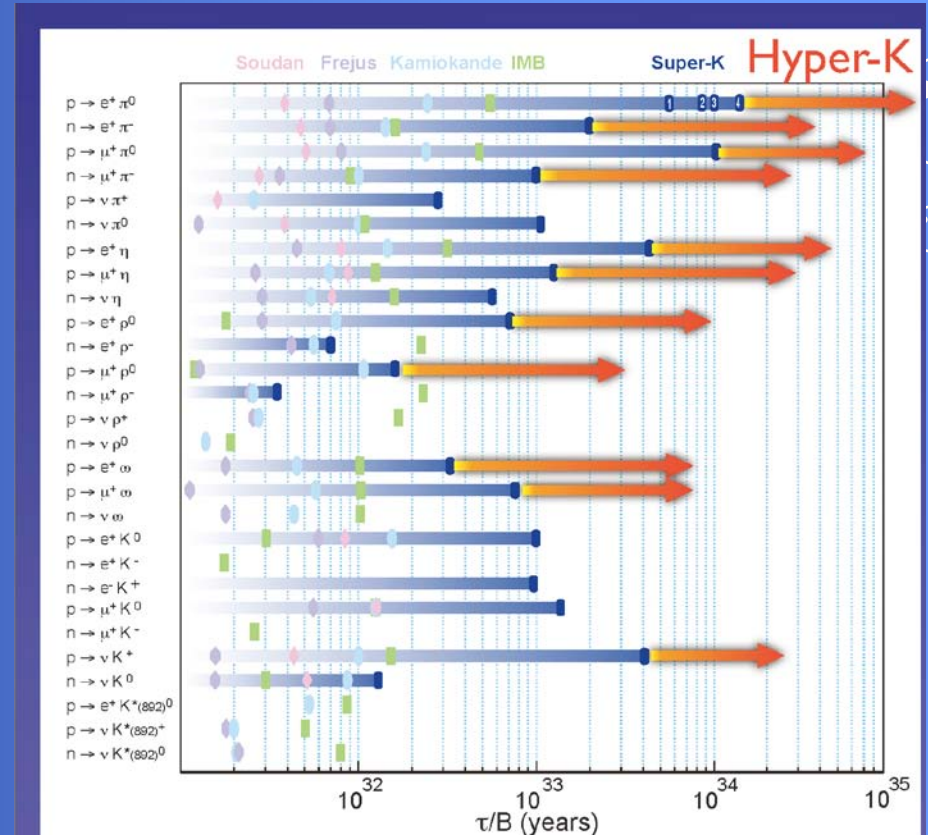
$\sim 30\%$  improved from LOI !

# Comparison with predictions



Can cover major models.

# Other modes



~ 10 times better than SK with 10 year run.

# Other physics potential

- $\delta_{CP}$  precision  $< 20^\circ$   
with 7.5 MW  $\cdot$  years ,  $s^2 2\theta_{13} > 0.03$  with MH known
- MH determination  $> 3\sigma$  CL  
with 10 yrs run,  $0.4 < s^2\theta_{23}$  and  $0.04 < s^2 2\theta_{13}$
- $\theta_{23}$  octant determination  $> 90\%$  CL  
with 10 yrs run,  $s^2 2\theta_{23} < 0.99$  and  $0.04 < s^2 2\theta_{13}$
- Solar  $\nu$  from  $^8B$  200 events/day  
with 7 MeV threshold
- Super Nova  $\nu$  170,000~200,000 events  
@ Galactic center (10kpc)

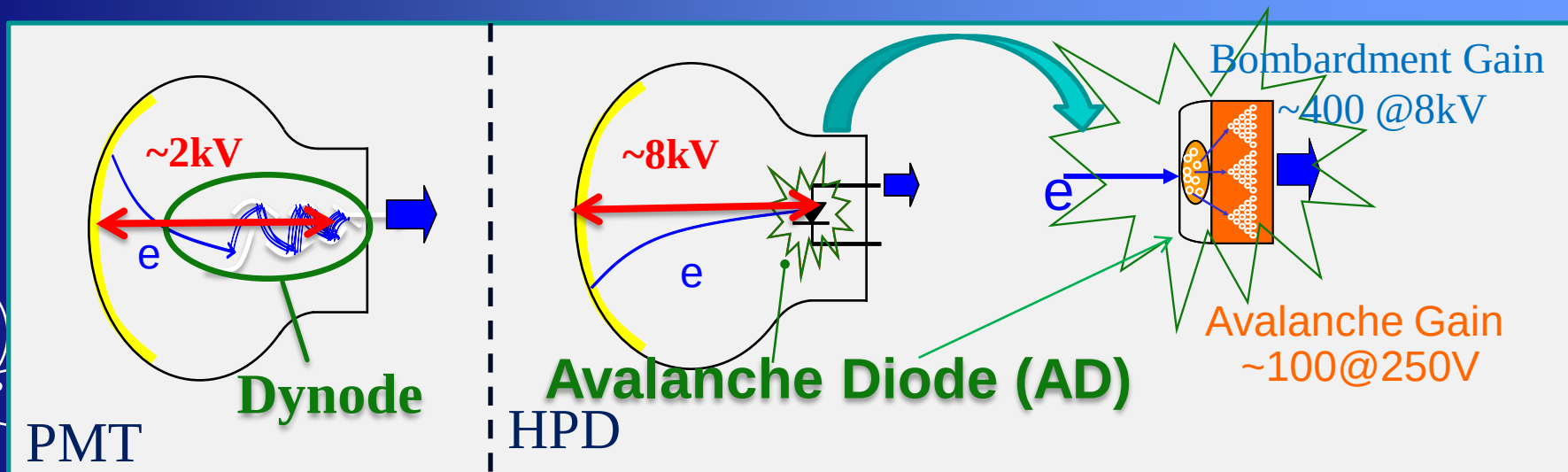
## 4. R&D status and plans

### Photo sensors:

#### 20 inch PMT in baseline design

- It is almost hand-made (even in electric parts).
- Expensive and not suitable for mass production.
- But we need 99,000 PMTs.

→ Other possibility: Hybrid Photo Detector (HPD)





# Advantage of HPD

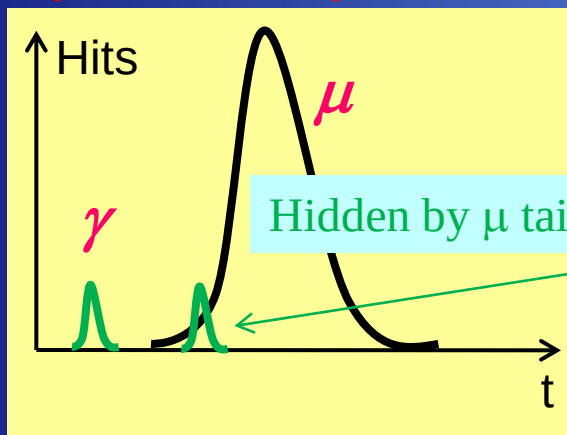
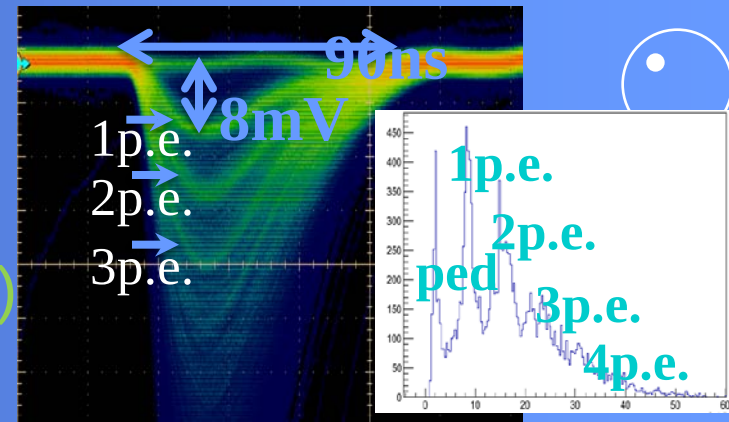
- **Avalanche Diode** is suitable for mass production.

→ lower cost.

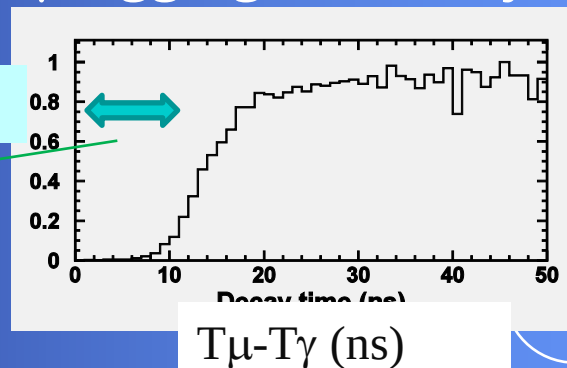
- Better photon counting.
- Better collection efficiency .  
(20 inch PMT ~75 % → HPD ~95%)

- Better timing resolution  
(20 inch PMT ~2.2 nsec → HPD ~1.1 ns in TTS)

→ Expect to improve nuclear  $\gamma$  tagging efficiency in  $p \rightarrow \bar{\nu} K^+$



$\gamma$  tagging efficiency



# HPD for Hyper-K

- Target: **20 inch HPD**, but no such a large one in the world.
  - ➔ Made **8 inch** with Hamamatu photonics as test module. Now checking basic performance.
- Never used in water ( $\sim 8\text{kV}$ )
  - ➔ Install to test tank (200 ton) in this summer. Check water proof, long term stability, e.t.c.

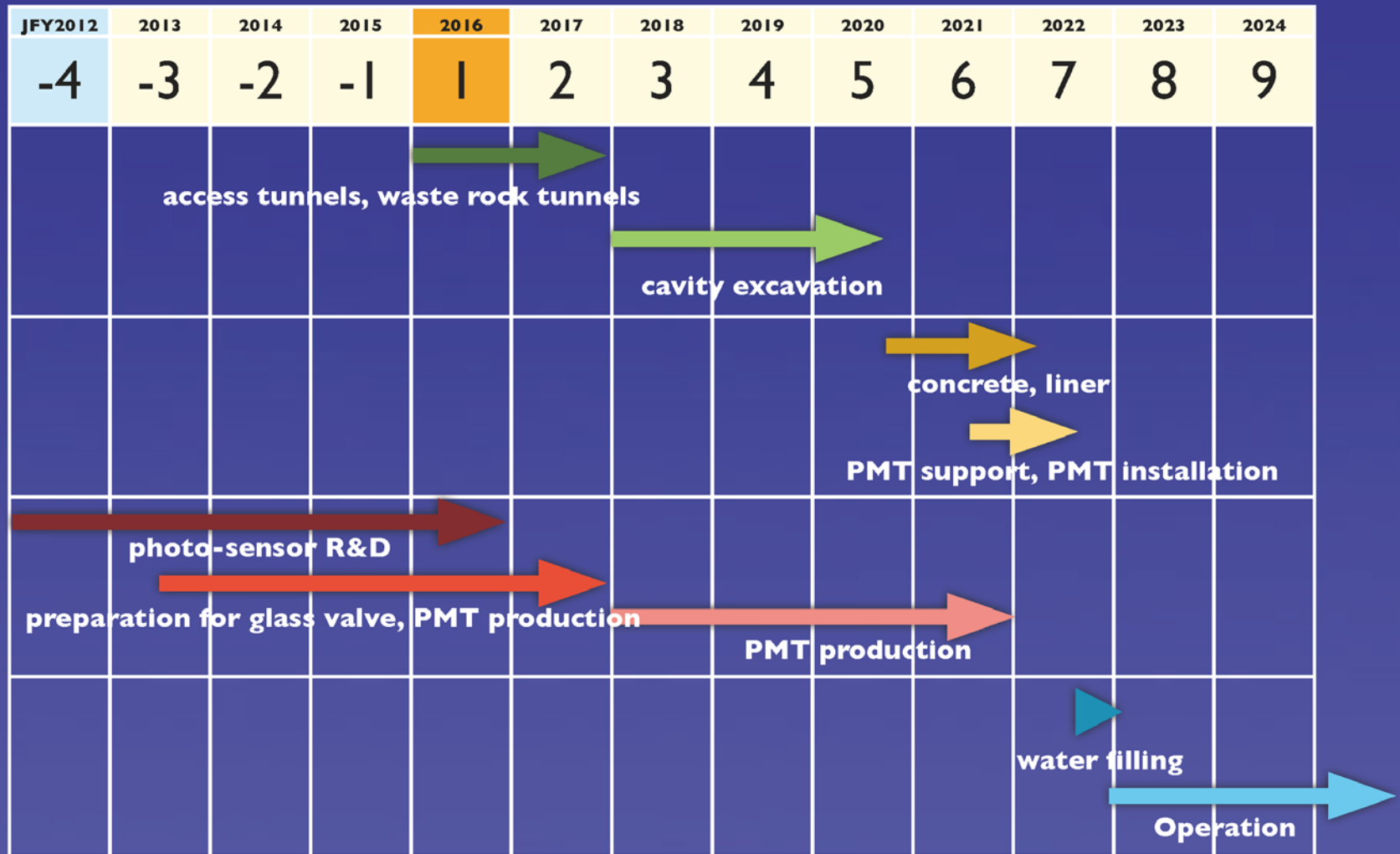
# Software development

- SK detector simulation : Based on GEANT3  
Reconstruction tool: written in Fortran
- Now start to making Hyper-K detector simulation based on GEANT4.
- New reconstruction algorithm written in C++:  
developing for SK and T2K analysis, also for Hyper-K.

**→ Evolving to improve physics reach!**

# Target Schedule

Construction start



assuming budget being approved from JPY2016

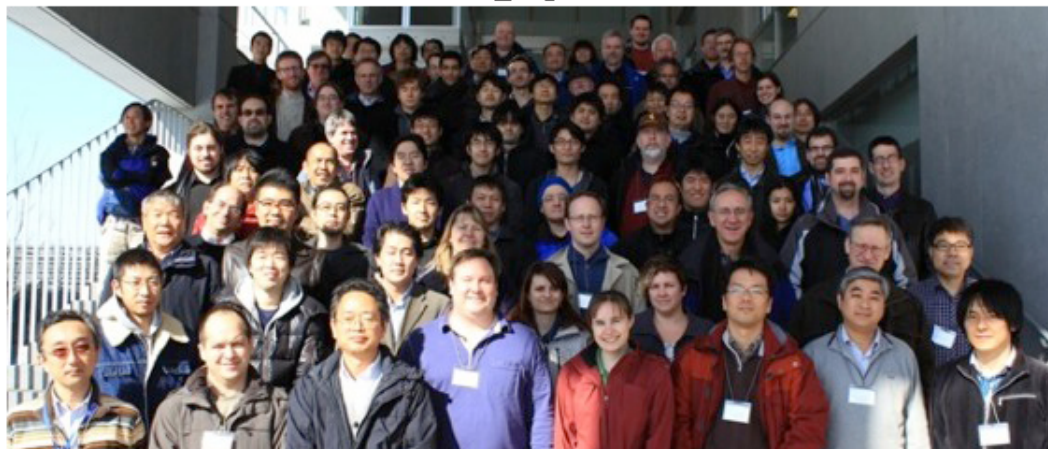
# International Hyper-K meetings

First meeting: Aug. 23-24, 2012



<http://indico.ipmu.jp/indico/conferenceTimeTable.py?confId=7>

Second meeting: Jan. 14-15, 2013



<http://indico.ipmu.jp/indico/conferenceTimeTable.py?confId=10>

Hyper-K is completely open to the international community

~100 participants for each of two meetings (~half from abroad)

International working group was formed

Current members from Japan, Canada, Spain, Switzerland, Russia, UK, and US

Next meeting: Jun. 21-22



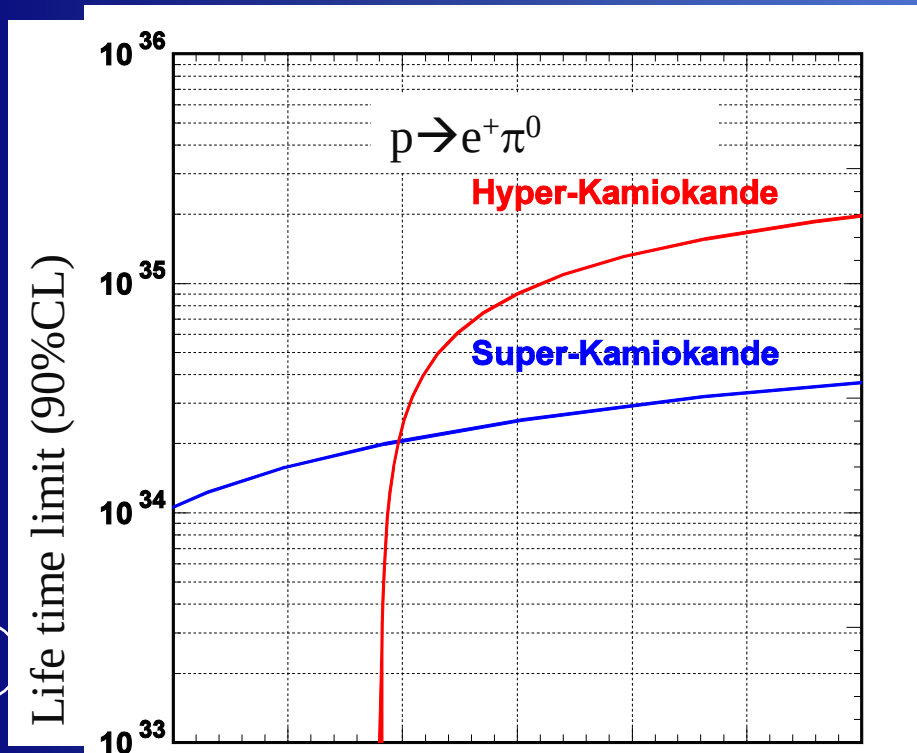
## 5. Summary

- We are planning to build Hyper-Kamiokande, 1Mton class water cherenkov detector.
- Hyper-Kamiokande can explore **one order longer nucleon lifetime** than Super-K.
- Now international working group is forming and **we welcome your join.**

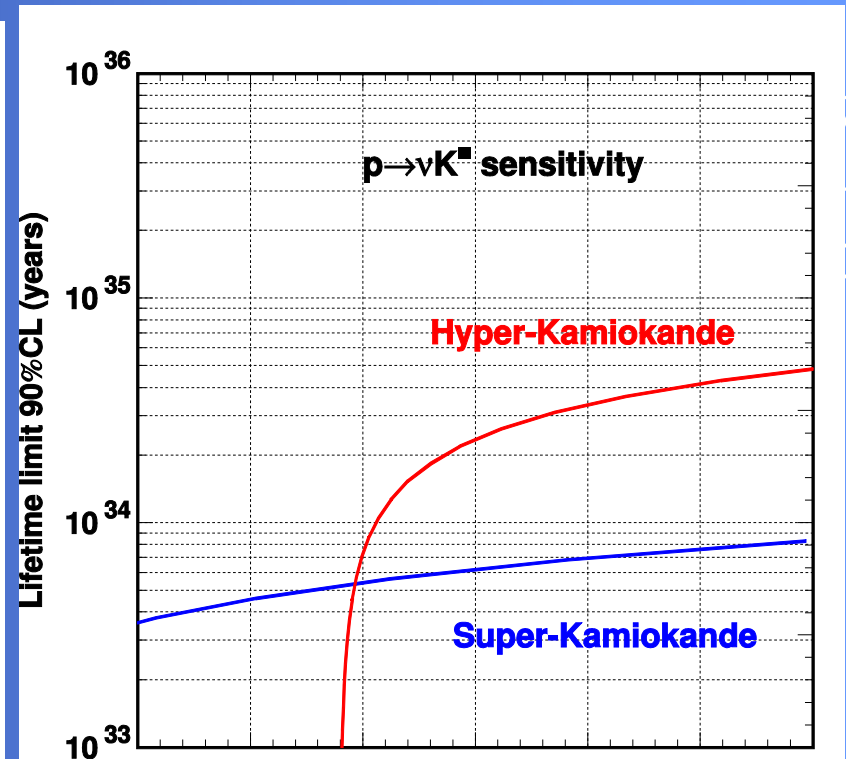


# Backup

# Comparison with Super-K



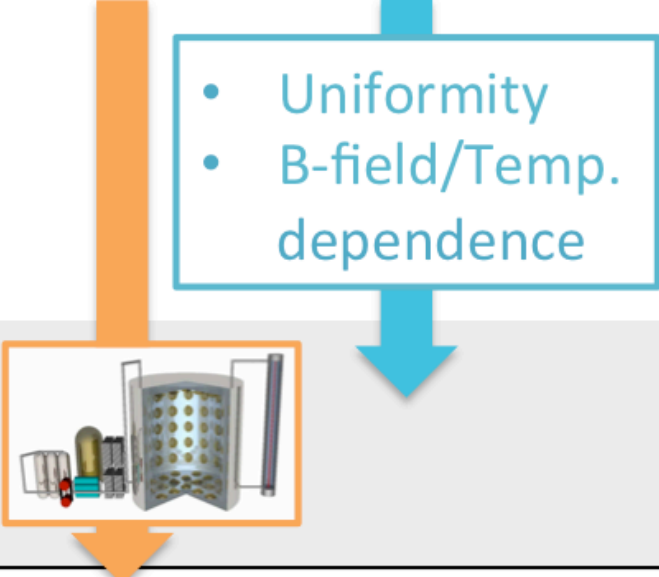
Running Year



Running Year

- HK Overtakes SK by  $\sim$ one year run.
- Explore one order longer lifetime than SK by 10 years run.

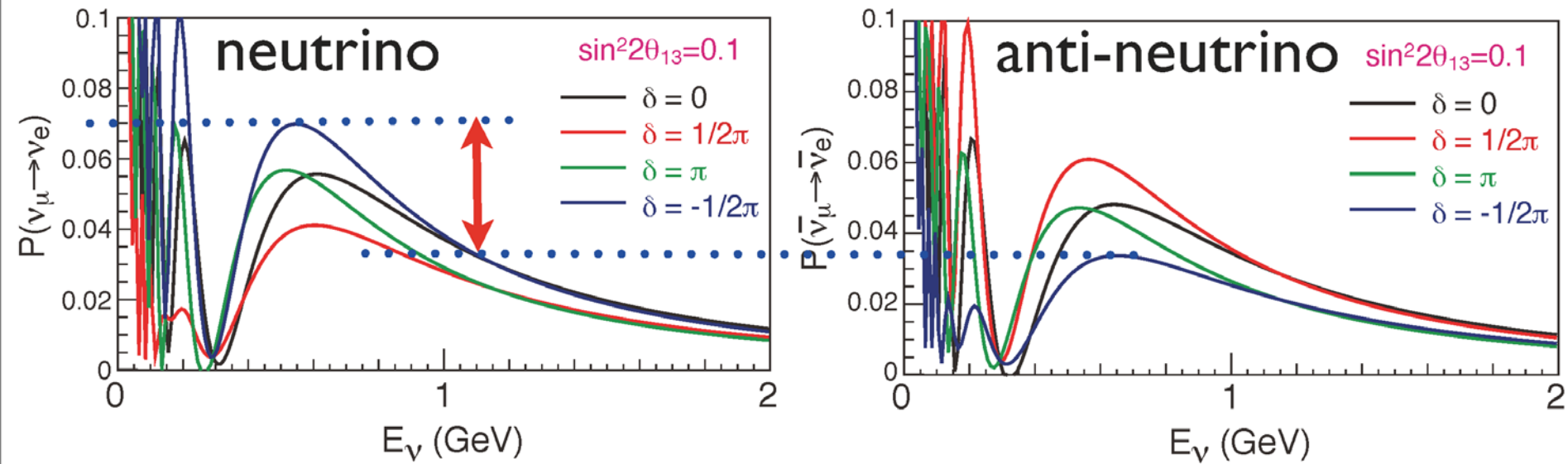
# Schedule of HPD R&D

| Year | Season. | 8 inch HPD                                                                                                                                                                             | 20 inch HPD                                |
|------|---------|----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|--------------------------------------------|
| 2013 | Spring  | <b>Proof test</b><br><b>200t tank</b><br>Performance Evaluation (detail)                                                                                                               | <b>Prototype</b><br>Performance Evaluation |
| 2014 |         |  <ul style="list-style-type: none"> <li>• Uniformity</li> <li>• B-field/Temp. dependence</li> </ul> |                                            |
| 2015 | Spring  |                                                                                                      | <b>Proof Test</b><br><b>in 200t tank</b>   |
| 2016 |         | <b>Determine HK photo sensor !</b>                                                                                                                                                     |                                            |

We plan to finish our HPD study until 2016.

# Measuring CP asymmetry w/ J-PARC $\nu$ beam

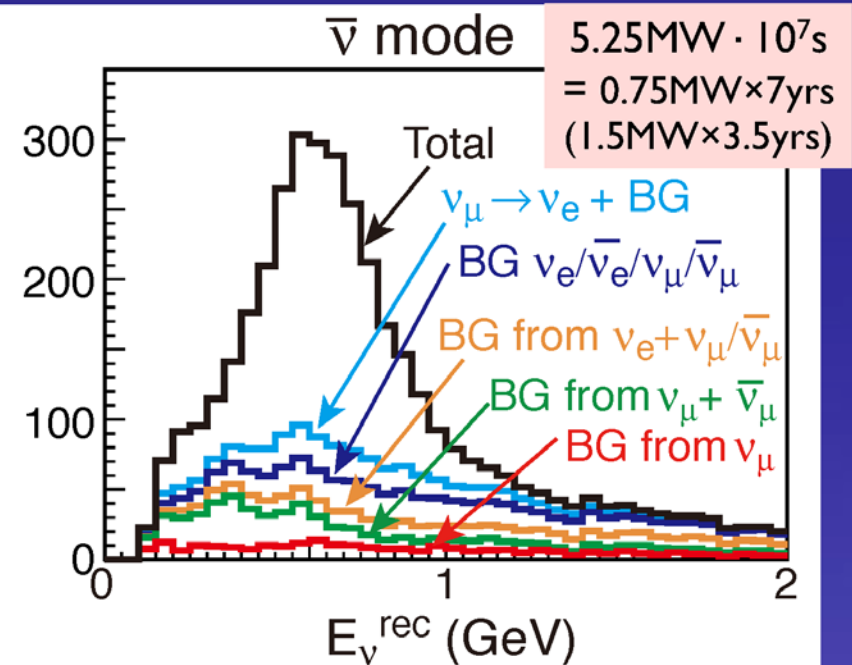
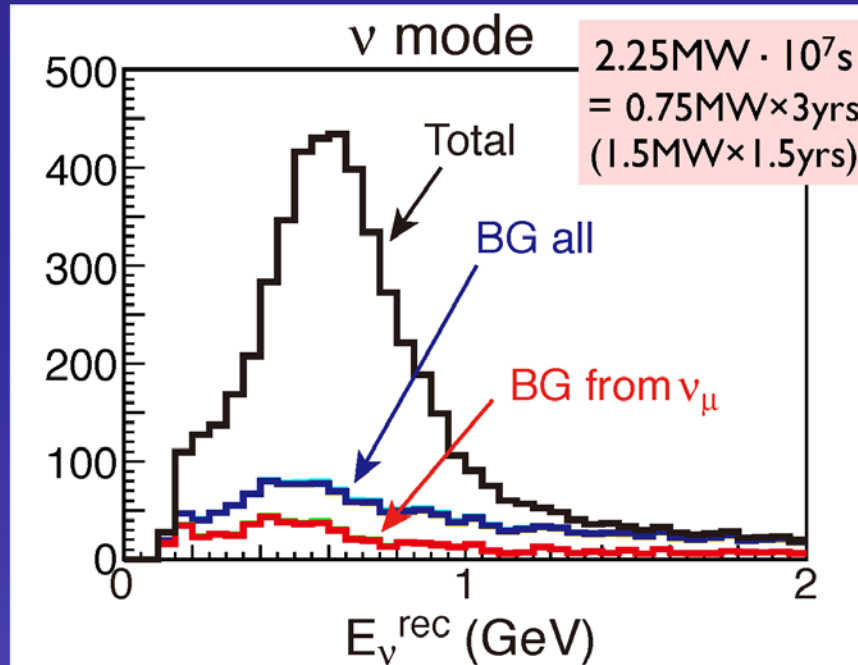
$P(\nu_\mu \rightarrow \nu_e)$  appearance probability  
(normal hierarchy)



- Comparison between  $P(\nu_\mu \rightarrow \nu_e)$  and  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

# $\nu_e$ candidate events after selection

$\sin^2 2\theta_{13}=0.1, \delta=0$ , normal MH



|                                          | Signal<br>( $\nu_\mu \rightarrow \nu_e$ CC) | Wrong sign<br>appearance | $\nu_\mu/\bar{\nu}_\mu$<br>CC | beam $\nu_e/\bar{\nu}_e$<br>contamination | NC  |
|------------------------------------------|---------------------------------------------|--------------------------|-------------------------------|-------------------------------------------|-----|
| $\nu$ (2.25MW · 10 <sup>7</sup> s)       | 3,560                                       | 46                       | 35                            | 880                                       | 649 |
| $\bar{\nu}$ (5.25MW · 10 <sup>7</sup> s) | 1,959                                       | 380                      | 23                            | 878                                       | 678 |

2000-4000 signal events expected for each of  $\nu$  and  $\bar{\nu}$

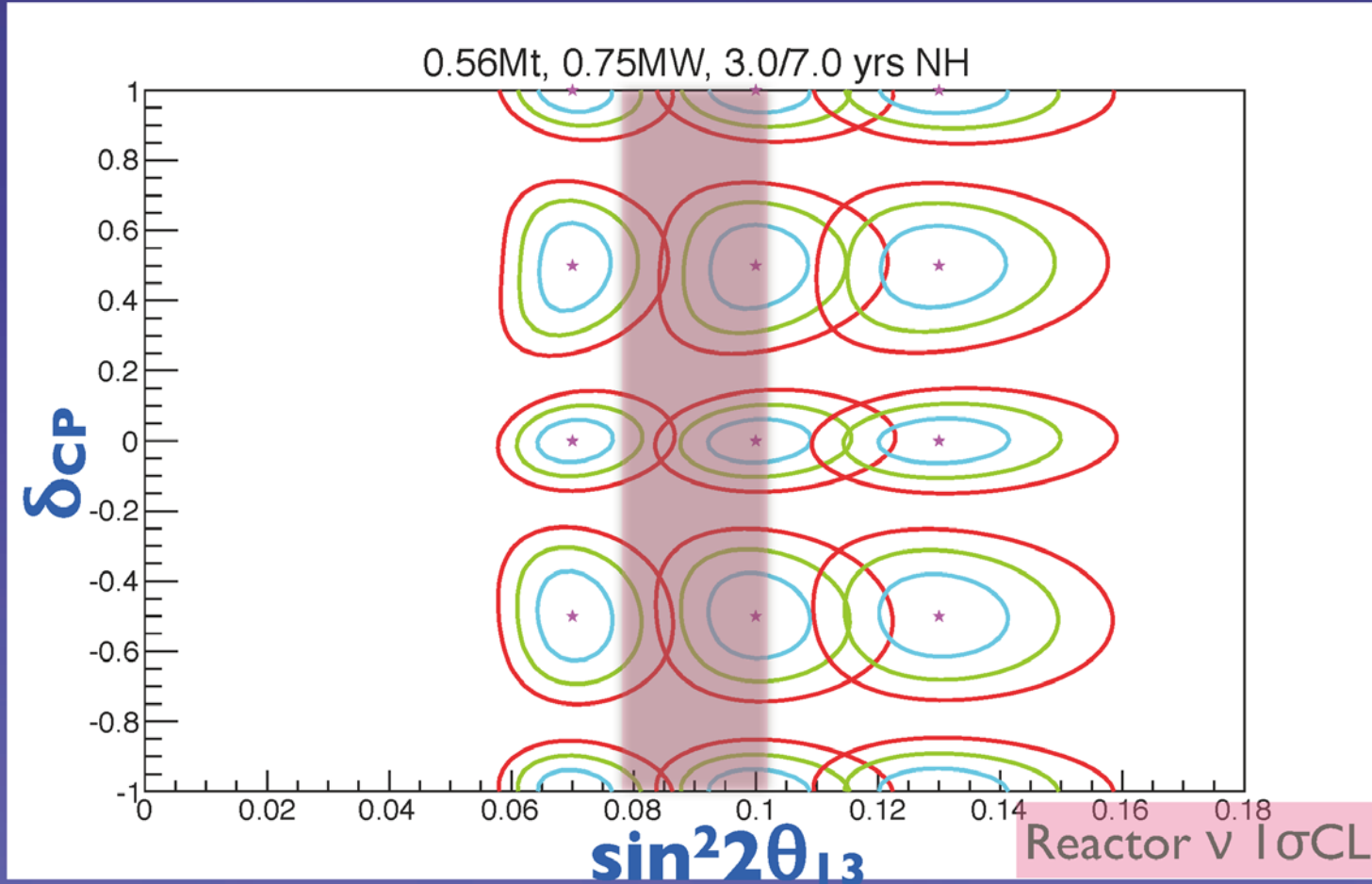
7.5MW · years

# Contours

Normal mass hierarchy (known)

5% systematics on signal,  $\nu_\mu$  BG,  $\nu_e$  BG,  $\nu/\bar{\nu}$

$\sin^2 2\theta_{23} = 1$



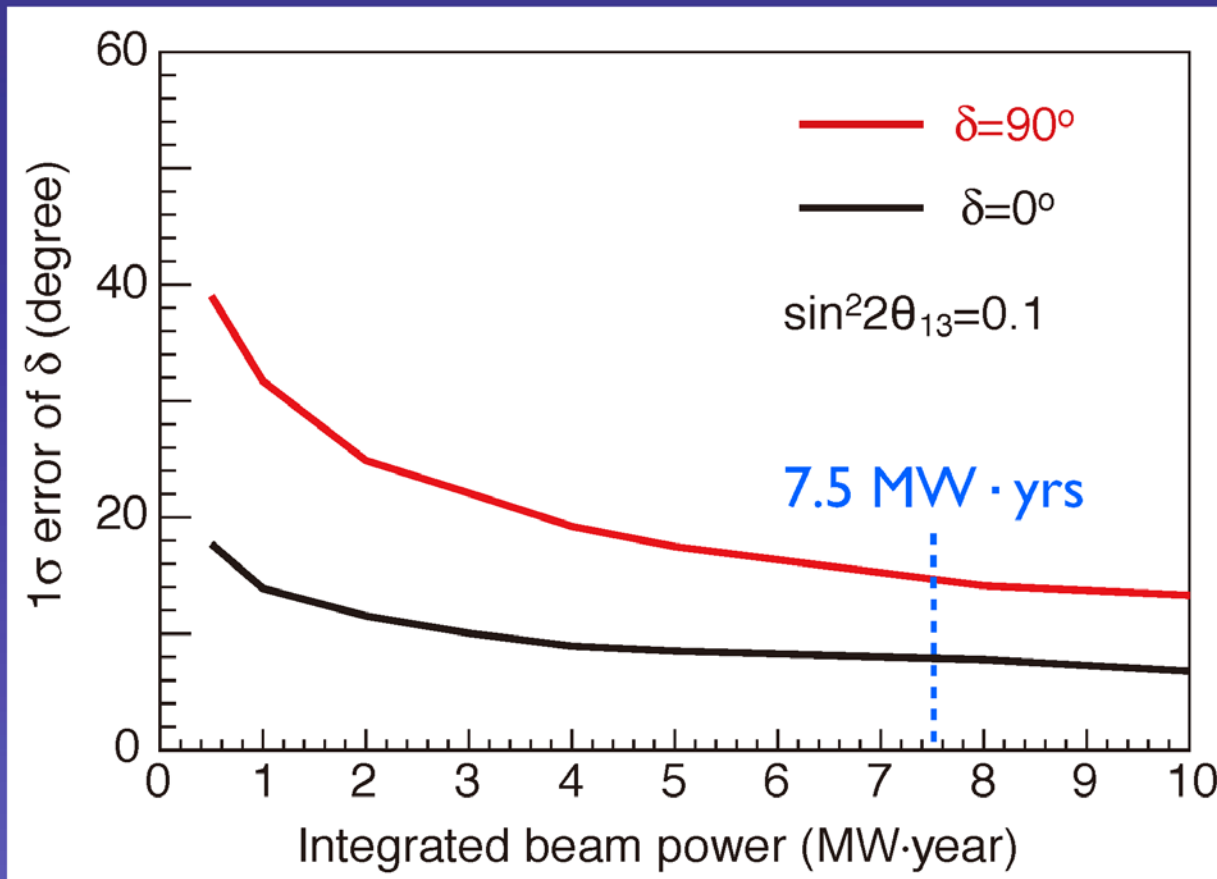
- Good sensitivity for CPV
- modest dependence on  $\theta_{13}$  value



# $\delta$ resolution

Normal mass hierarchy (known)

$$\sin^2 2\theta_{13} = 0.1$$



- ▶  $\delta$  precision  $< 20^\circ$  ( $\delta = 90^\circ$ )  
 $< 10^\circ$  ( $\delta = 0^\circ$ )
- ▶ modest dependence on  $\theta_{13}$

# 3-flavor oscillations in atmospheric $\nu$

NuclPhysB669,255(2003)

NuclPhysB680,479(2004)

$$\frac{\Phi(\nu_e)}{\Phi_0(\nu_e)} - 1 \approx P_2(r \cdot \cos^2 \theta_{23} - 1) \quad \text{Solar term}$$

$$-r \cdot \sin \tilde{\theta}_{13} \cdot \cos^2 \tilde{\theta}_{13} \cdot \sin 2\theta_{23} (\cos \delta \cdot R_2 - \sin \delta \cdot I_2)$$

$$+ 2 \sin^2 \tilde{\theta}_{13} (r \cdot \sin^2 \theta_{23} - 1)$$

Interference term ( $\delta\text{CP}$ )  
 $\theta_{13}$  resonance term

$r$  :  $\mu/e$  flux ratio ( $\sim 2$  at low energy)

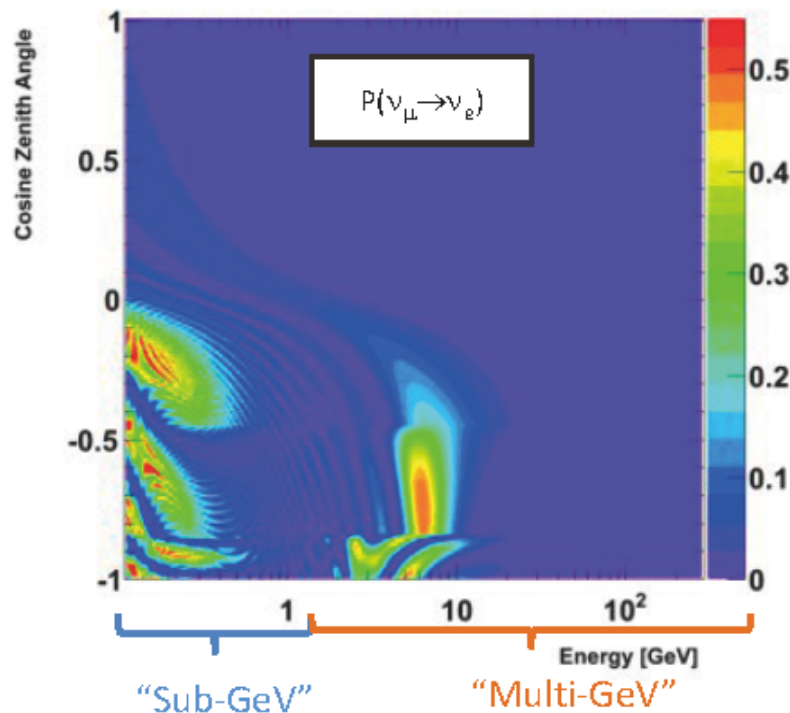
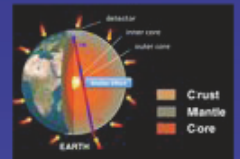
$P_2 = |A_{e\mu}|^2$  :  $2\nu$  transition probability  $\nu_e \rightarrow \nu_{\mu\tau}$  in matter

$$R_2 = \text{Re}(A_{ee}^* A_{e\mu})$$

$$I_2 = \text{Im}(A_{ee}^* A_{e\mu})$$

$A_{ee}$  : survival amplitude of the  $2\nu$  system

$A_{e\mu}$  : transition amplitude of the  $2\nu$  system

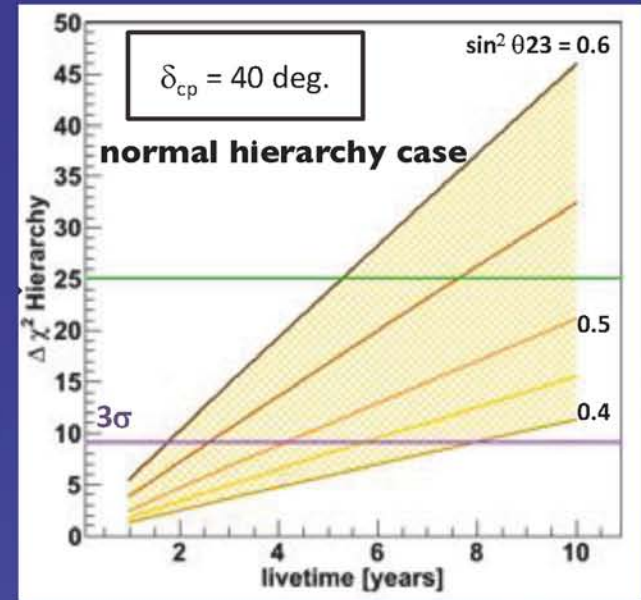
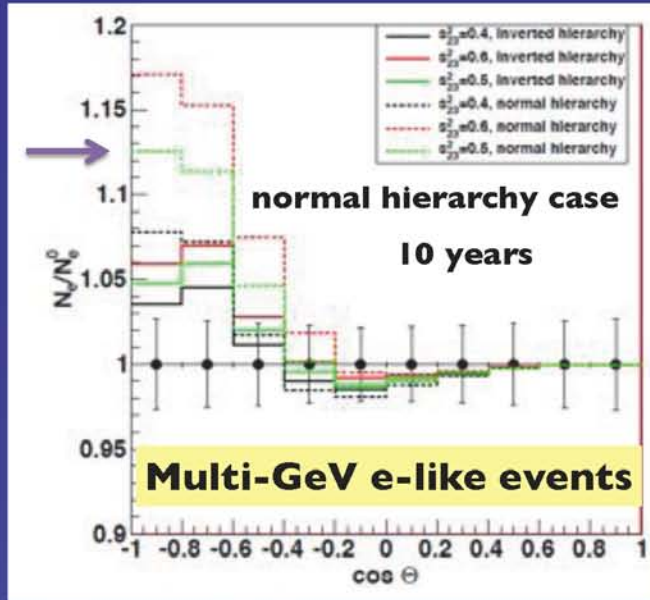


$\nu_e$  appearance (and  $\nu_{\mu}$  distortion) is expected due to MSW effect in the Earth's matter

- happens in  $\nu$  in the case of normal mass hierarchy
- in anti- $\nu$  in inverted mass hierarchy

Large  $\theta_{13}$  value gives us a good chance to discriminate mass hierarchy.

# Mass Hierarchy Sensitivity



- Sensitivity depends on  $\theta_{23}$ ,  $\delta$  and mass hierarch (a little).
- $3\sigma$  mass hierarchy determination for  $\sin^2 \theta_{23} > 0.42$  (0.43) in the case of normal (inverted) hierarchy.

